

Density @ Gravity

a) We have

$$d\rho = C e^{-mgz/kT} dz dx dy e^{-p^2/2mT} dp_x dp_y dp_z$$

- We note that the probability factorizes

$$1 = \int d\rho = C \int_0^\infty dz e^{-mgz/kT} \int_0^L dx dy \int e^{-p^2/2mT} d^3p$$

- Recognizing from the old result (in class)

$$\int d^3p e^{-p^2/2mT} \left(\frac{1}{2\pi m k_B T} \right)^{3/2} = 1$$

- We insert the factors

$$1 = C L^2 (2\pi m kT)^{3/2} \underbrace{\int_0^\infty dz e^{-mgz/kT}}_{= kT/mg} \underbrace{\int_0^L dx dy}_{= 1} \underbrace{\int \frac{d^3p e^{-p^2/2mT}}{(2\pi m k_B T)^{3/2}}}_{= 1}$$

$$1 = C (2\pi m kT)^{3/2} \left(\frac{kT}{mg} \right)$$

So

$$C = \frac{1}{(2\pi m kT)^{3/2}} \frac{1}{l} \quad \text{with } l \equiv \frac{kT}{mg}$$

• Then

$$d\mathcal{P} = \frac{e^{-z/l}}{l} dz \frac{dx dy}{L^2} \frac{e^{-p^2/2mkT} d^3p}{(2\pi mkT)^{3/2}}$$

The result factorizes

$$d\mathcal{P} = P(z) dz P(x,y) dx dy P(\vec{p}) d^3p$$

• To find $P(z)$ we only need to sum over x, y , and p_x, p_y, p_z which we don't want to know

$$d\mathcal{P}_z = \frac{e^{-z/l}}{l} dz \underbrace{\int \frac{dx dy}{L^2}}_{=1} \underbrace{\int d^3p \frac{e^{-p^2/2mkT}}{(2\pi mkT)^{3/2}}}_{=1}$$

So

$$d\mathcal{P}_z = \frac{e^{-z/l}}{l} dz \quad l \equiv \frac{kT}{mg}$$

b) So

$$\begin{aligned} \langle z \rangle &= \int_0^\infty dz z P(z) = \int_0^\infty dz z \frac{e^{-z/l}}{l} \\ &= l \int_0^\infty du u e^{-u} = l \cdot 1! = l \end{aligned}$$

Estimating

$$l = \frac{k_B T}{mg} = \frac{N_A k_B T}{N_A m g} = \frac{8.32 \text{ J} \times 300^\circ \text{K}}{(28 \text{ g})(9.8 \text{ m/s}^2)} = 9 \text{ km}$$

The molar mass of air is 28 g.

c) • The density is proportional to $P(z)$

$$n(z) \propto P(z) \propto e^{-z/l}$$

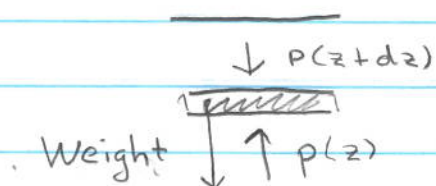
So, $n(z) = C e^{-z/l}$, at $z = 0$ $n(0) = n_0$

So

$$n(z) = n_0 e^{-z/l}$$

d) • Then

• the pressure difference
needs to support the air's
weight



$$\frac{dp}{dz} = -mg n(z)$$

So if $p = n(z) k_B T$

$$\frac{dp}{dz} = \frac{dn}{dz} k_B T$$

i.e.

$$\boxed{\frac{dn}{dz} = - \frac{mg}{kT} n}$$

For $n(z) = n_0 e^{-mgz/kT}$ this is satisfied,

Phase Space Density

Then

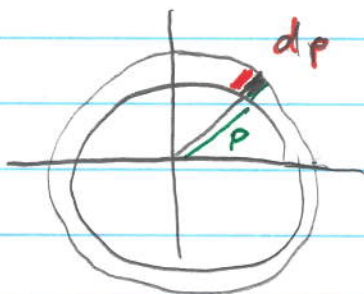
$$V_{ps} = \int_{[E, E+\delta E]} d^3r d^3p = L^3 V_p \quad \leftarrow \begin{array}{l} \text{momentum} \\ \text{space volume} \end{array}$$

