

# Physics 306: Thermal Physics

Final Exam

Stony Brook University

Spring 2023

## General Instructions:

You may use one page (front and back) of handwritten notes and a calculator. Graphing calculators are allowed. **No other materials may be used.**

# 1 Integrals

**Bose and Fermi:**

$$\int_0^\infty dx \frac{x}{e^x - 1} = \frac{\pi^2}{6} \quad (1)$$

$$\int_0^\infty dx \frac{x^2}{e^x - 1} = 2\zeta(3) \simeq 2.404 \quad (2)$$

$$\int_0^\infty dx \frac{x^3}{e^x - 1} = \frac{\pi^4}{15} \quad (3)$$

$$\int_0^\infty dx \frac{x^4}{e^x - 1} = 24\zeta(5) \simeq 24.88 \quad (4)$$

$$\int_0^\infty dx \frac{x^5}{e^x - 1} = \frac{8\pi^6}{63} \quad (5)$$

$$\int_0^\infty dx \frac{x}{e^x + 1} = \frac{\pi^2}{12} \quad (6)$$

$$\int_0^\infty dx \frac{x^2}{e^x + 1} = \frac{3}{2} \zeta(3) \simeq 1.80309 \quad (7)$$

$$\int_0^\infty dx \frac{x^3}{e^x + 1} = \frac{7\pi^4}{120} \quad (8)$$

$$\int_0^\infty dx \frac{x^4}{e^x + 1} = \frac{45}{2} \zeta(5) \simeq 23.33 \quad (9)$$

$$\int_0^\infty dx \frac{x^5}{e^x + 1} = \frac{31\pi^6}{252} \quad (10)$$

**Gamma Function:**

$$\Gamma(z) \equiv \int_0^\infty x^{z-1} e^{-x} dx \quad (11)$$

with specific results

$$\Gamma(z+1) = z\Gamma(z) \quad \Gamma(n) = (n-1)! \quad \Gamma\left(\frac{1}{2}\right) = \sqrt{\pi} \quad (12)$$

**Gaussian Integrals:**

$$I_n = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^\infty dx e^{-x^2/2} x^n \quad (13)$$

with specific results

$$I_0 = 1 \quad I_2 = 0 \quad I_4 = 3 \quad I_6 = 15 \quad (14)$$

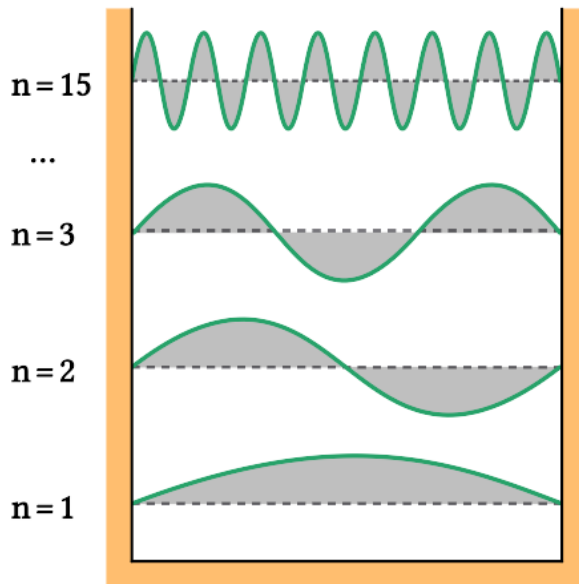
### Problem 1. A thermal particle in a box

Recall that a quantum mechanical particle in a one dimensional box of size  $L$  has energy levels

$$E_n = n^2 \epsilon_0, \quad \epsilon_0 = \frac{\hbar^2 \pi^2}{2mL^2}, \quad (15)$$

with quantum wave functions labelled by  $n$  that are shown below

$$n = 1, 2, 3, \dots \quad (16)$$



- (a) Approximately evaluate the partition function by including just the first two states  $n = 1$  and  $n = 2$ .
- (i) Explain why this approximation is valid at low temperatures, and define what is meant by *low* in this context (i.e. low compared to what?). Given an electron at room temperature, what is the range of the parameters for this problem where the temperature can be considered low?
- (b) Determine the mean energy with the same approximations as in (a).
- (c) In order for the approximation of part (a) to be valid, the probability to be in the  $n = 3$  state should be small. Determine this probability.  
*Hint:* You can use the partition function from (a) when evaluating this probability.
- (d) In the high temperature limit the sum over states  $n$  can be replaced with an integral. Evaluate the partition function in this limit and explain the result using classical reasoning.

