

- into modes

$$\vec{p} = \vec{h} \vec{k} = \vec{h} \left( \frac{\pi n_x}{L}, \frac{\pi n_y}{L}, \frac{\pi n_z}{L} \right)$$

- · Each mode is an independent subsystem sharing the available energy and particles. Temperature is a parameter describing how the energy is shared.

  M is a parameter describing how the particles are
  - The mean number of particles in a mode of single-particle energy E(p) is

$$n(s) = 1$$
 = if particles are  $e^{\beta(\epsilon(p)-\mu)} - 1$  bosons

$$\overline{n(p)} = \frac{1}{e^{\beta(\epsilon(p)-\mu)}+1}$$
 ( if particles are fermions

The energy in a mode is Ep=nE(p) and the mean energy of a mode is

$$E_p = n \, \epsilon(p)$$

## Black Body Radiation and the Photon Gas:



$$E(p) = p^2$$
 non-relativist particles  
2m

$$N = \sum_{\text{modes}} \frac{1}{e^{\beta(\epsilon(p)-\mu)} + 1}$$

## Photons

- · Photons are bosons.
- · E(p) = cp, since they are relativistic + massless.
- o  $\mu = 0$  since photons are easily created and destroyed
- The sum over modes / states was:

$$2\sum_{n \neq n}\sum_{n \neq n}$$

of photons gives an overall factor of 2:

$$\frac{\sum}{\text{modes}} \rightarrow 2 \int \frac{V d^3 p}{(2\pi h)^3}$$

$$N = 2V \int d^3p \qquad 1$$

$$(2\pi t)^3 e^{CP/kT} - 1$$

Then, 
$$P_0 = \frac{E_0 - kT}{c}$$
, and the typical vavelength is:

$$7_0 = \frac{1}{k} = \frac{1}{k}$$
 with  $3 = 2$ 

with 
$$\chi = 2$$

$$N = V \left(\frac{\rho_0}{\tau}\right)^3 = V \left(\frac{kT}{\tau}\right)^3 = V \left(\frac{kT}{\tau}\right)^3$$

So the density of photons is 
$$n_x = N/V$$

$$N_{g} = \left(\frac{N}{N}\right) = 0.244 \left(\frac{kT}{kC}\right)^{3} = 0.244$$