

Einstein Model of Solid

N atoms connected
by springs.



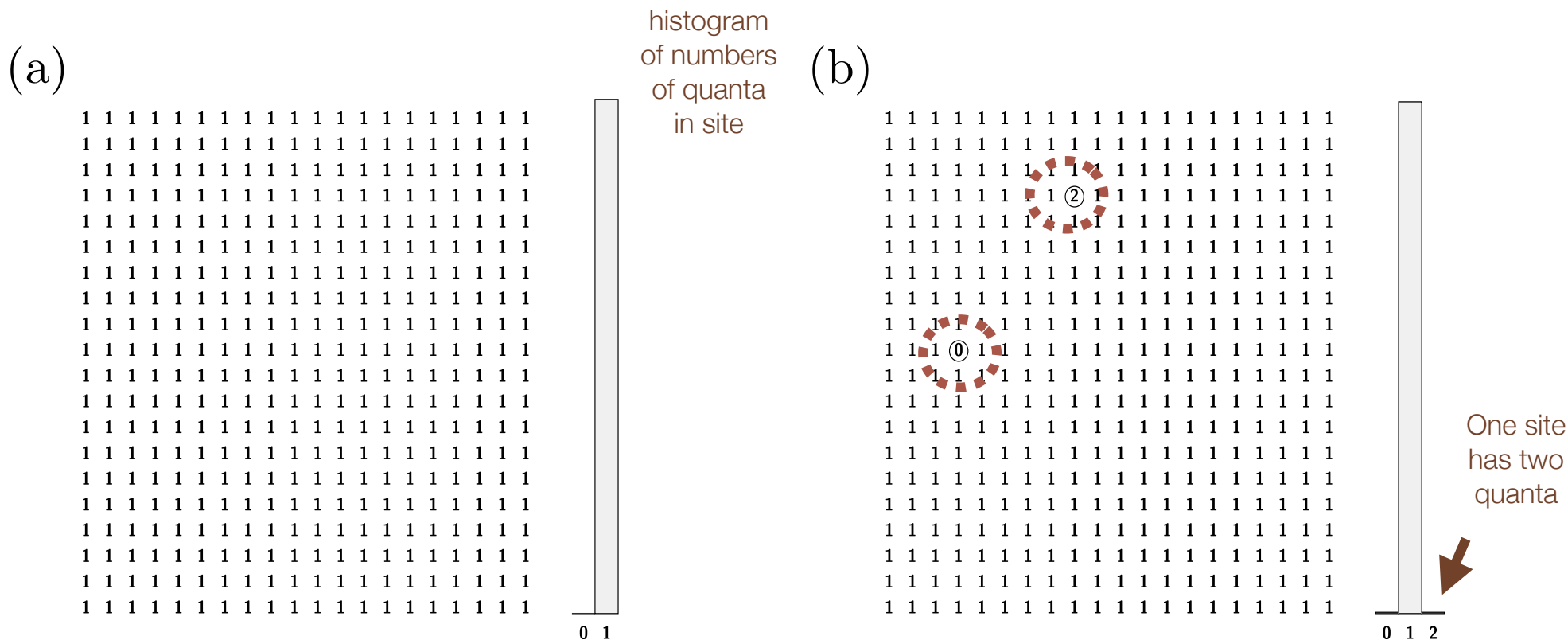
- As a simple picture of 1d solid, might consist of N oscillators. For simplicity take each of the oscillators to be independent
- Then the total Energy is $U = q \hbar \omega_0$.
 q is the total number of quanta of energy to be shared or partitioned amongst the N atoms
integer $q \gg 1$
- For definiteness take $\overset{N}{400}$ atoms and $\overset{q}{400}$ quanta of energy. $1 \text{ quanta} = \hbar \omega_0$
- One possible configuration is that each atom has one quanta of energy:

This is shown on the slide below.

- Now pick one oscillator at random and transfer one quanta' of energy to another randomly chosen site.

