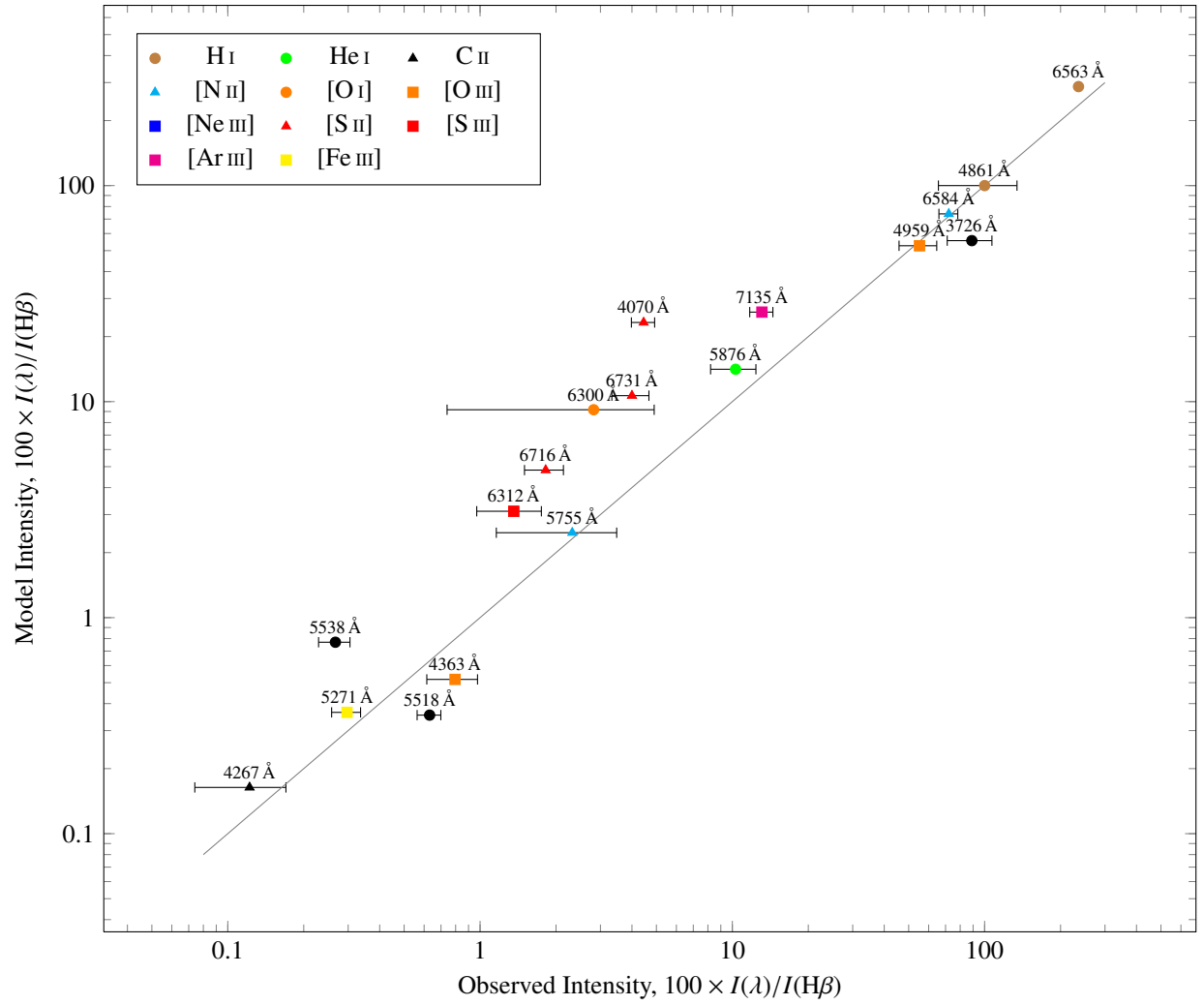# Model A: Total Flux

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 12.10$

