

most energetic

$$a) U = 2 \int \frac{V d^3 p}{(2\pi)^3} \frac{\epsilon}{e^{\beta \epsilon} - 1}$$

So writing $\epsilon = \hbar \omega = c p$

$$d^3 p = 4\pi p^2 dp = \frac{\hbar^3}{c^3} \omega^2 d\omega \cdot 4\pi$$

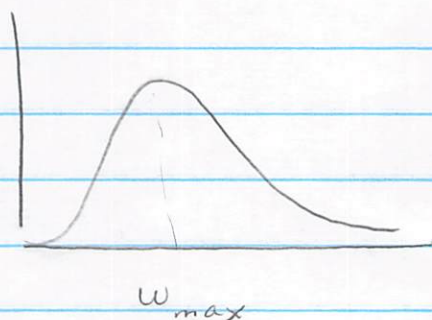
We find after algebra

$$\star U = \frac{\hbar}{\pi^2 c^3} \int_0^{\infty} \frac{\omega^3 d\omega}{e^{\beta \hbar \omega} - 1}$$

So

$$\boxed{\frac{dU}{d\omega} \propto \frac{\omega^3}{e^{\beta \hbar \omega} - 1}}$$

Plotting this, we find a maximum (see next page) at:



$$\beta \hbar \omega \approx 2.8$$

so

$$\hbar \omega = 2.8 k_B T = 2.8 \times 0.025 \text{ eV} \quad \begin{matrix} 6000^\circ \text{K} \\ 300^\circ \text{K} \end{matrix} = 1.4 \text{ eV}$$

b) Then

$$\omega = 2\pi f = \frac{2\pi c}{\lambda}$$

$$d\omega = d\lambda \left| \frac{d\omega}{d\lambda} \right| = d\lambda \left| -\frac{2\pi c}{\lambda^2} \right|$$

↑ note absolute value for unoriented integrals as discussed previously

So substituting into eq * above

$$u = \frac{h}{\pi^2 c^3} (2\pi c)^4 \int_0^\infty \frac{d\lambda}{\lambda^5} \frac{1}{e^{\beta h (2\pi)/\lambda} - 1}$$

Note $\beta h 2\pi/\lambda = \beta hc/\lambda$ so

$$\frac{du}{d\lambda} \propto \frac{1}{\lambda^5} \frac{1}{e^{\beta hc/\lambda} - 1}$$

Plotting this gives (see next page)

$$\beta \frac{hc}{\lambda} = 0.20 \quad \text{or} \quad \lambda \approx 5.0 \frac{hc}{k_B T}$$

infrared
↙

$$\text{So } \lambda = 5.0 \frac{1240 \text{ eV nm}}{.025 \text{ eV}} \frac{6000^\circ \text{K}}{300^\circ \text{K}} = 12,400 \text{ nm}$$

Density of States

a) So

$$d\mathcal{N} = V \frac{d^3 p}{(2\pi\hbar)^3} = \text{number of modes with momentum } \vec{p} = (p_x, p_y, p_z) \text{ in range}$$

$$[p_x; dp_x], [p_y; dp_y], [p_z; dp_z]$$

This means $p_x < p'_x < p_x + dp_x$:

$$d^3 p = 4\pi p^2 dp$$

And $p = \hbar k$, so

$$d\mathcal{N} = V \frac{4\pi}{(2\pi)^3} k^2 dk = \boxed{\frac{V}{2\pi^2} k^2 dk} \rightarrow \text{or } g(k) = \frac{V k^2}{2\pi^2}$$

= number of modes with k in range $k < k' < k + dk$

In two dimensions

$$d\mathcal{N} = \frac{A d^2 p}{(2\pi\hbar)^2} = A \frac{2\pi p dp}{(2\pi\hbar)^2} \quad \left. \vphantom{\frac{2\pi p dp}{(2\pi\hbar)^2}} \right) p = \hbar k$$

$$\boxed{d\mathcal{N} = \frac{1}{2\pi} A k dk}$$

or

$$\boxed{g(k) = A k / 2\pi}$$

So since

$$\mathcal{E}(p) = \frac{p^2}{2m} \quad d\mathcal{E} = \frac{p}{m} dp$$

Then in 3D

$$d\mathcal{N} = \frac{V}{2\pi^2} \frac{p^2 dp}{\hbar^3} = \frac{1}{2\pi^2} \frac{m p}{\hbar^3} d\mathcal{E} = \frac{1}{4\pi^2} \left(\frac{2m p}{\hbar^3} \right) d\mathcal{E}$$

$$= \boxed{\frac{V}{4\pi^2} \left(\frac{2m}{\hbar^2} \right)^{3/2} \sqrt{\mathcal{E}} d\mathcal{E} = g(\mathcal{E}) d\mathcal{E}}$$

