

Growing the Seeds of Cosmic Dawn: Environmental Impact on Stellar Mass Assembly

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UCDAVIS

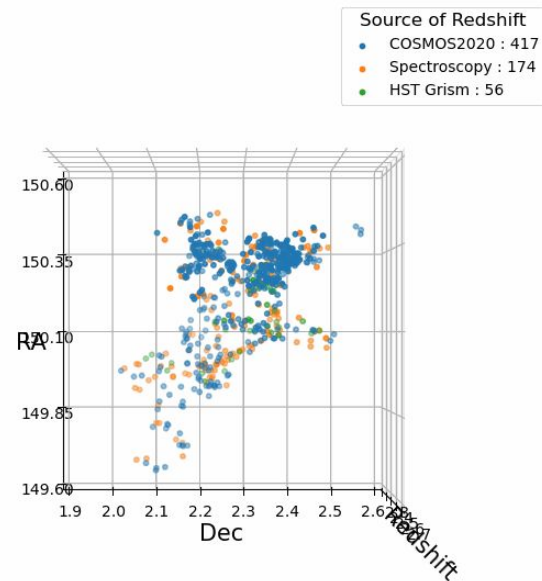
DEPARTMENT OF PHYSICS AND ASTRONOMY



Goals of the Study

Is stellar mass assembly different in overdense environments at high redshifts?

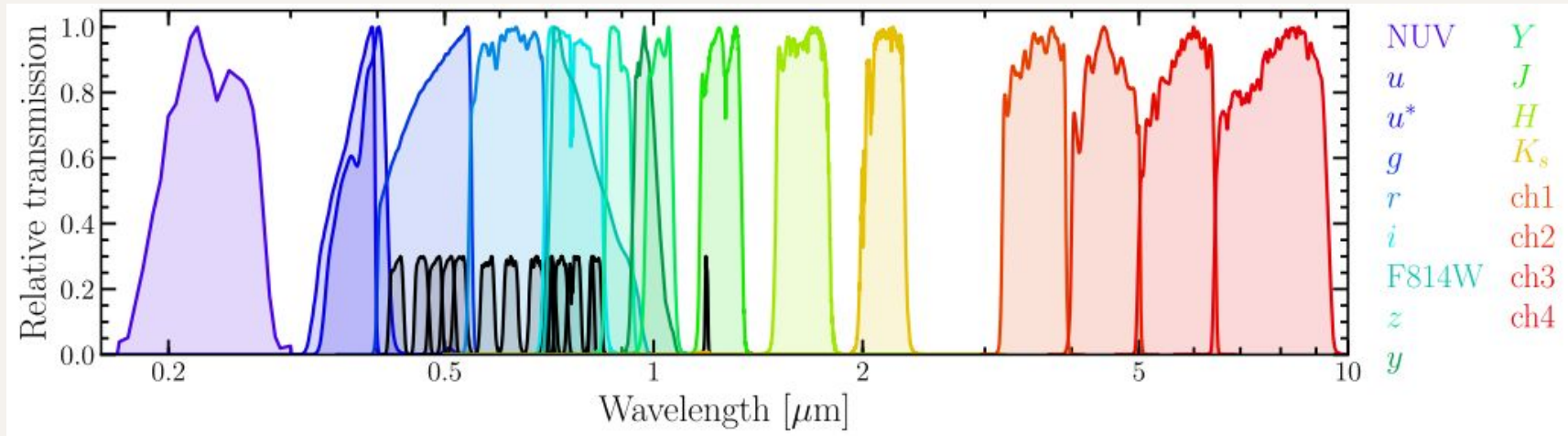
- Examine the stellar mass function (SMF) of Hyperion and a comparison field
- Determine if SFH of galaxies differ in overdense versus underdense regions



