dessn Documentation

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dessn

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CHAPTER

ONE

DESSN PACKAGE

Welcome to the **DESSN** code base.

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Submodules

dessn.model.edge module

class `dessn.model.edge.`**`Edge`**(*probability_of*, *given*)
Bases: `object`

An edge connection one or more parameters to one or more different parameters.

An edge is a connection between parameters (*not* Nodes), and signifies a joint probability distribution. That is, if in our mathematical definition of our model, we find the term $P(a, b|c, d, e)$, this would be represented by a single edge. Similarly, $P(a|b)P(b|c, d)$ would be two edges.

Parameters **probability_of** : str or list[str]

The dependent parameters. With the example $P(a, b|c, d)$, this input would be `['a','b']`.

given : str or list[str]

In the example $P(a, b|c, d)$, this input would be `['c','d']`.

`get_log_likelihood`(*data*)
Gets the log likelihood of this edge.

For example, if we had

$$P(a, b|c, d) = \frac{1}{\sqrt{2\pi}d} \exp\left(-\frac{(ab-c)^2}{d^2}\right),$$

we could implement this function as `return -np.log(np.sqrt(2*np.pi)*data['d']) - (data['a']*data['b'] - data['c'])**2/(data['d']**2)`

Returns float

the log likelihood given the supplied data and the model parametrisation.

class dessn.model.edge.**EdgeTransformation**(*inputs*, *transform_to*)
Bases: *dessn.model.edge.Edge*

This specialised edge is used to connect to transformation nodes.

A transformation edge does not give a likelihood, but - as it is a known transformation - returns a dictionary when *get_transformation* is invoked that is injected into the data dictionary given to regular edges.

See *LuminosityToAdjusted* for a simple example.

Parameters **inputs** : str or list[str]

The required parameter inputs to do the transformation.

transform_to : str or list[str]

The parameters that will be added to the data dictionary after transformation

get_log_likelihood(*data*)

get_transformation(*data*)
Calculates the new parameters from the given data

Returns dict

a dictionary containing a value for each parameter given in `transform_to`

dessn.model.model module

class dessn.model.model.**Model**(*model_name*)
Bases: `object`

A generalised model for use in arbitrary situations.

A model is, at heart, simply a collection of nodes and edges. Apart from simply being a container in which to place nodes and edges, the model is also responsible for figuring out how to connect edges (which map to parameters) with the right nodes, for sorting edges such that when an edge is evaluated all its required data has been generated by other nodes or edges, for managing the `emcee` running, and also for generating the visual PGMs.

It is thus a complex class, and I expect, as of writing this summary, it contains numerous bugs.

Parameters **model_name** : str

The model name, used for serialisation

add_edge(*edge*)
Adds an edge into the models collection of edges

add_node(*node*)
Adds a node into the models collection of nodes.

finalise()
Finalises the model.

This method runs consistency checks on the model (making sure there are not orphaned nodes, edges to parameters that do not exist, etc), and in doing so links the right edges to the right nodes and determines the order in which edges should be evaluated.

You can manually call this method after setting all nodes and edges to confirm as early as possible that the model is valid. If you do not call it manually, this method is invoked by the model when requesting concrete information, such as the PGM or model fits.

fit_model(*num_walkers=None*, *num_steps=5000*, *num_burn=3000*, *filename=None*, *save_interval=300*)

Uses `emcee` to fit the supplied model.

This method sets an emcee run using the `EnsembleSampler` and manual chain management to allow for very high dimension models. MPI running is detected automatically for less hassle, and chain progress is serialised to disk automatically for convenience.

This method works... but is still a work in progress

Parameters **num_walkers** : int, optional

The number of walkers to run. If not supplied, it defaults to eight times the model dimensionality

num_steps : int, optional

The number of steps to run

num_burn : int, optional

The number of steps to discard for burn in

filename : str, optional

If set, saves a corner plot to that filename in the top level plots directory.

save_interval : float

The amount of seconds between saving the chain to file. Setting to `None` disables serialisation.

Returns ndarray

The final flattened chain of dimensions `(num_dimensions, num_walkers * (num_steps - num_burn))`

fig

The corner plot figure returned from `corner.corner(...)`

get_pgm(*filename=None*)

Renders (and returns) a PGM of the current model.

