



Coordinating Filters for Faster Deep Neural Networks



Wei Wen¹, Cong Xu², Chunpeng Wu¹, Yandan Wang³, Yiran Chen¹, Hai Li¹ Duke University¹, HP Labs², University of Pittsburgh³

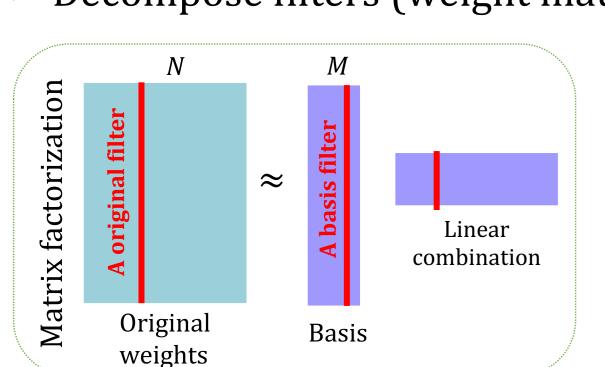
Background

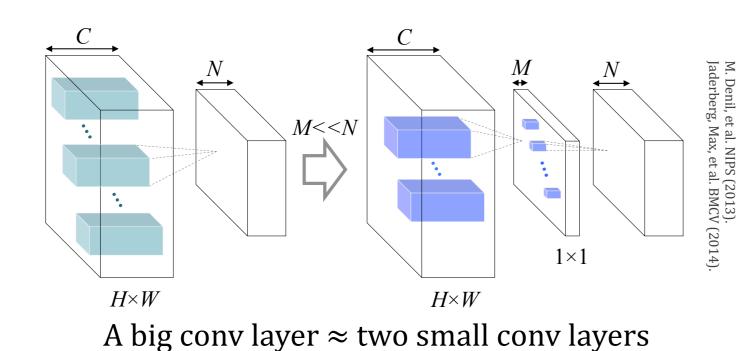
Goal

- ✓ Speedup the inference of Deep Neural Networks (DNNs)
- ✓ Focus on convolutional layers in deep neural networks

Low Rank Approximation (LRA) of Deep Neural Networks

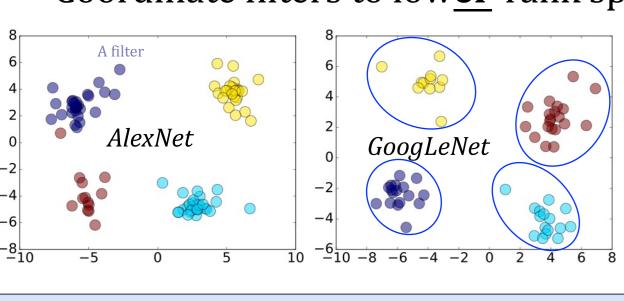
- ✓ Filters are redundant and highly correlated with each other
- ✓ Decompose filters (weight matrices) to low-rank space





This Work

✓ Coordinate filters to lower-rank space such that LRA has more compact DNNs



- Figure: Projected conv1 filters to 2D space by Linear Discriminant Analysis for visualization Goal: Coordinate a cluster of filters closer to each other (or even merge multiple clusters to one) An example: use each mean filter to approximate a cluster
- Closer filters in a cluster -> more accurate LRA
- Fewer clusters -> fewer mean filters (lower rank)

Method (Force Regularization)

Motivation

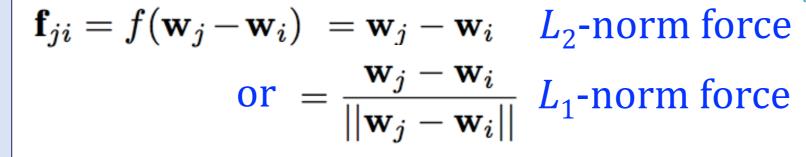
- ✓ Suppose the vector of a filter (\mathbf{W}_i) is a star in the universe
- ✓ There is pairwise gravity (\mathbf{f}_{ii}) between stars
- ✓ Gravity forces tend to pull stars closer
- ✓ Inertia resists stars to completely collapse