We get another possible state see slides

20x20 oscillators (sites), with 20x20 quanta of energy, one per site



Start with initial state (a), pick two oscillators randomly, and transfer an energy quanta between the two sites. You find (b).

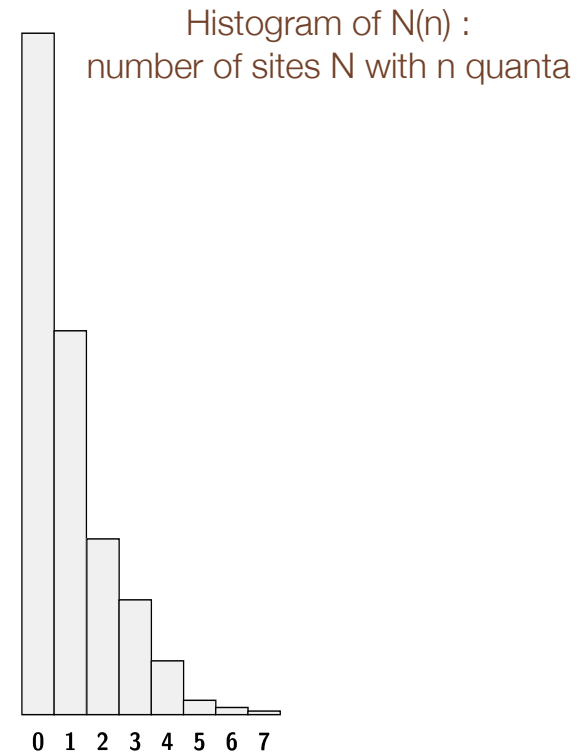
Now repeat the process: there are 10^{240} ways to share the energy

(c)

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0 0 1 0 1 1 0 0 0 0 1 3 1 2 0 0 0 1 0 0
0 1 0 5 1 4 0 1 1 0 2 0 1 0 0 0 1 3 1 0
0 3 0 1 1 0 1 0 1 2 3 0 0 1 2 4 1 0 3 2
2 1 2 4 3 4 0 0 1 1 0 4 0 1 0 2 1 1 1 0
1 2 0 0 1 0 1 0 4 0 0 0 0 0 0 1 2 0 0 0
0 1 1 1 0 4 0 1 0 2 2 1 3 1 0 0 3 0 0 0
1 0 0 0 0 2 0 0 2 0 6 0 3 1 3 0 2 1 1 0
2 2 4 1 2 0 0 0 0 1 3 0 2 0 0 0 2 1 3 2
3 0 0 2 1 1 2 0 0 0 0 0 0 0 1 0 0 0 1 0
1 3 1 1 0 0 0 0 3 0 1 0 1 0 0 0 0 2 0 0
2 1 0 1 0 1 2 0 4 1 0 1 0 2 1 1 1 1 1 2
1 0 0 0 0 0 1 4 2 2 2 0 1 0 0 2 0 0 1 1
0 3 0 1 1 0 0 0 1 0 0 3 2 0 0 2 2 2 0 3
5 2 0 0 1 0 0 2 1 0 0 0 1 0 0 1 0 3 0 3
1 1 0 3 0 0 1 4 1 0 2 0 0 6 3 0 1 0 1 3
0 1 1 0 2 0 0 4 1 3 2 0 0 0 0 2 1 0 2 0
1 4 1 0 3 0 2 1 1 0 3 1 1 0 3 1 3 0 2 0
5 0 3 1 7 2 2 0 0 1 0 0 1 1 1 0 0 0 0 3
0 0 5 0 0 1 0 1 0 2 2 1 0 4 3 3 0 0 1 0
0 0 0 0 0 1 0 1 0 0 0 0 1 0 4 1 0 1 1 1

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A typical distribution distribution is shown above: What is $N(n)$?