Parameters **filename** : str, optional

if the filename is set, the PGM is saved to file in the top level `plots` directory.

Returns `daft.PGM`

The `daft` PGM class, for further customisation if required.

dessn.model.node module

class `dessn.model.node.`**Node**(*node_name*, *names*, *labels*, *parameter_type*)

Bases: `object`

A node represented on a PGM model. Normally encapsulated by a single parameter, or several related parameters.

The Node class can essentially be thought of as a wrapper around a parameter or variable in your model. However, as some parameters are highly related (for example, flux and flux error), Nodes allow you to declare multiple parameters.

This class is an abstract class, and cannot be directly instantiated. Instead, instantiate one of the provided subclasses, as detailed below.

Parameters **node_name** : str

The node name, only used when plotting on a PGM

names : str or list[str]

The model parameter encapsulated by the node, or list of model parameters

labels : str or list[str]

Latex ready labels for the given names. Used in the PGM and corner plots.

parameter_type : *NodeType*

The type of subclass. Informs the model how to utilise the node.

class dessn.model.node.**NodeLatent**(*node_name*, *names*, *labels*)
Bases: *dessn.model.node.Node*

A node representing a latent, or hidden, variable in our model.

Given infinitely powerful computers, these nodes would not be necessary, for they represent marginalisation over unknown / kidden / latent parameters in the model, and we would simple integrate them out when computing the likelihood probability. However, this is not the case, and it is more efficient to simply incorporate latent parameters into our model and essentially marginalise over them using Monte Carlo integration. We thus trade explicit numerical integration in each step of our calculation for increased dimensionality.

For examples on why and how to use latent parameters, see the examples beginning in *Example*.

Parameters **node_name** : str

The node name, only used when plotting on a PGM

names : str or list[str]

The model parameter encapsulated by the node, or list of model parameters

labels : str or list[str]

Latex ready labels for the given names. Used in the PGM and corner plots.

get_num_latent()
The number of latent parameters to include in the model.

Running MCMC requires knowing the dimensionality of our model, which means knowing how many latent parameters (realisations of an underlying hidden distribution) we require.

For example, if we observe a hundred supernova drawn from an underlying supernova distribution, we would have to realise a hundred latent variables - one per data point.

Returns int

the number of latent parameters required by this node

class dessn.model.node.**NodeObserved**(*node_name*, *names*, *labels*, *datas*)
Bases: *dessn.model.node.Node*

A node representing one or more observed variables

This node is used for all observables in the model. In addition to a normal node, it also contains data, which can be in arbitrary format. This data is what is given to the incoming and outgoing node edges to calculate likelihoods.

Parameters **node_name** : str

The node name, only used when plotting on a PGM

names : str or list[str]

The model parameter encapsulated by the node, or list of model parameters

labels : str or list[str]

Latex ready labels for the given names. Used in the PGM and corner plots.

datas : object or list[obj]

One data object for each supplied parameter name. **Must** be the same length as names if names is a list.

get_data()
Returns a dictionary containing keys of the parameter names and values of the parameter data object

class dessn.model.node.**NodeTransformation**(*node_name*, *names*, *labels*)
Bases: *dessn.model.node.Node*

A node representing a variable transformation.

This node essentially represents latent variables which are fully determined - their probability is given by a delta function. Examples of this might be the luminosity distance, as it is known exactly when given cosmology and redshift. Or it might represent a conversion between observed flux and actual flux, given we have a well defined flux correction.

On a PGM, this node would be represented by a point, not an ellipse.

Note that this node declares all associated parameters to be transformation parameters, although the transformation functions themselves are defined by the edges into and out of this node.

Parameters **node_name** : str

The node name, only used when plotting on a PGM

names : str or list[str]

The model parameter encapsulated by the node, or list of model parameters

labels : str or list[str]

Latex ready labels for the given names. Used in the PGM and corner plots.

class dessn.model.node.**NodeType**
Bases: enum.Enum

LATENT = <NodeType.LATENT: 3>

OBSERVED = <NodeType.OBSERVED: 2>

TRANSFORMATION = <NodeType.TRANSFORMATION: 4>

UNDERLYING = <NodeType.UNDERLYING: 1>

class dessn.model.node.**NodeUnderlying**(*node_name*, *names*, *labels*)
Bases: *dessn.model.node.Node*

A node representing an underlying parameter in your model.

