

Detecting Asteroids with Neural Networks

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Outline

- ▶ What's the goal?
- ▶ What's the data?
- ▶ Getting started
- ▶ Building a feature set
- ▶ Building the neural network
- ▶ Training the network
- ▶ Results

The goal

Build and train a neural network to correctly identify asteroids in astrophotography data.

So...

How do we do it?

So...

How do we *really* do it?

The data

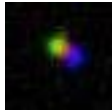
The Sloan Digital Sky Survey:

- ▶ "One of the most ambitious and influential surveys in the history of astronomy."
- ▶ Approx 35% of sky
- ▶ Largest uniform survey of the sky yet accomplished
- ▶ Data is freely available online
- ▶ Each image is 922x680 pixels

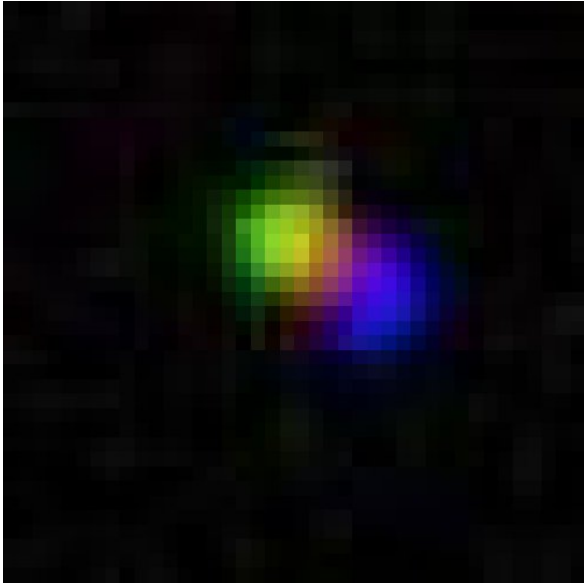




An example asteroid



An example asteroid

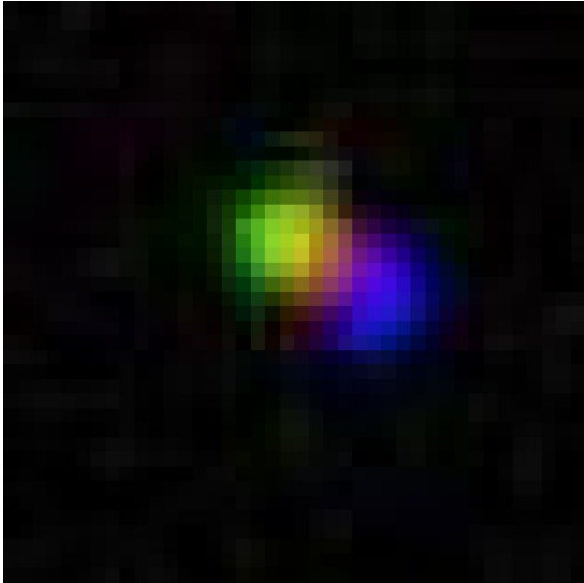


How does this work?

This exploits a property of CCDs:

- ▶ SDSS telescopes use five different filters
- ▶ They are not simultaneous
- ▶ Moving objects appear in different locations
- ▶ Always the same order

An example asteroid



Getting started

Getting the initial training data:

- ▶ Small tool to extract potential candidates from full-scale images
- ▶ Extremely naïve, approx 100:5 false positives to actual positives
- ▶ Very low false negatives (approx 1:1000)
- ▶ Incredibly slow (complex scan of 100Ks of potentials)
- ▶ Manual classification, somewhat slow
- ▶ Yields approx 250 valid items, 500 invalid items

The feature set

Good ideas for features:

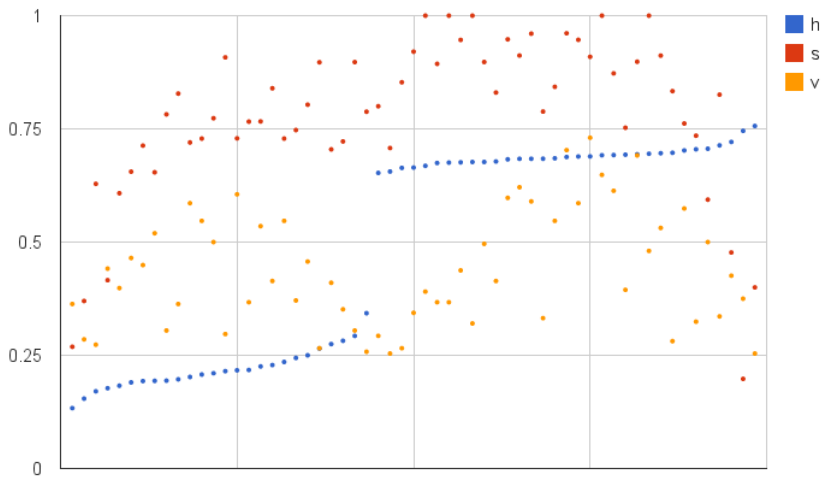
- ▶ Ratio valid hues to non-valid hues
- ▶ Best possible cluster collinearity
- ▶ Best possible average cluster distance

Feature: Ratio valid hues to non-valid hues

The goal here is to match the colors, a.k.a. “hues”:

- ▶ First step: convert to HSV space
- ▶ For pixels in the valid value-spectrum ($0.25 < v < 0.90$)
- ▶ How many are within 2 standard deviations from an optimal value?
- ▶ What's the ratio to ones that aren't?

An example HSV plot

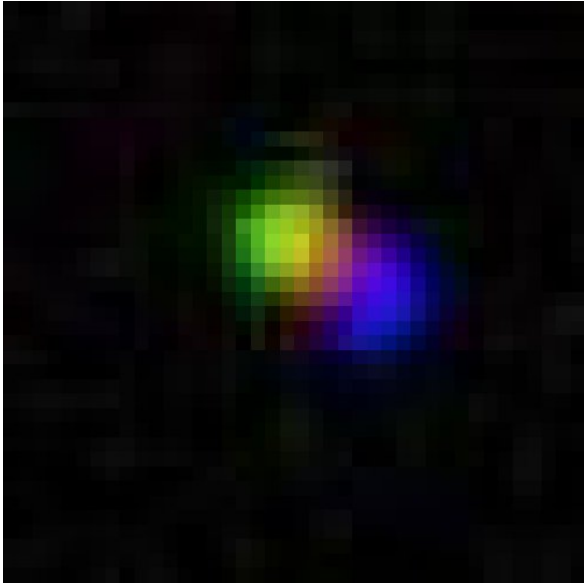


Feature: Best possible cluster collinearity

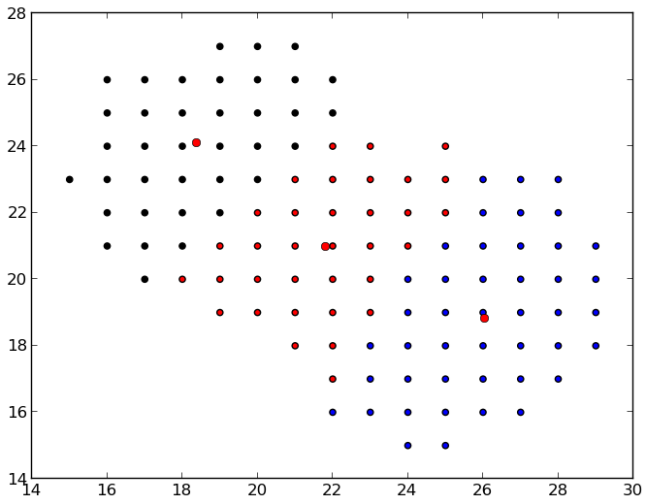
k-means clustering

- ▶ Using the valid hues from the previous feature
- ▶ Attempts to cluster n points into k groups
- ▶ Here, $k = 3$
- ▶ Produces three centroids