- Repeating the process, the system is equally likely to be in any of its accessible microstates. The system is thermalized:

$$P_m = C$$

↑ probability to be in a microstate m

Normalizing, then $\sum_m P_m = 1$ or $C \sum_m 1 = 1$

or

$$C \Omega(E) = 1 \quad \text{and}$$

$$P_m = \frac{1}{\Omega(E)}$$

↑ this letterally counts all states

probability dist

This [^] is known as the micro-canonical ensemble

- The entropy of the ensemble is

$$S = k_B \ln \Omega = -k_B \ln P_m$$

Micro-canonical only

- For the problem at hand, you will show for $N = q = 400$, that $\Omega = 10^{240}$. So

$$\frac{S}{k_B} = 240 \ln 10 \approx 554.9$$

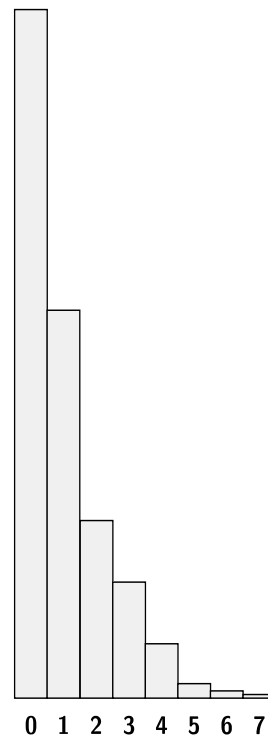
You will count the states and find $\Omega = 10^{240}$ in homework.

What is $N(n)$?

(c)

0	0	1	0	1	1	0	0	0	0	1	3	1	2	0	0	0	1	0	0
0	1	0	5	1	4	0	1	1	0	2	0	1	0	0	0	1	3	1	0
0	3	0	1	1	0	1	0	1	2	3	0	0	1	2	4	1	0	3	2
2	1	2	4	3	4	0	0	1	1	0	4	0	1	0	2	1	1	1	0
1	2	0	0	1	0	1	0	4	0	0	0	0	0	0	1	2	0	0	0
0	1	1	1	0	4	1	0	2	2	1	3	1	0	0	3	0	0	0	0
1	0	0	0	0	2	0	0	2	0	6	0	3	1	3	0	2	1	1	0
2	2	4	1	2	0	0	0	0	1	3	0	2	0	0	0	2	1	3	2
3	0	0	2	1	1	2	0	0	0	0	0	0	0	1	0	0	0	1	0
1	3	1	1	0	0	0	0	3	0	1	0	1	0	0	0	0	2	0	0
2	1	0	1	0	1	2	0	4	1	0	1	0	2	1	1	1	1	1	2
1	0	0	0	0	0	1	4	2	2	2	0	1	0	0	2	0	0	1	1
0	3	0	1	1	0	0	0	1	0	0	3	2	0	0	2	2	2	0	3
5	2	0	0	1	0	0	2	1	0	0	0	1	0	0	1	0	3	0	3
1	1	0	3	0	0	1	4	1	0	2	0	0	6	3	0	1	0	1	3
0	1	1	0	2	0	0	4	1	3	2	0	0	0	0	2	1	0	2	0
1	4	1	0	3	0	2	1	1	0	3	1	1	0	3	1	3	0	2	0
5	0	3	1	7	2	2	0	0	1	0	0	1	1	1	0	0	0	0	3
0	0	5	0	0	1	0	1	0	2	2	1	0	4	3	3	0	0	1	0
0	0	0	0	0	1	0	1	0	0	0	0	1	0	4	1	0	1	1	1

