Final Project Write-Up

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Solar system small body analysis

The Presentation

I have used Python and Jupyter notebooks throughout the course since those are tools that I use daily in my own work. I made a risky decision to change course at "the 11th hour". Jupyter notesbooks has some frustrating limitations when it comes to high quality presentations such as a lack of built-in code hiding and variable templating within Markdown sections. The Python package Pandas draws much of its inspiration from R dataframes in the first place, so it was a natural transition. I am very pleased with the results, and will definitely be using R-Studio going forward for business presentations. The website r-tutor.com authored by Chi Yau was a tremendous help in my learning curve. I have turned off cell echoing to preserve the calculations used in the presentation.

My data set

The data presented here is a sample collected from NASA's JPL Small-Body Database. They represent smaller objects in the Solar System such as comets, asteroids, and moons. I chose this data set because I am curious about space science, and this is an opportunity to learn more about the subject while fulfilling the requirements of the assignment. Please forgive any inaccurate statements about orbital mechanics! There are **22,458 cases** in my sample.

Variables

I learned what most of these variables mean from Holli Riebeek's article on NASA's Earth Observatory Website (Riebeek, 2009).

Title	Description	Type
full_name	Full Name	string
a	[au] semi-major axis	float64
e	eccentricity	float64
i	[deg] inclination	float64
w	[deg] perihelion degree	float64
om	[deg] node degree (longitude of the ascending node)	float64
ma	[deg] M mean anomaly degree	float64
q	[au] perihelion distance	float64
ad	[au] Q aphelion distance	float64
per_y	period years	float64
H	absolute magnitude	float64
moid	[au] Earth MOID (minimum orbit intersection distance)	float64
$\operatorname{orbit_id}$	Orbit solution ID	string
class	Orbit class	string

All of the float64 data types are quantitative variables, and the string data types orbit_id and class are categorical.

Orbit Classes

Below is a table of orbit classes. There is a wealth of information on the CNEOS website that provides much more detail than I can here.

class	name	num
$\overline{\text{AMO}}$	Amor	8267
APO	Apollo	12308
ATE	Aten	1684
ETc	Encke-type Comet	1
HTC	Halley-type Comet	30
IEO	Atira	21
JFC	Jupiter-family Comet*	10
$_{ m JFc}$	Jupiter-family Comet	137

JFC -> Jupiter-family comet, classical definition (P < 20 y). JFc -> Jupiter-family comet, as defined by Levison and Duncan (2 < TJupiter < 3).

```
mydata = read.csv("../neo.csv")
summary(mydata)
```

Data summary

```
##
                full_name
                                      a
                                                           е
##
            (1979 XB):
                                          0.5554
                                                            :0.002846
                               Min.
                                                    Min.
                           1
##
            (1982 YA):
                           1
                               1st Qu.:
                                          1.3160
                                                    1st Qu.:0.315925
##
            (1983 LC):
                                          1.7232
                                                    Median: 0.459597
                               Median:
                           1
##
            (1986 NA):
                                          1.8372
                                                            :0.446875
                           1
                               Mean
                                                    Mean
##
            (1988 NE):
                           1
                               3rd Qu.:
                                          2.2082
                                                    3rd Qu.:0.572447
##
            (1989 AZ):
                           1
                               Max.
                                       :453.0215
                                                            :0.997982
                                                    Max.
    (Other)
                      :22452
##
##
                                                     om
                                                                         ma
                                  : 0.0079
##
              0.01352
                                                      : 0.026
                                                                          :-15.57
    Min.
            :
                          Min.
                                              Min.
                                                                  Min.
                          1st Qu.: 92.8879
##
    1st Qu.:
              4.62325
                                              1st Qu.: 82.717
                                                                  1st Qu.: 54.08
    Median :
               8.95086
                          Median: 184.5539
                                              Median: 172.094
                                                                  Median :164.37
##
##
    Mean
           : 12.67136
                                  :182.1138
                                              Mean
                                                      :173.066
                                                                  Mean
                                                                          :172.14
                          Mean
##
    3rd Qu.: 17.85681
                          3rd Qu.:271.5291
                                               3rd Qu.:255.628
                                                                  3rd Qu.:289.22
##
    Max.
            :172.51374
                          Max.
                                  :359.9820
                                              Max.
                                                      :359.978
                                                                  Max.
                                                                          :360.00
##
##
                                                                        Η
                              ad
                                                 per_y
            :0.02832
##
    Min.
                       Min.
                                  0.6538
                                            Min.
                                                        0.414
                                                                 Min.
                                                                         : 9.40
    1st Qu.:0.78427
                                  1.7179
                                                        1.510
##
                        1st Qu.:
                                            1st Qu.:
                                                                 1st Qu.:20.50
##
    Median : 0.96387
                       Median :
                                  2.5009
                                            Median:
                                                        2.262
                                                                 Median :23.00
##
    Mean
            :0.91446
                                  2.7599
                                                        3.323
                                                                         :22.82
                       Mean
                                            Mean
                                                                 Mean
    3rd Qu.:1.06960
                        3rd Qu.:
                                  3.4212
                                            3rd Qu.:
                                                        3.281
                                                                 3rd Qu.:25.10
##
    Max.
            :1.29999
                       Max.
                               :905.1288
                                            Max.
                                                    :9642.430
                                                                 Max.
                                                                         :33.20
##
                                                                 NA's
                                                                         :189
##
         moid
                           orbit_id
                                              class
            :0.00000
                                  797
                                         APO
                                                 :12308
##
    Min.
                        JPL 3
                               :
    1st Qu.:0.01601
                        JPL 6
                                                 : 8267
##
                                  726
                                         OMA
```

```
Median :0.05349
##
                       JPL 4
                                  722
                                         ATE
                                                 : 1684
##
           :0.09443
                       JPL 1
                                  715
                                         JFc
                                                    137
    Mean
##
    3rd Qu.:0.14457
                       8
                                  698
                                         HTC
                                                     30
            :0.70772
                       9
                                  693
                                         IE0
                                                     21
##
    Max.
    NA's
            :14
                       (Other):18107
                                         (Other):
                                                     11
```

