

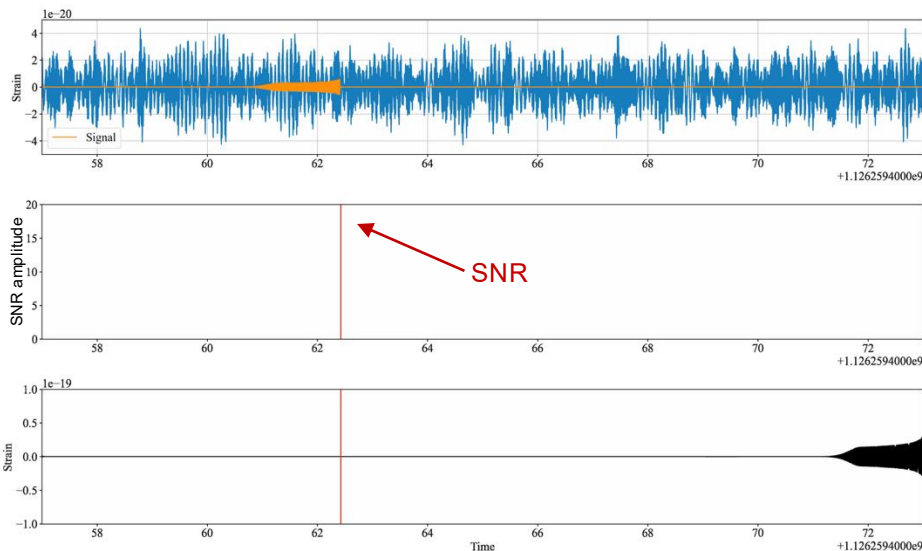


# Neural network classifiers for distinguishing signals from instrumental noise

Melissa Lopez  
ICERM 2025

# Modelled searches: matched filtering (MF) for CBC

What is matched filtering?



Unknown signal

\*

Template  
Bank  
(simulated GW)

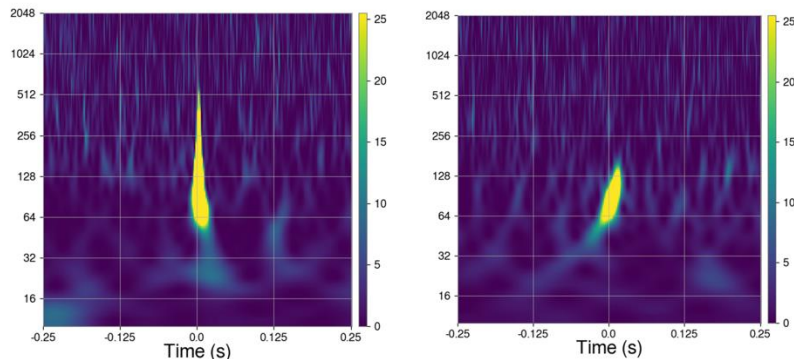
→

Triggers

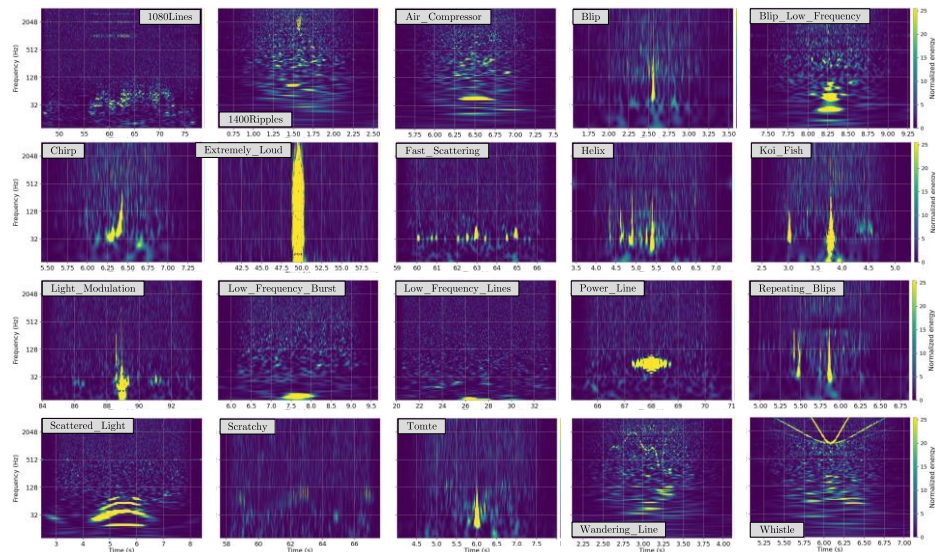
**Idea:** unknown signals generate *multiple* triggers. Can we find *patterns* with Machine Learning?