- The momenta are confined to a spherical shell

$$2mE < p^2 < 2m(E + \delta E)$$

- The volume of this shell is

$$V_p = 4\pi p^2 dp$$



Since $p^2 = 2mE$

$$2p dp = 2m dE$$

$$dp = \frac{m dE}{(2mE)^{1/2}} = (2mE)^{1/2} \frac{dE}{2E}$$

So

$$V_p = 4\pi (2mE)^{3/2} \frac{dE}{2E} \quad \text{and}$$

$$V_{ps} = L^3 (2mE)^{3/2} 4\pi \frac{dE}{2E}$$

Oscillator

• Then

$$a) \quad \sum_n P_n = \sum_n C e^{-E_n/kT} = C \sum_n e^{-E_n/kT} = 1$$

$\uparrow C = \frac{1}{Z}$

• So

$$Z = \sum_n e^{-E_n \beta}$$

$$Z = \sum_n e^{-\beta \hbar \omega_0 n}$$

we use $\frac{1}{1-u} = 1 + u + u^2 + \dots$

$$b) \quad Z = \sum_{n=0}^{\infty} u^n = \frac{1}{1-u} = \frac{1}{1 - e^{-\beta \hbar \omega_0}}$$

Then

$$P_n = \frac{1}{Z} e^{-n \hbar \omega_0 / kT} = e^{-n \hbar \omega_0 \beta} (1 - e^{-\hbar \omega_0 \beta})$$

• The plot is telling me that for low temperatures the oscillator is most likely in the ground state, while for high temperature is in its vibrational states

c) Then we see that when

$$\frac{h\nu_0}{k_B T} \sim 1$$

for instance when $h\nu_0 = k_B T$
60% of atoms are in ground
state and 40%

- The atoms will begin to vibrate. are in vibrational states

$$h\nu_0 = hc k_0 = 197 \text{ eV nm } 4400 \frac{1}{\text{nm}}$$

$$h\nu_0 = 0.086 \text{ eV}$$

- At room temperature $k_B T$ is 0.025 eV
So

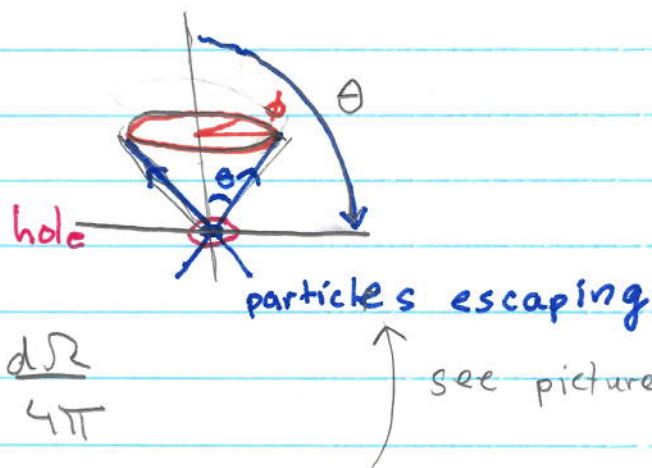
$$T = \frac{0.086 \text{ eV}}{k_B} = \frac{0.086 \text{ eV}}{0.025 \text{ eV}} \cdot 300^\circ \text{K}$$

i.e. $\frac{1}{k_B} = \frac{300^\circ \text{K}}{0.025 \text{ eV}}$

$$T \approx 1000^\circ \text{K}$$

Escaping

a)



$$d\Phi = n P(v) dv \cdot v \cos\theta \cdot \frac{d\Omega}{4\pi}$$

• Integrating from $\theta \in [0, \pi/2]$ and $\phi \in [0, 2\pi]$, and v :

$$\Phi = \int_0^\infty n P(v) dv \cdot v \int_0^{\pi/2} \cos\theta \sin\theta d\theta \int_0^{2\pi} \frac{d\phi}{4\pi}$$

$$= n \langle v \rangle \left[\frac{-\cos^2\theta}{2} \right]_0^{\pi/2} \cdot \frac{1}{2}$$

$$\Phi = n \langle v \rangle \left(\frac{1}{2} \right) \left(\frac{1}{2} \right)$$

$$\boxed{\Phi = n \langle v \rangle \frac{1}{4}}$$

b) Then we want

$$\langle \cos\theta \rangle$$

The, $d\Phi \propto \cos\theta \sin\theta d\theta$, and this serves as the probability distribution for escaping particles.

