## Function GET\_MW - single-thread version

built-in abundance tables

Calling syntax:

Function parameters:

- 0. Lparms 5-element long integer array of dimensions and global (for all voxels) integer parameters (see below).
- 1. Rparms 3-element double array of global (for all voxels) real parameters (see below).
- 2. Parms array of LOS parameters,  $15 \times Nz$  elements, double. Parms[\*, i] represents the parameters for *i*th voxel (see below).
- 3. T\_arr array of temperatures where DEM/DDM are specified, NT elements, double, in K. The temperature grid is assumed to be the same in all voxels, and the same for both DEM and DDM.
- 4. DEM\_arr array of DEMs, NT × Nz, double, in cm<sup>-6</sup> K<sup>-1</sup>. DEM\_arr[\*, i] represents the DEM for ith voxel.
- 5. DDM\_arr array of DDMs, NT × Nz, double, in cm<sup>-3</sup> K<sup>-1</sup>. DDM\_arr[\*, i] represents the DDM for ith voxel.
- 6. RL input/output array,  $7 \times Nf$ , double. RL[\*, i] corresponds to *i*th frequency (see below).

Array of dimensions and global integer parameters Lparms:

Lparms = [Nz, Nf, NT, DEM\_key, DDM\_key]

- 0. Nz number of voxels along LOS;
- 1. Nf number of frequencies in the spectrum;
- 2. NT number of temperatures in the T\_arr array; must be  $\geq 2$  otherwise DEM/DEM are ignored;
  - 3. DEM\_key global DEM on/off key.
    - a. 0: DEM is enabled: it can be used in all or some voxels, depending on the local DEM on/off keys (see below).
    - b.  $\neq$  0: DEM is disabled for all voxels, regardless of the local DEM on/off keys.
  - 4. DDM\_key global DDM on/off key: same as above, but for DDM.

Array of global real parameters Rparms:

Rparms =  $[S, f_0, \Delta f]$ 

- 0. S visible source area, in cm<sup>2</sup>.
- 1.  $f_0$  starting frequency of the spectrum, in Hz:
  - a. is used, only if  $f_0 > 0$ ;
  - b. if  $f_0 \le 0$ , the frequencies are taken from the RL[0, \*] array.
- 2.  $\Delta f$  logarithmic frequency step (is used only if  $f_0 > 0$ ).

Array of parameters Parms (for a single voxel, 15 parameters):

- 0. Parms[0] =  $\Delta z$  voxel length, in cm.
- 1. Parms[1] =  $T_0$  plasma temperature, in K (is used if DEM or DDM are not specified).
- 2. Parms[2] =  $n_0$  either electron concentration or total atomic concentration (depending on other parameters), in cm<sup>-3</sup> (is used if DEM or DDM are not specified).
  - 3. Parms[3] = B magnetic field strength, in G.
  - 4. Parms[4] =  $\theta$  viewing angle, in degrees.
  - 5. Parms[5] =  $\varphi$  magnetic field azimuthal angle, in degrees.
  - 6. Parms[6] emission mechanism flag (rounded to the nearest integer):
    - a. 0: all emission mechanisms (gyroresonance + free-free + contribution of neutrals) are included;
    - b. 1: gyroresonance is off;
    - c. 2: free-free is off;
    - d. 4: contribution of neutrals is off;
    - e. 8: even if DEM and/or DDM are present, the free-free and gyroresonance emissions are computed using the isothermal approximation with the electron concentration and temperature derived from the DDM or DEM (from the DDM, if both are specified).

Several flags can be combined by usual or bitwise summation: e.g.,  $mechanism\ flag = 2 + 4\ turns\ off\ both\ free-free\ and\ contribution\ of\ neutrals,\ etc.$ 

- 7. Parms[7] =  $s_{\text{max}}$  maximum cyclotron harmonic number.
- 8. Parms[8] =  $n_p$  proton concentration, in cm<sup>-3</sup> (is used if DEM or DDM are not specified, and the temperature is low).
  - 9. Parms[9] =  $n_{\rm HI}$  neutral hydrogen concentration, in cm<sup>-3</sup>.
  - 10. Parms[10] =  $n_{\text{HeI}}$  neutral helium concentration, in cm<sup>-3</sup>.
  - 11. Parms[11] local DEM on/off key:
    - a. 0: DEM is used (provided that  $NT \ge 2$  and DEM is enabled by the global key);
    - b.  $\neq$  0: DEM in this voxel is ignored even if it is specified;  $T_0$  and  $n_0$  are used instead.
  - 12. Parms[12] local DDM on/off key: same as above, but for DDM.
  - 13. Parms[13] element abundance model:
    - a. 0: coronal (default);
    - b. 1: photospheric (Caffau);
    - c. 2: photospheric (Scott).
- 14. Parms[14] = Vox\_ID voxel type (coronal / chromospheric / etc.), currently ignored.

Input/output array RL:

0. First row (RL[0, \*]) – emission frequencies, in GHz. On input, this array is used if  $f_0$  = Rparms[1]  $\leq$  0; otherwise, the frequencies are computed using the  $f_0$  and  $\Delta f$  parameters:  $f_1 = f_0 10^{\Delta f}$ ,  $f_2 = f_1 10^{\Delta f}$ , etc. On output, this array contains the computed or pre-defined emission frequencies.

