

# Optimization of an Active Spherical Coded Mask Gamma-ray Imager

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# Outline

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- Planar configuration
  - The High Efficiency Multi-modal Imager (HEMI)
- Spherical configuration
  - Portable Radiation Imaging Spectroscopy and Mapping (PRISM)
- Mask optimization
- Results
- Conclusions

# Gamma-ray Imaging

## Motivation

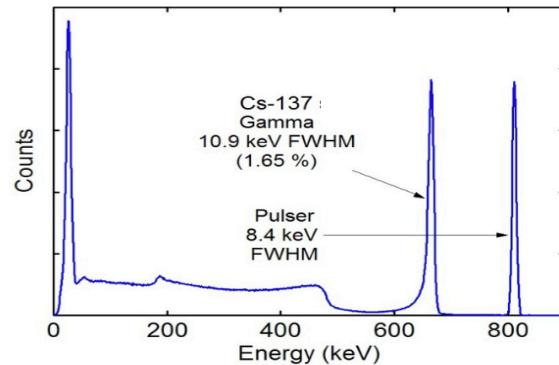
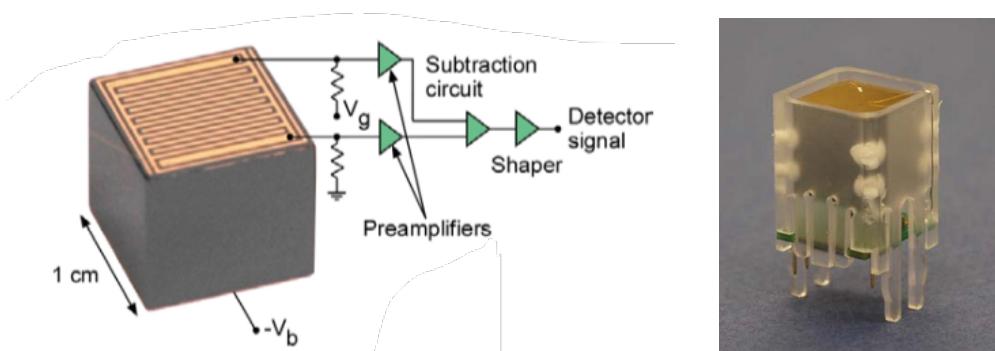
Detect, locate, identify, and characterize radioactive sources in the field. Ability to search for weak sources in complex environments and efficiently map radiation fields.

## Need

Hand-held, portable 3D imaging system with high efficiency, wide field-of-view, broad energy sensitivity, and high energy resolution.

## Approach

Multiple room-temperature operated  $\text{cm}^3$  CdZnTe (CZT) coplanar grid (CPG) detectors arranged to facilitate coded aperture and Compton imaging modalities.



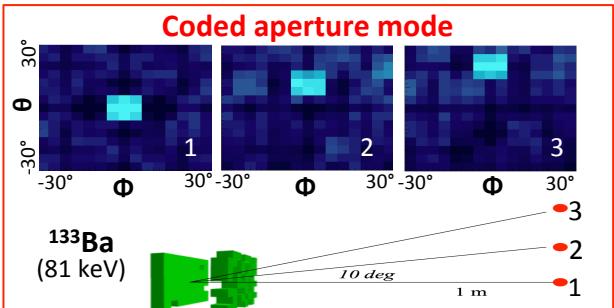
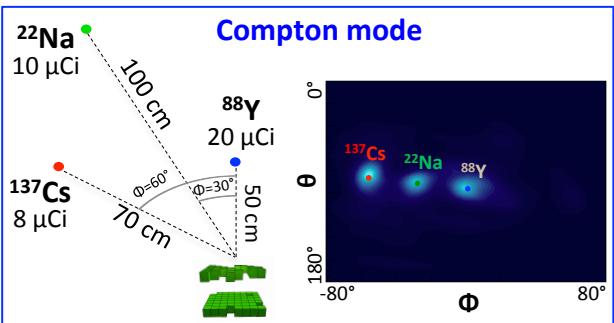
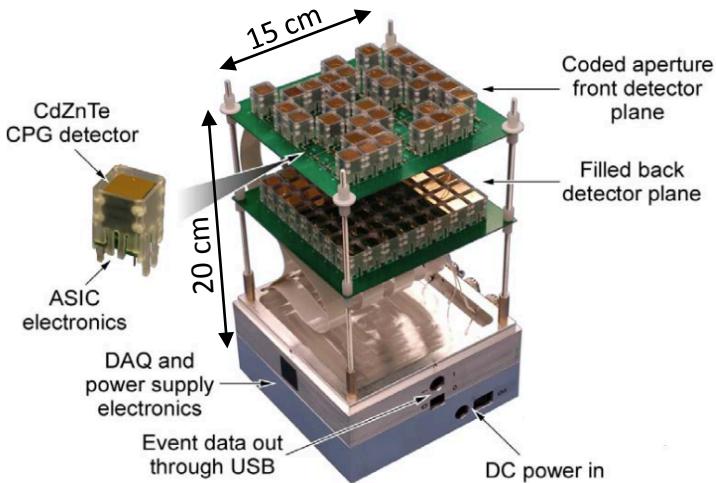
# Planar Arrangement