On the PGM, these nodes would be at the very top, and would represent the variables we are trying to fit for, such as Ω_M.

These nodes are required to implement the abstract method `get_log_prior`

Parameters **node_name** : str

The node name, only used when plotting on a PGM

names : str or list[str]

The model parameter encapsulated by the node, or list of model parameters

labels : str or list[str]

Latex ready labels for the given names. Used in the PGM and corner plots.

get_log_prior(*data*)

Returns the log prior for the parameter.

Parameters **data** : dic

A dictionary containing all data and the model parameters being tested at a given step in the MCMC chain. For this class, if the class was instantiated with a name of "omega_m", the input dictionary would have the key "omega_m", and the value of "omega_m" at that particular step in your chain.

Returns float

the log prior probability given the current value of the parameters

1.1.3 dessn.simple package

This module is designed to give a step by step overview of a very simplified example Bayesian model.

The basic example model is laid out in the parent class *Example*, and there are three implementations. The first implementation, *ExampleIntegral*, shows how the problem might be approached in a simple model, where numerical integration is simply done as part of the likelihood calculation.

However, if there are multiple latent parameters, we get polynomial growth of the number of numerical integrations we have to do, and so this does not scale well at all.

This leads us to the implementation in *ExampleLatent*, where we use the MCMC algorithm to essentially do Monte Carlo integration via marginalisation. This means we do not need to perform the numerical integration in the likelihood calculation, however the cost of doing so is increase dimensionality of our MCMC.

Finally, the *ExampleModel* implementation shows how the *ExampleLatent* class might be written to make use of Nodes. This is done in preparation for more complicated models, which will have more than one layer and needs to be configurable.

Subpackages

dessn.simple.modelbased package

I have placed the class based example for implementing the simplified model into its own module, so that the documentation generating for the `simple` module does not get cluttered with all the small classes this module will have.

The primary class to look at in code is the *ExampleModel* class.

I should finally note that in order to demonstrate parameter transformations, I have modified the model used in the previous two examples (*ExampleIntegral* and *ExampleLatent*) to also include a luminosity transformation,

where I simply halve the luminosity before converting it to flux. Physically, this could represent a perfect 50% mirror absorption on the primary telescope mirror.

Submodules

dessn.simple.modelbased.exampleModel module

class dessn.simple.modelbased.exampleModel.**ExampleModel**

Bases: *dessn.model.model.Model*

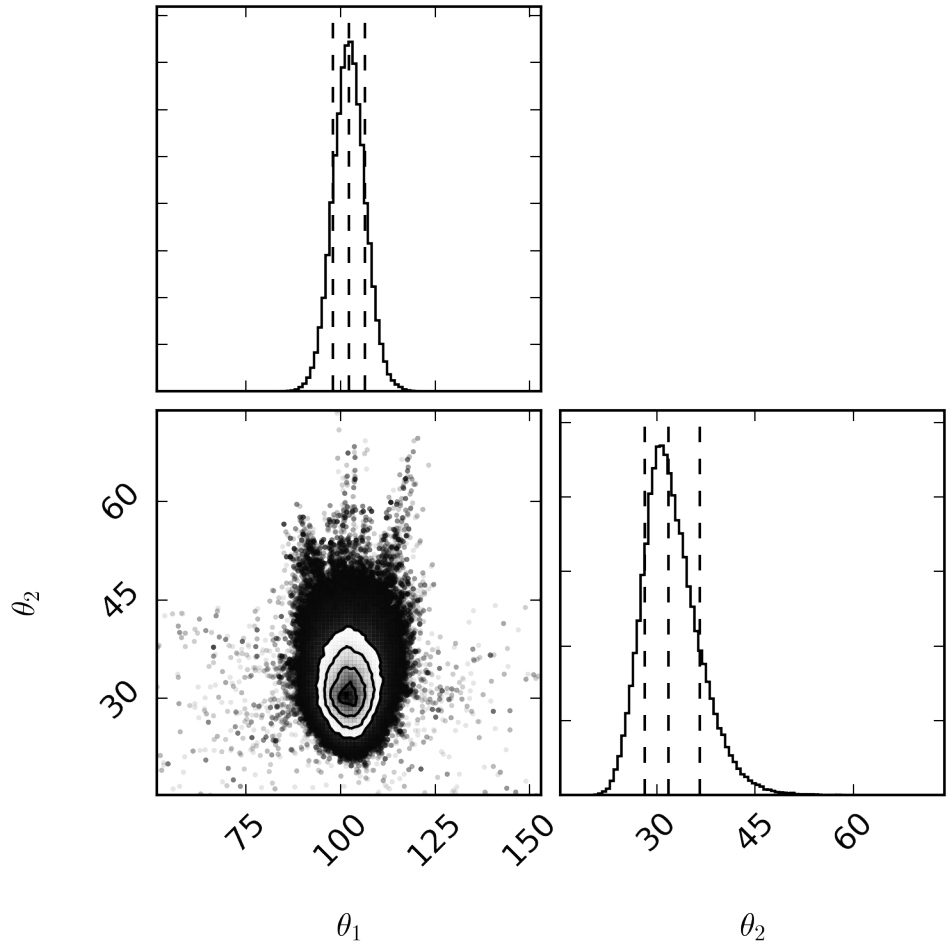
An implementation of *ExampleLatent* using classes instead of procedural code.

