Physics 306: Thermal Physics

Midterm Exam Stony Brook University

Spring 2024

General Instructions:

You may use one page (front and back) of handwritten notes and a calculator. Graphing calculators are allowed. No other materials may be used.

1 Integrals

Gamma Function:

$$\Gamma(z) \equiv \int_0^\infty x^{z-1} e^{-x} \mathrm{d}x \tag{1}$$

with specific results

$$\Gamma(z+1) = z\Gamma(z)$$
 $\Gamma(n) = (n-1)!$ $\Gamma(\frac{1}{2}) = \sqrt{\pi}$ (2)

Gaussian Integrals:

$$I_n = \frac{1}{\sqrt{2\pi}\sigma} \int_{-\infty}^{\infty} \mathrm{d}x \, e^{-x^2/2\sigma^2} x^n \tag{3}$$

with specific results

$$I_0 = 1 \quad I_2 = \sigma^2 \quad I_4 = 3\sigma^4 \quad I_6 = 15\sigma^6$$
 (4)

Problem 1. Quick sums

Using techniques from the quantum harmonic oscillator, evaluate the following sums in explicit form:

$$S_0 = \sum_{n=0}^{\infty} q^n$$
 $S_1 = \sum_{n=0}^{\infty} nq^n$. (5)

Here q is a real number that is less than unity and greater than 0.

Solution

The partition function of the harmonic oscillator are these sums. In the harmonic oscillator the energy levels are

$$\epsilon_n = n\hbar\omega_0 \tag{6}$$

Thus partition function is the geometric series with $q \equiv e^{-\beta\hbar\omega_0}$:

$$Z = \sum_{n} e^{-n\beta\hbar\omega_0} = 1 + (e^{-\beta\hbar\omega_0}) + (e^{-\beta\hbar\omega_0})^2 + \dots = \frac{1}{1 - e^{-\beta\hbar\omega_0}}$$
 (7)

This implies

$$S_0 = 1 + q + q^2 + \dots = \frac{1}{1 - q}$$
 (8)

i.e. the geometric series.

The mean energy of the harmonic oscillator is:

$$\frac{\langle \epsilon \rangle}{\hbar \omega_0} = \frac{1}{Z} \sum_{n} n e^{-n\beta\hbar\omega_0} = \frac{e^{-\beta\hbar\omega_0}}{1 - e^{-\beta\hbar\omega_0}} \tag{9}$$

This implies

$$\frac{\sum_{n} nq^{n}}{\sum_{n} q^{n}} = \frac{q}{1-q} \tag{10}$$

So using the first sum we find

$$\sum_{n} nq^{n} = \frac{q}{(1-q)^{2}} \tag{11}$$

Alternate: One can also differentiate¹:

$$q\frac{\partial S_0}{\partial q} = q\sum_n nq^{n-1} = S_1 \tag{12}$$

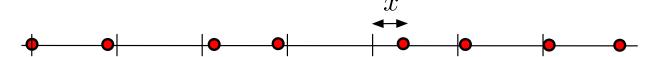
So using S_0 one has

$$S_1 = q \frac{\partial}{\partial q} \left(\frac{1}{1 - q} \right) = \frac{q}{(1 - q)^2} \tag{13}$$

¹This step is essentially the same as differentiating the partition function with respect to β in the statistical mechanics problem

Problem 2. Vibrations of a chain of atoms

A one dimensional wire at temperature T, consists of N independent atoms of mass m "living" at separate sites as shown below. The atoms oscillate harmonically around their individual site centers with vibrational potential energy of $V(x) = \frac{1}{2}k_0x^2$ for an atom displaced from by x from the cite center. Here k_0 is the spring constant.



- (a) Treat the harmonic oscillators as independent and classical, and assume that the motion is only in the x direction. What is the energy and specific heat of the substance? How would your result change if the atoms could move in both the x and y directions?
- (b) Consider a single classical oscillator of part (a) with the motion only in the x direction. What is the mean and standard deviation of the vibrational potential energy of the oscillator?
- (c) Now, imagine that the oscillators are quantum as opposed to classical.
 - (i) What is the probability that an atom is vibrating with at least two units of vibrational quanta $2\hbar\omega_0$?

