

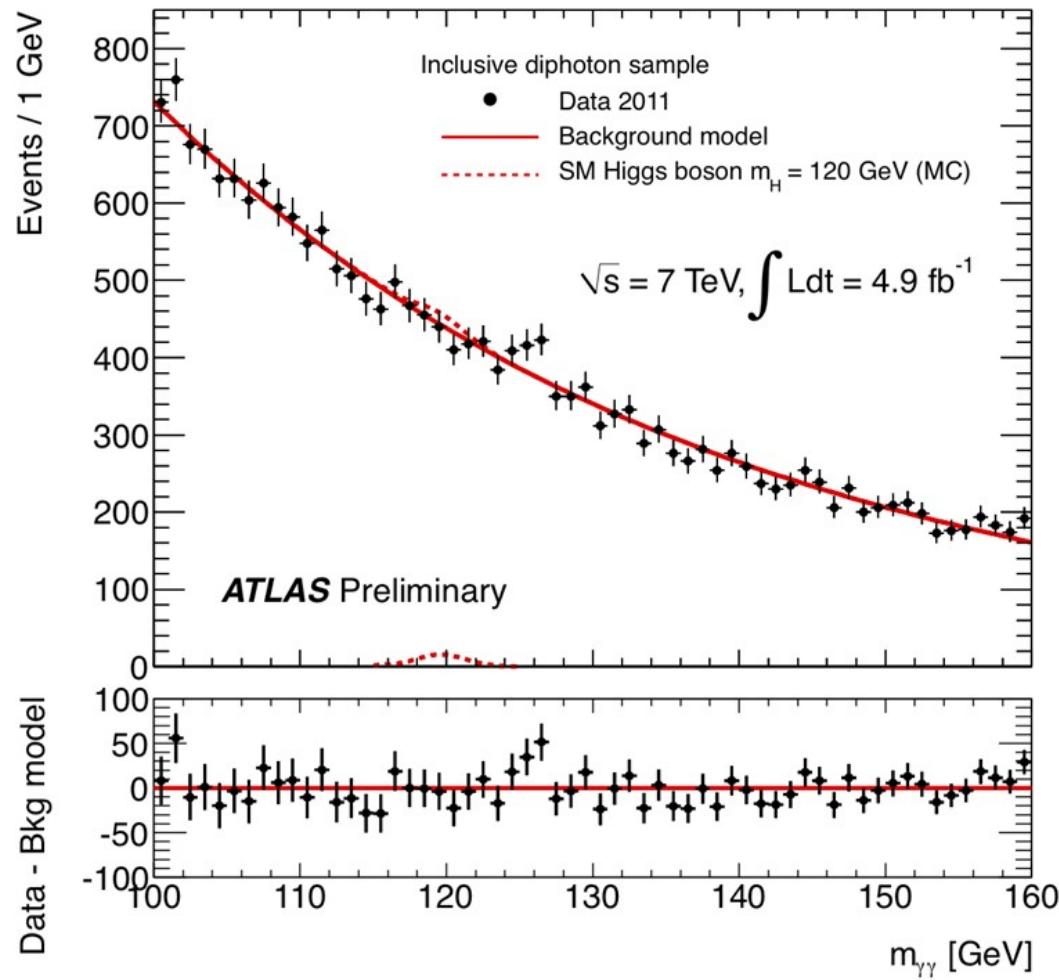
# Bayesian model comparison

ICIC Data Analysis Workshop 2016

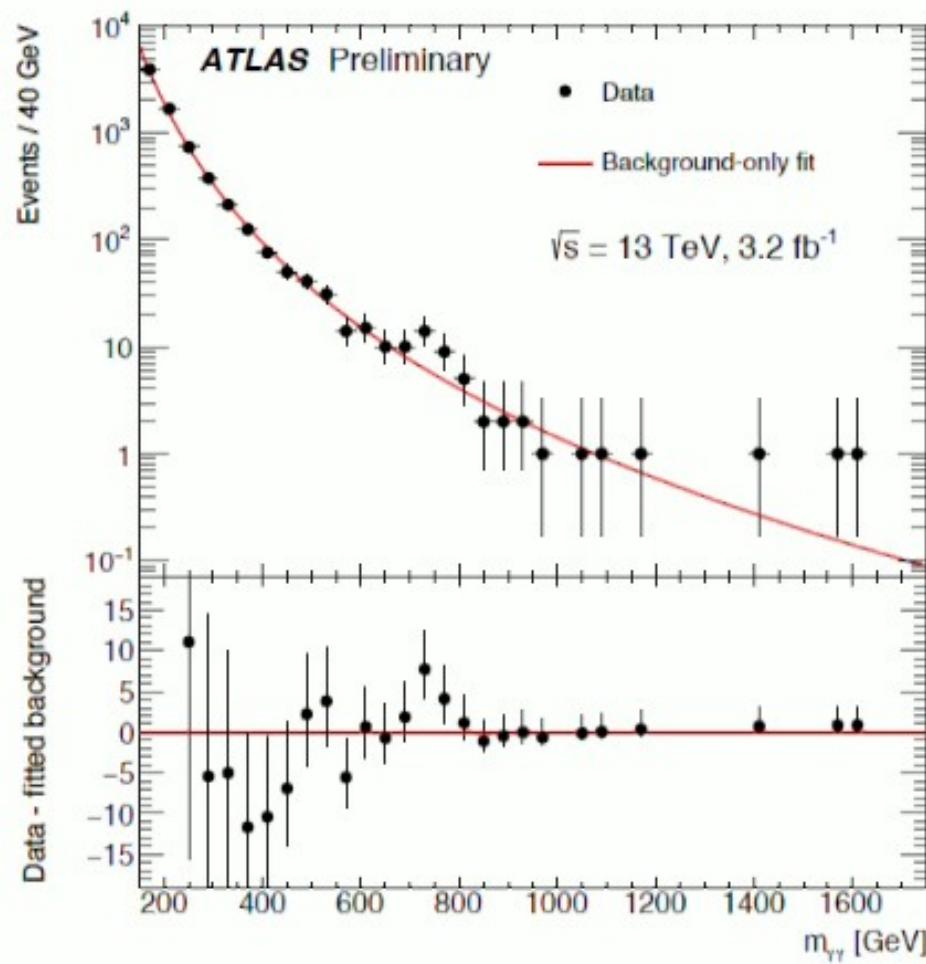
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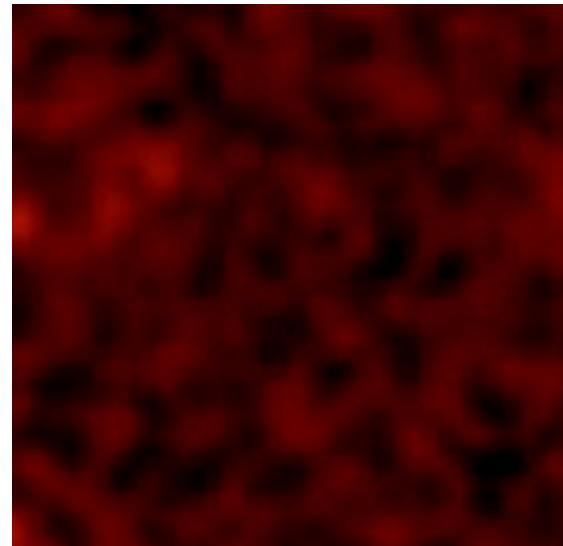
# Typical questions



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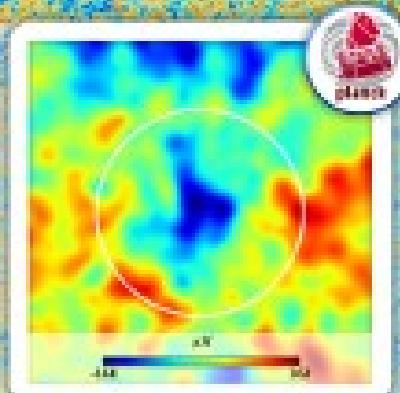
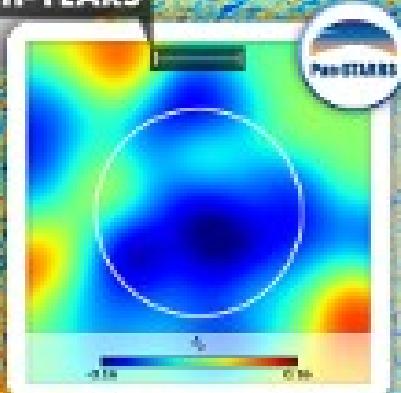


# Typical questions



# Typical questions

1 BILLION LIGHT-YEARS

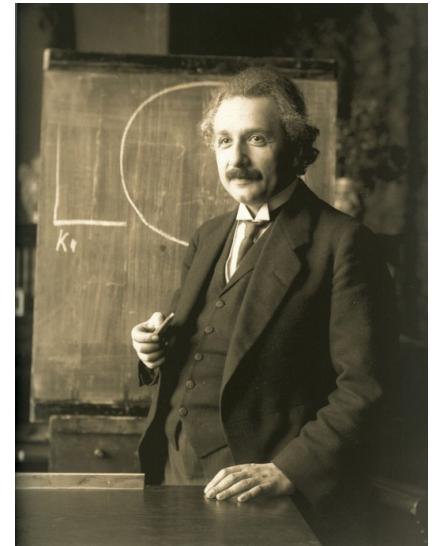


ANDROMEDA

MOON

# Typical questions

$$S_{EH} = \frac{1}{16\pi G} \int \sqrt{-g} (R - 2\Lambda) d^4x$$



$$S = \int d^4x \sqrt{-g} \mathcal{L}_H$$

$$\mathcal{L}_H = \mathcal{L}_2 + \mathcal{L}_3 + \mathcal{L}_4 + \mathcal{L}_5$$

$$\mathcal{L}_2 = K(\phi, X),$$

$$\mathcal{L}_3 = -G_3(\phi, X)\square\phi,$$

$$\mathcal{L}_4 = G_4(\phi, X)R + G_{4,X}[(\square\phi)^2 - (\nabla_\mu\nabla_\nu\phi)(\nabla^\mu\nabla^\nu\phi)]$$

$$\begin{aligned} \mathcal{L}_5 = & G_5(\phi, X)G_{\mu\nu}(\nabla^\mu\nabla^\nu\phi) - \frac{1}{6}G_{5,X}[(\square\phi)^3 - 3(\square\phi)(\nabla_\mu\nabla_\nu\phi)(\nabla^\mu\nabla^\nu\phi) \\ & + 2(\nabla^\mu\nabla_\alpha\phi)(\nabla^\alpha\nabla_\beta\phi)(\nabla^\beta\nabla_\mu\phi)] \end{aligned}$$