Eccentricity

The first quantitative variable I would like to explore is labeled e for eccentricity. This variable takes on values between 0 and 1 where 0 is a perfectly circular orbit and close to 1 flattens the orbit into line (Air Command and Staff College Space Research Electives Seminars, 2009). The distribution appears to be close to **normal** and slightly **skewed to the right**.

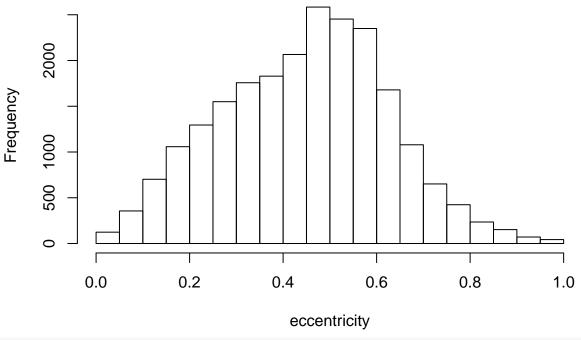
```
e_mean = round(mean(mydata$e), digits = 3)
e_median = round(median(mydata$e), digits = 3)
```

The **mean** is 0.447.

The **median** is 0.46.

```
hist(mydata$e, main="Histogram of Eccentricity", xlab = "eccentricity")
```

Histogram of Eccentricity



```
summary(mydata$e)

# Outliers Q3 + 1.5 * IQR > value < Q1 - 1.5 * IQR

q1 = quantile(mydata$e, 0.25)

q3 = quantile(mydata$e, 0.75)

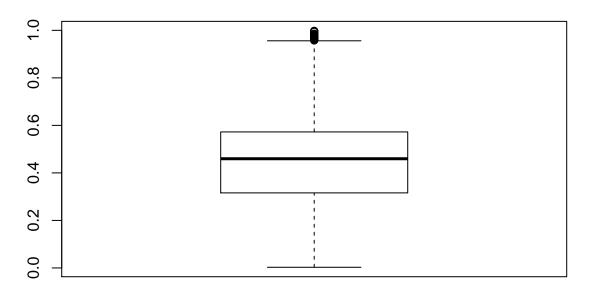
iqr = q3 - q1
low_outliers <- mydata[which(mydata$e < q1 - (1.5 * iqr)),]
high_outliers <- mydata[which(mydata$e > q3 + (1.5 * iqr)),]
```

The \mathbf{IQR} is 0.257.

