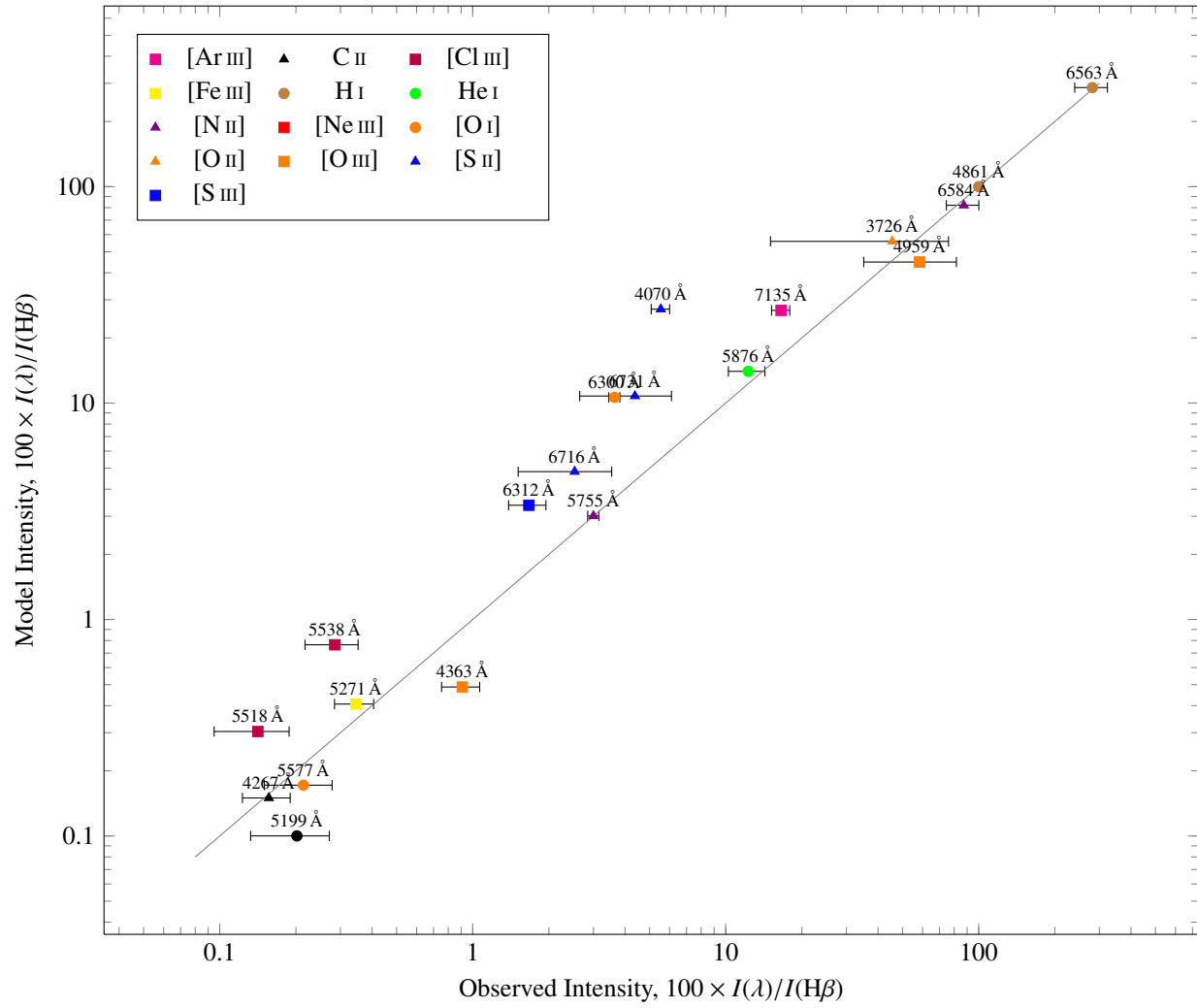


Model A: Esteban

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 12.10$