An example asteroid



k -means clustering of the same asteroid



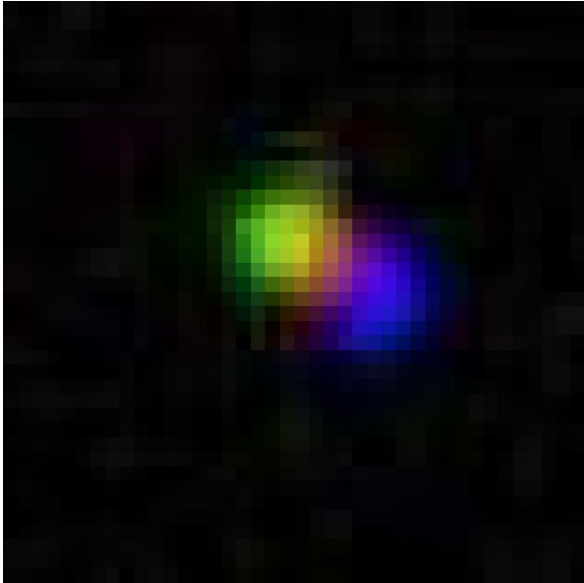
Feature: Best possible cluster collinearity

Collinearity:

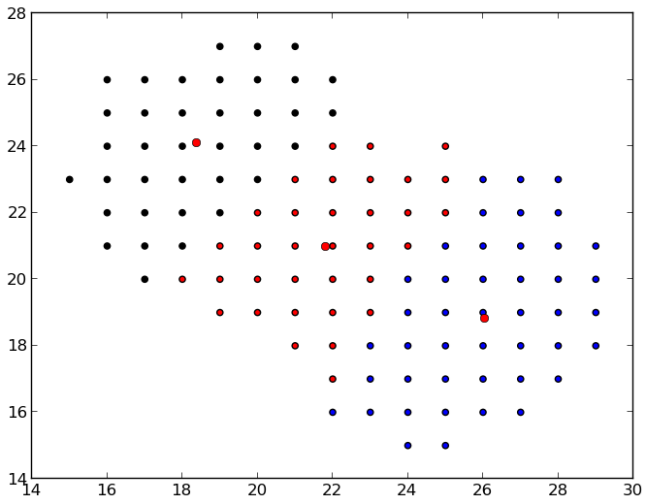
- ▶ The property of a set of points which lie on the same line
- ▶ Iterate the k -means clustering approx. 20 times
- ▶ The resulting metric is the ratio between the actual collinearity and the maximum potential collinearity
- ▶ Given points a , b , and c :

$$colin = |(c.x - a.x) * (b.y - a.y) + (c.y - a.y) * (a.x - b.x)|$$

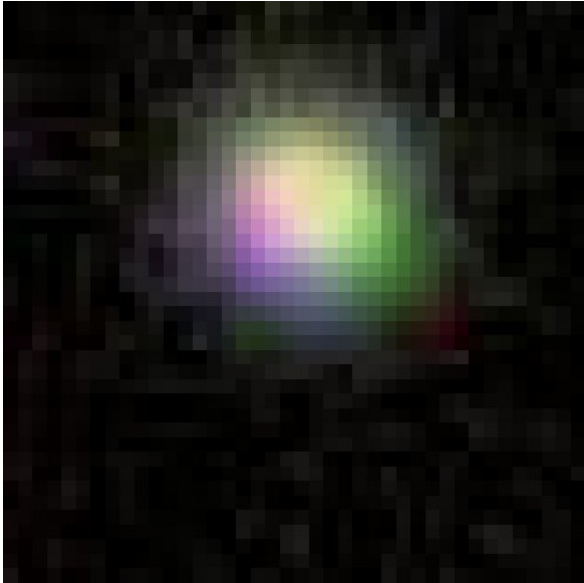
An example asteroid



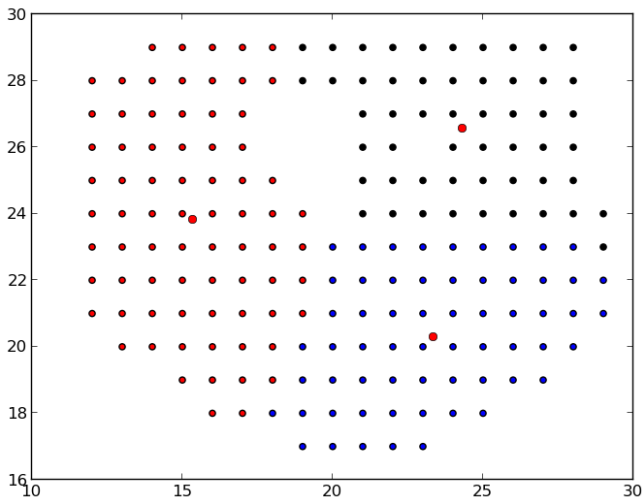
k -means clustering of the same asteroid