Data

1. COSMOS2020 photometry

→ Weaver et al., 2022



Data

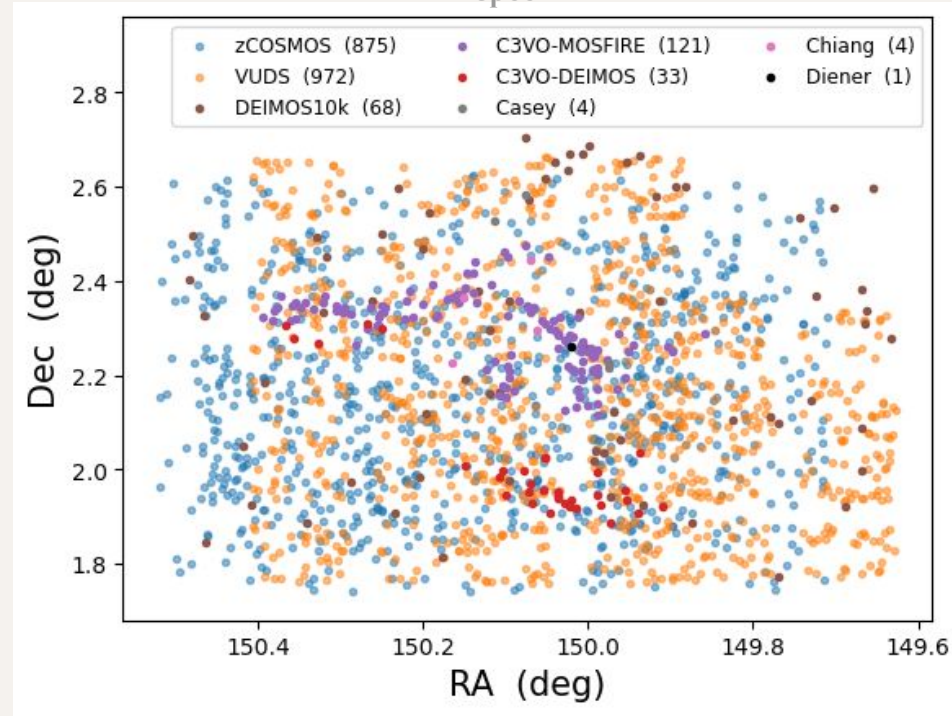
1. COSMOS2020 photometry

2. Ground-based spectroscopy

→ 2078 usable redshifts in range

$$2 \leq z_{\text{spec}} \leq 3$$

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Data

1. COSMOS2020 photometry

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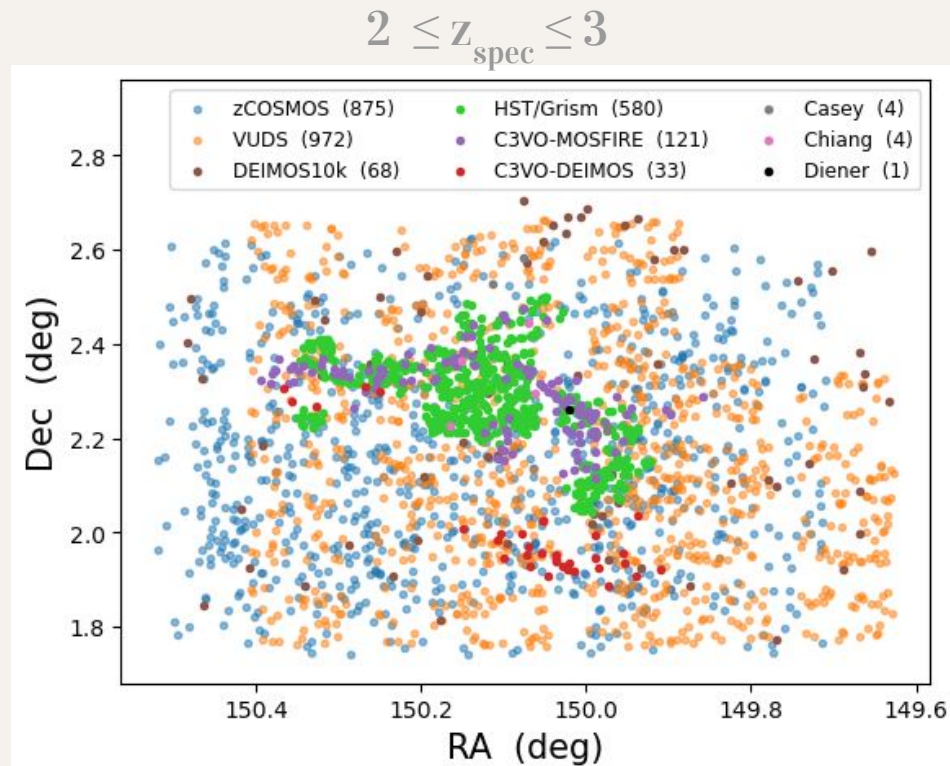
→ 2078 usable redshifts in range