# Transient noise burst (glitches)

- Caused by instruments or environment (known or unknown)
- Diminish scientific data available
- Hinder GW detection (mask and/or mimic)

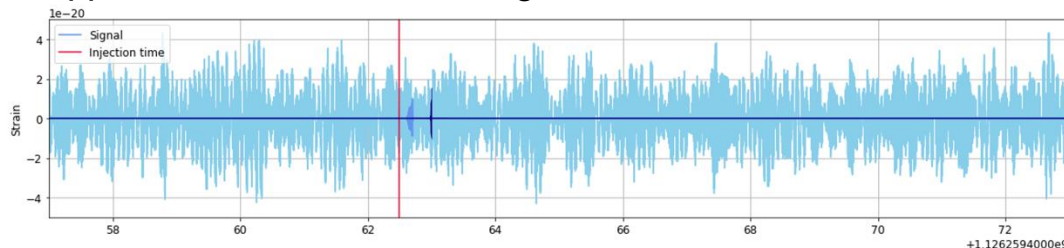


Example of a blip glitch (left) and a intermediate-mass black hole (right)

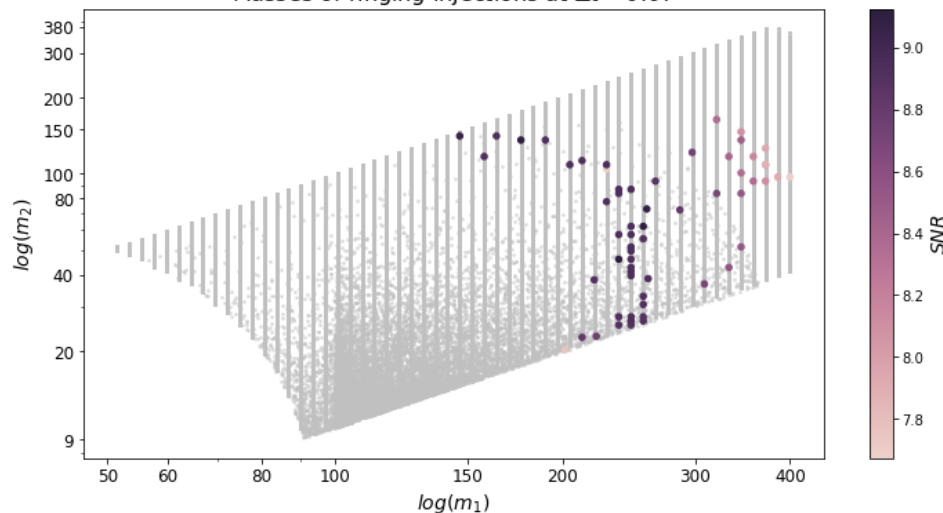


# A simulated GW through a detection pipeline

$\Delta t$ : time when trigger happened – time when GW signal was added to the noise