In 2D:

grouping it like this is motivated by units:

$$d\mathcal{N}_p = \frac{A d^2 p}{(2\pi \hbar)^2}$$

$$\frac{1}{\lambda_{typ}} \sim \frac{p_{typ}}{\hbar} \sim \left(\frac{2m}{\hbar^2} \right)^{1/2} \mathcal{E}^{1/2}$$

Integrating over the angles of p we have

$$d\mathcal{N}_p = \frac{A p dp}{2\pi \hbar^2}$$

now with

$$\mathcal{E} = \frac{p^2}{2m} \quad d\mathcal{E} = \frac{p dp}{m}$$

$$d\mathcal{N}_{\mathcal{E}} = \frac{A m d\mathcal{E}}{2\pi \hbar^2}$$

$$\boxed{d\mathcal{N}_{\mathcal{E}} = \frac{A}{4\pi} \left(\frac{2m}{\hbar^2} \right) d\mathcal{E}}$$

and

$$\Phi_G = \int_0^{\infty} g(\epsilon) d\epsilon - k_B T \ln(1 + e^{-\beta(\epsilon - \mu)})$$

★ Then for photon spin of photons (or polarizations) = 2

$$d\mathcal{N}_{\text{modes}} = \int_{\text{angles}} 2 \frac{V d^3p}{(2\pi\hbar)^3} = \frac{1}{\pi^2 \hbar^3} V p^2 dp$$

Now $\epsilon = cp$ so

$$d\mathcal{N} = \frac{1}{\pi^2} \frac{V}{(\hbar c)^3} \epsilon^2 d\epsilon = g(\epsilon) d\epsilon$$

And

$$\Phi_G = \frac{V}{\pi^2 (\hbar c)^3} \int_0^{\infty} \epsilon^2 k_B T \ln(1 - e^{-\beta(\epsilon - \mu)})$$

So

$$\Phi_G = -pV \quad \text{So}$$

$$pV = - \frac{1}{\pi^2 (\hbar c)^3} \int_0^{\infty} \epsilon^2 k_B T \ln(1 - e^{-\beta(\epsilon - \mu)}) d\epsilon$$

c) Then the free energy of one mode is

$$\Phi_G^\epsilon = -k_B T \ln \frac{1}{1 - e^{-\beta(\epsilon - \mu)}}$$

where $2_p = \frac{1}{1 - e^{-\beta(\epsilon - \mu)}}$ is the grand partition function of one mode for a boson

So

$$\Phi_G^{\text{tot}} = \sum_{\text{modes}} \Phi_G^\epsilon$$

sum over

By definition the \sum modes becomes an integral over the mode density, $g(\epsilon) d\epsilon$

$$\sum_{\text{modes}} \dots = \int g(\epsilon) d\epsilon \dots$$

So

$$\Phi_G = \int_0^\infty g(\epsilon) k_B T \ln (1 - e^{-\beta(\epsilon - \mu)})$$

For a fermion

$$2_p = 1 + e^{-\beta(\epsilon - \mu)}$$

Entropy / Photon

- So setting $\epsilon = \hbar\omega$ the results of the previous problem can be written

$$pV = \frac{-1}{\pi^2 (\hbar c)^3} V \hbar^3 \int_0^\infty \omega^2 d\omega \underbrace{k_B T \ln(1 - e^{-\beta \hbar \omega})}_u$$

- Now integrate by parts

$$dv = \omega^2 d\omega$$

$$u = k_B T \ln(1 - e^{-\beta \hbar \omega})$$

$$v = \frac{1}{3} \omega^3$$

$$du = k_B T \frac{e^{-\beta \hbar \omega}}{1 - e^{-\beta \hbar \omega}} \beta \hbar d\omega$$

$$= \frac{\hbar d\omega}{e^{\beta \hbar \omega} - 1}$$

So

$$\int_0^\infty u dv = uv \Big|_0^\infty - \int_0^\infty v du$$



this is zero since $u \Big|_0 = 0$ and $v \Big|_\infty = 0$

- And so

$$pV = \frac{V}{\pi^2 c^3} \hbar \int_0^\infty \frac{1}{3} \omega^3 d\omega \frac{\hbar}{e^{\beta \hbar \omega} - 1}$$