## High Efficiency Multi-modal Imager (HEMI)<sup>†</sup>

- 96 CZT CPG detectors ( $1 \text{ cm}^3$ ,  $\sim 2\%$  FWHM at 662 keV).
- 32 in **active coded** front plane, 64 in fully populated back plane).
- Coded aperture (30 keV to 400 keV) and Compton imaging (300 keV to 3 MeV).
- $\sim 10^\circ$  resolution at 662 keV (Compton) and 186 keV (CA).
- < 10 lbs  $\rightarrow$  Hand-portable. Also placed on UAVs.

$\rightarrow$  **Limited field-of-view in both modalities, especially coded aperture.**

$\rightarrow$  **GOAL: Redesign for a  $4\pi$  FOV with coded aperture.**



<sup>†</sup> M. Galloway, *Characterization and Applications of a CdZnTe-Based Gamma-Ray Imager*, Ph.D. Thesis, 2014.

# Spherical Arrangement

Rearrange the detectors into an **active spherical coded configuration** to enable **omnidirectional  $4\pi$  coded aperture imaging**.

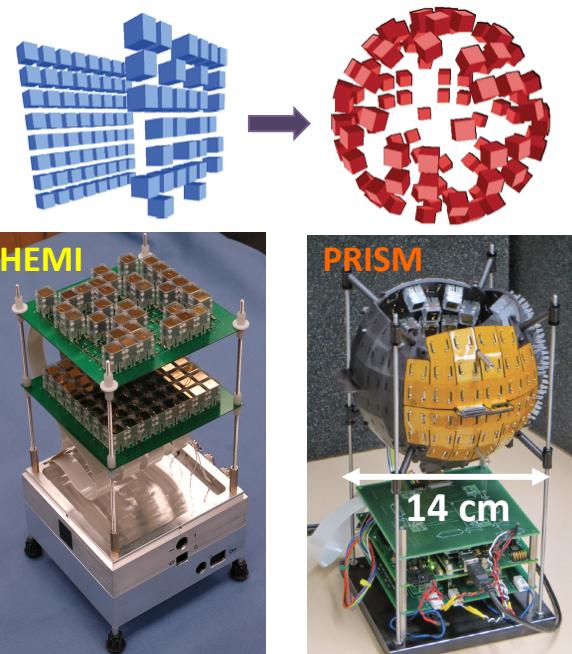
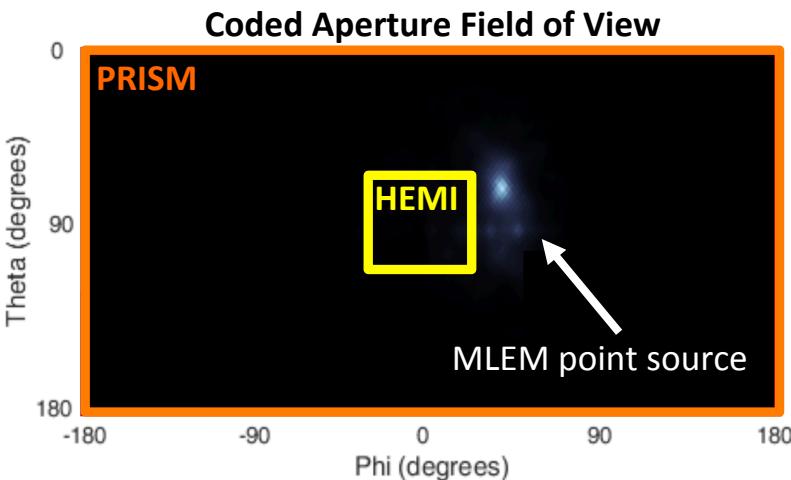
Feasibility study conducted for a spherical housing with **192** available detector locations.