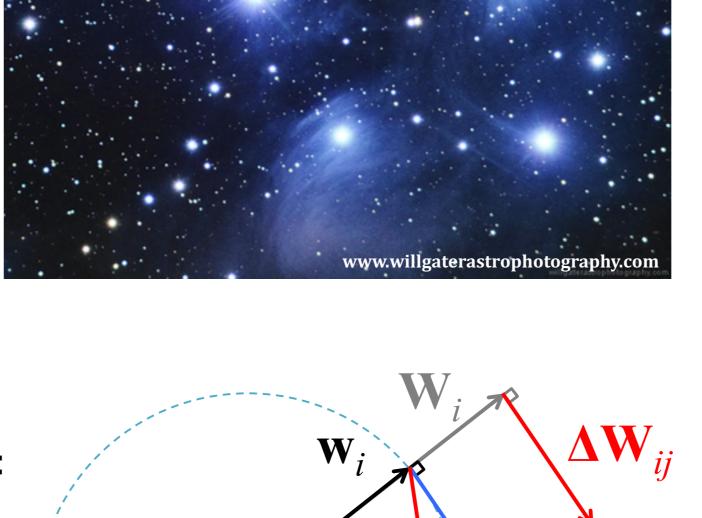
Force Regularization

✓ Introduce additional gradients in Stochastic Gradient Descent (SGD):

$$\Delta \mathbf{W}_i = \sum_{j=1}^N \Delta \mathbf{W}_{ij} = ||\mathbf{W}_i|| \sum_{j=1}^N \left(\mathbf{f}_{ji} - \mathbf{f}_{ji}\mathbf{w}_i^T\mathbf{w}_i \right)$$

Forces from all other stars/filters





SGD Training with Force Regularization

✓ Filters are updated by both loss function gradients and force gradients:

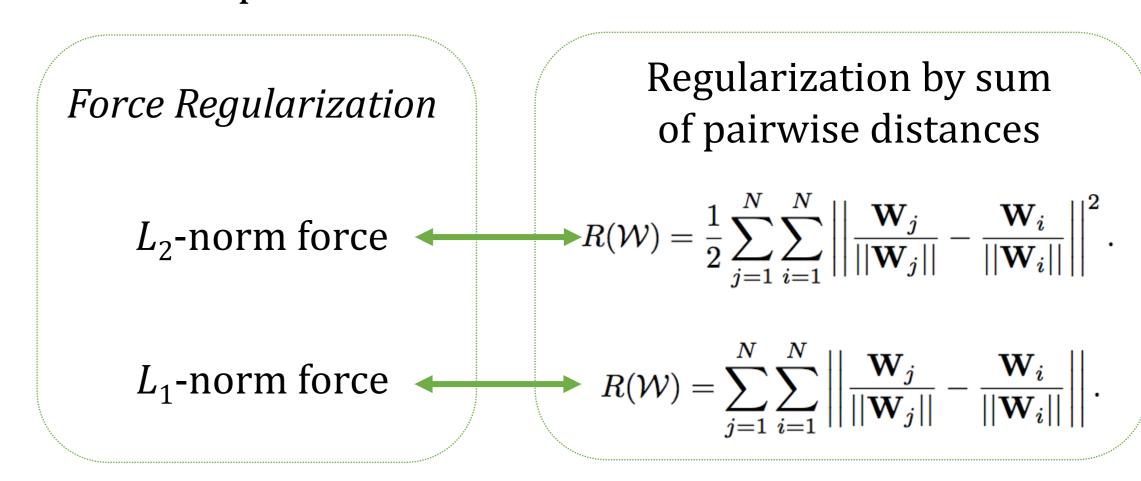
$$\mathbf{W}_{i} \leftarrow \mathbf{W}_{i} - \eta \cdot \left(\frac{\partial E(\mathcal{W})}{\partial \mathbf{W}_{i}} - \lambda_{s} \cdot \Delta \mathbf{W}_{i}\right) \quad \Delta \mathbf{W}_{i} = \sum_{j=1}^{N} \Delta \mathbf{W}_{ij} = ||\mathbf{W}_{i}|| \sum_{j=1}^{N} (\mathbf{f}_{ji} - \mathbf{f}_{ji} \mathbf{w}_{i}^{T} \mathbf{w}_{i})$$