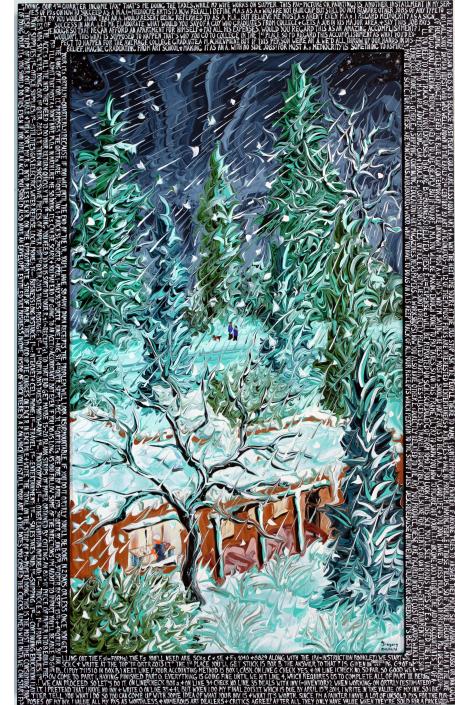


Image credit: Horndeski, Gregory W.

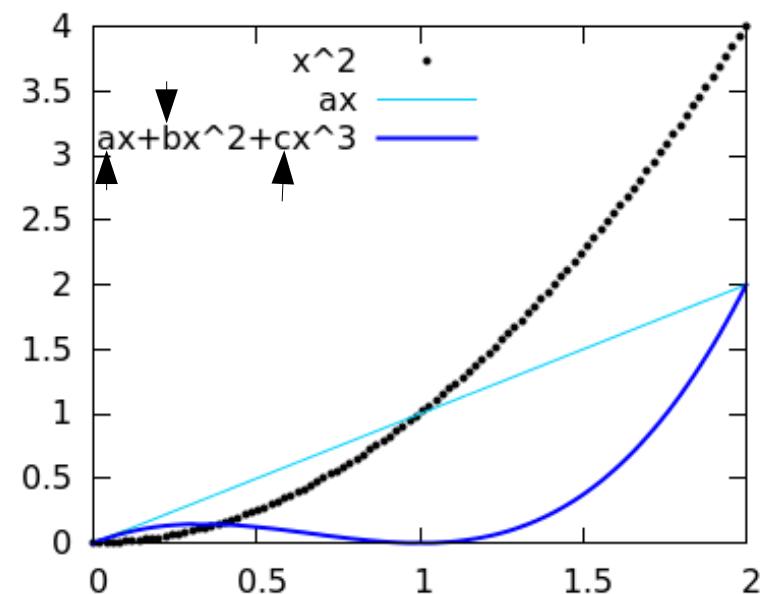
# The evidence

- Normalization constant in parameter inference
- **The** quantity for model comparison

$$P(\boldsymbol{\theta}_M | \mathbf{X}) = \frac{\mathcal{P}(\boldsymbol{\theta}_M) L(\mathbf{X} | \boldsymbol{\theta}_M)}{\varepsilon}$$

$$\varepsilon = L(\mathbf{X} | M_1)$$

$$\varepsilon = \int L(\mathbf{X} | \boldsymbol{\theta}_M) \mathcal{P}(\boldsymbol{\theta}_M) d^n \boldsymbol{\theta}$$



- It balances the goodness of fit against the number of parameters.  
'Occam's razor'.
  - It avoids (extreme) overfitting.

# Toy Model

$$\varepsilon = \int L(\mathbf{X} | \boldsymbol{\theta}_M) \mathcal{P}(\boldsymbol{\theta}_M) d^n\theta = L(\boldsymbol{\theta}') \mathcal{P}(\boldsymbol{\theta}') \Delta L$$

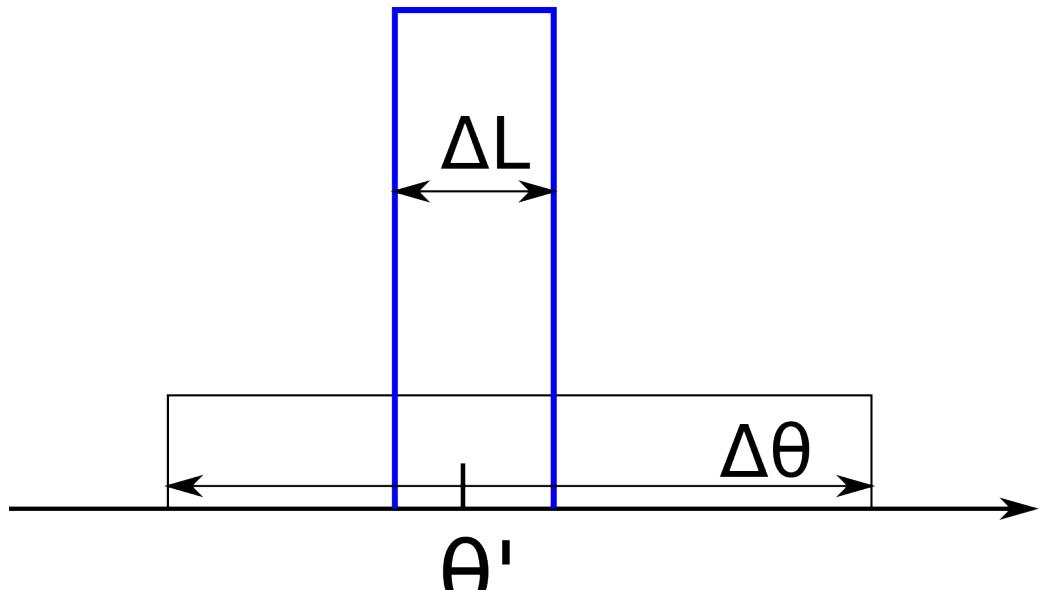
$$= L(\boldsymbol{\theta}') \frac{\Delta L}{\Delta \mathcal{P}}$$