## → Portable Radiation Imaging Spectroscopy and Mapping (PRISM)

- 6 modular (identical) faces, flexible circuit design
- $\leq 2\%$  FWHM at 662 keV
- $< 10^\circ$  resolution at 186 and 662 keV
- Depth-of-interaction (DOI) readout

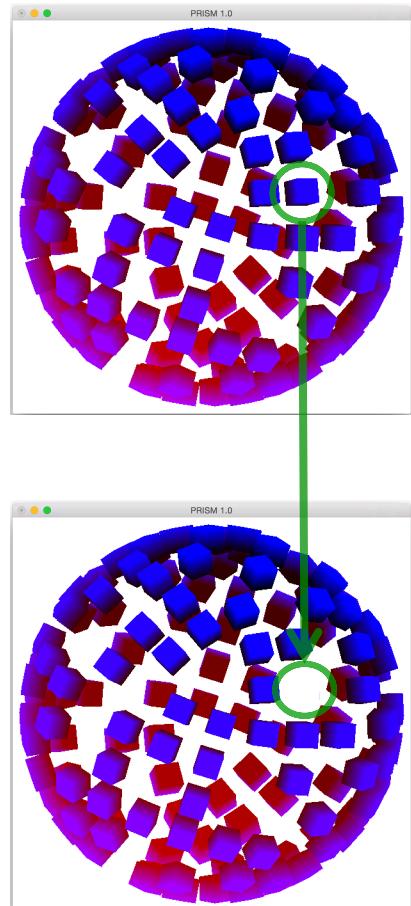
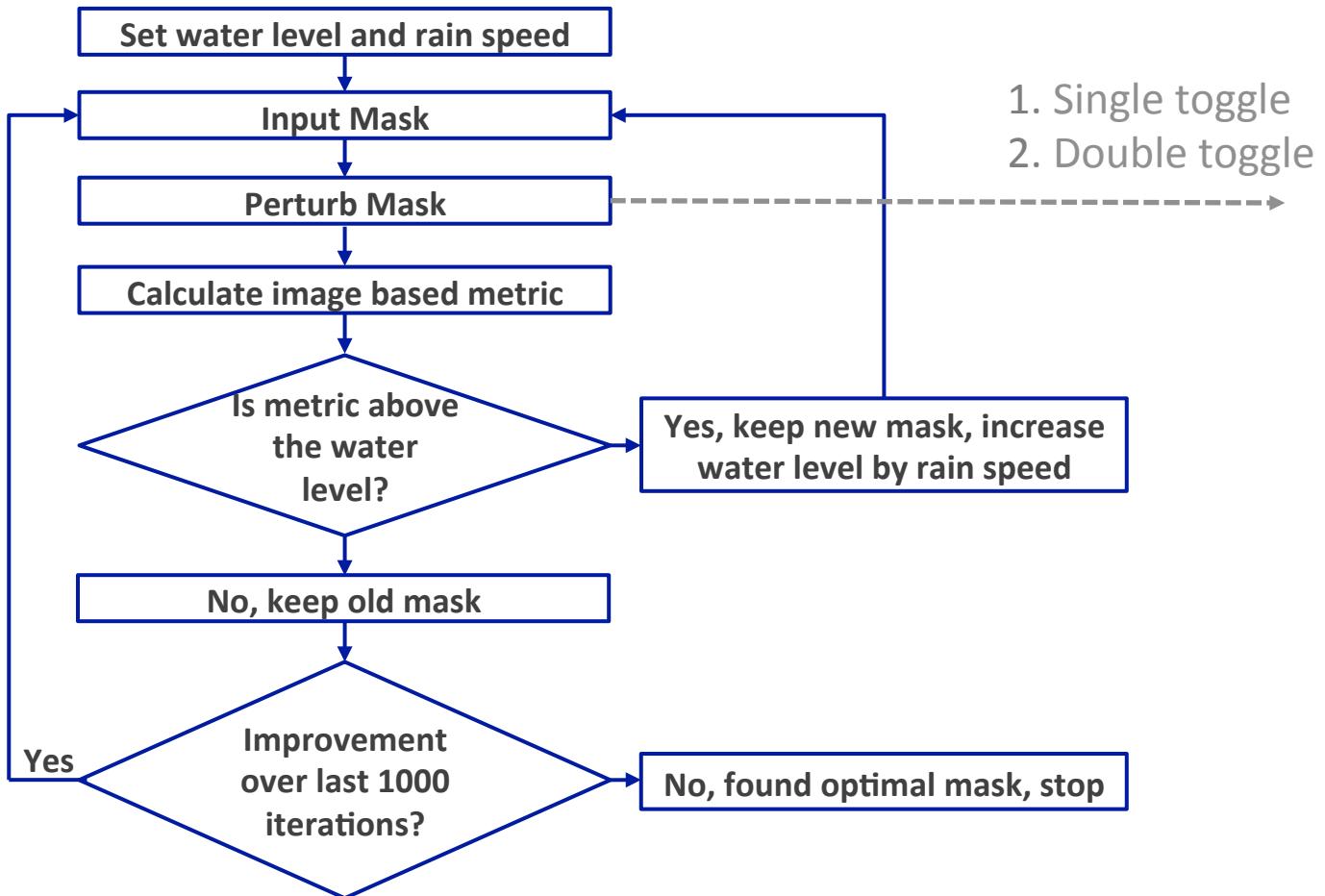