The model is set up by declaring nodes, the edges between nodes, and then calling `finalise` on the model to verify its correctness.

This is the primary class in this package, and you can see that other classes inherit from either *Node* or from *Edge*.

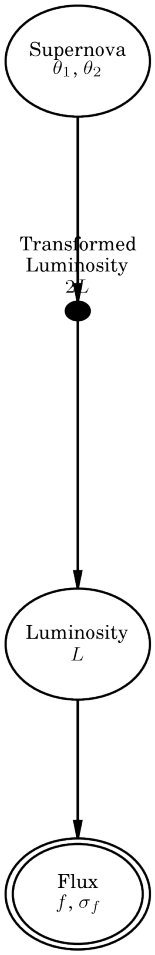
I leave the documentation for `Node`s and `:class:`.Edge`s to those classes, and encourage viewing the code directly to understand exactly what is happening.

Running this file in python first generates a PGM of the model, and then runs `emcee` and creates a corner plot:



class dessn.simple.modelbased.exampleModel.**FluxToLuminosity**

Bases: *dessn.model.edge.Edge*


Supernova
θ_1, θ_2
Transformed
Luminosity
$2L$
Luminosity
L
Flux
f, σ_f

get_log_likelihood(*data*)

class dessn.simple.modelbased.exampleModel.**LatentLuminosity**(*n=100*)
Bases: *dessn.model.node.NodeLatent*

get_num_latent()

class dessn.simple.modelbased.exampleModel.**LuminosityToAdjusted**
Bases: *dessn.model.edge.EdgeTransformation*

get_transformation(*data*)

class dessn.simple.modelbased.exampleModel.**LuminosityToSupernovaDistribution**
Bases: *dessn.model.edge.Edge*

get_log_likelihood(*data*)

class dessn.simple.modelbased.exampleModel.**ObservedFlux**(*n=100*)
Bases: *dessn.model.node.NodeObserved*

class dessn.simple.modelbased.exampleModel.**UnderlyingSupernovaDistribution**
Bases: *dessn.model.node.NodeUnderlying*

get_log_prior(*data*)

class dessn.simple.modelbased.exampleModel.**UselessTransformation**
Bases: *dessn.model.node.NodeTransformation*

Submodules

dessn.simple.example module

class dessn.simple.example.**Example**(*n=30*, *theta_1=100.0*, *theta_2=20.0*)
Bases: object

Setting up the math for some examples.

Let us assume that we are observing supernova that are drawn from an underlying supernova distribution parameterised by θ, where the supernova itself simply a luminosity L. We measure the luminosity of multiple supernovas, giving us an array of measurements D. If we wish to recover the underlying distribution of supernovas from our measurements, we wish to find $P(\theta|D)$, which is given by

$$P(\theta|D) \propto P(D|\theta)P(\theta)$$

Note that in the above equation, we realise that $P(D|L) = \prod_{i=1}^{N} P(D_i|L_i)$ as our measurements are independent. The likelihood $P(D|\theta)$ is given by

$$P(D|\theta) = \prod_{i=1}^{N} \int_{-\infty}^{\infty} P(D_i|L_i)P(L_i|\theta)dL_i$$