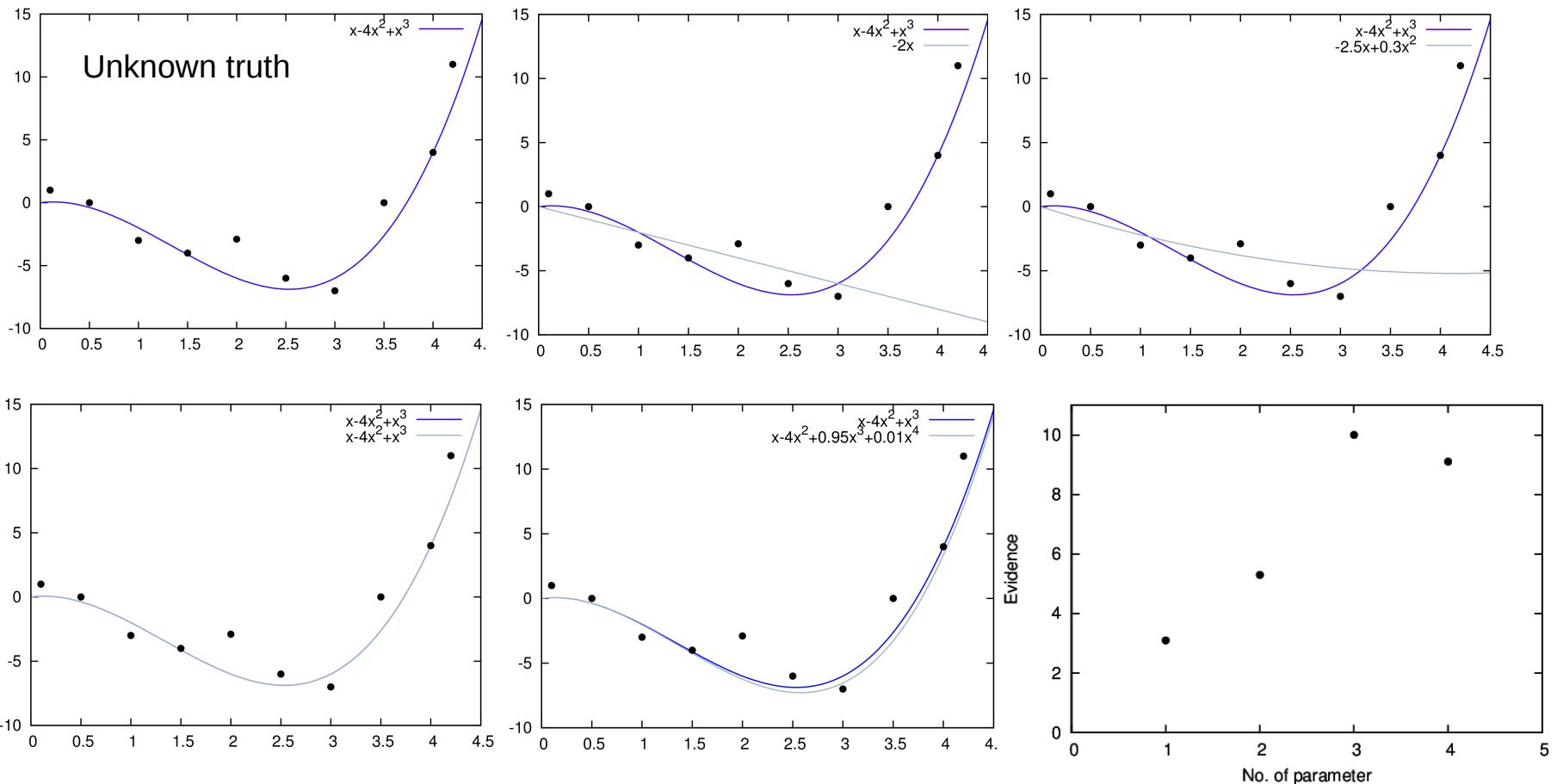
$$L \propto \exp(-\frac{1}{2}\chi^2)$$



Will always decrease with number of parameters.



# Polynomial example

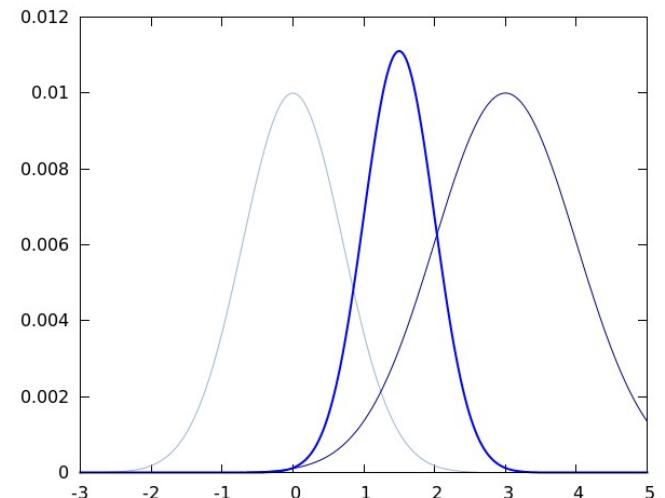


$$L(\theta') \frac{\Delta L}{\Delta \mathcal{P}}$$

$$L \propto \exp(-\frac{1}{2}\chi^2)$$

# A word on priors in $\varepsilon = \int L(\mathbf{X}|\boldsymbol{\theta}_M) \mathcal{P}(\boldsymbol{\theta}_M) d^n\theta$