**What is the optimal number and configuration of the detectors on the sphere?**



# Optimization Procedure

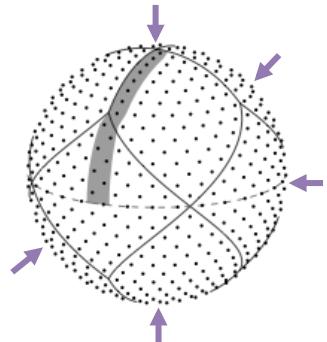
→ The Great Deluge algorithm<sup>†</sup>:



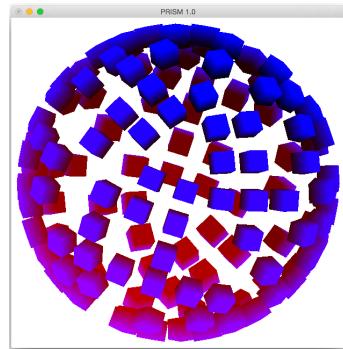
<sup>†</sup> G. Dueck, *Journal of Computational Physics*, vol. 104, pp. 86–92, 1993.

# System Response Generation

- Image space pixelized into 3072 far-field equal area pixels with *HEALPix*<sup>†</sup>.
- Geant4 simulation of point sources at all source angles → **O(min)**
- Graphics-based program in OpenGL to perform mask projection  
**(no scattering, zero-energy)** → < 1.5 s