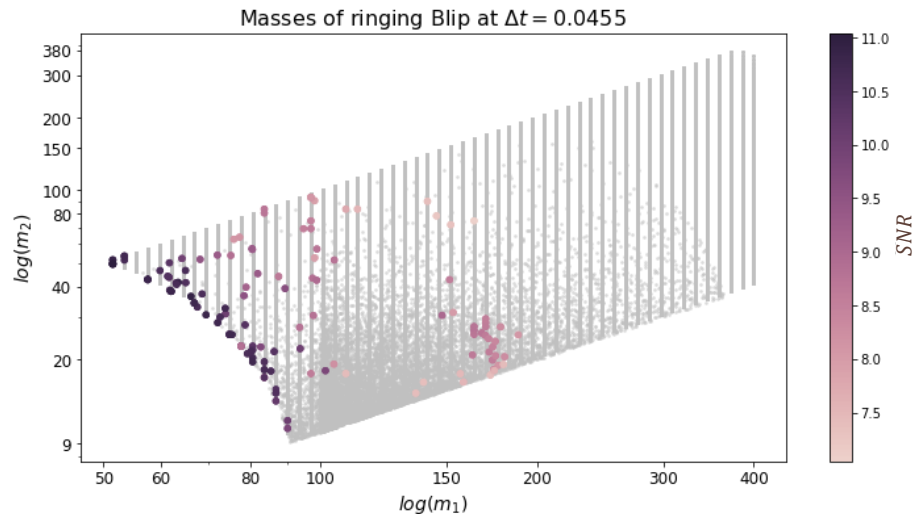
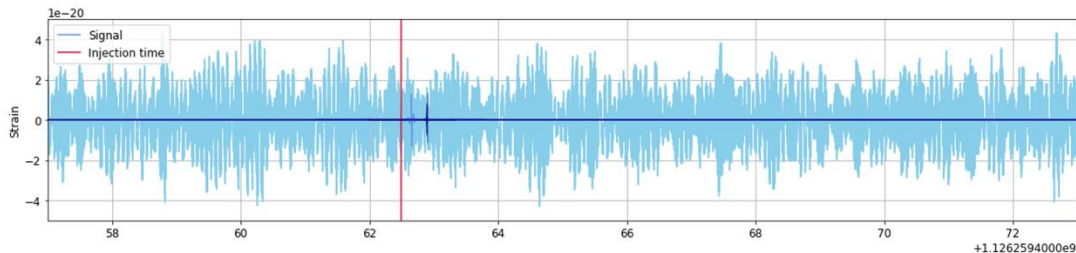