- Theory or physics driven priors
  - $\Omega_m \in [0, 1]$ , Mass > 0
- Data driven priors & combination of experiments
  - Prior = old data
  - Likelihood = new data
  - Posterior = old and new data
- Subjective & informative priors
  - 'Only an unstated prior is a bad prior.'
- Objective & 'uninformative' priors
  - Maximize KL-divergence  $D_{\text{KL}}(P||Q) = \int_{-\infty}^{\infty} p(x) \log \frac{p(x)}{q(x)} dx$
  - Exploit symmetry groups: Haar-measures and invariant 'volumes'
  - Reparameterization independence (Jeffreys priors)  $\pi_{IJ}(\boldsymbol{\mu}, \boldsymbol{\Sigma}) = |\boldsymbol{\Sigma}|^{-(p+1)/2}$
- Frequentist matching priors  $\frac{d}{d\theta}(\pi(\theta) I^{-1/2}(\theta)) = 0$



# Model comparison

Have:  $\varepsilon = L(\mathbf{X} | M_1)$

Want:  $L(M_1 | \mathbf{X})$

Bayes' theorem:

$$L(M_1 | \mathbf{X}) = L(\mathbf{X} | M_1) \frac{\mathcal{P}(M_1)}{\mathcal{P}(\mathbf{X})}$$

?

# Model comparison

Get rid off the prior probability for the data by taking a ratio:

$$\frac{L(M_1 | \mathbf{X})}{L(M_2 | \mathbf{X})} = \frac{\mathcal{P}(M_1)L(\mathbf{X}|M_1)}{\mathcal{P}(M_2)L(\mathbf{X}|M_2)}$$

$$= \frac{\mathcal{P}(M_1)}{\mathcal{P}(M_2)} \frac{\varepsilon_1}{\varepsilon_2}$$

→ Bayes factor:  $> 1$  prefers  $M_1$   
 $< 1$  prefers  $M_2$

Where:

$$\varepsilon = \int L(\mathbf{X} | \boldsymbol{\theta}_M) \mathcal{P}(\boldsymbol{\theta}_M) d^n \theta$$

# Magnitude of B

- Bayes factor = evidence<sub>1</sub>/evidence<sub>2</sub>.
- Without loss of generality:  $\epsilon_1 = b\epsilon_2$

- Then:

$$B_{12} = \frac{1}{b} \quad \text{and} \quad B_{21} = \frac{b}{1}$$

decisiveness asymptotes to zero      vs.      decisiveness grows linearly

- Ergo: Introduce ln for measure of decisiveness:

$$\ln(B_{12}) = \ln(1) - \ln(b)$$

$$\ln(B_{21}) = \ln(b) \quad \rightarrow \text{now } B_{12} \text{ and } B_{21} \text{ are treated equally}$$

# Calibration on the Jeffreys scale

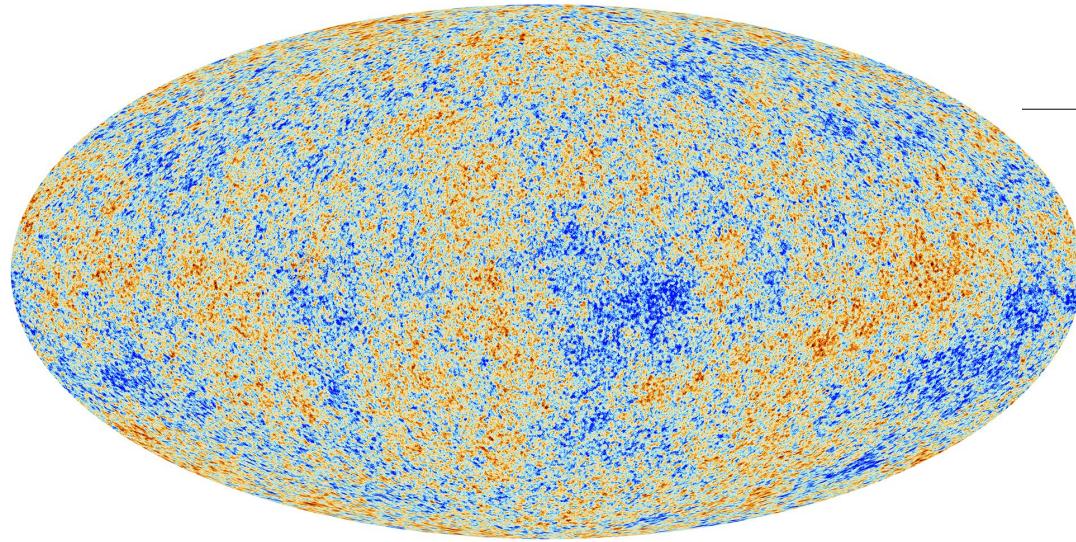
Table 6.1: Jeffreys scale

$ \log\left(\frac{\varepsilon(M_1)}{\varepsilon(M_2)}\right) $	odds	interpretation	prob. of favoured model
$\leq 1.0$	3:1	better data is needed	$\leq 0.75$
$\leq 2.5$	12:1	weak evidence	0.923
$\leq 5.0$	$\leq 150:1$	moderate evidence	0.993
$\geq 5.0$	$> 150:1$	strong evidence	$> 0.993$