First we make the integral dimensionless

$$u = \beta \hbar \omega \quad du = \beta \hbar d\omega$$

Then we have

$$pV = \frac{V}{\pi^2 c^3} \hbar \frac{1}{3} \frac{1}{(\beta \hbar)^4} \int_0^{\infty} \frac{u^4}{e^u - 1}$$

$$= V \left(\frac{k_B T}{\hbar c} \right)^3 k_B T \left[\frac{1}{3\pi^2} \int_0^{\infty} \frac{u^4}{e^u - 1} du \right]$$

The integral is

$$\frac{\pi^4}{15} \quad (\text{See attached Table})$$

And so we find

$$pV = V \left(\frac{k_B T}{\hbar c} \right)^3 k_B T \frac{\pi^2}{45}$$

Note

$$d\Phi_G = -SdT - Nd\mu - pdV$$

$$\text{With } \Phi_G = -p(T, \mu) V$$

So

$$\frac{\partial p}{\partial T} = S$$

So since $pV = CT^4$
 $S =$

$$S = \frac{\partial p}{\partial T} = 4CT^3 = 4 \frac{CT^4}{T} = 4 \frac{pV}{T}$$

So

$$U - TS + pV = \mu N$$

$$\frac{S}{Nk_B} = \frac{4\pi^2/15}{0.245}$$

Then

$$= 3.6$$

$$U = -pV + TS$$

$$U = -pV + 4pV = 3pV$$

So

$$\frac{U}{V} = 3p \quad \text{as before}$$

Fluctuations

$$Z = \sum_s e^{-\beta(\epsilon_s - \mu N_s)} = \sum_s e^{-\beta \epsilon_s} e^{\beta \mu N_s}$$

Then

$$\langle N \rangle = \frac{1}{Z} \frac{1}{\beta} \frac{\partial}{\partial \mu} Z = \frac{\sum_s e^{-\beta \epsilon_s} e^{\beta \mu N_s} N_s}{Z}$$

$$\underline{\langle N \rangle = \sum_s P_s N_s}$$

Here we have recognized that the probability to be in a state with energy ϵ_s and number N_s is

$$P_s = \frac{e^{-\beta(\epsilon_s - \mu N_s)}}{Z}$$

Similarly

$$\langle N^2 \rangle = \sum_s P_s N_s^2 = \frac{1}{Z} \sum_s e^{-\beta \epsilon_s} e^{\beta \mu N_s} N_s^2$$

No

$$\left(\frac{1}{\beta} \frac{\partial}{\partial \mu} \right) \left(\frac{1}{\beta} \frac{\partial}{\partial \mu} \right) e^{\beta \mu N_s} = e^{\beta \mu N_s} N_s^2$$

So

$$\sigma_N^2 = \frac{1}{\beta} \frac{\partial}{\partial \mu} \left(-\frac{1}{\beta} \frac{\partial \bar{\Phi}}{\partial \mu} \right)$$

$$\boxed{\sigma_N^2 = \frac{1}{\beta} \frac{\partial \bar{N}}{\partial \mu}}$$

★ This formula should be compared to the variance in the energy

$$\sigma_E^2 = \frac{\partial^2}{\partial \beta^2} (\ln Z)$$

$$\sigma_E^2 = -\frac{\partial \bar{E}}{\partial \beta} \quad \text{note} \quad \bar{\sigma}_E = -\frac{\partial \ln Z}{\partial \beta}$$

b) So

$$Z_{BE} = \sum_{n=0}^{\infty} e^{-\beta(n\varepsilon - \mu n)} = \frac{1}{1 - e^{-\beta(\varepsilon - \mu)}}$$

The terms where

$n \neq 0, 1$ are small only if $e^{-\beta(\varepsilon - \mu)} \ll 1$

For instance Look at the contribution from two particles in a mode to Z :

$$e^{-\beta(2\varepsilon - \mu 2)} = e^{-2\beta(\varepsilon - \mu)} = (e^{-\beta(\varepsilon - \mu)})^2$$

So

$$\underline{\langle N^2 \rangle = \frac{1}{2} \left(\frac{1}{\beta} \frac{\partial}{\partial \mu} \right)^2 2}$$