Solution

(a) We use the equipartition theorem. Each oscillator has Hamiltonian

$$E = \frac{p^2}{2m} + \frac{1}{2}k_0x^2 \tag{14}$$

There are two quadratic forms. Thus the mean energy is

$$U = NkT (15)$$

where N is the number of particles. The specific heat is $C_V = \partial U/\partial T = Nk$. In two dimensions one would have

$$E = \frac{p_x^2}{2m} + \frac{p_y^2}{2m} + \frac{1}{2}k_0(x^2 + y^2)$$
 (16)

which doubles the number of quadratic forms leading to

$$U = 2NkT C_V = 2Nk (17)$$

(b) The potential energy is $V \equiv \frac{1}{2}k_0x^2$. The mean potential energy is given again by the equipartition theorem

$$\langle V \rangle = \left\langle \frac{1}{2} k_0 x^2 \right\rangle = \frac{1}{2} k_B T \tag{18}$$

Indeed, the equipartition theorem says that the average of each quadratic form in the energy function (Hamiltonian) is $\frac{1}{2}k_BT$. Then variance is

$$\langle \delta V^2 \rangle = \langle V^2 \rangle - \langle V \rangle^2 = \frac{1}{4} k_0^2 \langle x^4 \rangle - \langle V \rangle^2$$
 (19)

So we need to compute $\langle x^4 \rangle$.

The probability of x and p is

$$\mathscr{P}_{x,p} = Ce^{-\beta(p^2/2m + V(x))} dxdp \tag{20}$$

Since the energy is a sum of a part that depends on momentum plus a part that depends only on space, V(x), the probability distribution will also factorize P(x, p) dx dp = P(p)dp P(x) dx. The part that depends on x is

$$\mathscr{P}_x = P(x)dx = Ce^{-\beta V(x)}dx = Ce^{-\beta k_0 x^2/2} = Ce^{-x^2/2\sigma^2}dx,$$
(21)

where we defined $\sigma^2 \equiv 1/\beta k_0$. The probability distribution is Gaussian, which we studied in detail. Then using the Gaussian integrals provided at the front of the exam

$$C = \frac{1}{\sqrt{2\pi\sigma^2}} \qquad \langle x^4 \rangle = 3\sigma^4 \,. \tag{22}$$

So assembling the ingredients

$$\langle V^2 \rangle - \langle V \rangle^2 = \left[\left(\frac{1}{4} k_0^2 \right) 3 \left(\frac{1}{\beta k_0} \right)^2 \right] - \left[\frac{1}{2} k_B T \right]^2 = \frac{1}{2} (k_B T)^2$$
 (23)

(c) For the quantum harmonic oscillator, the probability of having n quanta of energy is

$$P_n = \frac{e^{-n\beta\hbar\omega_0}}{Z} = e^{-n\hbar\omega_0}(1 - e^{-\hbar\omega_0})$$
(24)

To find the probability of having at least two quanta, i.e. not n = 0 or n = 1, we subtract from unity the probability of having n = 0 and n = 1, i.e.

$$P_{n\geq 2} = 1 - P_0 - P_1 = 1 - (1 + e^{-\hbar\omega_0})(1 - e^{-\hbar\omega_0}) = e^{-2\hbar\omega_0}$$
(25)

Problem 3. Relativistic particles

Consider a classical mono-atomic ideal gas in two dimensions at temperature T. Note however, that these particles have zero mass and move at the speed of light, c. Compared to the non-relativistic case, the difference is that the energy of a massless particle is related to its momentum via $\varepsilon(p) = cp$ instead of $\varepsilon(p) = p^2/2m$. Otherwise, statistical mechanics and notions of phase space are unchanged.

- (a) Determine the probability of finding a particle with momentum (magnitude) between p and p + dp, P(p)dp.
- (b) Determine the mean de Broglie wavelength $\lambda \equiv h/p$ of the particles.
- (c) Determine the probability of finding a particle with de Broglie wavelength between λ and $\lambda + d\lambda$?

Solution

(a) We have
$$d\mathscr{P}_{p} = Ce^{-\beta\epsilon} dp_x dp_y = Ce^{-\beta cp} dp_x dp_y \tag{26}$$

Since we are asking about the magnitude we integrate over the angle as done in homework

$$d\mathscr{P}_p = Ce^{-\beta cp} \, 2\pi p \, dp \tag{27}$$

Normalizing

$$1 = 2\pi C \int_0^\infty e^{-\beta cp} p dp \tag{28}$$

$$=2\pi C \frac{1}{(\beta c)^2} \int_0^\infty e^{-u} u du \tag{29}$$

$$=2\pi C \frac{1}{(\beta c)^2} \Gamma(2) \tag{30}$$

Using that $\Gamma(2) = 1$ we have finally $C = (\beta c)^2/2\pi$ and thus

$$d\mathscr{P}_p = (\beta c)^2 e^{-\beta cp} \, p \, dp \,. \tag{31}$$

(b) We compute

$$\langle \lambda \rangle = \int_0^\infty P(p)dp \, \frac{h}{p} \tag{32}$$

$$= \int_0^\infty (\beta c)^2 e^{-\beta cp} p \,\mathrm{d}p \,\frac{h}{p} \tag{33}$$