## Example:

- Dark Energy Survey (DES) SV data
- WL analysis: flat LCDM vs. LCDM + curvature
- $\pi(\Omega_k) = \text{uniform}[-0.2, 0.2]$
- $\ln(B) = 0.17 \pm 0.09$     Sellentin & Heavens (2016)

# Model selection in the CMB



CMB = photons, in gravitational potentials of all particle species

$$\gamma + p + n + e + DM + 3\xi_{rel}$$

Planck :  $\nu!$

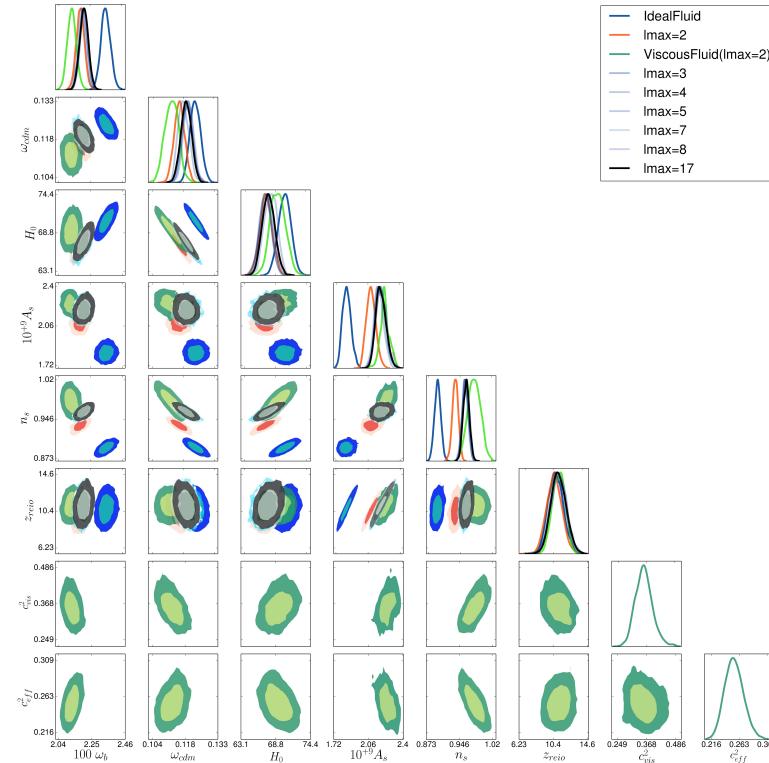
But are these neutrinos? Or just any relativistic fluid?