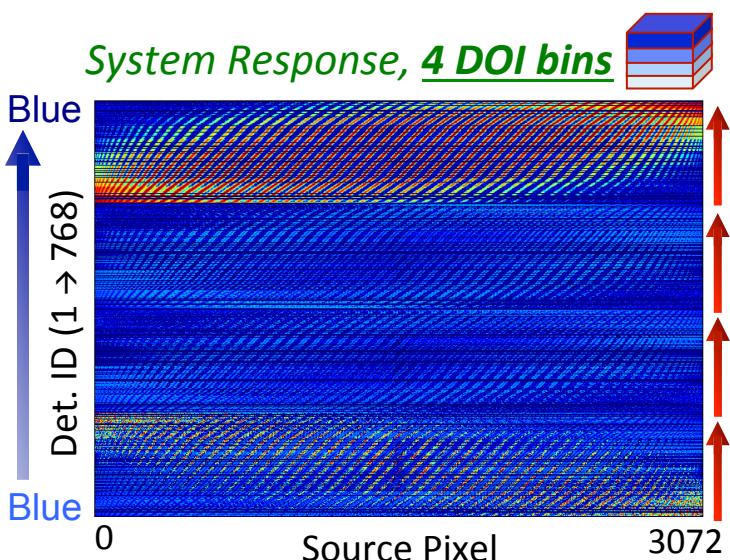
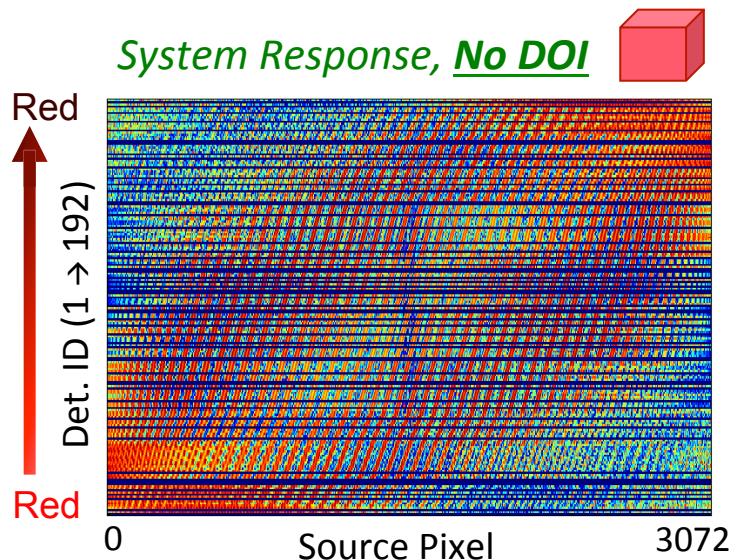


OpenGL Display



Red = Det. ID  
Blue = Depth

Histogram pixels at  
each camera angle



<sup>†</sup> Hierarchical Equal Area isoLatitude Pixelization of a sphere: K. Górski et al., *The Astrophysical Journal*, vol. 662, no. 2, pp. 759–771, 2005.

# Optimization Results

- PRISM is reconfigurable → can optimize to variety of different scenarios, each using different metrics.
- To develop the framework, we optimize imaging accuracy and uniform sensitivity.

$$\sigma^2 \left( \sum_{i=1}^{192} Y_{i,k} \right) + \sum_k B(\mu_k, k)$$

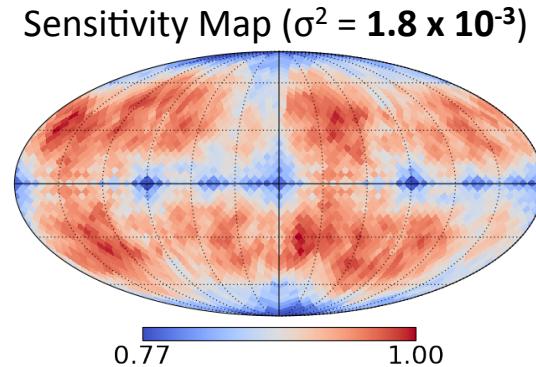
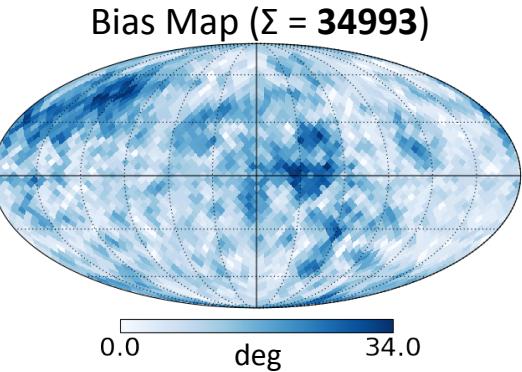
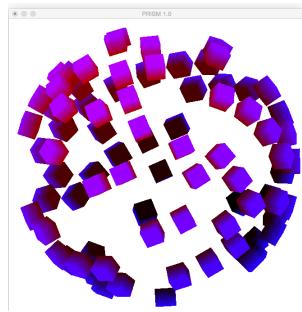
$k = 3072$  source angles