$$d\mathcal{P} = C \cos\theta \sin\theta d\theta$$

So

$$\langle \cos\theta \rangle = \frac{\int_0^{\pi/2} \cos\theta \times \cos\theta \sin\theta d\theta}{\int_0^{\pi/2} \cos\theta \sin\theta d\theta}$$

$$C = \frac{1}{\int_0^{\pi/2} \cos\theta \sin\theta d\theta}$$

$$= \frac{-\frac{1}{3} \cos^3\theta \Big|_0^{\pi/2}}{-\frac{1}{2} \cos^2\theta \Big|_0^{\pi/2}} = \frac{\frac{1}{3}}{\frac{1}{2}} = \frac{2}{3}$$

c) Then

$$\frac{dN}{dt} = -A \Phi$$

area of hole
number escaping per area per time

$$= -A \frac{1}{4} \frac{N}{V} \left(\frac{8kT}{\pi m} \right)^{1/2} = -\frac{N}{\tau}$$

↑
↑

n
 $\langle v \rangle$

where $\tau = \frac{V}{A} \left(\frac{2\pi m}{kT} \right)^{1/2}$

So

$$A = \frac{V}{\tau} \left(\frac{2\pi m}{kT} \right)^{1/2}$$

- So using $N_A m_{\text{air}} = 28g$ $N_A k_B = 8.32 J$

$V = 1L$ we have

$$A = \frac{(1L)}{3600s} \left(\frac{2\pi \cdot 28g}{8.32J \times 300^\circ K} \right)^{1/2}$$

$$= 2.33 \times 10^{-9} m^2$$

- The size is $A = L^2$, $L = 48 \mu m$ ^{microns}

2D World

$$a) \quad \left| \frac{\partial (v_x, v_y)}{\partial (v, \theta)} \right| = \begin{pmatrix} \frac{\partial v_x}{\partial v} & \frac{\partial v_x}{\partial \theta} \\ \frac{\partial v_y}{\partial v} & \frac{\partial v_y}{\partial \theta} \end{pmatrix} = \begin{pmatrix} \cos \theta & -v \sin \theta \\ \sin \theta & v \cos \theta \end{pmatrix}$$

- The determinant is

$$v \cos^2 \theta + v \sin^2 \theta = v$$

i.e.

$$\left| \frac{\partial (v_x, v_y)}{\partial (v, \theta)} \right| dv d\theta = v dv d\theta$$

$$b) \quad \frac{\partial (x, y)}{\partial (r, \theta)} = \begin{pmatrix} \cos \theta & -r \sin \theta \\ \sin \theta & r \cos \theta \end{pmatrix}$$

- $\vec{e}_r dr = dr (\cos \theta \hat{i} + \sin \theta \hat{j})$, $|\vec{e}_r dr| = dr$
- $\vec{e}_\theta d\theta = d\theta (-r \cos \theta \hat{i} + r \sin \theta \hat{j})$, $|\vec{e}_\theta d\theta| = r d\theta$

$$\vec{e}_r \cdot \vec{e}_\theta = 0$$

- $\vec{e}_r \cdot d\vec{r}$ is the displacement \wedge caused by increasing r to $r + dr$ (see θ example)
- $\vec{e}_\theta d\theta$ is the displacement caused by increasing θ to $\theta + d\theta$, for example

$$d\vec{r} = \vec{e}_\theta d\theta = \frac{\partial \vec{r}}{\partial \theta} d\theta$$

$$\text{where } \vec{r} = x \hat{i} + y \hat{j} + z \hat{k}$$

d) • We have

$$d\mathcal{P} = C e^{-mv_x^2/2kT} dv_x e^{-mv_y^2/2kT} dv_y$$

- The constant comes from

$$\int d\mathcal{P} = 1, \quad \text{we know } \int dx \frac{e^{-x^2/2\sigma^2}}{(2\pi\sigma^2)^{1/2}} = 1$$

- Look at $e^{-V_x^2/2\sigma^2}$ with $\sigma^2 \equiv \frac{kT}{m}$. So

$$d\mathcal{P}_{\vec{v}} = \frac{e^{-V_x^2/2\sigma^2}}{(2\pi\sigma^2)^{1/2}} dv_x \frac{e^{-V_y^2/2\sigma^2}}{(2\pi\sigma^2)^{1/2}} dv_y$$

So since $dv_x dv_y = v dv d\theta$ and

$$\boxed{d\mathcal{P}_{v,\theta} = \frac{e^{-v^2/2\sigma^2}}{2\pi\sigma^2} v dv d\theta}$$

distribution of
speed and
angle

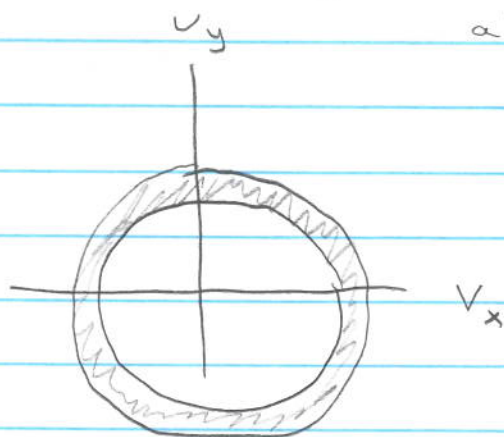
Then integrating over $\theta \in [0, 2\pi]$:

$$d\mathcal{P}_v = \int_{\theta \in [0, 2\pi]} d\mathcal{P}_{v,\theta} = \frac{e^{-v^2/2\sigma^2}}{2\pi\sigma^2} v dv \int_0^{2\pi} d\theta$$

$$\boxed{d\mathcal{P}_v = \frac{e^{-v^2/2\sigma^2}}{\sigma^2} v dv}$$

distribution of speed v

We are integrating over a shell of configurations
all of which have 'speed'
between $v + dv$



- (e) S_0

$$\left\langle \frac{1}{2} m v^2 \right\rangle = \frac{1}{2} m \int_0^{\infty} \frac{dv}{\sigma^2} e^{-v^2/2\sigma^2} v \cdot v^2$$

- Changing vars, writing $u = v/\sigma$ we have

$$\begin{aligned} \left\langle \frac{1}{2} m v^2 \right\rangle &= \frac{1}{2} m \sigma^2 \int_0^{\infty} \frac{dv}{\sigma} e^{-v^2/2\sigma^2} \frac{v}{\sigma} \frac{v^2}{\sigma^2} \\ &= \frac{1}{2} m \sigma^2 \int_0^{\infty} du e^{-u^2/2} u^3 \end{aligned}$$

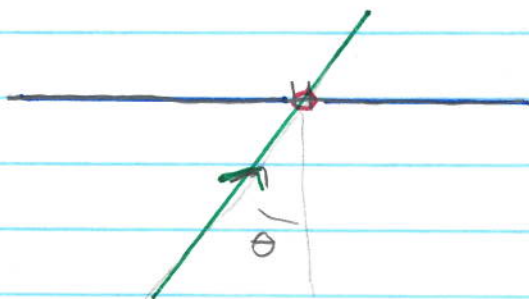
- This is very nearly the Γ fcn, let $x = \frac{u^2}{2}$
the $dx = u du$

$$\begin{aligned} \left\langle \frac{1}{2} m v^2 \right\rangle &= \frac{1}{2} m \sigma^2 \underbrace{\int_0^{\infty} dx e^{-x} \cdot 2x}_{2 \cdot 1!} \\ &= m \sigma^2 \end{aligned}$$

$$= \cancel{m} \frac{k_B T}{\cancel{m}} = k_B T = 2 \times \frac{1}{2} k_B T$$

- This is the equipartition theorem in 2D
The molecules have 2 dof, motion in x and y.

f)



We have simply (see lecture for more discussion)

$$d\Phi = n P(v) v \cos\theta \frac{v dv d\theta}{2\pi}$$

- Compute 2D and 3D

$$dv_x dv_y = v dv d\theta \quad (2D)$$

$$dv_x dv_y dv_z = v^2 dv \sin\theta d\theta d\phi \quad (3D)$$

- For particles uniformly distributed on the circle, we have

$$d\mathcal{P} = \frac{d\theta}{2\pi} \quad (2D)$$

While for the sphere, we have

$$d\mathcal{P} = \frac{d\Omega}{4\pi} \quad (3D)$$

- So in 2D

$$d\Phi = n P(v) v \cos\theta \frac{dv d\theta}{2\pi}$$

Then total flux is

$$\boxed{\Phi = \int_v \int_{\theta \in [-\pi/2, \pi/2]} n P(v) v \cos\theta \frac{dv d\theta}{2\pi}}$$



$$= n \langle v \rangle \times \int_{-\pi/2}^{\pi/2} \cos\theta \frac{d\theta}{2\pi} = \boxed{\frac{1}{\pi} n \langle v \rangle}$$

- Similarly ^{see lecture!}

Δp_x = impulse

$d\Phi$ ← flux

$$\boxed{P = \int_v \int_{\theta \in [-\pi/2, \pi/2]} (2mv \cos\theta) n P(v) v \cos\theta \frac{dv d\theta}{2\pi}}$$

$$= 2mn \int_0^\infty P(v) v^2 dv \int_{-\pi/2}^{\pi/2} \cos^2\theta \frac{d\theta}{2\pi}$$

$$= 2mn \langle v^2 \rangle \cdot \frac{1}{4}$$

$$= \boxed{\frac{1}{2} mn \langle v^2 \rangle}$$