An example non-asteroid



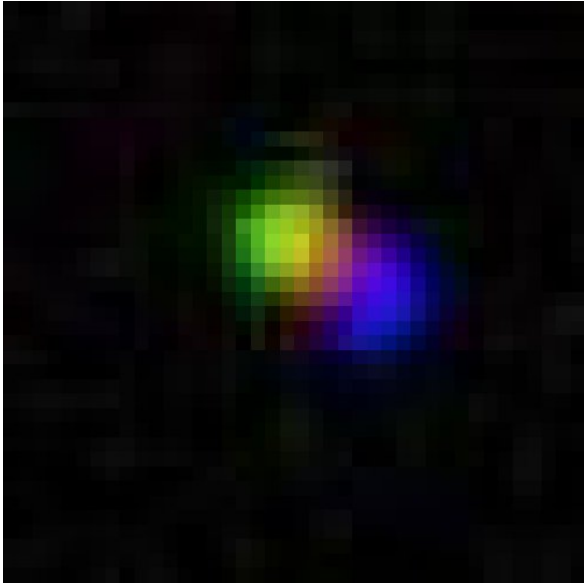
k -means clustering of the same non-asteroid



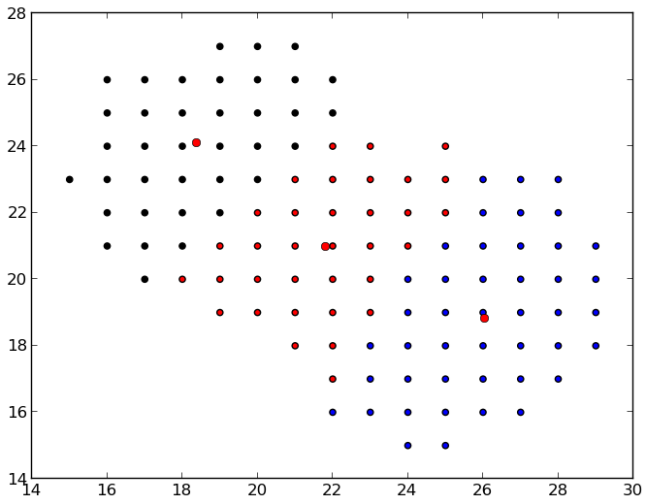
Feature: Best possible average cluster distance

- ▶ Using the same k -means clusters from the previous features
- ▶ What is the average distance from any point in a cluster to the center of the cluster?

