## Derivation of the Maxwell-Boltzmann distribution

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## Introduction

The Maxwell-Boltzmann distribution of particle speeds is derived using the kinetic theory of gases.

## Derivation

Particle motion is random and so the velocity components are independent of one another. Following Maxwell, the probability some particle has velocities between  $v_x$  and  $v_x + dv_x$ ,  $v_y$  and  $v_y + dv_y$  and between  $v_z$  and  $v_z + dv_z$  is then:

$$p_{v_x}(v_x)p_{v_y}(v_y)p_{v_z}(v_z)dv_xdv_ydv_z \tag{1}$$

Since the motion of the particles is random velocity distributions are symmetric with respect to the origin and therefore it follows that the probability some particle has velocities between  $v_x$  and  $v_x + dv_x$ ,  $v_y$  and  $v_y + dv_y$  and between  $v_z$  and  $v_z + dv_z$  should depend only on speed v:

$$p_{v_x}(v_x)p_{v_y}(v_y)p_{v_z}(v_z) = p(v)$$
(2)

With speed v given by:

$$v = \sqrt{v_x^2 + v_y^2 + v_z^2} \tag{3}$$

It is reasonable to assume that the velocity components are normally distributed. Based on the reasoning and the requirement that solutions are finite one can deduce that the solution of (2) is:

$$Ce^{-Kv_x^2}Ce^{-Kv_y^2}Ce^{-Kv_z^2} = C^3e^{-Kv^2}$$
(4)

With:

$$p_{v_i}(v_i) = Ce^{-Kv_i^2} \tag{5}$$

And:

$$p(v) = C^3 e^{-Kv^2} (6)$$

To facilitate the following steps of the derivation a change of variables is made:

$$x \leftarrow v_x$$

$$y \leftarrow v_y$$

$$z \leftarrow v_z$$

$$r \leftarrow v$$
(7)

Equation (2) then becomes:

$$p_x(x)p_y(y)p_z(z) = p(r)$$
(8)

To determine the constant C the integral of (1) over the whole velocity space is equated with 1 and the resulting equation is solved for C:

$$1 = \int_{-\infty}^{\infty} p_x(x) p_y(y) p_z(z) dx dy dz$$

$$= C^3 \int_{-\infty}^{\infty} e^{-K(x^2 + y^2 + z^2)} dx dy dz$$
(9)

The integral in (9) is more readily computed in spherical coordinates:

$$1 = C^{3} \int_{0}^{2\pi} d\theta \int_{0}^{\pi} \sin(\phi) d\phi \int_{0}^{\infty} e^{-Kr^{2}} r^{2} dr$$

$$= 4\pi C^{3} \int_{0}^{\infty} r^{2} e^{-Kr^{2}} dr = \int_{0}^{\infty} \rho_{r} dr$$

$$= C^{3} \left(\frac{\pi}{K}\right)^{3/2}$$
(10)

Where  $\phi$  is the polar angle,  $\theta$  the azimutal angle and  $\rho_r$  the probability density. The constant C is therefore:

$$C = \sqrt{\frac{K}{\pi}} \tag{11}$$

Which means that the probability density  $\rho_r$  is:

$$\rho_r = 4\pi r^2 \left(\frac{K}{\pi}\right)^{3/2} e^{-Kr^2} \tag{12}$$

The average squared speed of the system of particles  $\langle r^2 \rangle$  is then:

$$\langle r^2 \rangle = \int_0^\infty r^2 \rho_r dr = \frac{3}{2K} \tag{13}$$

From kinetic theory and experimental observations the temperature and average particle speed of the system are related by:

$$kT = \frac{1}{3}m\langle r^2 \rangle \tag{14}$$

With m the mass of a particle and k the Boltzmann constant. Which means that the constant K is equal to:

$$K = \frac{m}{2kT} \tag{15}$$

The probability density  $\rho_r$  then becomes:

$$\rho_r = \left(\frac{m}{2kT\pi}\right)^{3/2} e^{-\frac{mr^2}{2kT}} 4\pi r^2 \tag{16}$$

Converting back to the original variables  $(v \leftarrow r)$  equation (16) becomes:

$$\rho_v = \left(\frac{m}{2kT\pi}\right)^{3/2} e^{-\frac{mv^2}{2kT}} 4\pi v^2 \tag{17}$$

Which is the Maxwell-Boltzmann distribution.