

## The variance from partition fens

a) Note;

$$(1) \quad Z = \sum_s e^{-\beta \epsilon_s}$$

$$(2) \quad - \frac{\partial Z}{\partial \beta} = \sum_s e^{-\beta \epsilon_s} \epsilon_s$$

$$(3) \quad + \frac{\partial^2 Z}{\partial \beta^2} = \sum_s e^{-\beta \epsilon_s} \epsilon_s^2$$

So since  $P_s = e^{-\beta \epsilon_s} / Z$

$$\langle \epsilon \rangle = \sum_s P_s \epsilon_s = \frac{1}{Z} - \frac{\partial Z}{\partial \beta} = - \frac{\partial}{\partial \beta} \log Z$$

$$\langle \epsilon^2 \rangle = \sum_s P_s \epsilon_s^2 = \frac{1}{Z} \frac{\partial^2 Z}{\partial \beta^2}$$

b) Now we have from (a)

$$\langle \epsilon^2 \rangle - \langle \epsilon \rangle^2 = \frac{1}{Z} \frac{\partial^2 Z}{\partial \beta^2} - \left( \frac{1}{Z} \frac{\partial Z}{\partial \beta} \right)^2$$

So we have to check that

these are the same

$$\frac{\partial^2}{\partial \beta^2} \log Z = \frac{\partial}{\partial \beta} \left( \frac{1}{Z} \frac{\partial Z}{\partial \beta} \right) = - \frac{1}{Z^2} \left( \frac{\partial Z}{\partial \beta} \right)^2 + \frac{1}{Z} \frac{\partial^2 Z}{\partial \beta^2}$$

$$\text{So } \langle \epsilon^2 \rangle - \langle \epsilon \rangle^2 = \frac{\partial^2}{\partial \beta^2} \log Z$$

Now since

$$\langle \epsilon \rangle = - \frac{\partial \log Z}{\partial \beta}$$

$$- \frac{\partial \langle \epsilon \rangle}{\partial \beta} = \frac{\partial^2 \log Z}{\partial \beta^2} = \langle \epsilon^2 \rangle - \langle \epsilon \rangle^2 \quad \checkmark$$

(c) We have from last week the mean energy

$$\langle \epsilon \rangle = \frac{h\omega}{e^{\beta h\omega} - 1}$$

Differentiating:

$$\langle \epsilon^2 \rangle - \langle \epsilon \rangle^2 = - \frac{\partial \langle \epsilon \rangle}{\partial \beta} = - \frac{\partial}{\partial \beta} \frac{h\omega}{e^{\beta h\omega} - 1}$$

$$= \frac{h\omega e^{\beta h\omega}}{(e^{\beta h\omega} - 1)^2} h\omega$$

$$= (h\omega)^2 \frac{e^{\beta h\omega}}{(e^{\beta h\omega} - 1)^2} = (h\omega)^2 \frac{e^{-\beta h\omega}}{(1 - e^{-\beta h\omega})^2}$$

(d) At low temperatures  $e^{-\beta h\omega} \ll 1$ . So in this expression we may approximate

$$\langle \epsilon^2 \rangle - \langle \epsilon \rangle^2 \approx (h\omega)^2 e^{-\beta h\omega}$$

e)

The probability of having  $n$  quanta is

$$P_n = \frac{e^{-n\beta\hbar\omega}}{Z} = e^{-\beta n\hbar\omega} (1 - e^{-\beta\hbar\omega})$$

Thus for  $e^{-\beta\hbar\omega} \ll 1$  we have:

- $P_0 \approx (1 - e^{-\beta\hbar\omega}) \approx 1 - \text{small}$

- $P_1 \approx e^{-\beta\hbar\omega} (1 + O(e^{-\beta\hbar\omega})) \approx \text{small}$

- $P_2 = (e^{-\beta\hbar\omega})^2 (1 + O(e^{-\beta\hbar\omega})) \approx 0$

↑ this is  $(\text{small})^2$  and can be dropped  
Similarly,  $P_3 \approx e^{-3\beta\hbar\omega} \approx (e^{-\beta\hbar\omega})^3 \approx 0$

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$$\langle \mathcal{E} \rangle = P_0 \cdot 0 + P_1 \hbar\omega = e^{-\beta\hbar\omega} \hbar\omega$$

$$\langle \mathcal{E}^2 \rangle = P_0 \cdot 0 + P_1 (\hbar\omega)^2 = e^{-\beta\hbar\omega} (\hbar\omega)^2$$

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$$\langle \mathcal{E}^2 \rangle - \langle \mathcal{E} \rangle^2 = e^{-\beta\hbar\omega} (\hbar\omega)^2 + O(e^{-2\beta\hbar\omega})$$

$$\approx e^{-\beta\hbar\omega} (\hbar\omega)^2$$

↑

This agrees with part d)

## Einstein Solid

- a) There are  $3N$  oscillators. The Hamiltonian of the oscillator is

$$H = \frac{p^2}{2m} + \frac{1}{2} m \omega_0^2 x^2$$

two quadratic forms

So,  $\bar{E} = 3N \times 2 \times \frac{1}{2} kT \approx 3NkT$

Now  $C_V = \partial E / \partial T$ .