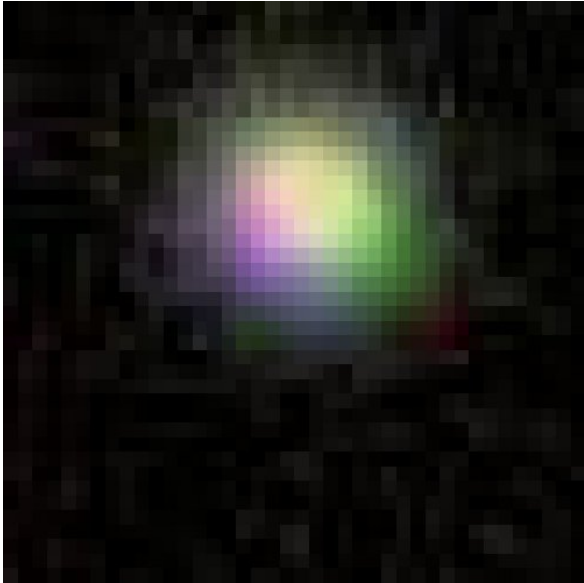
An example asteroid



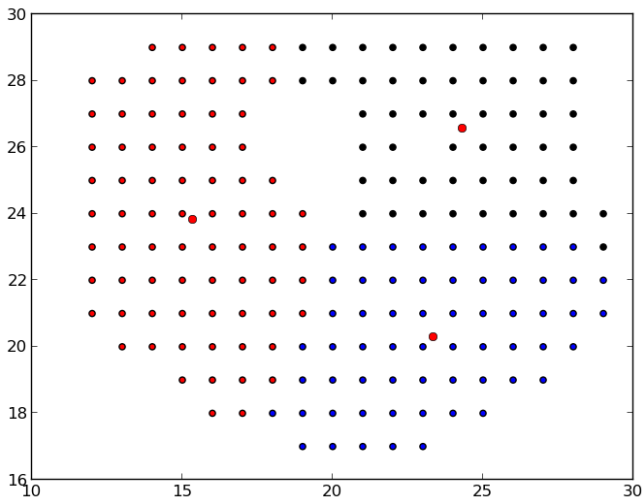
k -means clustering of the same asteroid



An example non-asteroid



k -means clustering of the same non-asteroid



A comparison of all three features

	Hue Ratio	Collinearity	Cluster distance
Asteroid	0.687	0.046	0.432
Non-asteroid	0.376	0.388	0.557

- ▶ We see that the for a valid asteroid:
 - ▶ The hue ratio is much higher
 - ▶ The colinearity metric is much lower
 - ▶ The mean cluster distance is smaller

Ok... where's the AI?

This type of classification is extremely well suited for a neural network:

- ▶ We have a clear set of training data
- ▶ The output is either affirmative (1) or negative (0)
- ▶ Each of the input features can be resolved to a 0 \rightarrow 1 metric
- ▶ There is a small amount of input features which can accurately define an item
- ▶ Neural network activation will be much faster than almost any algorithm we can come up with

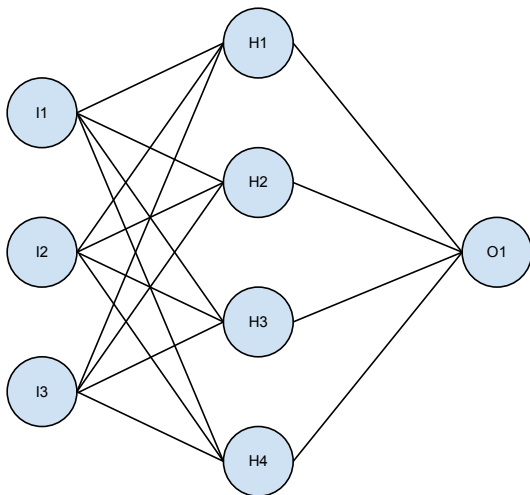
Building the neural network

The resulting neural network:

- ▶ Use supervised learning
- ▶ Uses a backpropagation trainer
- ▶ Three layers:
 - ▶ Input layer
 - ▶ Single hidden layer
 - ▶ Output layer
- ▶ Total of 8 “neurons”:
 - ▶ 3 input neurons (hue ratio, collinearity metric, distance metric)
 - ▶ 4 hidden neurons
 - ▶ 1 output neuron (1 if valid asteroid, 0 if invalid)
- ▶ Learning rate of 0.01, momentum of 0.99

Building the neural network

The resulting neural network:



Training the network

- ▶ Approx 250 valid items
- ▶ Approx 500 invalid items
- ▶ Trained for 5,000 iterations
- ▶ Took approx. 3 hours
- ▶ Probably could have gotten by with less iterations

Results

Trial	Found Valid	Actual Valid	Total	False positive
Trial 1	8	5	190	37.50%
Trial 2	23	21	286	8.70%
Trial 3	54	46	955	14.81%

Results

Trial	Found Invalid	Actual Invalid	Total	False negative
Trial 1	182	182	190	0.00%
Trial 2	263	262	286	0.38%
Trial 3	901	892	955	1.00%

Conclusion

- ▶ Using a neural network allows us to do it faster, and more accurately
- ▶ Need to spend time coming up with good features for the data
- ▶ When paired with human validation, the process would become very quick and very accurate

References

- ▶ <http://www.sdss.org/>
- ▶ <http://pybrain.org/>
- ▶ http://en.wikipedia.org/wiki/Sloan_Digital_Sky_Survey
- ▶ <http://en.wikipedia.org/wiki/Collinearity>
- ▶ http://en.wikipedia.org/wiki/K-means_clustering