$$2 \leq z_{\text{spec}} \leq 3$$

3. *HST* grism spectroscopy

→ 580 usable redshifts in range

$$2 \leq z_{\text{spec}} \leq 3$$



General Method

Generate 100 Monte Carlo Realizations

General Method

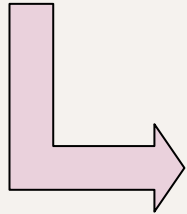
Generate 100 Monte Carlo Realizations

Draw Redshifts

General Method

Generate 100 Monte Carlo Realizations

Draw Redshifts



**Fit Stellar Mass
with LePhare**

General Method

Generate 100 Monte Carlo Realizations

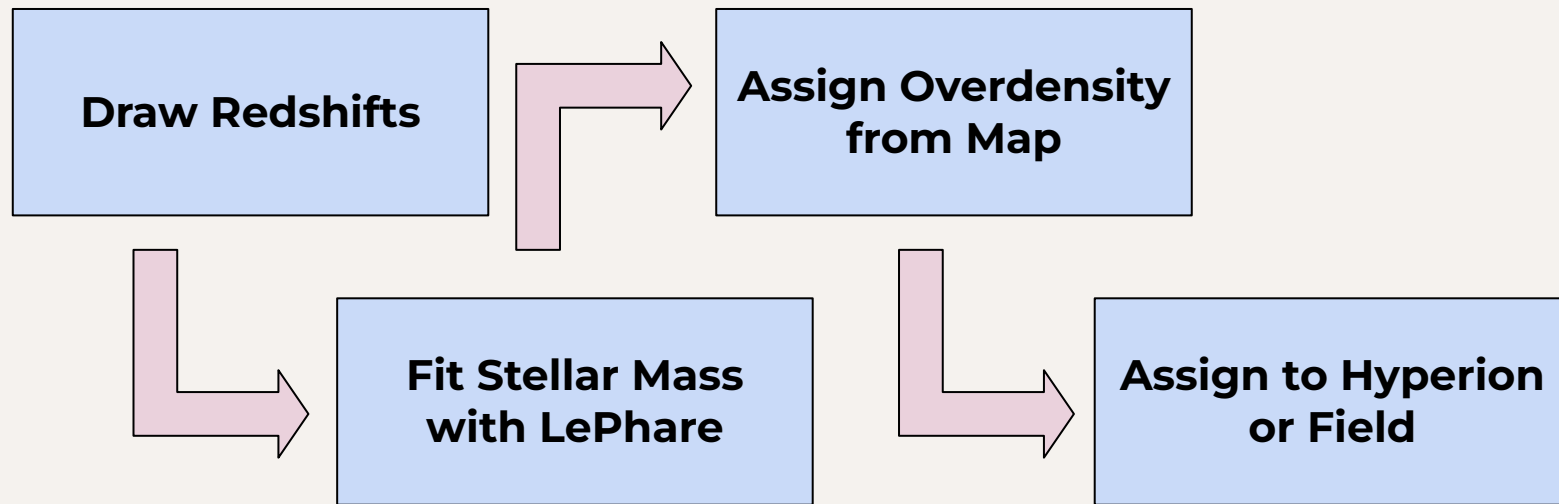
Draw Redshifts

**Assign Overdensity
from Map**

**Fit Stellar Mass
with LePhare**

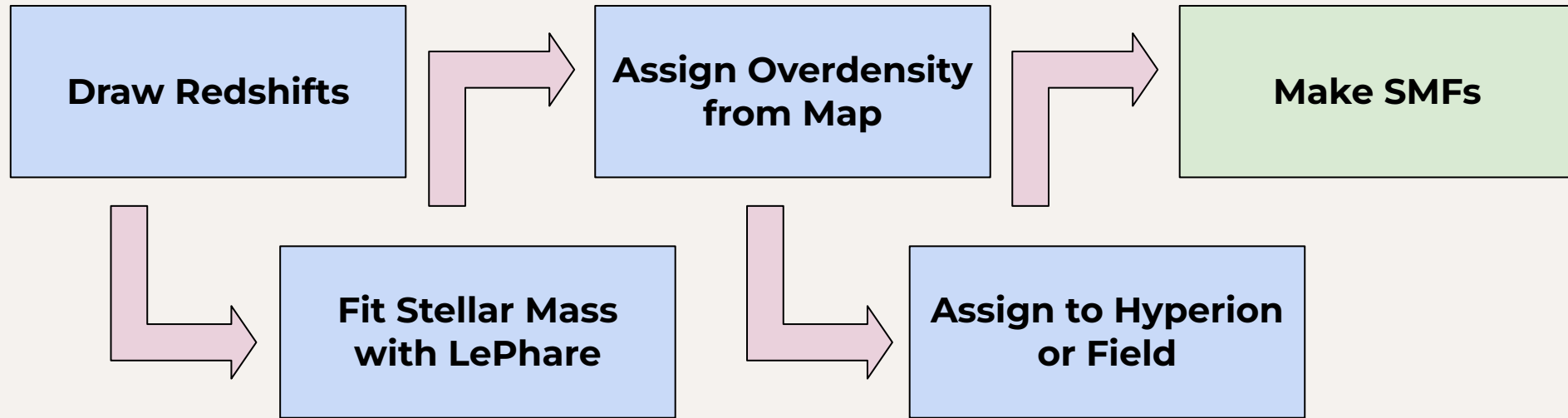
General Method

Generate 100 Monte Carlo Realizations



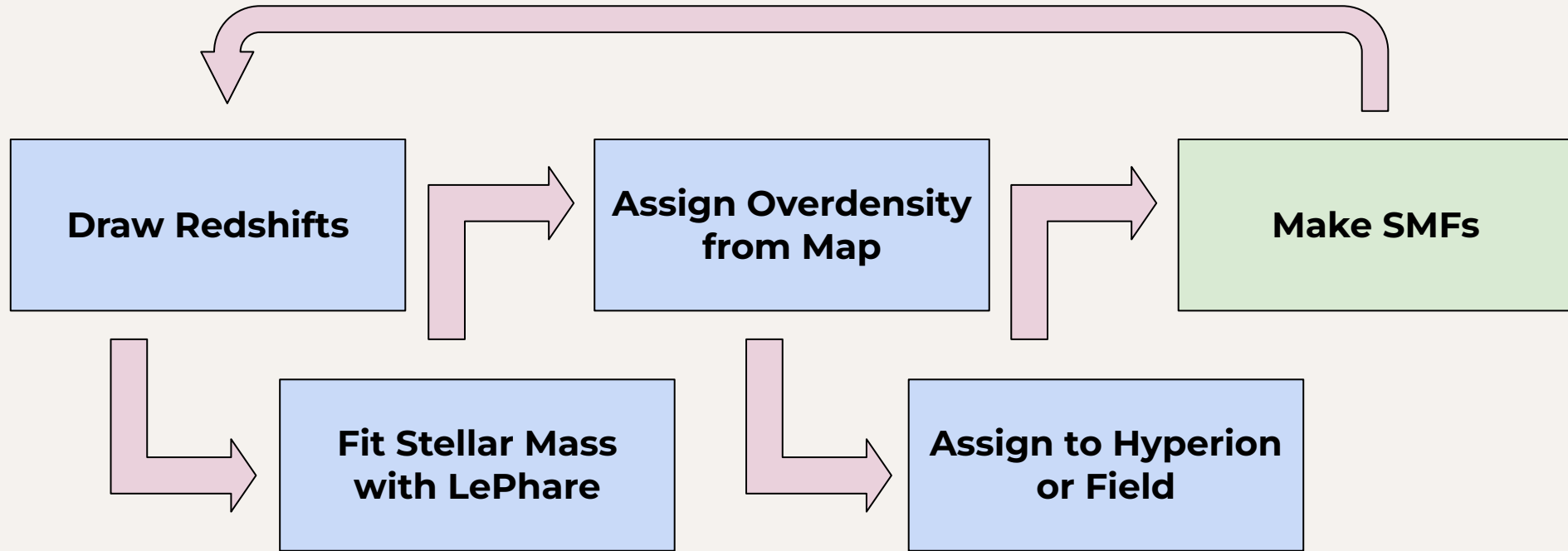
General Method

Generate 100 Monte Carlo Realizations

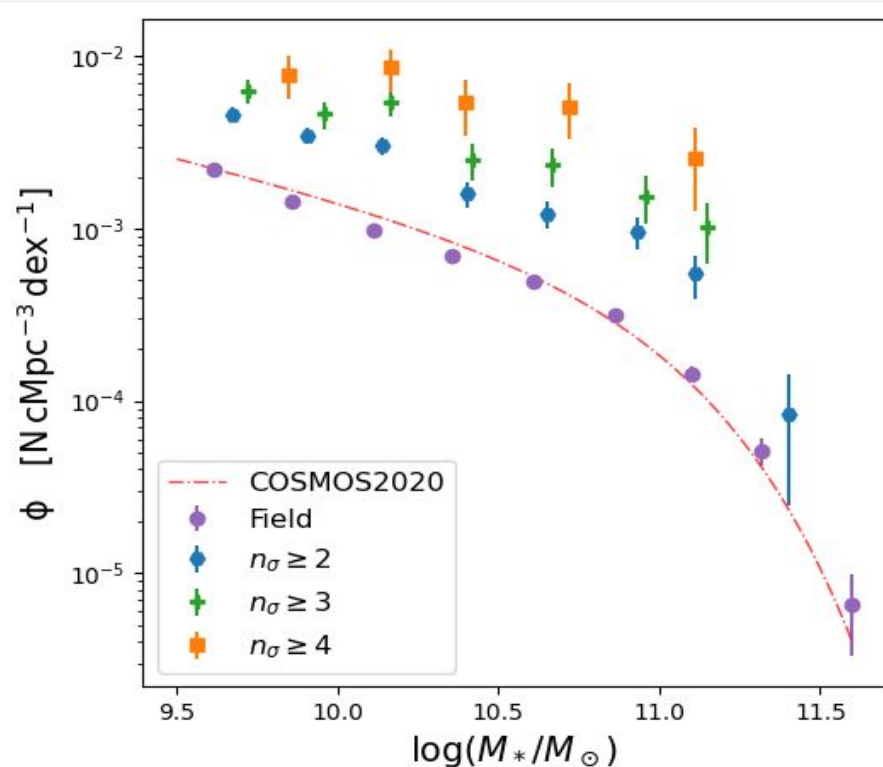


General Method

Generate 100 Monte Carlo Realizations

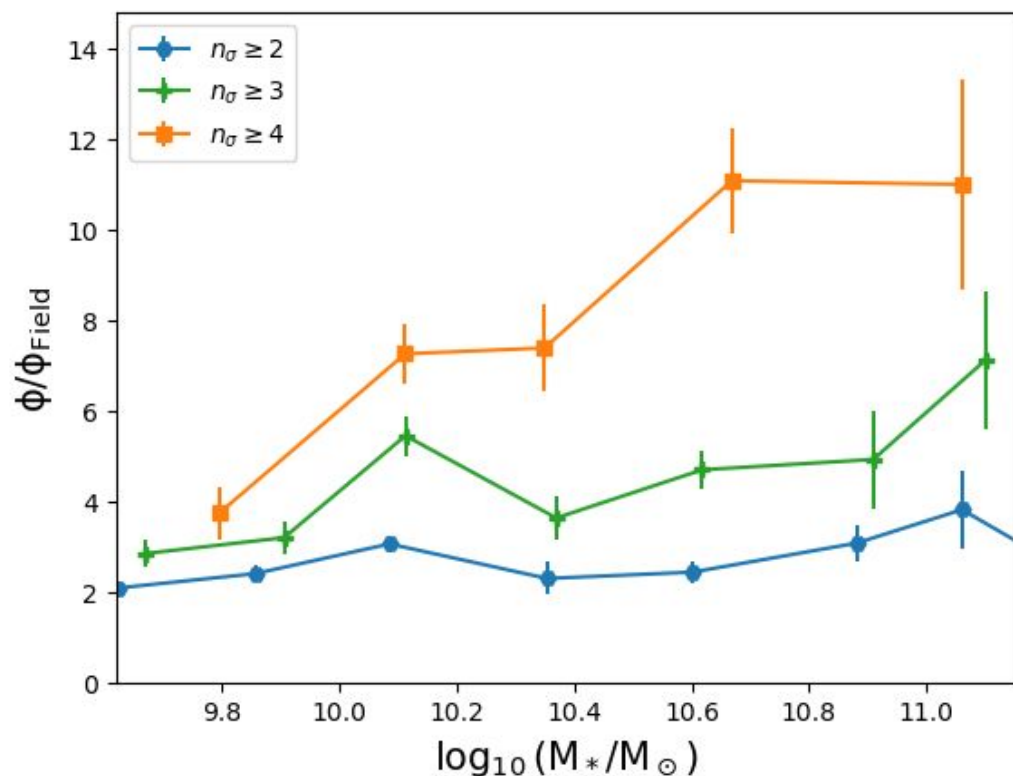


Stellar Mass Functions



- Compare 5 SMFs:
 - The combined field sample
 - Three different overdensity thresholds of Hyperion
 - The COSMOS2020 field for $2.0 \leq z \leq 2.5$ (Weaver et al., 2023)

Normalized SMFs



There is an abundance of massive galaxies in the most overdense regions

Could mean:

- Galaxies form at earlier epochs in these regions
- Galaxies experience enhanced SFRs in overdense regions
- Mergers are more frequent and driving up stellar mass (see talk by Finn Giddings).

Stellar Mass Functions