$Y_{i,k}$  = counts in detector  $i$  from source angle  $k$

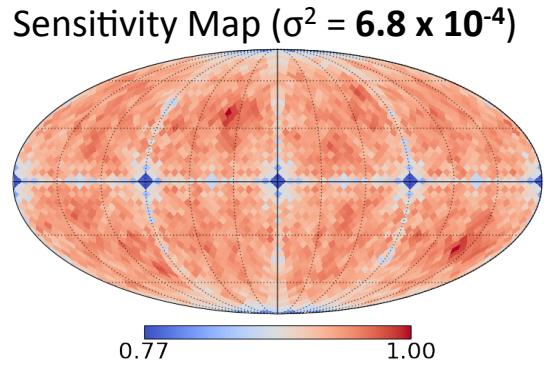
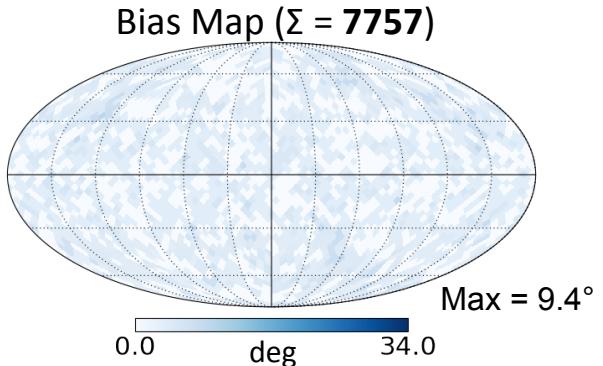
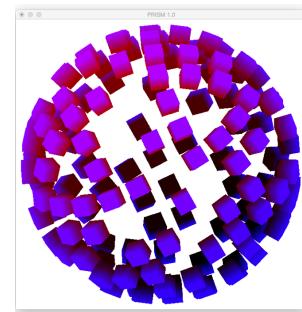
$\sigma^2()$  = variance

$B(\mu_k, k)$  = angle between mean pixel and true pixel

Starting  
(Random)  
82 detectors  
43% populated



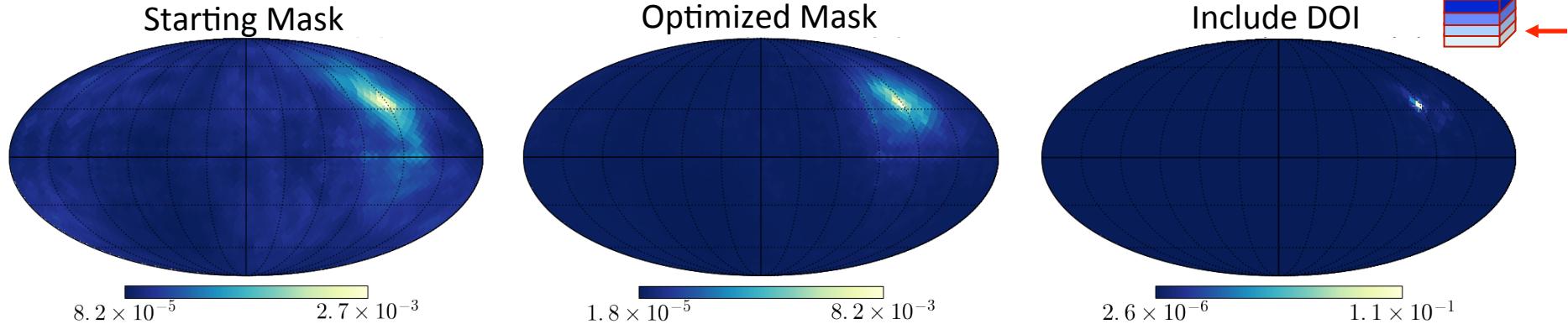
Optimized  
135 detectors  
70% populated