$$=h\beta c \int_0^\infty (\beta c)e^{-\beta cp}\mathrm{d}p \tag{34}$$

$$=h\beta c\tag{35}$$

(c) Changing variables $p = h/\lambda$, using

$$dp = \left| \frac{dp}{d\lambda} \right| d\lambda = \frac{h}{\lambda^2} d\lambda \tag{36}$$

we find

$$d\mathscr{P}_{\lambda} = (h\beta c)^2 e^{-h\beta c/\lambda} \frac{d\lambda}{\lambda^3}$$
(37)

Problem 4. Partition function of three level system

Consider an ensemble at temperature T of N independent "atoms", each of which have the three energy levels shown below.

$$E_2 = 5\Delta$$

$$E_1 = \Delta$$

$$E_0 = 0$$

- (a) Determine the partition function and mean energy of an atom as a function of temperature.
- (b) Determine the probabilities to be in the first excited and second excited states, P_1 and P_2 , and qualitatively sketch these probabilities versus temperature on the same graph, from very low to very high temperatures:
 - (i) Explain the qualitative features of your graph by pointing to specific terms in your equations.
 - (ii) Determine a Taylor series expansion for the probability P_2 at high temperature, including the leading and first subleading terms. Explain the value of the leading term physically. Taking $\Delta = 0.1 \, \mathrm{eV}$, estimate the temperature (in Kelvin) when the Taylor series becomes approximately valid.

Solution

(a) We have

$$Z = \sum_{n} e^{-\beta \epsilon_n} = 1 + e^{-\beta \Delta} + e^{-5\beta \Delta}$$
(38)

The mean energy is

$$\langle \epsilon \rangle = \frac{-1}{Z} \frac{\partial Z}{\partial \beta} = \frac{\Delta e^{-\beta \Delta} + 5\Delta e^{-5\beta \Delta}}{1 + e^{-\beta \Delta} + e^{-5\beta \Delta}}$$
(39)

(b) We have

$$P_1 = \frac{e^{-\beta\Delta}}{1 + e^{-\beta\Delta} + e^{-5\beta\Delta}} \qquad P_2 = \frac{e^{-5\beta\Delta}}{1 + e^{-\beta\Delta} + e^{-5\beta\Delta}} \tag{40}$$

(i) At low temperatures, $e^{-\beta\Delta} \to 0$, and thus $P_1 \to 0$ and $P_2 \to 0$. At high temperatures we have

$$P_1 \to \frac{1}{1+1+1} = \frac{1}{3} \qquad P_2 \to \frac{1}{1+1+1} = \frac{1}{3}$$
 (41)

This is the expected result. At high temperatures the spacing between the levels $\sim \Delta$ is negligible compared to kT. Thus in the high temperature limit it is as if all the levels have the same energy² and thus the same probabilities. Your graphs should transition between 0 and 1/3.

(ii) We call $x = \beta \Delta \ll 1$ and have

$$P_2 = \frac{e^{-5x}}{1 + e^{-x} + e^{-5x}} \simeq \frac{1 - 5x}{1 + (1 - x) + (1 - 5x)}$$
(42)

$$\simeq \frac{1 - 5x}{3 - 6x} \tag{43}$$

We should now factor out the leading result

$$P_2 \simeq \frac{1}{3} \frac{1 - 5x}{1 - 2x} \simeq \frac{1}{3} (1 - 5x)(1 + 2x) \simeq \frac{1}{3} (1 - 3x + \mathcal{O}(x^2))$$
 (44)

Based on the first term in the series expansion we require that

$$3x \ll 1 \tag{45}$$

So the series would begin to work for

$$x = \frac{\Delta}{kT} \sim \frac{1}{3} \tag{46}$$

So roughly when

$$T \sim \frac{3\Delta}{k_B} \simeq \frac{0.3 \,\text{eV}}{0.025 \,\text{eV}/300^{\circ} \text{K}} \sim 3600 \,^{\circ} \text{K}$$
 (47)

Though any rough estimate of a couple thousand degrees would have been acceptable.

Take for instance $\Delta = 10^{-6}$ in some units and kT = 1 in the same units. Then all levels are essentially at energy zero i.e. the energy levels are $0, 10^{-6}, 5 \times 10^{-6} \simeq 0, 0, 0$.