So

$$\sigma_N^2 = \frac{1}{2} \left(\frac{1}{\beta} \frac{\partial}{\partial \mu} \right)^2 2 - \left(\frac{1}{2} \frac{1}{\beta} \frac{\partial 2}{\partial \mu} \right)^2$$

$$\underline{\sigma_N^2 = \left(\frac{1}{\beta} \frac{\partial}{\partial \mu} \right)^2 \ln 2}$$

$$\text{Now } \ln 2 = -\frac{\bar{\Phi}_G}{k_B T} = \frac{pV}{k_B T}$$

$$\sigma_N^2 = k_B T V \frac{\partial^2 p}{\partial \mu^2}$$

Note also

$$d\bar{\Phi}_G = -SdT - N d\mu - p dV$$

$$\frac{\partial \bar{\Phi}}{\partial \mu} = N$$

The classical approximation assumes this is small compared to having one particle per mode:

$$(e^{-\beta(\epsilon-\mu)})^2 \ll e^{-\beta(\epsilon-\mu)}$$

So we have

$$(1) \quad n_{BE} = \frac{1}{e^{\beta(\epsilon-\mu)} - 1} = \frac{e^{-\beta(\epsilon-\mu)}}{1 - e^{-\beta(\epsilon-\mu)}} \\ \approx e^{-\beta(\epsilon-\mu)} (1 + e^{-\beta(\epsilon-\mu)} + \dots)$$

Similarly

$$(2) \quad \Phi_G = k_B T \ln(1 - e^{-\beta(\epsilon-\mu)})$$

use $\ln(1+x) \approx x - \frac{x^2}{2}$

$$\approx -k_B T e^{-\beta(\epsilon-\mu)} - \frac{1}{2} k_B T (e^{-\beta(\epsilon-\mu)})^2$$

c) Then

$$pV = \sum_{\text{modes}} -\Phi_G^{\text{mode}} \quad \left. \vphantom{\sum_{\text{modes}}} \right\} \text{use Eq. (2)} \\ = \int \frac{V d^3p}{(2\pi\hbar)^3} k_B T e^{-\beta(\epsilon-\mu)} \left(1 + \frac{1}{2} e^{-\beta(\epsilon-\mu)} \right)$$

Performing the integral

$$pV = V \int \frac{d^3p}{(2\pi\hbar)^3} k_B T e^{-P^2/2mkT} e^{\beta\mu}$$
$$\left[1 + \frac{1}{2} e^{-P^2/2mkT} e^{\beta\mu} \right]$$

$$= V \int \frac{d^3p}{(2\pi\hbar)^3} k_B T e^{-P^2/2mkT} e^{\beta\mu}$$

$$+ \frac{V}{2} \int \frac{d^3p}{(2\pi\hbar)^3} k_B T e^{-2P^2/2mkT} e^{2\beta\mu}$$

The Integrals are

$$I_1 = \int \frac{d^3p}{(2\pi\hbar)^3} e^{-P^2/2mkT} = n_Q \equiv \frac{(2\pi m k_B T)^{3/2}}{h^3}$$

$$I_2 = \int \frac{d^3p}{(2\pi\hbar)^3} e^{-2P^2/2mkT} = \frac{n_Q}{2\sqrt{2}} = \frac{(2\pi (m/2) k_B T)^{3/2}}{h^3}$$

So

$$pV = V k_B T e^{\beta\mu} n_Q \left[1 + \frac{e^{\beta\mu}}{4\sqrt{2}} \right]$$

Thus

$$p = k_B T e^{\beta\mu} n_Q \quad \text{at leading order}$$

S₀

$$\bar{N} = - \left(\frac{\partial \bar{F}_G}{\partial \mu} \right)_T = \frac{\partial (pV)}{\partial (\mu)}$$

$$\bar{N} = V k_B T e^{\beta \mu} \beta n_Q$$

$$\left(\frac{\bar{N}}{V} \right) = e^{\beta \mu} n_Q$$

← this agrees with HW/2 prob 2.

d)

$$\sigma_N^2 = \frac{1}{\beta} \frac{\partial \bar{N}}{\partial \mu} = V \frac{1}{\beta} \frac{\partial (e^{\beta \mu})}{\partial \mu} n_Q$$

$$\sigma_N^2 = V e^{\beta \mu} n_Q = \bar{N}$$

S₀

$$\frac{\sigma_N^2}{\langle N \rangle} = \frac{1}{\langle N \rangle}$$