Masses of ringing injections at  $\Delta t = 0.07$



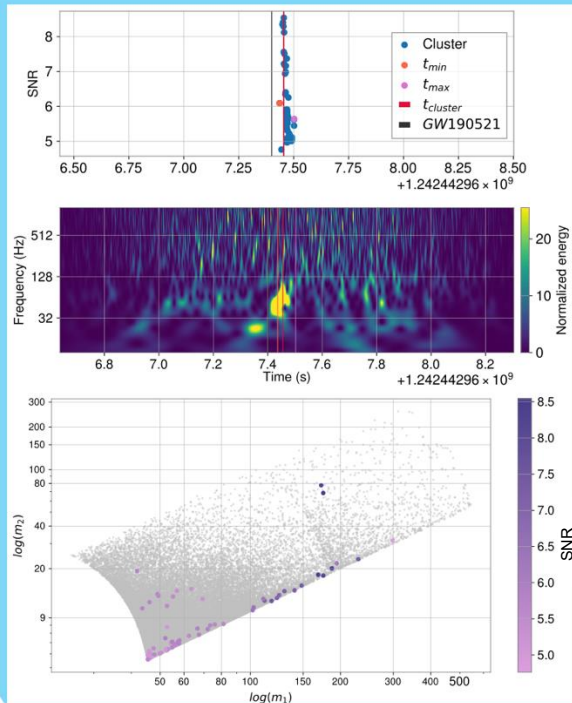
# A glitch through a detection pipeline

$\Delta t$ : time when trigger happened – time when glitch happened

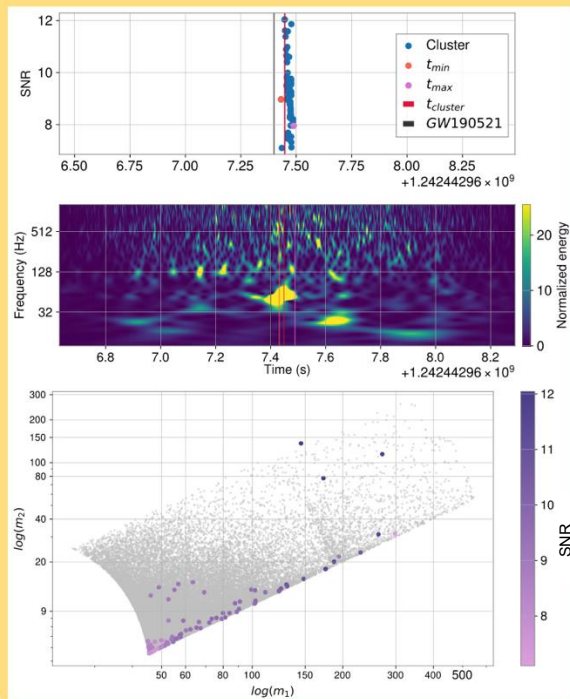


# GW190521

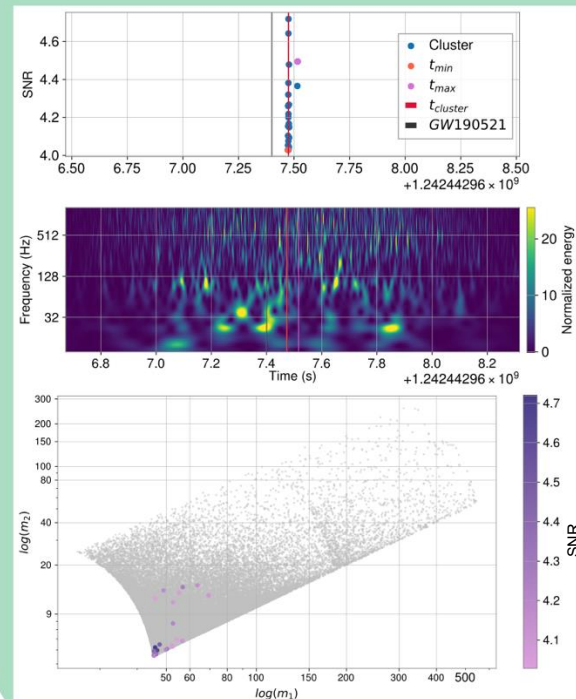
## LIGO Hanford (H1)



## LIGO Livingston (L1)



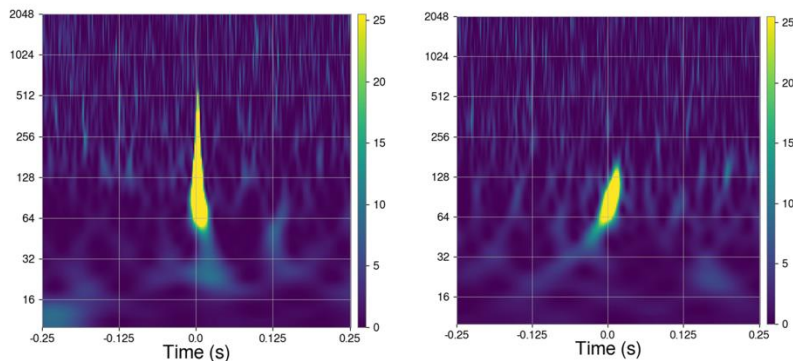
## Virgo (V1)



# Motivation

**Context:** intermediate-mass black holes (IMBH) are the missing link between stellar black holes and supermassive black holes, but they are hard to detect!

**Idea:** use triggers from matched filtering (free information) from detection algorithms to learn the background (glitches) and foreground (GW signals) with ML



Example of a blip glitch (left) and a IMBH (right)

- MF searches use *strict* conditions for detection.
- Can we *relax* the search with the interpolation ability of ML?