Abundance set

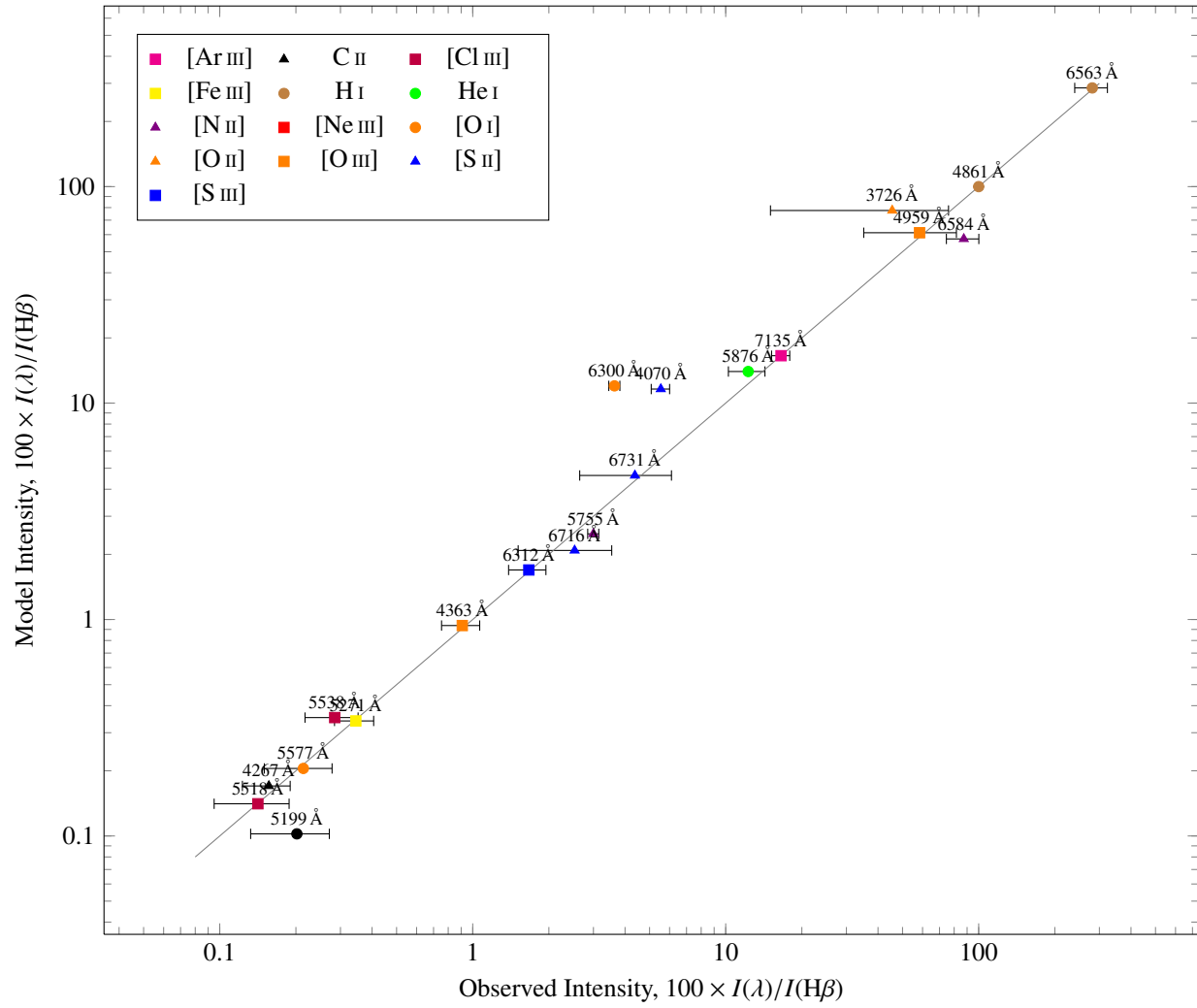


Model Q: Tsamis final

