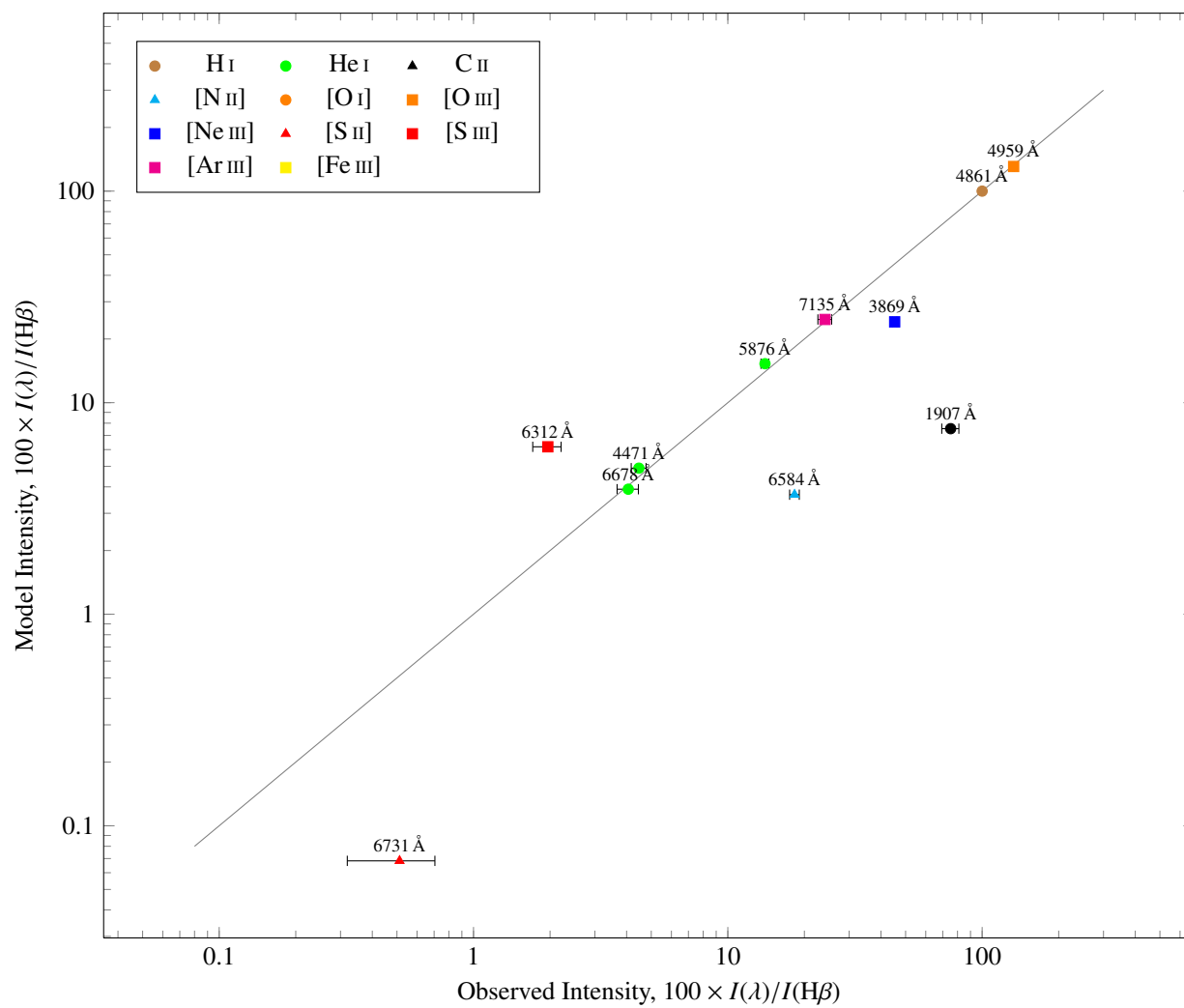


Model A: Baseline model, FOS obs

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 14.54$

Abundance set Cloudy Orion

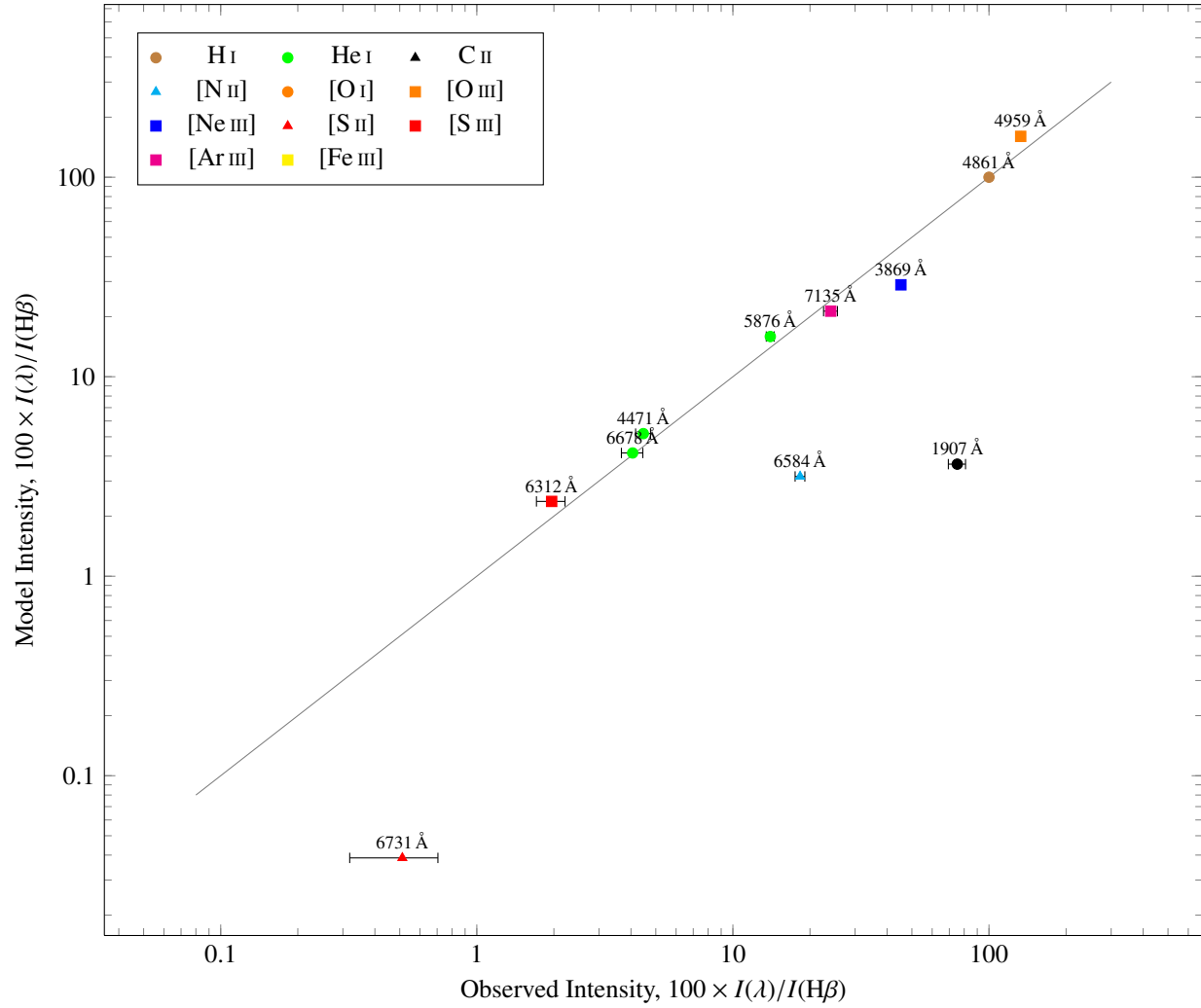