# Multi-class classification demo

Demo data from PhysRevD.111.103020 - arXiv 2412.17169

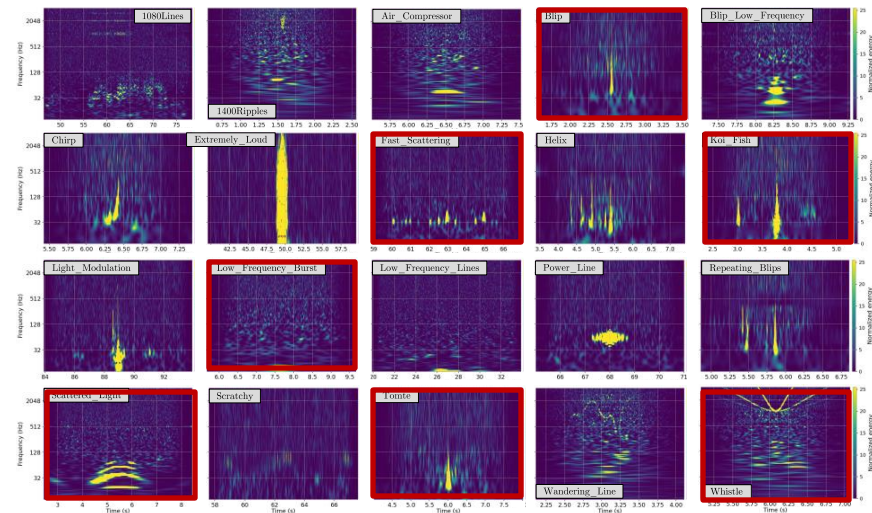
**Task:** Distinguish IMBH from different glitch classes in single detector → we have 3 detectors!

**Algorithm:** Multi-layer perceptron (MLP)

**Input:** Adding time is hard, so let's simplify the problem. Each template is defined by  $m_1, m_2, s_{1z}, s_{2z}, \chi^2, SNR$ . We weight average by SNR to get the feature vector

$$\mu(m_1, m_2, s_{1z}, s_{2z}, \chi^2, SNR)$$

**Output:** class probability



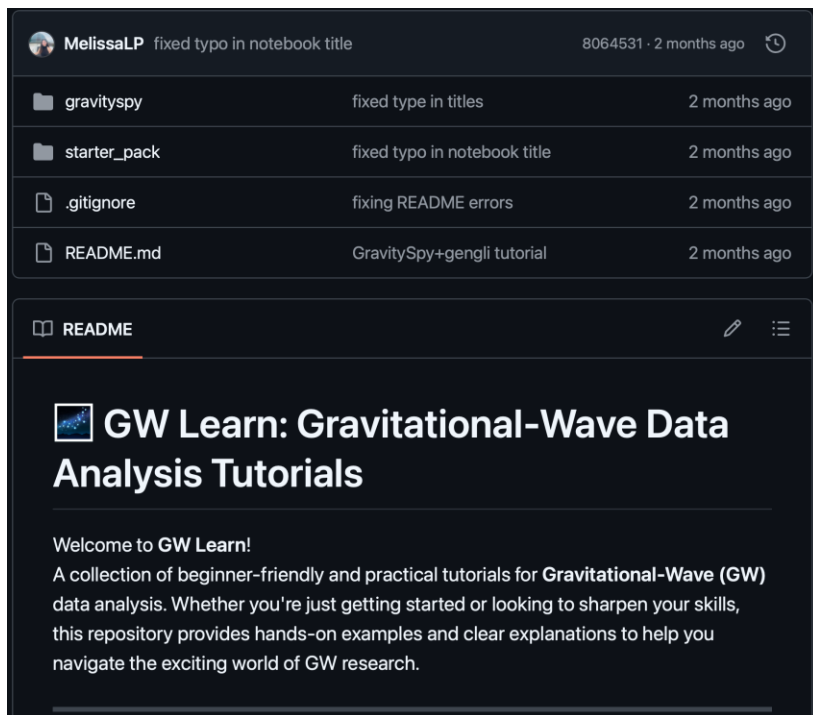
**Idea:** MLP differentiates **6 classes**: 5 different types of background (glitches) and single foreground (GW signals). It uses only **6 parameters** in **single detector**



# About today's tutorial



Promotion time! [gw\\_learn tutorials](#)



**MelissaLP** fixed typo in notebook title 8064531 · 2 months ago

File	Description	Time
gravityspy	fixed type in titles	2 months ago
starter_pack	fixed typo in notebook title	2 months ago
.gitignore	fixing README errors	2 months ago
README.md	GravitySpy+gengli tutorial	2 months ago

**README**

## GW Learn: Gravitational-Wave Data Analysis Tutorials

Welcome to **GW Learn**!