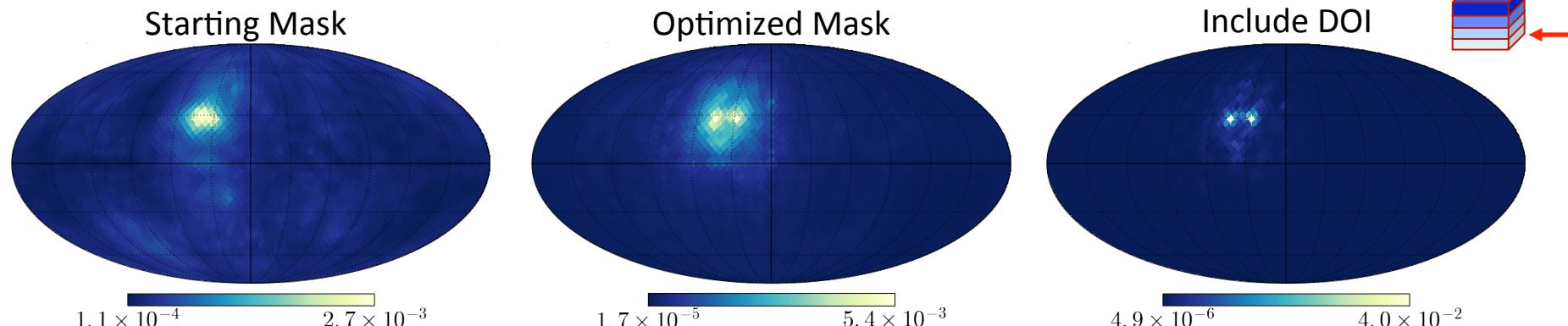
# Image Reconstruction and DOI

MLEM reconstruction (20 iterations) of far-field, zero-energy, point source in  $4\pi$ .

## Single source:



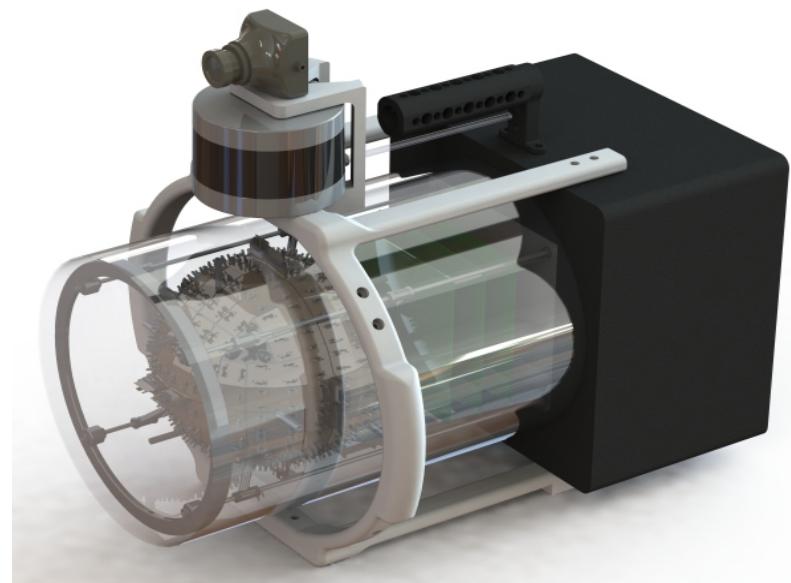
## Two sources (15° separation):



# Conclusions

- **Successful demonstration of  $4\pi$  active spherical coded aperture imaging.**
- A simulation and optimization framework has been developed to optimize the number and configuration of detectors on the sphere.
- Depth-of-interaction significantly improves coded aperture imaging (and should improve Compton imaging as well).

- Many more metrics exist and are being explored.
- DOI and energy being integrated into the optimization to observe effects on the optimal mask pattern.



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# Thank you for your attention.

## Questions?