Model B: Tsamis abundances, FOS obs

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 14.54$

Abundance set Tsamis paper

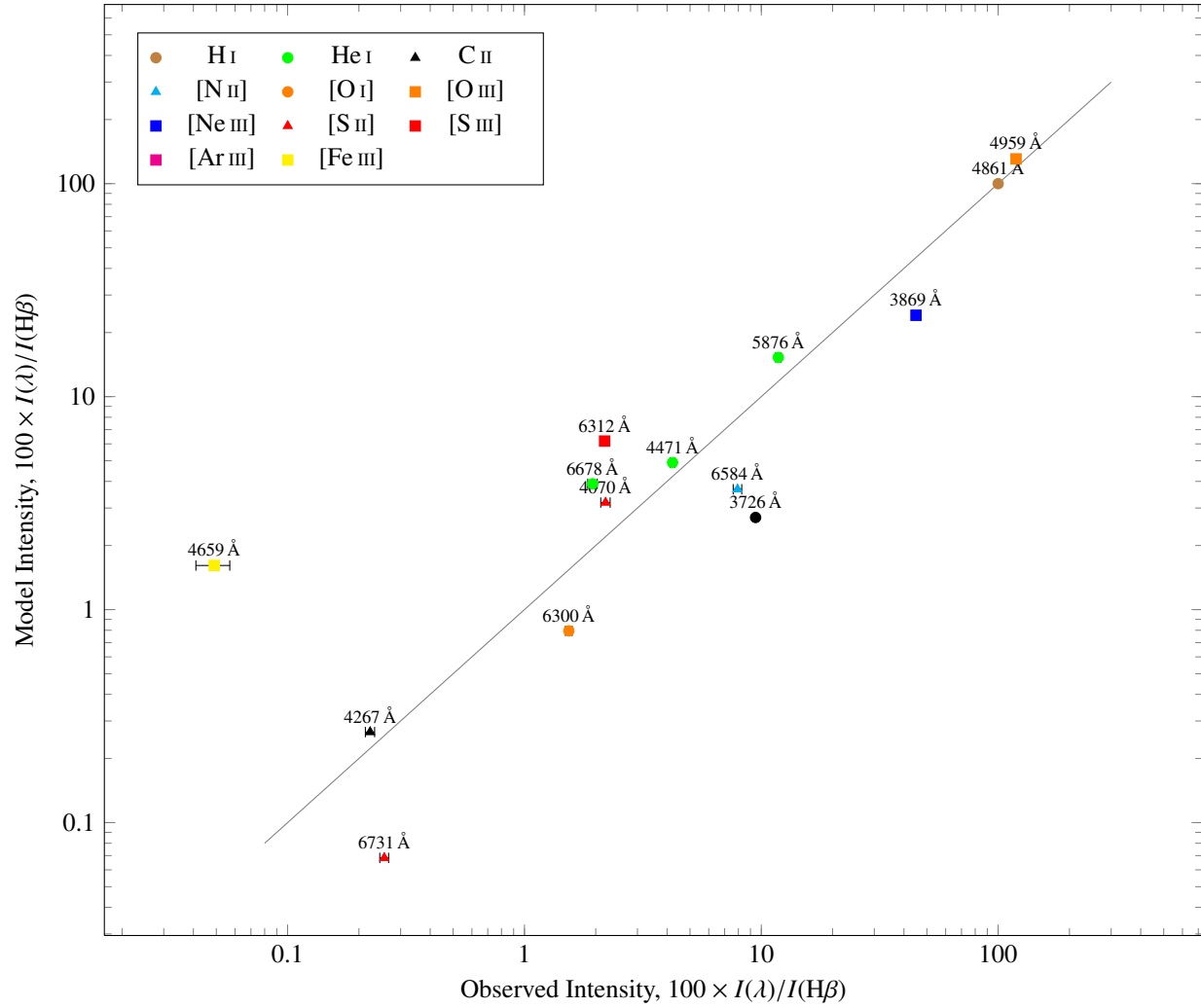


Model A: Baseline model, VLT obs

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 14.54$

Abundance set Cloudy Orion



Model B: Tsamis abundances, VLT obs

Spectrum WMBasic, 39 000 K

Flux $\log_{10} \Phi = 14.54$

Abundance set Tsamis paper