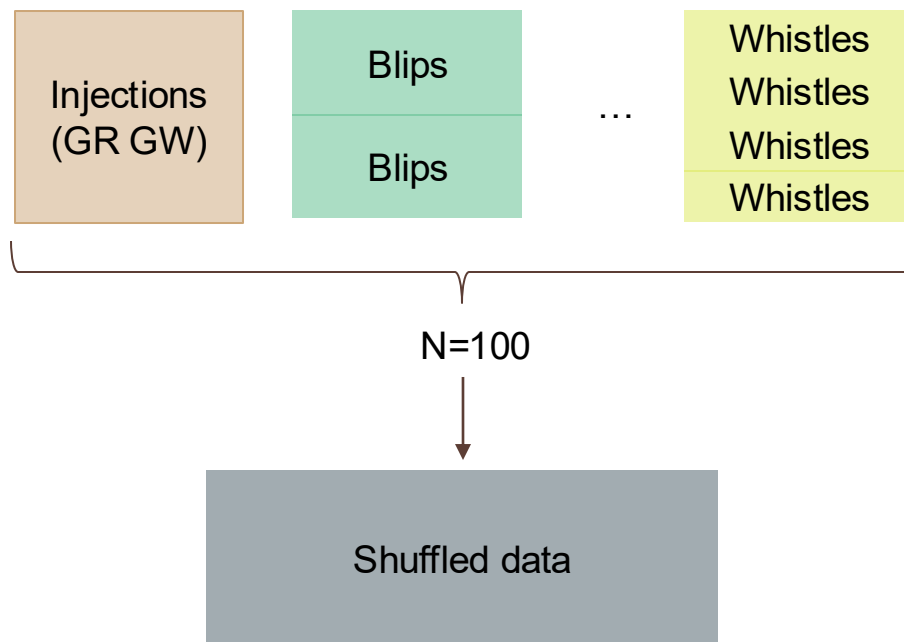
A collection of beginner-friendly and practical tutorials for **Gravitational-Wave (GW)** data analysis. Whether you're just getting started or looking to sharpen your skills, this repository provides hands-on examples and clear explanations to help you navigate the exciting world of GW research.

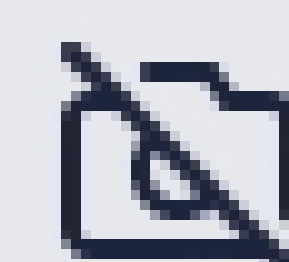
Access tutorial of today:

<https://shorturl.at/vxMH4>

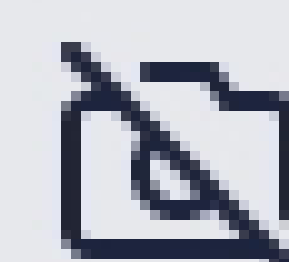
# Dealing with imbalanced data

1. Accounting for imbalanced data (bootstrapping with replacement)





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# Exercise 1: What is bootstrapping and why is it important?

This is a trick question --  
E. T. Jaynes dixit

I'm not sure — I'm a physicist  
and I've ever heard of  
bootstrapping. Is it like  
dimensionality reduction or  
some kind of data compression  
technique?

Resampling

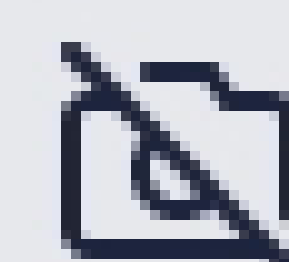
Computing  
uncertainties when you  
don't know what else to  
do

resampling

A way to estimate the unknown  
sampling/data distribution  
using the current data to derive  
estimators for target (usually  
Frequentist) statistics.