$$\text{Minimize error} \qquad \text{Reduce ranks}$$

$$(Inertia) \qquad (Gravity)$$

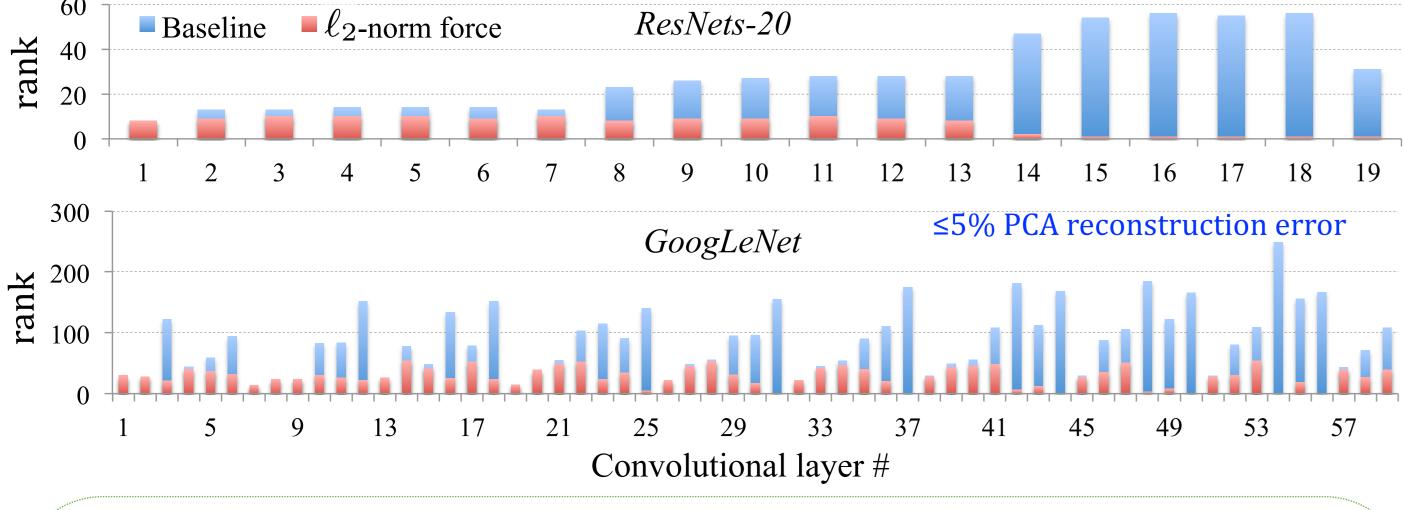
Intuitive Force Regularization has strong mathematical implications

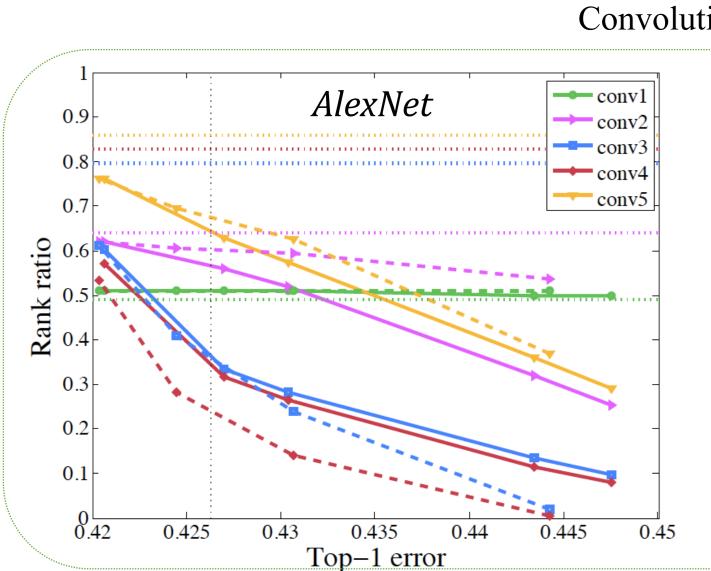
- ✓ Two types of regularization have
 - the same gradient direction, but
 - different step sizes