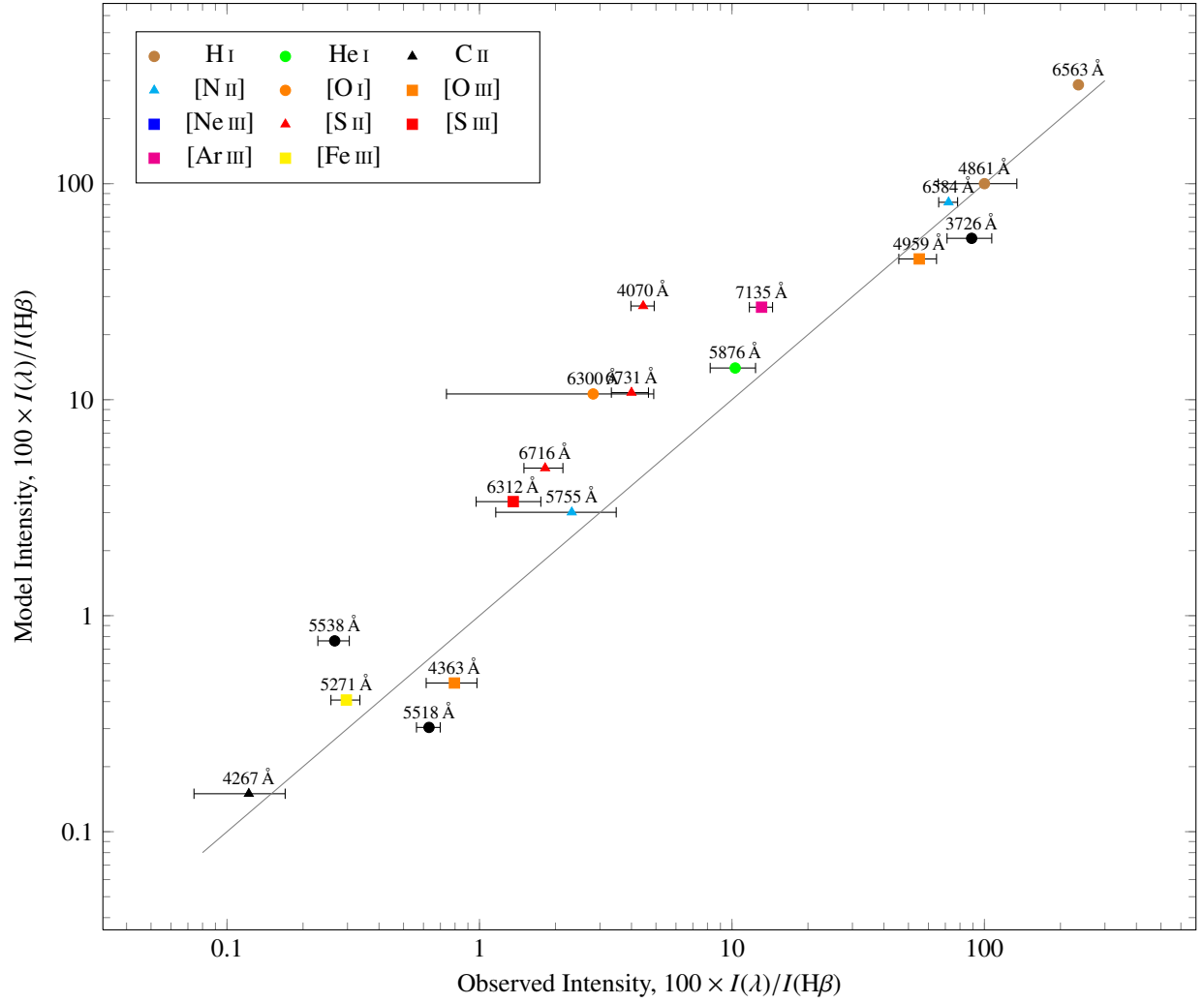**Abundance set** Esteban et al. (2004), M42, $t^2 = 0.002$

Model B: 1.6 x 1.6 aperture, c=(1.0, 0.0)