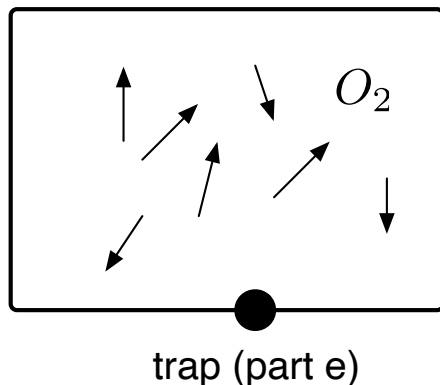
## Problem 2. A classical relativistic gas

In class we studied a non-relativistic classical mono-atomic gas of  $N$  particles in a volume  $V$  at temperature  $T$  using partition functions and the canonical ensemble. The energy was  $\epsilon = \frac{p^2}{2m}$ . The same methods can be used for a classical ultra-relativistic gas. The only difference is the energy is now,  $\epsilon(p) = cp$ . Consider a three dimensional classical relativistic gas consisting of  $N$  particles in a volume  $V$  at temperature  $T$ .

- Determine the partition function and the free energy.
- Determine the pressure. What are the similarities and differences to the non-relativistic case?
- Determine the entropy of the gas? What are the similarities and differences to the non-relativistic case?

## Problem 3. $O_2$ in equilibrium

Consider a classical ideal gas in volume  $V$  at temperature  $T$  consisting of  $N$  molecules of diatomic oxygen. The moment of inertia of the  $O_2$  molecule is  $I$ .



- Use the equipartition theorem to determine the energy of the gas and the specific heat at constant pressure. You may treat the rotational motion classically and neglect vibrations.
- Determine the free energy of the gas and show that the pressure satisfies the ideal gas law,  $PV = NkT$ .
- Determine the chemical potential,  $\mu$ , and the “fugacity”,  $z_g \equiv e^{\beta\mu}$ , as a function of temperature and pressure of the gas.
- Estimate  $z_g$  at room temperature and standard pressure. Assume  $\hbar^2/2I = 0.00018 \text{ eV}$ . Note that for  $O_2$  the “quantum pressure” is<sup>1</sup>

$$p_Q \equiv k_B T n_Q \equiv \frac{k_B T}{\lambda_{\text{th}}^3} \equiv T \left( \frac{2\pi M k T}{h^2} \right)^{3/2} = 6.8 \times 10^6 \text{ bar}. \quad (17)$$

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<sup>1</sup> $\lambda_{\text{th}}$  is the thermal de Broglie wavelength discussed in class.

A trap on the wall of the container can absorb up to two  $O_2$  molecules from the room. The energy of the trap is lowered from its unoccupied state by  $-\epsilon_0$  if one  $O_2$  is absorbed, and  $-2\epsilon_0$  if two  $O_2$  are absorbed.

- (e) What is the mean number of  $O_2$  molecules that are absorbed by the trap? Express your result as a function of the  $u \equiv z_g e^{\beta\epsilon_0}$ .
- (f) For the conditions of part (d), estimate the binding energy  $\epsilon_0$  where the mean number of  $O_2$  molecules in the trap is of order unity.

## Problem 4. Neutrino Gas

A neutrino is like a photon – it is neutral, massless, and it has two spin states, But a neutrino is a fermion not a boson, and the modes can be in two possible states, either unoccupied or occupied by a particle, with corresponding energies 0 and  $\epsilon$ . Consider a gas of neutrinos at temperature  $T$  and in a cubic box of volume  $V = L^3$ .

- (a) Consider a single mode in the container with single particle energy  $\epsilon$ . Determine the grand partition function for the mode and derive the Fermi-Dirac expression for the mean number of particles in the mode:

$$n_{FD}(\epsilon) = \frac{1}{e^{\beta(\epsilon-\mu)} + 1} \quad (18)$$

What are the probabilities of finding the mode unoccupied and occupied respectively?

- (b) What is the energy and number of the neutrinos in the box? Explain each step carefully and assume  $\mu = 0$ .
- (c) Determine the average de Broglie wavelength  $\lambda \equiv h/p$  of the neutrino by integrating over the momenta.
- (d) What is the number of neutrinos striking the walls per unit time?