Experiments

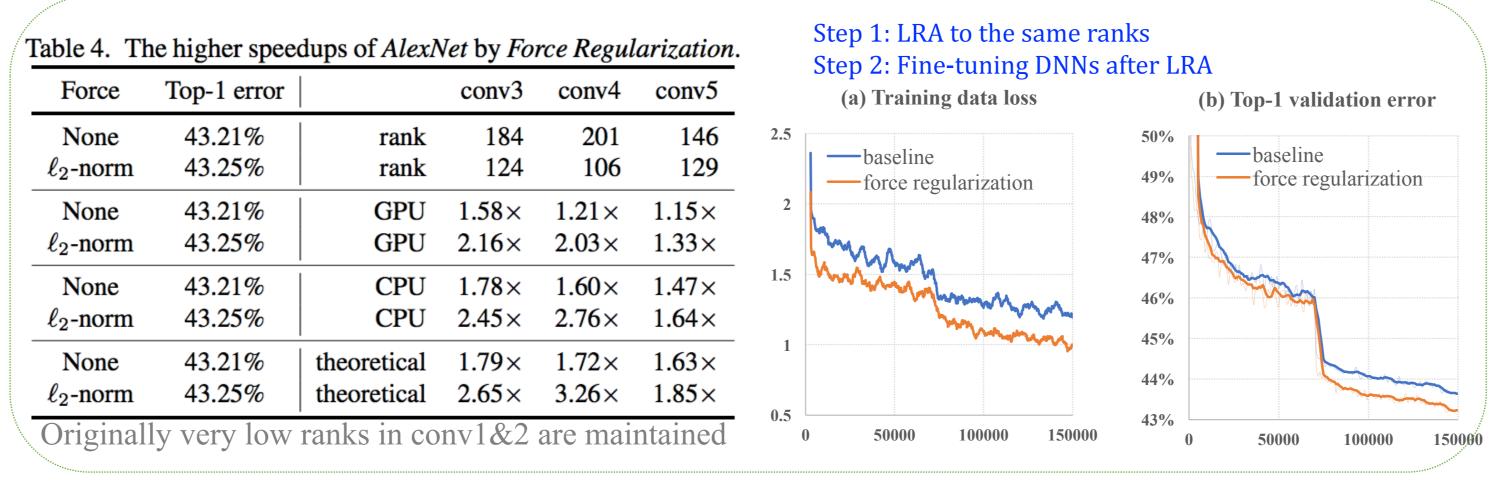
Coordinating DNNs to lower-rank space by Force Regularization