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 12.10$

Abundance set Esteban et al. (2004), M42, $t^2 = 0.002$
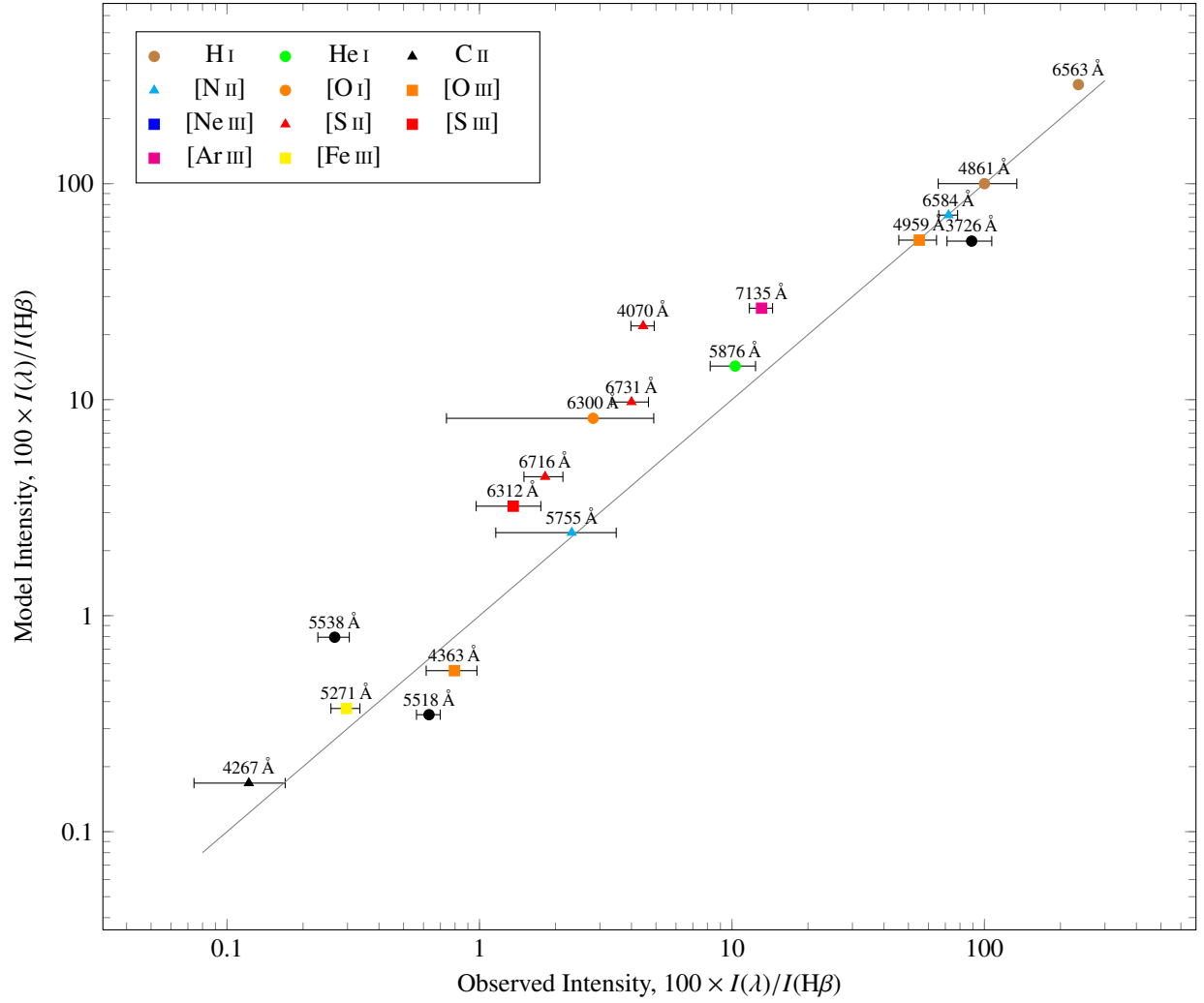
Model C: 1.6 x 1.6 aperture, c=(1.5, 0.0)

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 12.10$

Abundance set Esteban et al. (2004), M42, $t^2 = 0.002$

Model D: 1.6 x 1.6 aperture, c=(2.0, 0.0)

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 12.10$

Abundance set Esteban et al. (2004), M42, $t^2 = 0.002$