## Model comparison:

Neutrinos vs. ideal fluid:  $\ln(B) \approx 10$

Neutrinos vs. viscous fluid:  $\ln(B) \approx 10.5$

+ parameter constraints as a side effect

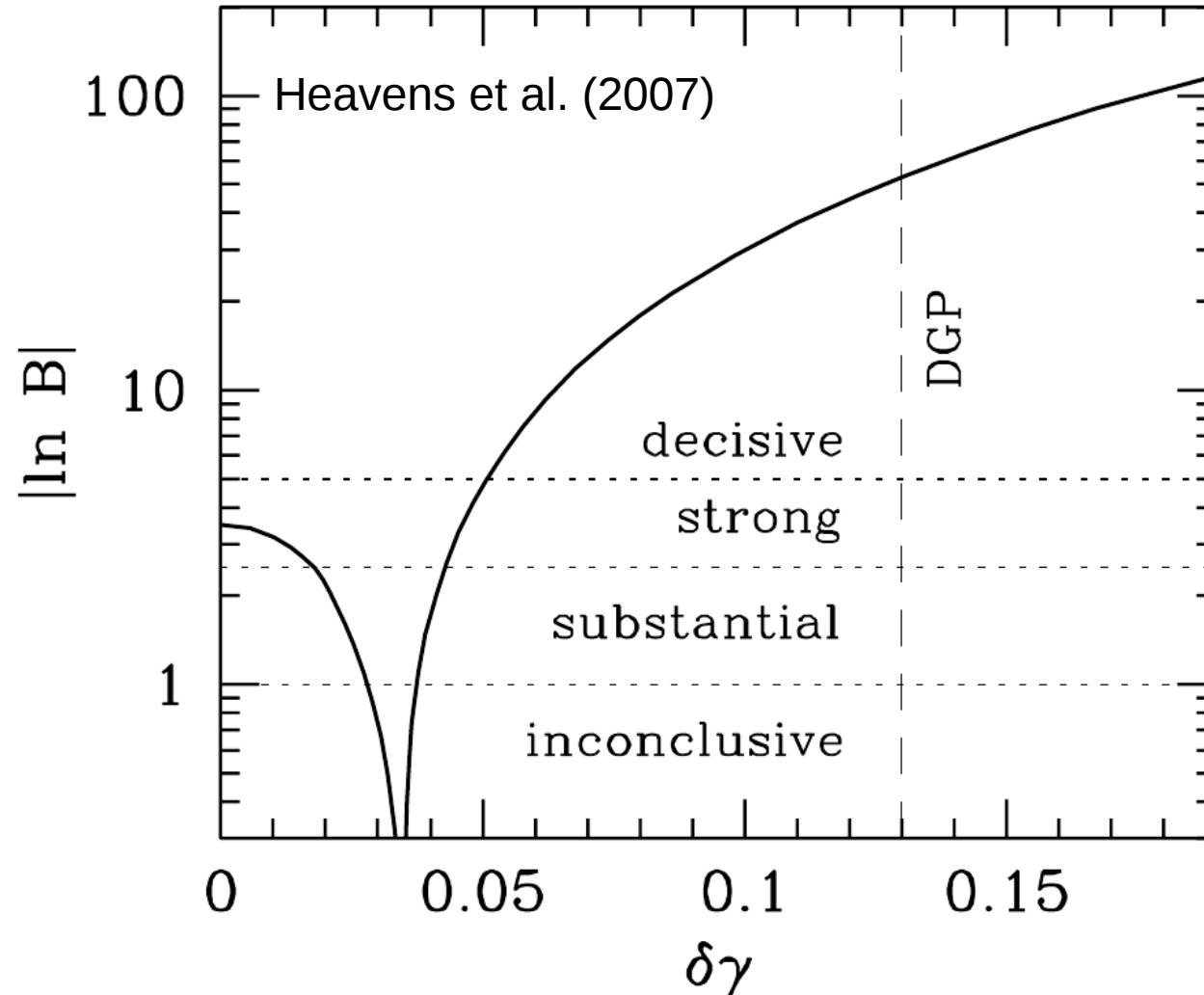


# Expected support for models

- Single data realization:  $B_{01} = \frac{\epsilon_0}{\epsilon_1} = \frac{\int L_0(\vec{x}|\vec{\theta}_{M_0})\mathcal{P}_0(\vec{\theta}_{M_0})d^{n_0}\theta}{\int L_1(\vec{x}|\vec{\theta}_{M_1})\mathcal{P}_1(\vec{\theta}_{M_1})d^{n_1}\theta}$
- Know statistical properties of data → calculate expected likelihood (even without having real data at all)

$$B_{01}^{exp} = \frac{\int \langle L_0(\vec{x}|\vec{\theta}_{M_0}) \rangle \mathcal{P}_0(\vec{\theta}_{M_0}) d^{n_0}\theta}{\int \langle L_1(\vec{x}|\vec{\theta}_{M_1}) \rangle \mathcal{P}_1(\vec{\theta}_{M_1}) d^{n_1}\theta}$$

# Expected support for models



$$M_0: g(a) = \exp \left\{ \int_0^a \frac{da'}{a'} [\Omega_m(a')^\gamma - 1] \right\} \quad M_1: \Omega_m^{\gamma + \delta\gamma}$$

# Nested Models

- Imagine  $M_1$  uses all parameters  $\vec{\theta}$  of  $M_0$  but introduces some extra parameters  $\vec{\psi}$
- **Nested model:** for  $\vec{\psi} = \vec{\psi}_0$  have  $M_1 \rightarrow M_0$
- Examples:
  - wCDM  $\rightarrow$  LambdaCDM for  $w = -1$
  - Curved LambdaCDM  $\rightarrow$  flat LambdaCDM for  $k = 0$
  - Rainy day  $\rightarrow$  sunny day for rain = 0

# Savage-Dickey Density Ratio

- SDDR is an approximate Bayes factor for nested models
- The full Bayes factor is  $B_{01} = \frac{\epsilon_0}{\epsilon_1} = \frac{\int L_0(\vec{x}|\vec{\theta}_{M_0})\mathcal{P}_0(\vec{\theta}_{M_0})d^{n_0}\theta}{\int L_1(\vec{x}|\vec{\theta}_{M_1})\mathcal{P}_1(\vec{\theta}_{M_1})d^{n_1}\theta}$
- For nested models:  $L_0(\vec{x}|\vec{\theta}_{M_0}) = L_1(\vec{x}|\vec{\theta}_{M_0}, \vec{\psi} = \vec{\psi}_0)$
- Insert into Bayes factor:
$$B_{01} = \frac{\int L_1(\vec{x}|\vec{\theta}_{M_0}, \vec{\psi} = \vec{\psi}_0)\mathcal{P}_0(\vec{\theta}_{M_0})d^{n_0}\theta}{\int L_1(\vec{x}|\vec{\theta}_{M_0}, \vec{\psi})\mathcal{P}_1(\vec{\theta}_{M_0}, \vec{\psi})d^{n_0}\theta d^n\psi}$$
- Now need to care about the priors.