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 12.10$

Abundance set

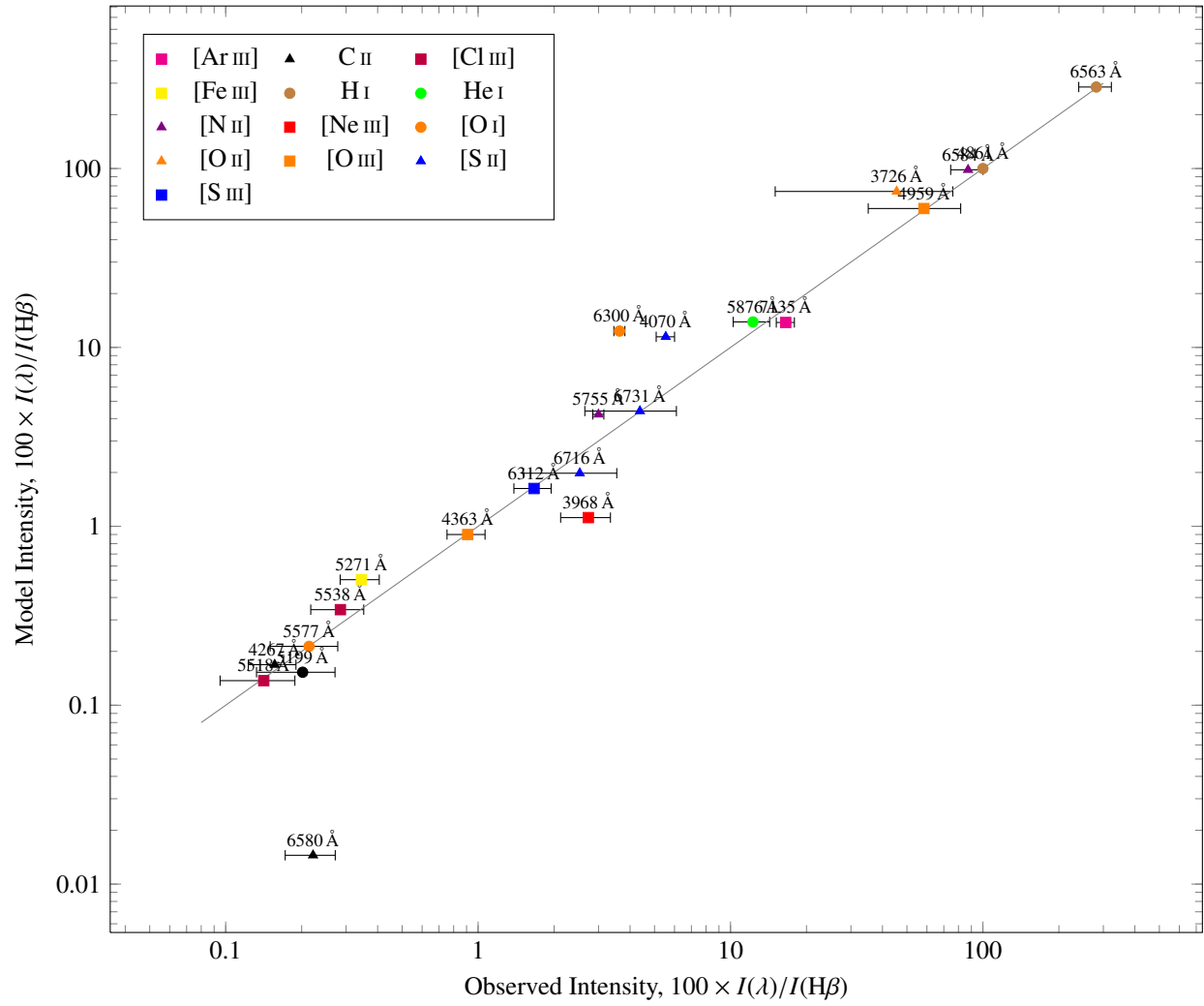


Model R:

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 12.10$

Abundance set

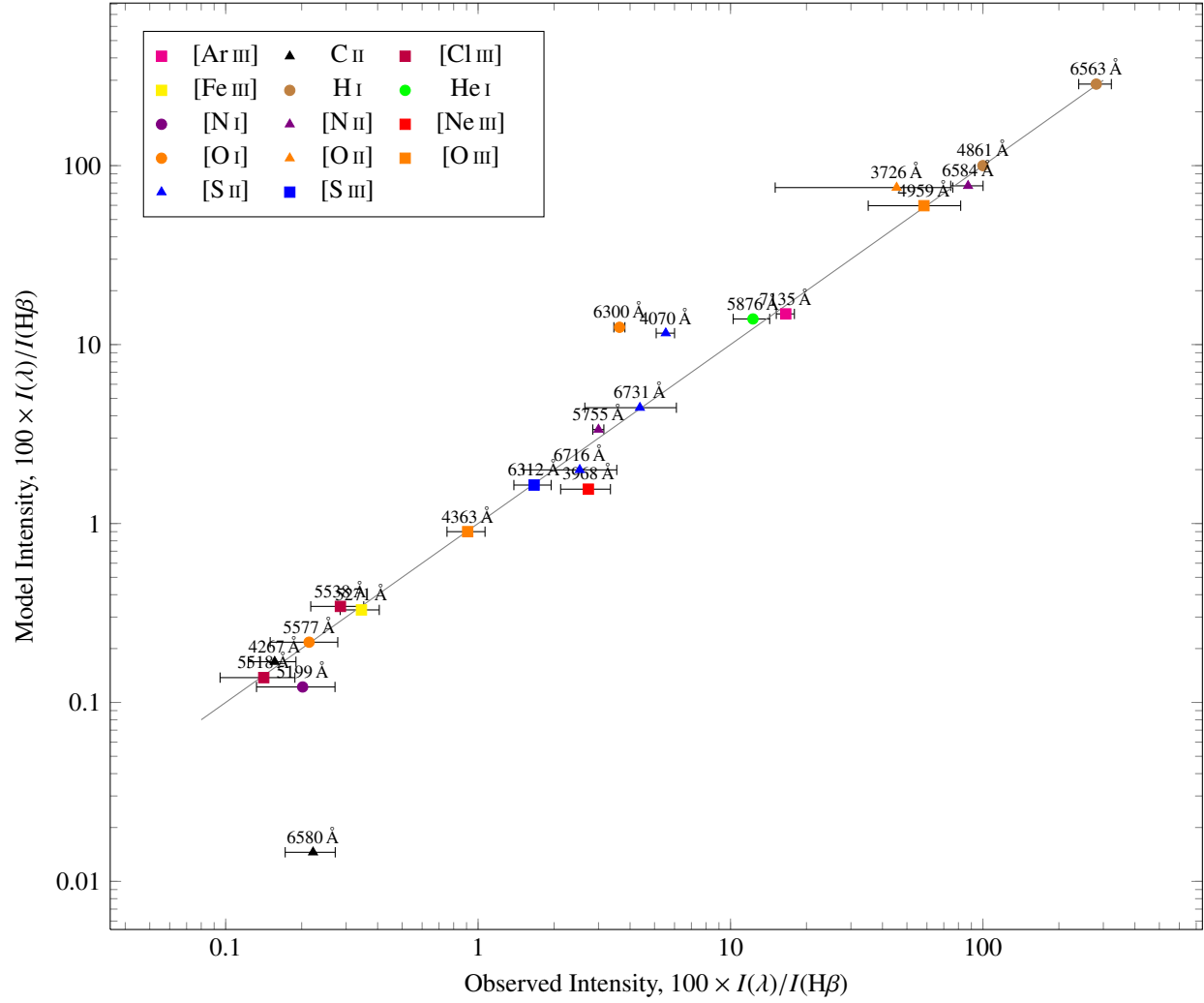


Model S:

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 12.10$

Abundance set

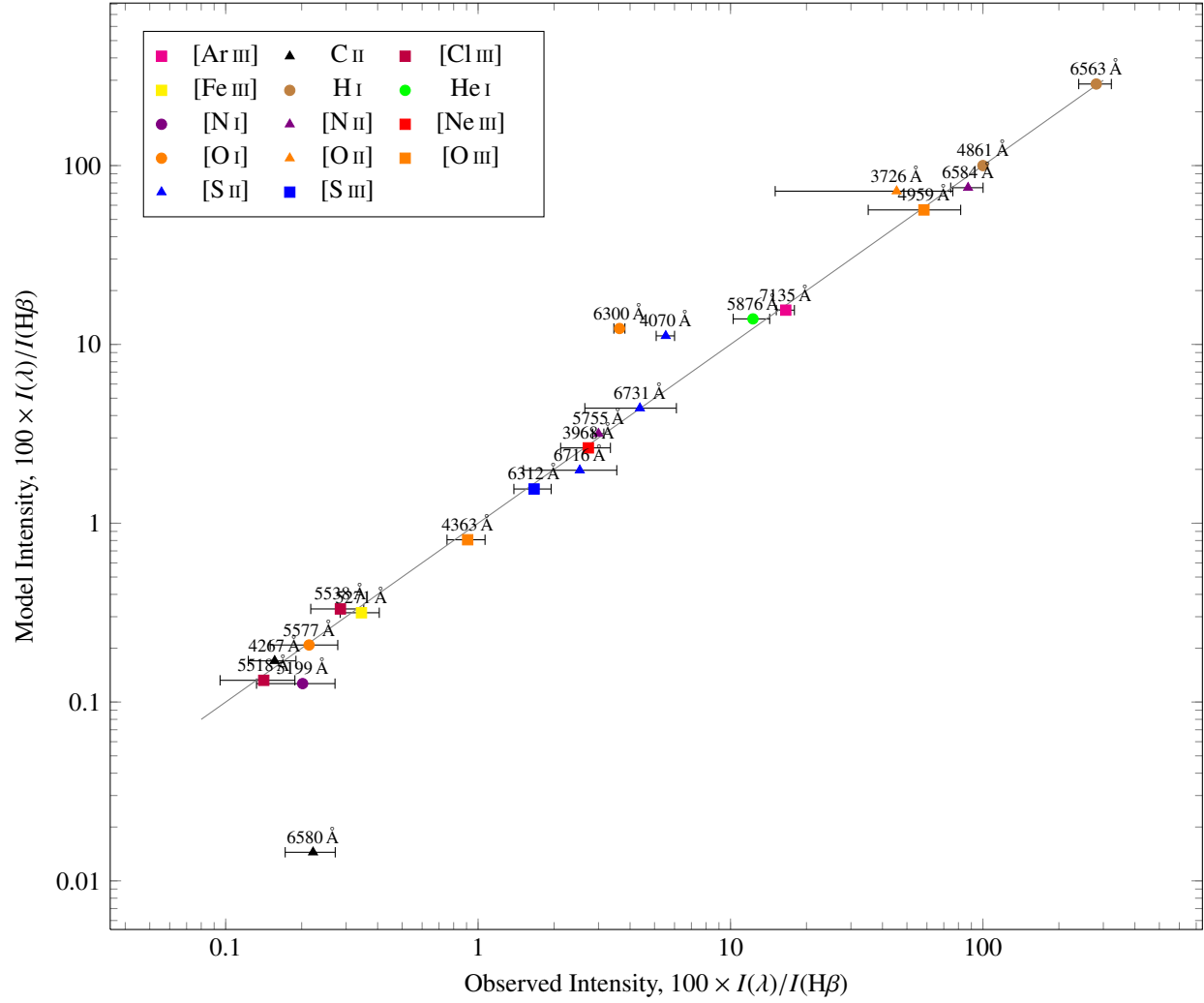


Model T:

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 12.10$

Abundance set



Model U:

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 12.10$

Abundance set