We now have two distributions to characterise. Let us assume both are gaussian, that is our observed luminosity x_i has gaussian error σ_i from the actual supernova luminosity, and the supernova luminosity is drawn from an underlying gaussian distribution parameterised by θ.

$$P(D_i|L_i) = \frac{1}{\sqrt{2\pi}\sigma_i} \exp\left(-\frac{(x_i - L_i)^2}{\sigma_i^2}\right)$$
$$P(L_i|\theta) = \frac{1}{\sqrt{2\pi}\theta_2} \exp\left(-\frac{(L_i - \theta_1)^2}{\theta_2^2}\right)$$

This gives us a likelihood of

$$P(D|\theta) = \prod_{i=1}^{N} \frac{1}{2\pi\theta_2\sigma_i} \int_{-\infty}^{\infty} \exp\left(-\frac{(x_i - L_i)^2}{\sigma_i^2} - \frac{(L_i - \theta_1)^2}{\theta_2^2}\right) dL_i$$

Working in log space for as much as possible will assist in numerical precision, so we can rewrite this as

$$\log\left(P(D|\theta)\right) = \sum_{i=1}^{N} \left[\log\left(\int_{-\infty}^{\infty} \exp\left(-\frac{(x_i - L_i)^2}{\sigma_i^2} - \frac{(L_i - \theta_1)^2}{\theta_2^2}\right) dL_i\right) - \log(2\pi\theta_2\sigma_i)\right]$$

Parameters **n** : int, optional

The number of supernova to 'observe'

theta_1 : float, optional

The mean of the underlying supernova luminosity distribution

theta_2 : float, optional

The standard deviation of the underlying supernova luminosity distribution

do_emcee(*nwalkers=None*, *nburn=None*, *nsteps=None*)
Abstract method to configure the emcee parameters

get_likelihood(*theta*, *data*, *error*)
Abstract method to return the log likelihood

get_posterior(*theta*, *data*, *error*)
Gives the log posterior probability given the supplied input parameters.

Parameters **theta** : array of model parameters

data : array of length *n*

An array of observed luminosities

error : array of length *n*

An array of observed luminosity errors

Returns float

the log posterior probability

get_prior(*theta*)
Get the log prior probability given the input.

The prior distribution is currently implemented as flat prior.

Parameters **theta** : array of model parameters

Returns float

the log prior probability

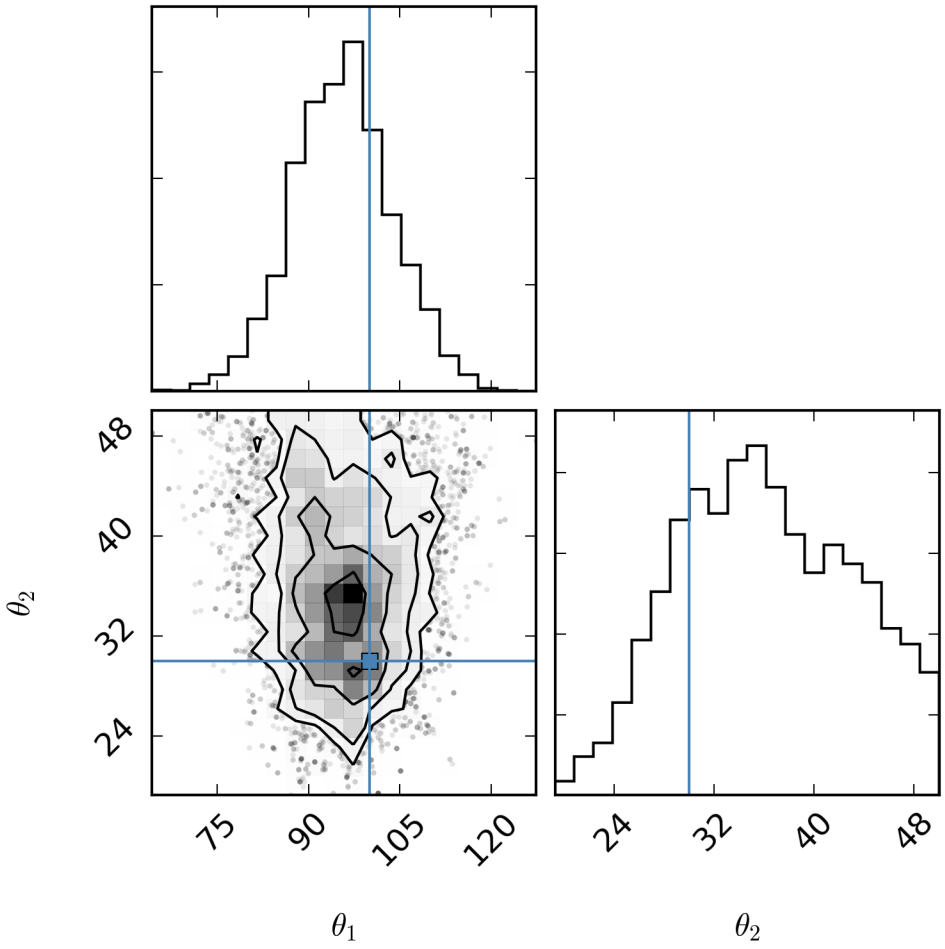
plot_observations()
Plot the observations and observation distribution.