Histogram of energies
in units of $\epsilon_0 = \hbar\omega_0$



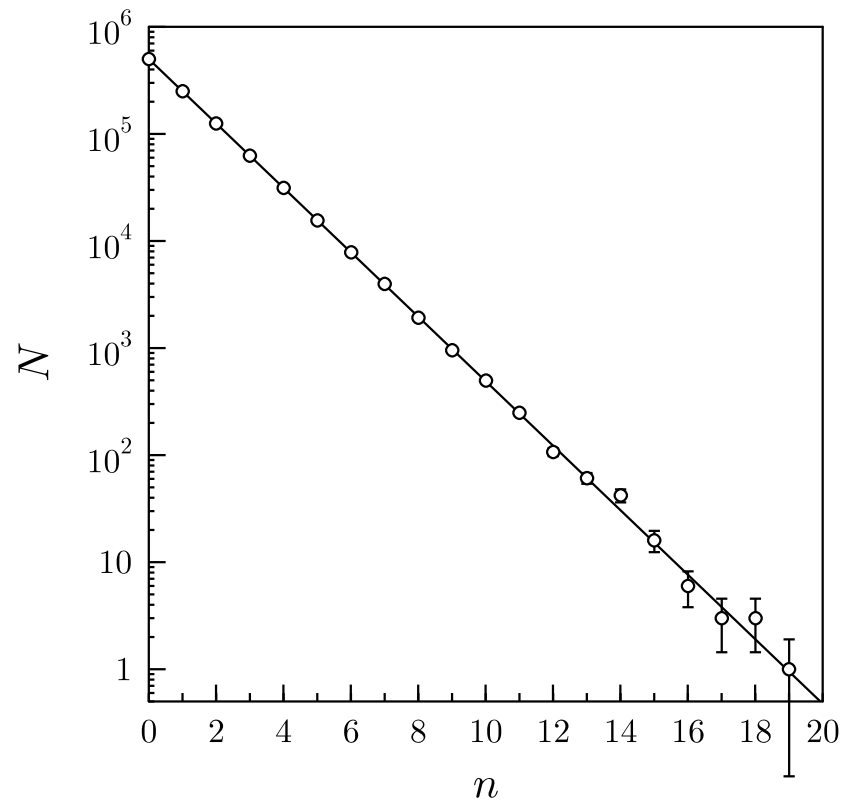
Pick a site:
The remaining sites are the reservoir

Expect the probability for a site
to have n quanta to be

$$P(\epsilon_n) \propto e^{-\beta\epsilon_n} = e^{-n\beta\epsilon_0}$$

The histogram $N(n)$ is
the number of sites with n
quanta, and should be P_n up to
normalization

Numerical verification: number of sites, $N(n)$, with n quanta on 1000x1000 grid



What you are seeing (on a log scale) is

$$N(n) = C_1 e^{-C_2 n}$$

The log of $N(n)$ is

$$\ln N(n) = \ln C_1 - C_2 n$$

The constant $C_2 = \beta \epsilon_0$ determines the temperature

- Then the probability for a site / oscillator to have energy ϵ_n is (see slide)

$$P_n \propto e^{-\epsilon_n/kT}$$

or

$$P_N = \frac{1}{Z} e^{-\epsilon_n/kT}$$

with

$$Z = \sum_{n=0}^{\infty} e^{-\epsilon_n/kT}$$

$$= \sum_{n=0}^{\infty} e^{-n\hbar\omega_0/kT}$$

$$Z = \frac{1}{1 - e^{-\hbar\omega_0/kT}}$$

magenta is general
green is for SHO

see homework

- The mean energy is:

$$\langle E \rangle = N \langle \epsilon \rangle = N \frac{\hbar\omega_0}{e^{\hbar\omega_0/kT} - 1}$$

calculated energy

input energy

- Now for the problem at hand $N=200$ $E=200 \hbar\omega_0$. The temperature is adjusted so that the mean energy agrees with the input energy

$$\langle E \rangle = 200 \hbar\omega_0$$

$$200 \hbar\omega_0 = 200 \frac{\hbar\omega_0}{e^{\hbar\omega_0/kT} - 1}$$

or

$$e^{\hbar\omega_0/kT} = 2 \Rightarrow \frac{\hbar\omega_0}{\ln 2} = k_B T$$

operational

This example shows how the temperature is defined ^ in the canonical ensemble $P_n \propto e^{-\epsilon_n/kT}$. T is adjusted to reproduce the mean energy of the full system.