The standard deviation s is 3.943.

boxplot(mydata\$e, main="Eccentricity Boxplot")

Eccentricity Boxplot



Five-number summary

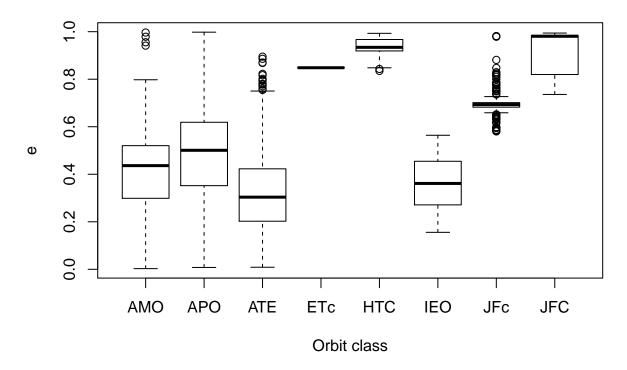
Variable	Value
minimum	0.003
Q_1	0.316
median	0.46
Q_3	0.572
\max imum	0.998

Eccentricity by orbit class

According to NASA, Near-Earth objects (NEO) have orbits that put them in proximity to Earth's orbit. It is not much surprise to see similar IQR's in these classes (AMO, APO, ATE, and IEO). The Atiras class (IEO) is noteworthy for its lack of extreme values. NEO's in this class have orbits contained entirely within Earth's own orbit (NASA).

boxplot(mydata\$e~mydata\$class, main="Eccentricity", xlab="Orbit class", ylab="e")

Eccentricity



Semi-major axis

The other quantitative variable I would like to explore is the semi-major axis. This is the larger of the two axis of an ellipse (for e > 0) (Air Command and Staff College Space Research Electives Seminars, 2009). This distribution has a log-normal shape and is **skewed to the right** with a much longer tail than the previous graph. Given the spread of this data, please note the *logarithmic scale* where indicated.

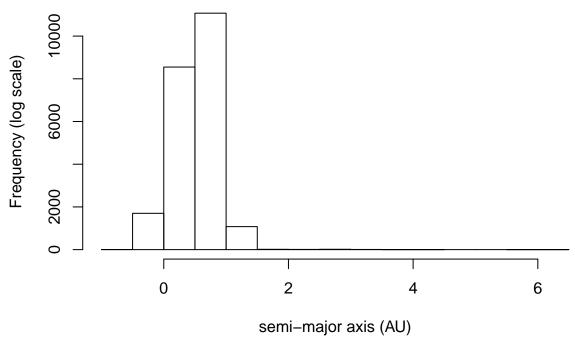
```
a_mean = round(mean(mydata$a), digits = 3)
a_median = round(median(mydata$a), digits = 3)
```

The **mean** is 1.837.

The **median** is 1.723.

hist(log(mydata\$a), main="Histogram of Semi-Major Axis", xlab = "semi-major axis (AU)", ylab="Frequency

Histogram of Semi-Major Axis



```
summary(mydata$a)

# Outliers Q3 + 1.5 * IQR > value < Q1 - 1.5 * IQR

q1 = quantile(mydata$a, 0.25)

q3 = quantile(mydata$a, 0.75)

iqr = q3 - q1
low_outliers <- mydata[which(mydata$a < q1 - (1.5 * iqr)),]
high_outliers <- mydata[which(mydata$a > q3 + (1.5 * iqr)),]
```

The \mathbf{IQR} is 0.892.

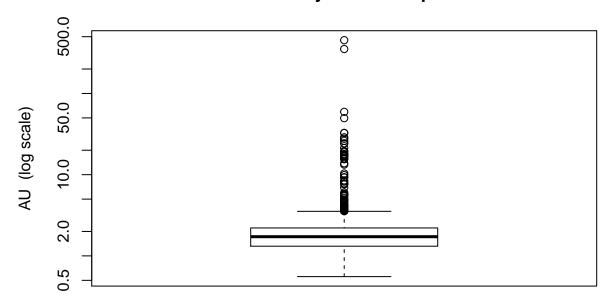
The standard deviation s is 3.943.

There are 0 **outlier** values less than -0.022

There are 98 ${f outlier}$ values greater than 3.546

boxplot(mydata\$a, main="Semi-major axis Boxplot", log="y", xlab="semi-major axis (AU)", ylab="AU (log

Semi-major axis Boxplot



semi-major axis (AU)

Five-number summary

Variable	Value
minimum	0.555
Q_1	1.316
median	1.723
Q_3	2.208
$\max $	453.022

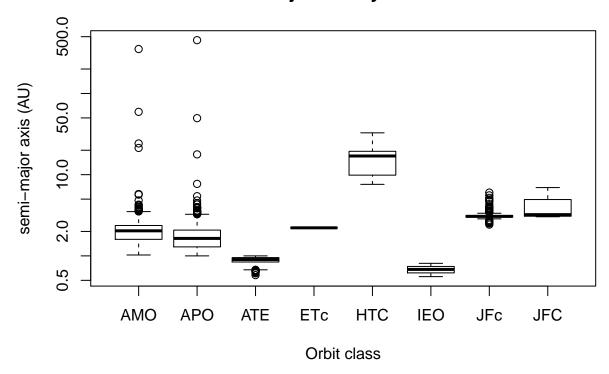
Semi-major axis by orbit class

Aten (ATE) and Atira (IEO) class bodies have semi-major axes smaller than 1 AU by definition. I am surprised that there are no outliers in the Jupiter and Haley-family classes. This may be due to observational practices or other constraints that I am not familiar with.