dessn.simple.exampleIntegral module

class dessn.simple.exampleIntegral.**ExampleIntegral**(*n=10*, *theta_1=100.0*, *theta_2=30.0*)
Bases: *dessn.simple.example.Example*

An example implementation using integration over a latent parameter.

Building off the math from *Example* Creating this class will set up observations from an underlying distribution. Invoke `emcee` by calling the object. In this example, we perform the marginalisation inside the likelihood calculation, which gives us dimensionality only of two (the length of the θ array). However, this is at the expense of performing the marginalisation over dL_i, as this requires computing n integrals for each step in the MCMC.



Parameters **n** : int, optional

The number of supernova to 'observe'

theta_1 : float, optional

The mean of the underlying supernova luminosity distribution

theta_2 : float, optional

The standard deviation of the underlying supernova luminosity distribution

do_emcee (*nwalkers=20*, *nburn=2500*, *nsteps=3000*)

Run the *emcee* chain and produce a corner plot.

Saves a png image of the corner plot to plots/exampleIntegration.png.

Parameters **nwalkers** : int, optional

The number of walkers to use. Minimum of four.

nburn : int, optional

The burn in period of the chains.

nsteps : int, optional

The number of steps to run

get_likelihood(*theta*, *data*, *error*)

Gets the log likelihood given the supplied input parameters.

Parameters **theta** : array of size 2

An array representing $[\theta_1, \theta_2]$

data : array of length *n*

An array of observed luminosities

error : array of length *n*

An array of observed luminosity errors

Returns float

the log likelihood probability

dessn.simple.exampleLatent module

class dessn.simple.exampleLatent.**ExampleLatent**(*n=30*, *theta_1=100.0*, *theta_2=20.0*)

Bases: *dessn.simple.example.Example*

An example implementation using marginalisation over latent parameters.

Building off the math from *Example*, instead of performing the integration numerically in the computation of the likelihood, we can instead use Monte Carlo integration by simply setting the latent parameters $\vec{L}$ as free parameters, giving us

$$\log\left(P(D|\theta, \vec{L})\right) = -\sum_{i=1}^{N}\left[\frac{(x_i - L_i)^2}{\sigma_i^2} + \frac{(L_i - \theta_1)^2}{\theta_2^2} + \log(2\pi\theta_2\sigma_i)\right]$$

Creating this class will set up observations from an underlying distribution. Invoke emcee by calling the object. In this example, we marginalise over L_i after running our MCMC, and so we no longer have to compute integrals in our chain, but instead have dimensionality of $2 + n$, where n are the number of observations.

Parameters **n** : int, optional

The number of supernova to 'observe'