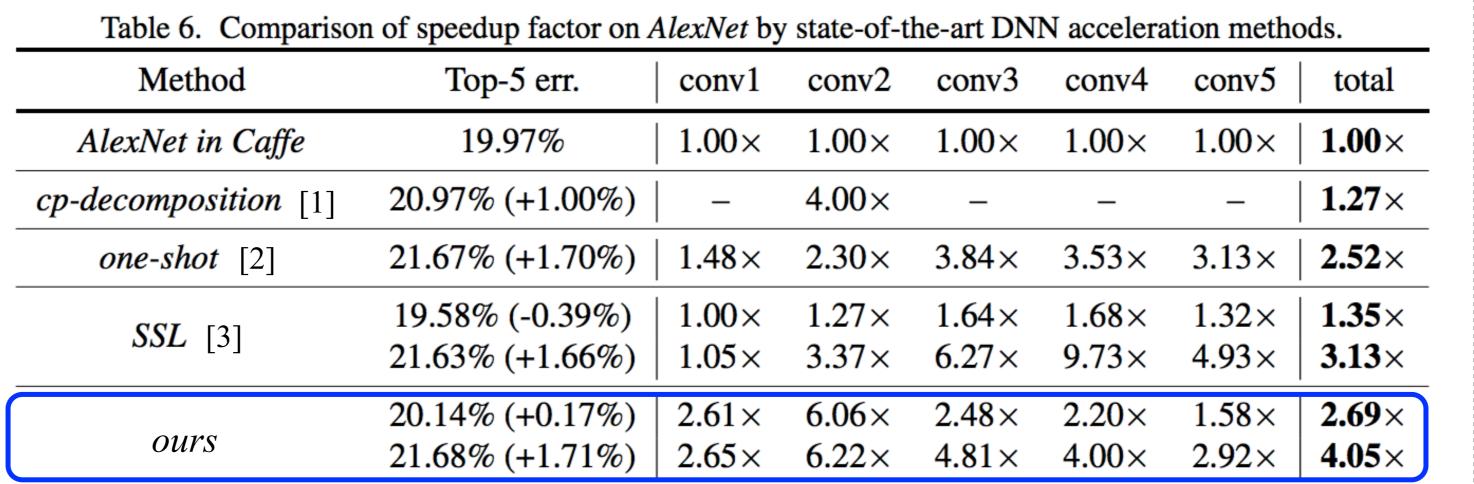




- ✓ ≤ 5% PCA reconstruction error
- ✓ Horizontal dotted lines: baseline ranks
- ✓ Vertical dotted line: baseline error
- ✓ Solid curves: L_2 -norm force
- ✓ Dashed curves: L_1 -norm force
- ✓ Control λ_s to make trade-off
- ✓ Reduce ranks without accuracy loss

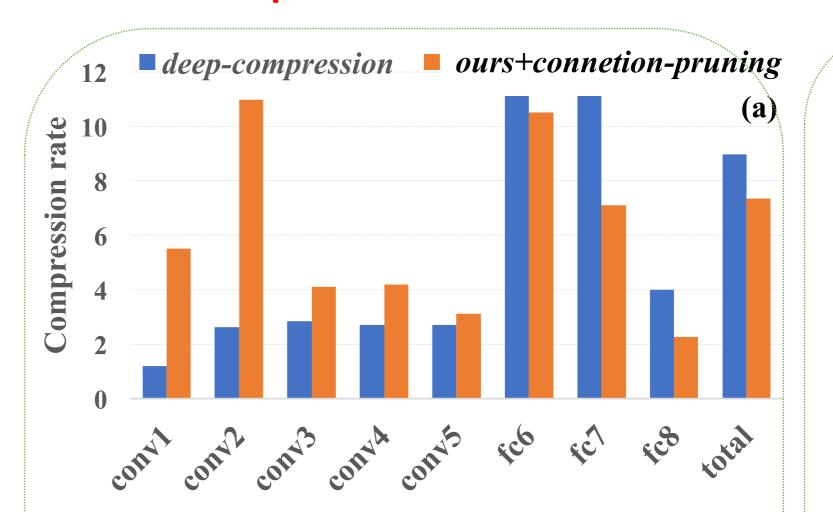
Speedup by Force Regularization





[1] V. Lebedev, et al. ICLR 2015; [2] Y.-D. Kim, et al. ICLR 2016; [3] W. Wen, et al. NIPS 2016

Lower rank + sparse DNNs



- ✓ *deep-compression*: S. Han, et al., NIPS 2015 (only counting compression from connection pruning)
- ✓ Non-structurally sparse DNNs
- ✓ Higher compression in conv layers for computation saving
- ✓ Comparable total compression rate ✓ Higher speedup (\sim 2.7x)
- ✓ conv3_s: 1st small conv3 after LRA ✓ conv3_f: 2nd small conv3 after LRA ✓ Ours can work with SSL for potentially higher speedup

conv3 5 conv4 5 conv5 5 conv3 f conv4 f

✓ SSL: W. Wen, et al., NIPS 2016

✓ Structurally sparse DNNs

Acknowledgments

This work was supported in part by NSF CCF-1744082. Any opinions, findings and conclusions or recommendations expressed in this material are those of the authors and do not necessarily reflect the views of NSF or their contractors.