

DOC 221 Dinámica orbital y control de actitud

Problems Lecture ADCS - IXA

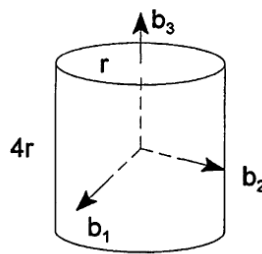
Problem 1:

Consider a cylinder of mass m , radius r , length $4r$ (assume uniform mass distribution).

The body is spinning with initial conditions

$$\vec{\omega}(t=0) = \Omega_0 \left(\frac{3}{5} \vec{b}_1 + \frac{4}{5} \vec{b}_3 \right).$$

Solve for $\omega(t)$.



Problem 2:

A spacecraft is launched into a low Earth orbit. The spacecraft principal moments of inertia are $I_x = 98 \text{ kg} \cdot \text{m}^2$, $I_y = 102 \text{ kg} \cdot \text{m}^2$ and $I_z = 150 \text{ kg} \cdot \text{m}^2$. For stability, the launch vehicle deploys the spacecraft such that it is in a major axis spin when released, with $\omega_z = 0.5 \text{ rad/s}$. Because no deployment is perfect, the spacecraft also has some angular velocity about the other two principle axes, given by $\omega_x = 0.1 \text{ rad/s}$ and $\omega_y = 0.02 \text{ rad/s}$. Making appropriate approximations:

- Describe the resulting spacecraft attitude motion if there are no disturbance torques.
- Determine the nutation angle.
- Determine the precession rate.

Problem 3:

Explore the polhode trajectory (intersection of kinetic energy ellipsoid and angular momentum ellipsoid) for the torque-free rotation for the two particular body geometries where the body is either axisymmetric with $I_1 = I_2$, or the body principal inertias are all equal with $I_1 = I_2 = I_3$. This last condition occurs if the body is a homogenous cube and the body frame B is a principal coordinate frame, or the body is a homogenous sphere with any body fixed frame B.

Problem 4:

The principal inertias of a rigid satellite are given by

$$I_1 = 210 \text{ kg}\cdot\text{m}^2, \quad I_2 = 200 \text{ kg}\cdot\text{m}^2 \quad \text{and} \quad I_3 = 118 \text{ kg}\cdot\text{m}^2.$$

At time t_0 the body angular velocity vector is $\omega = (0.05, 0.02, -0.02)^T \text{ rad/s}$. Numerically solve the resulting torque-free motion for 100 s and plot the resulting attitude in terms of the 3-2-1 Euler angles. Assume the initial Euler angles to be zeros. The kinematic equation for the 3-2-1 Euler angles is given by

$$\begin{bmatrix} \dot{\theta}_1 \\ \dot{\theta}_2 \\ \dot{\theta}_3 \end{bmatrix} = \frac{1}{\cos \theta_2} \begin{bmatrix} \cos \theta_2 & \sin \theta_1 \sin \theta_2 & \cos \theta_1 \sin \theta_2 \\ 0 & \cos \theta_1 \cos \theta_2 & -\sin \theta_1 \cos \theta_2 \\ 0 & \sin \theta_1 & \cos \theta_1 \end{bmatrix} \begin{bmatrix} \omega_1 \\ \omega_2 \\ \omega_3 \end{bmatrix}$$