Other rows – emission intensities, as observed from the Earth, in sfu:

- 1. RL[1, \*] left polarization, weak mode coupling;
- 2. RL[2, \*] right polarization, weak mode coupling;
- 3. RL[3, \*] left polarization, strong mode coupling;
- 4. RL[4, \*] right polarization, strong mode coupling;
- 5. RL[5, \*] left polarization, exact mode coupling.
- 6. RL[6, \*] right polarization, exact mode coupling.

On input, these arrays specify the emission intensities at the start of the line-of-sight; on output, they contain the emission intensities at the end of the line-of-sight.

Return value: currently, -1 if the input was incorrect (incorrect number of parameters); 0 otherwise.

## Function GET\_MW - single-thread version

user-defined abundance tables

Calling syntax:

Function parameters:

- 0. Lparms 8-element long integer array of dimensions and global (for all voxels) integer parameters (see below).
- 1-5. Rparms, Parms, T\_arr, DEM\_arr, DDM\_arr same as in the version with built-in abundance tables; for Parms, see a note below.
- 6. fzeta\_arr array of frequencies where the  $\zeta$ -function is specified, Nf\_zeta elements, double, in Hz. The frequency grid is assumed to be the same in all voxels and for all supplied abundance sets.
- 7. Tzeta\_arr array of temperatures where the  $\zeta$ -function is specified, NT\_zeta elements, double, in K. The temperature grid is assumed to be the same in all voxels and for all supplied abundance sets.
- 8. zeta\_arr array of  $\zeta$ -function values, Nf\_zeta × NT\_zeta × N\_zeta elements, double. This array is the same for all voxels. zeta\_arr[\*, \*, m] represents the 2D  $\zeta$ -function table for mth abundance set.
  - 9. RL same as in the version with built-in abundance tables.

Array of dimensions and global integer parameters Lparms:

Lparms = [Nz, Nf, NT, DEM\_key, DDM\_key, Nf\_zeta, NT\_zeta, N\_zeta]

- 0-4. Nz, Nf, NT, DEM\_key, DDM\_key same as in the version with built-in abundance tables.
  - 5. Nf\_zeta number of frequencies where the  $\zeta$ -function is specified.
  - 6. NT\_zeta number of temperatures where the  $\zeta$ -function is specified.
  - 7. N\_zeta number of supplied 2D  $\zeta$ -function tables (abundance sets).

Array of parameters Parms (for a single voxel):

Most of parameters are the same as in the version with built-in abundance tables, except Parms[13] that specifies the element abundance model index m, so that the 2D  $\zeta$ -function table given by zeta\_arr[\*, \*, m] is used in this voxel.

Return value: currently, -1 if the input was incorrect (incorrect number of parameters); 0 otherwise.

## Function GET\_MW\_SLICE - multi-thread version

built-in abundance tables

Calling syntax:

res = call\_external(libname, 'GET\_MW\_SLICE', Lparms\_M, Rparms\_M, \$
Parms\_M, T\_arr, DEM\_arr\_M, DDM\_arrM, RL\_M)

Function parameters:

- 0. Lparms\_M 6-element long integer array of dimensions and global (for all voxels and all LOSs) integer parameters (see below).
- 1. Rparms\_M array of real parameters common for all voxels within each LOS, 3 × Npix, double (see below).
- 2. Parms\_M array of voxel parameters,  $15 \times Nz \times Npix$  elements, double (see below).
- 3. T\_arr array of temperatures where DEM/DDM are specified, NT elements, double, in K. This parameter is the same as in the GET\_MW function: the temperature grid is assumed to be the same in all voxels and all LOSs, and the same for both DEM and DDM.
  - 4. DEM\_arr\_M array of DEMs, NT × Nz × Npix, double, in cm<sup>-6</sup> K<sup>-1</sup> (see below).
- 5. DDM\_arr\_M array of DDMs, NT × Nz × Npix, double, in cm<sup>-3</sup>  $K^{-1}$  (see below).
  - 6. RL\_M input/output array,  $7 \times Nf \times Npix$ , double (see below).

Array of dimensions and global integer parameters Lparms\_M: Lparms\_M = [Npix, Nz, Nf, NT, DEM\_key, DDM\_key]

0. Npix – number of LOSs.

Other elements ( $1^{st}$  to  $5^{th}$ ) are the same as the  $0^{th}$  to  $4^{th}$  elements of the Lparms array in the GET MW function. In particular:

- a. all LOSs have the same number of voxels Nz;
- b. the number of frequencies Nf is the same for all LOSs;
- c. the global DEM and DDM on/off keys are related to all voxels within all LOSs.

Other parameters: sub-arrays Rparms\_M[\*, i], Parms\_M[\*, \*, i], DEM\_arr\_M[\*, \*, i], DDM\_arr\_M[\*, \*, i] and RL\_M[\*, \*, i] correspond respectively to the parameters Rparms, Parms, DEM\_arr, DDM\_arr and RL of the single-thread GET\_MW function, for ith LOS.

Return value: currently, -1 if the input was incorrect (incorrect number of parameters); 0 otherwise.

Function GET\_MW\_SLICE – multi-thread version for user-defined abundance tables: not implemented yet.