# Savage-Dickey Density Ratio

- Bayes factor:

$$B_{01} = \frac{\int L_1(\vec{x}|\vec{\theta}_{M_0}, \vec{\psi} = \vec{\psi}_0) \mathcal{P}_0(\vec{\theta}_{M_0}) d^{n_0} \theta}{\int L_1(\vec{x}|\vec{\theta}_{M_0}, \vec{\psi}) \mathcal{P}_1(\vec{\theta}_{M_0}, \vec{\psi}) d^{n_0} \theta d^n \psi}$$

- Make **extra assumption** for priors:  $\mathcal{P}_1(\vec{\theta}_{M_0}|\vec{\psi} = \vec{\psi}_0) = a \mathcal{P}_0(\vec{\theta}_{M_0})$
- Insert into Bayes factor:

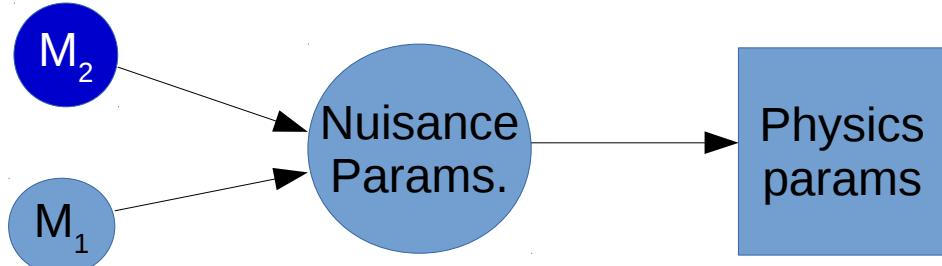
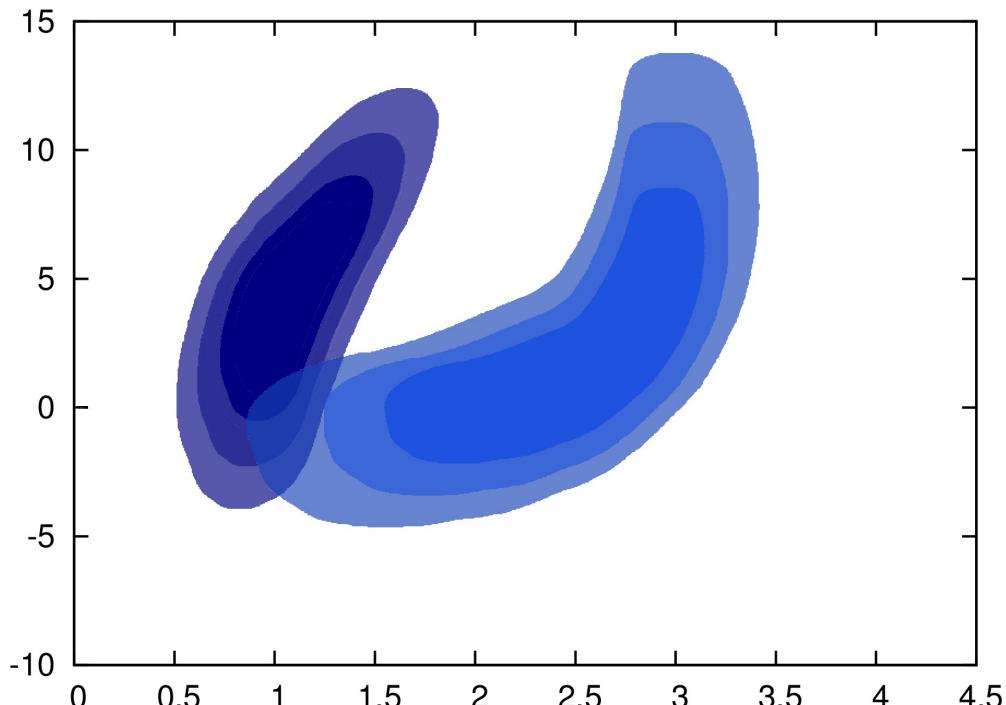
$$B_{01} \approx a \frac{\int L_1(\vec{x}|\vec{\theta}_{M_0}, \vec{\psi} = \vec{\psi}_0) \mathcal{P}_1(\vec{\theta}_{M_0}|\psi = \psi_0) d^{n_0} \theta}{\int L_1(\vec{x}|\vec{\theta}_{M_0}, \vec{\psi}) \mathcal{P}_1(\vec{\theta}_{M_0}, \vec{\psi}) d^{n_0} \theta d^n \psi}$$

- Leading to the Savage-Dickey Density Ratio:  $B_{01} \approx a \frac{P_1(\vec{\psi} = \vec{\psi}_0|\vec{x})}{P_1(\vec{\psi} = \vec{\psi}_0)}$

**Example** from Dirian et al.(2016):  $B_{\Lambda(\Lambda+i)} \equiv \frac{P(d|\mathcal{M}_\Lambda)}{P(d|\mathcal{M}_{\Lambda+i})} = \frac{P(\Omega_{X_i} = 0|d, \mathcal{M}_{\Lambda+i})}{P(\Omega_{X_i} = 0|\mathcal{M}_{\Lambda+i})}$

- Plan ahead, use Nested Sampling not MCMC to get B + param. constraints
- If too late: MCMC+SDDR+importance sampling approximate B (**excercise**)

# Model averaging



- Imagine two models explain the same effect. None is 'better' than the other, as given by B.
- Weak lensing: Intrinsic alignment model?
- Structure formation: Press-Schechter mass function or Sheth-Torman or Jenkins et al. or...?

$$P(\vec{\theta}|\vec{x}) \propto \sum_i P(\vec{\theta}|\vec{x}, M_i) P(M_i|\vec{x})$$

- Includes model uncertainty into parameter uncertainty.

# Summary

- Bayesians compare models by evidence ratios
- Balance goodness of fit against number of parameters
- Samplers exist that give parameter constraints and evidences ( $\rightarrow$  JP's lecture)
- Savage-Dickey Density Ratio may or may not be of relevance to you in case of nested models...
- ... depending on your attitude towards priors (subjective/objective).
- Model comparison is prior dependent.