For reference, Earth is 1 AU from the sun, and Pluto is around 40 AU from the sun. Some of these bodies travel very far outside of our planetary system!

boxplot(mydata\$a~mydata\$class, main="Semi-major axis by orbit class", xlab="Orbit class", ylab="semi-ma

Semi-major axis by orbit class

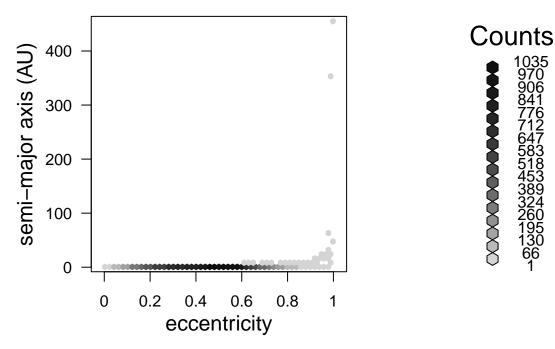


Linear correlation

Scatter plots can lose their usefulness with too many data points. I read about a plot called a "hex bin", and I feel this did a much better job of allowing the data to tell its story ("Quick-R: Scatterplots," n.d.).

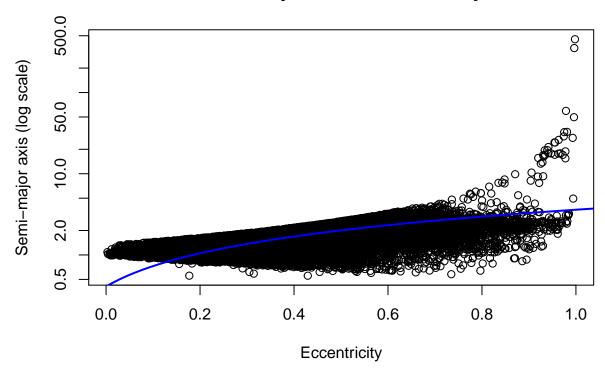
```
library(hexbin)
# plot(mydata$e, mydata$a, log="y", ylab="semi-major axis", xlab="eccentricity")
hbin <- hexbin(mydata$e, mydata$a, xbins=50)
plot(hbin, main="Semi-major axis vs Eccentricity", ylab="semi-major axis (AU)\n", xlab="eccentricity")</pre>
```

Semi-major axis vs Eccentricity



I struggled with the decision to use a logarithmic scale on a graph depicting a best-fit "line". Ultimately, I feel using this scale reveales much more about the data than a clumb of color along the x-axis.

Semi-major axis vs Eccentricity



The correlation coefficient r is 0.145. One possible confounding, or lurking variable is i or inclination. Inclination exhibits a similar correlation coefficient and clustering of values. The r-value is low, but it appears from the clustering in the above graph that one would have a decent probability of finding satellites near the

regression line, especially for larger values of eccentricity. I am surprised that orbits with large semi-major axes tend to be highly elliptic as well.

Summary

I do not have the physics background and experience to do this data justice. There are many more variables and relationships that could be explored much further. I am finishing this assignment with many more questions than answers about this data set, but it was an exciting opporunity to explore a new subject! I incurred a bit of extra work by making the switch to R, but I am very happy with the results, and I will be able to use this experience on the job in the near future.

References

- Air Command and Staff College Space Research Electives Seminars. (2009, September 1). AU-18 Space Primer. Retrieved March 29, 2020, from https://www.airuniversity.af.edu/Portals/10/AUPress/Books/AU-18.PDF
- NEO Basics. (n.d.). Retrieved March 29, 2020, from https://cneos.jpl.nasa.gov/about/neo groups.html
- Quick-R: Scatterplots. (n.d.). Retrieved March 29, 2020, from https://www.statmethods.net/graphs/scatterplot.html
- Riebeek, H. (2009, September 4). Catalog of Earth Satellite Orbits. Retrieved March 29, 2020, from https://earthobservatory.nasa.gov/features/OrbitsCatalog

Additional sites used:

- https://ssd.jpl.nasa.gov/sbdb_query.cgi
- https://cneos.jpl.nasa.gov/glossary/