- The Einstein model gives a successful phenomenology of solids at constant vol

$$dQ = dE + p dV$$

So at constant volume $dV = 0$

$$\left(\frac{\partial E}{\partial T}\right)_V = \frac{dQ}{dT} \leftarrow \text{this is } C_V$$

So

$$C_V = \left(\frac{\partial E}{\partial T}\right)_V = N (\beta \hbar \omega_0)^2 \frac{e^{-\beta \hbar \omega_0}}{(1 - e^{-\beta \hbar \omega_0})^2}$$



You will do this in homework! In fact you did it already!

We did it here in 1D.

In 3D the only change is

$$C_V = 3N (\beta \hbar \omega_0)^2 \frac{e^{-\beta \hbar \omega_0}}{(1 - e^{-\beta \hbar \omega_0})^2}$$

3 for 3

dimensions, each atom can oscillate in x, y, z

Specific Heat
of Einstein
model!

This is a one parameter model for the specific heats of solids. $\hbar \omega_0$ is adjusted to reproduce the data

- The Specific Heat of silver is shown on the slide. It is reasonably fit by taking $\hbar\omega_0 = 0.013 \text{ eV}$

- In the high temperature limit find

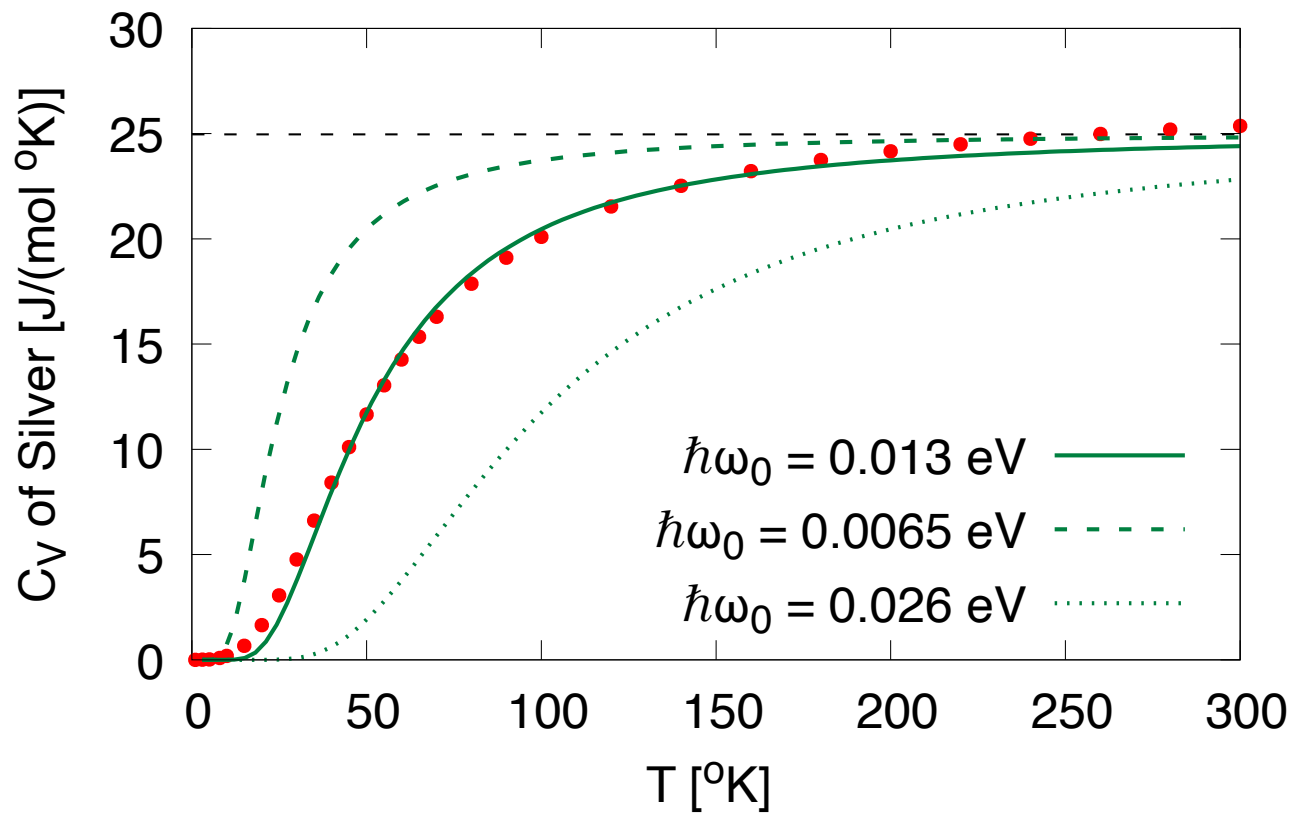
$$C_v = 3Nk_B$$

↑ or for one mole $N = N_A$, $C_v^{\text{lim}} = 3R$

Indeed at high temperatures the specific heats of solids approach $3R$ (see slide)

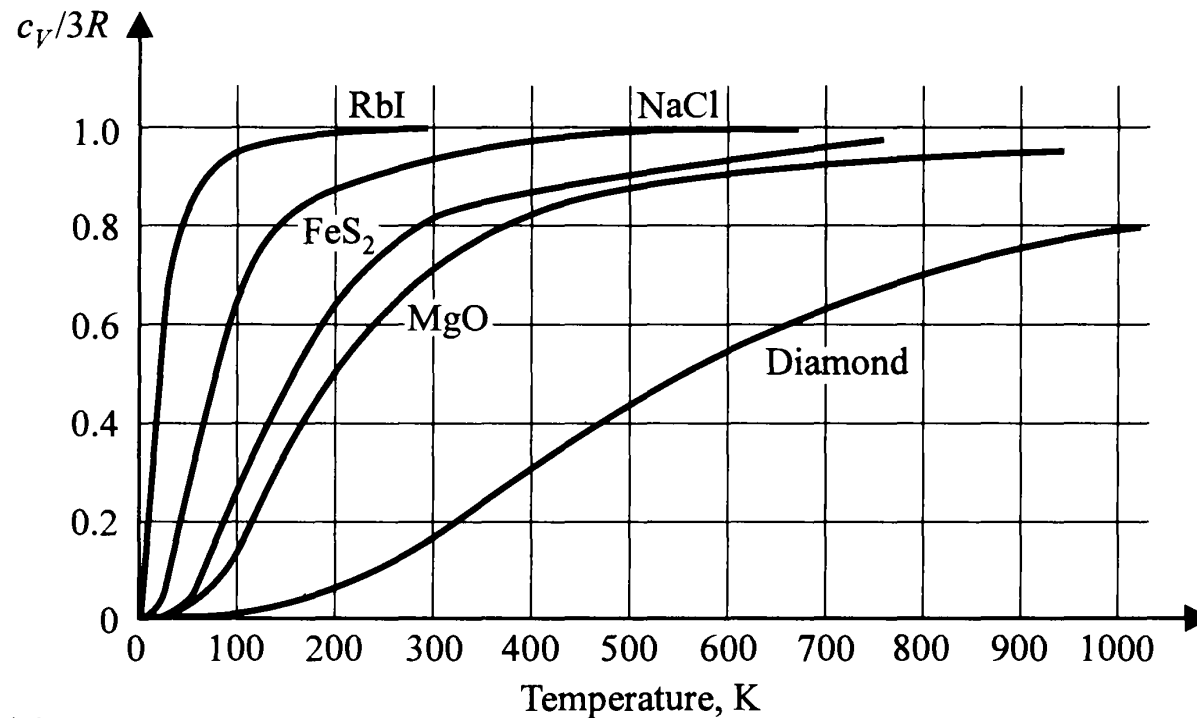
- Classically one finds $C_v = 3R$, and that the result is independent of temperature. Getting the specific heat to drop with temperature was a great early success of quantum mechanics

Specific Heat of Silver



Approaches $3R$
at high temperature
where the dynamics
is classical

Specific Heats of Solids: (Taken from Zemansky and Dittman)



Approaches $3R$
at high temperature
where the dynamics
is classical

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The general shape of these curves agrees with the Einstein Model!