So for  $N = 1 \text{ mol}$

$N_A k_B = R$ , and so

$$C_V^{1\text{mol}} = 3R$$

b)  $E = 3N \frac{\hbar \omega}{e^{\beta \hbar \omega} - 1}$        $\bar{n} = \frac{1}{e^{\beta \hbar \omega} - 1}$

- c) Solving for  $\beta$  from  $\bar{n}$  (to make contact with later work in the next problem)

$$\frac{1}{\bar{n}} = e^{\beta \hbar \omega} - 1 \Rightarrow e^{\beta \hbar \omega} = 1 + \frac{1}{\bar{n}}$$

So

$$\beta \hbar \omega = \ln \left( \frac{1 + \bar{n}}{\bar{n}} \right) \Rightarrow \frac{1}{kT} = \frac{1}{\hbar \omega} \ln \left( \frac{1 + \bar{n}}{\bar{n}} \right)$$

d) We know that

$$C_v = \left( \frac{\partial E}{\partial T} \right) \quad \text{with} \quad E = 3N \frac{\hbar \omega}{e^{\beta \hbar \omega} - 1}$$

We use that for any quantity  $X$

$$\frac{\partial X}{\partial T} = \frac{\partial X}{\partial \beta} \frac{\partial \beta}{\partial T} = \frac{\partial X}{\partial \beta} \frac{\partial}{\partial T} \left( \frac{1}{kT} \right)$$

$$\frac{\partial X}{\partial T} = -k\beta^2 \frac{\partial X}{\partial \beta}$$

So

$$C_v = -k\beta^2 \frac{\partial}{\partial \beta} \left( \frac{\hbar \omega}{e^{\beta \hbar \omega} - 1} \right)$$

Differentiating

$$C_v = 3Nk\beta^2 \frac{(\hbar \omega)^2 e^{\beta \hbar \omega}}{(e^{\beta \hbar \omega} - 1)^2}$$

For 1 mol  $N = N_A$   $N_A k_B = R$

$$C_v^{1\text{mol}} = 3R \frac{(\beta \hbar \omega)^2 e^{\beta \hbar \omega}}{(e^{\beta \hbar \omega} - 1)^2}$$



In the high temperature limit the mean # of quanta in the oscillator gets larger and larger. In this limit the dynamics should be classical. (see previous homework)

Expanding for  $\beta\hbar\omega \ll 1$  (high temperature), we approximate:

$$e^{\beta\hbar\omega} \approx 1 + \beta\hbar\omega$$

And

$$C_v^{lm} \approx \frac{3R(\beta\hbar\omega)^2}{(1 + \beta\hbar\omega - 1)^2}$$

$$C_v^{lm} \approx 3R$$

↑ this agrees with part (a) as it should

e) The hard materials have a larger  $\omega_0 = \sqrt{k/m}$ . They have, therefore, a larger spring constant,  $k$ . Because  $\omega_0$  is higher for diamond  $C_v$  will approach the classical limit  $3R$  only at very high temperatures  $kT \gg \hbar\omega_0$ , when the number of vibrational quanta is large.

## Einstein Model : Microcanonical Ensemble

$$\boxed{e)} \quad \Omega = \frac{(N+q-1)!}{q! (N-1)!} \sim \frac{(N+q)!}{q! N!} \quad \text{Using Sterling}$$

$$\ln \Omega \approx (N+q) \ln(N+q) - \cancel{(N+q)} - q \ln q + \cancel{q} - N \ln N + \cancel{N}$$

we note that the mean occupation number is  
 $\bar{n} \equiv q/N$

$$\ln \Omega = N \left[ (1+\bar{n}) \ln N(1+\bar{n}) - \bar{n} \ln N\bar{n} - \ln N \right]$$

Looking at this carefully we collect all the  $\ln N$  terms and the remaining;

$$\ln \Omega = N \left[ \cancel{(1+\bar{n}) \ln N} - \cancel{\bar{n} \ln N} - \cancel{\ln N} \right] \quad \text{cancel}$$
$$+ N \left[ (1+\bar{n}) \ln(1+\bar{n}) - \bar{n} \ln \bar{n} \right]$$

so

$$\Omega = e^{N \left[ (1+\bar{n}) \ln(1+\bar{n}) - \bar{n} \ln \bar{n} \right]}$$

$$S = k \ln \Omega = Nk \left[ (1+\bar{n}) \ln(1+\bar{n}) - \bar{n} \ln \bar{n} \right]$$

So

$$f) \frac{1}{T} = \frac{\partial S}{\partial E} = \frac{\partial S}{\partial \bar{n}} \frac{\partial \bar{n}}{\partial E}$$

$$E = N \bar{n} \hbar \omega_0$$

$$\frac{\partial E}{\partial \bar{n}} = N \hbar \omega_0$$

$$= \frac{1}{N \hbar \omega_0} \frac{\partial S}{\partial \bar{n}}$$

So

$$\frac{\partial S}{\partial \bar{n}} = Nk \left[ \ln(1 + \bar{n}) + (1 + \bar{n}) \frac{1}{(1 + \bar{n})} - \ln \bar{n} - \bar{n} \frac{1}{\bar{n}} \right]$$

$$= Nk \ln \left( \frac{1 + \bar{n}}{\bar{n}} \right)$$

So

$$\frac{1}{T} = \frac{k}{\hbar \omega_0} \ln \left( \frac{1 + \bar{n}}{\bar{n}} \right)$$

g) We can invert this

$$\frac{\hbar \omega_0}{kT} = \ln \left( \frac{1 + \bar{n}}{\bar{n}} \right) \Rightarrow e^{\hbar \omega_0 / kT} = 1 + \frac{1}{\bar{n}}$$

$$\bar{n} = \frac{1}{e^{\hbar \omega_0 / kT} - 1}$$

← this agrees with the canonical approach discussed above.