When you want to report  
uncertainties but don't have  
the ability/compute/will to draw  
additional samples from the  
distribution of interest.

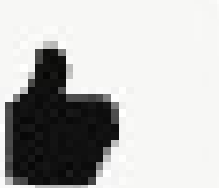
I think it's important for  
uncertainty  
quantification



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//  
Quote

– *Who*





## Exercise 2: What are the main problems with this network?

Too shallow

Not wide enough

too small.

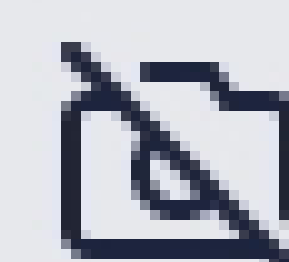
Data are not normalized

It's very small — it only has two layers. Maybe include more layers and also include a variable learning rate, as well as dropout. Maybe it's overfitting? I noticed the accuracy goes up only to .75.

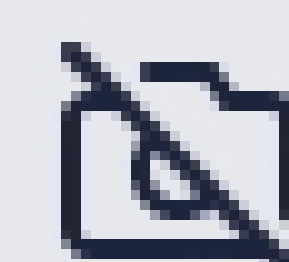
Input data not scaled appropriately

Neural network architecture is a "prior" over the structure you expect in your data. MLP assumes just a general curve, and with few parameters/layers the functions are probably too simple.

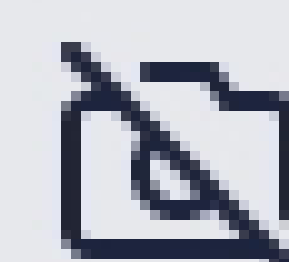
Change the activation to RELU may better performance.



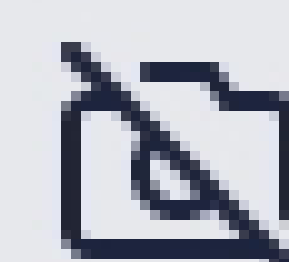
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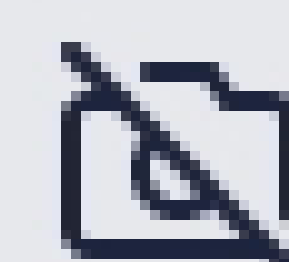
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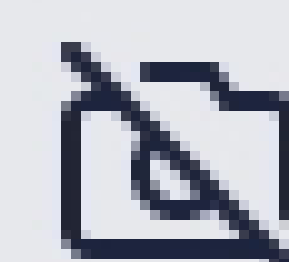
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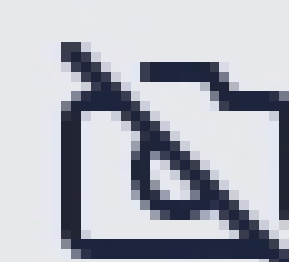


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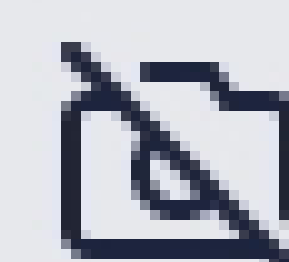


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# What would you like to talk about? Short link: <https://shorturl.at/Blil6>

Incorporating diagnostic data channels from GW detectors in searches directly by assimilating them using ML stages, and using that to discriminate signals from noise transients invariantly

How to design a search pipeline around ML in practice?

Calibration uncertainty reduction using ML

At which SNR levels will template-banks-based methods become too computationally expensive? Will they at all?

How would we actually verify a detection of an unmodeled signal by a ML algorithm?

What can traditional searches learn from ML?

Is an end to end joint model for noise transients and astrophysical sources feasible with ML?

how often does the distribution of noise change during the run roughly? Or how often do ML models need to be trained/recalibrated during the obs run?

What would you like to talk about? Short link: <https://shorturl.at/Blil6>

ML for  
unmodelled/unknown  
signals