theta_1 : float, optional

The mean of the underlying supernova luminosity distribution

theta_2 : float, optional

The standard deviation of the underlying supernova luminosity distribution

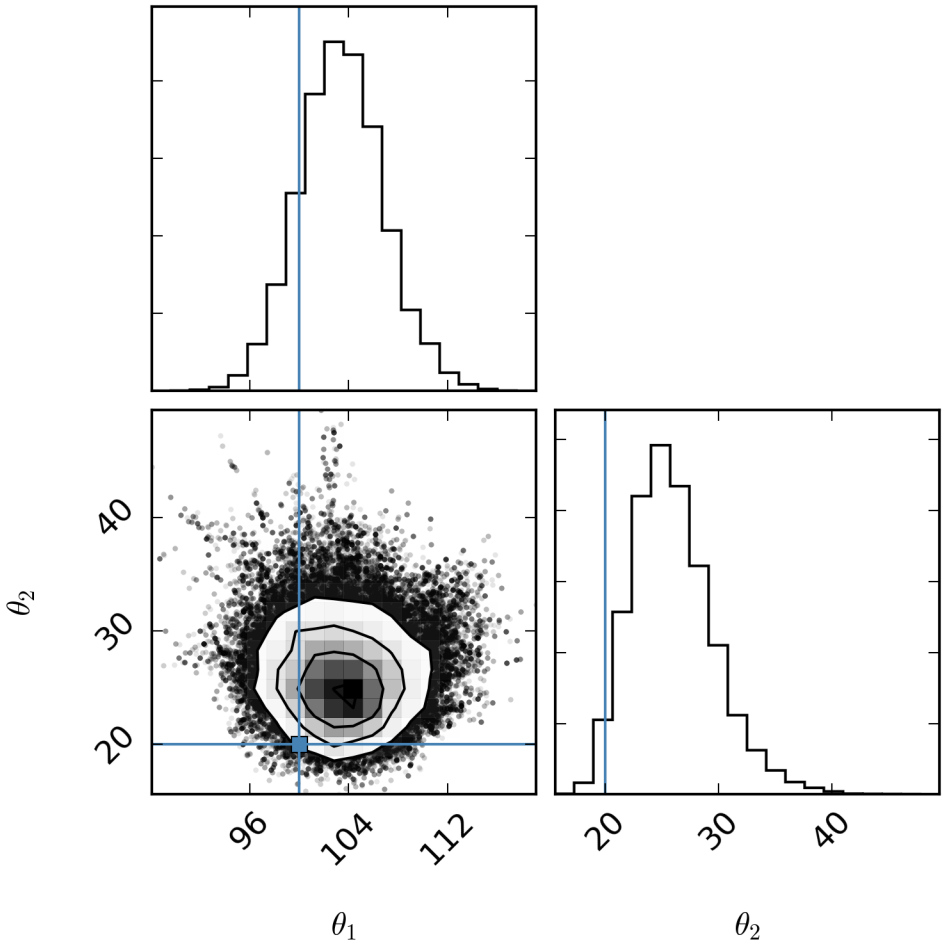
do_emcee(*nwalkers=500*, *nburn=2000*, *nsteps=2500*)

Run the *emcee* chain and produce a corner plot.

Saves a png image of the corner plot to plots/exampleLatent.png.

Parameters **nwalkers** : int, optional

The number of walkers to use.


40
30
20
θ_2
96
104
112
θ_1
20
30
40
θ_2

nburn : int, optional

The burn in period of the chains.

nsteps : int, optional

The number of steps to run

get_likelihood(*theta*, *data*, *error*)

Gets the log likelihood given the supplied input parameters.

Parameters **theta** : array of length $2 + n$

An array representing $[\theta_1, \theta_2, \vec{L}]$

data : array of length n

An array of observed luminosities

error : array of length n

An array of observed luminosity errors

Returns float

the log likelihood probability

1.1.4 dessn.simulation package

Submodules

dessn.simulation.observationFactory module

class dessn.simulation.observationFactory.**ObservationFactory**(***kwargs*)

Bases: object

check_kwargs()

get_observations(*num*)

Still needs massive refactoring

dessn.simulation.simulation module

class dessn.simulation.simulation.**Simulation**

Bases: object

get_simulation(*num_trans=30*)

1.1.5 dessn.utility package

Submodules

dessn.utility.hdemcee module

class dessn.utility.hdemcee.**EmceeWrapper**(*sampler*, *filename=None*)

Bases: object

get_results()

run_chain(*num_steps*, *num_burn*, *num_walkers*, *num_dim*, *start=None*, *save_interval=300*, *save_dim=None*)

dessn.utility.newtonian module

class dessn.utility.newtonian.**NewtonianPosition**(*nodes*, *edges*, *top=None*, *bottom=None*)

Bases: object

fit(*plot=False*)

iterate(*p*, *v*)

plot(*p*, *i*)

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