

# Latent space representations of galaxy images

and their relationship with galactic properties.

---

Nagoya University B4

Dillon Loh

Takeuchi Tsutomu, Suchetha Cooray, Iwasaki Daiki

# Flow of today's presentation

## 1. **Introduction**

- Galactic Morphologies
- Morphological vs Physical Properties

## 2. **Data**

- Data Source
- Preprocessing

## 3. **Method**

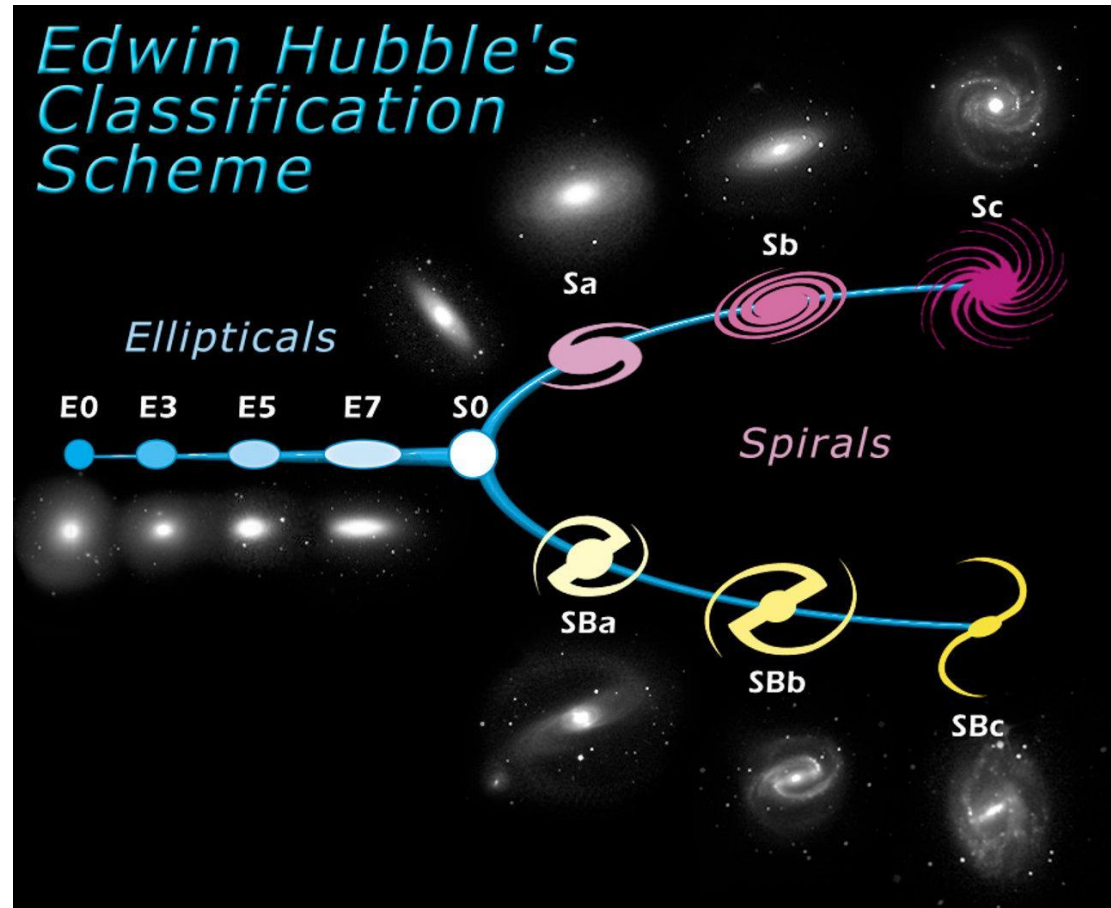
- What are Neural Networks?
- What are Variational Autoencoders?

## 4. **Application to Astrophysics**

## 5. **Conclusion/Future Work**

# Galaxy Morphologies

- Galaxies are systems of stars and interstellar matter.
  - Come in all sorts of shapes and have defining features
- Morphology of a galaxy refers to its shape/features (i.e. stellar mass distribution)
- Galaxies that do not fall under Ellipticals or Spirals are called 'Irregulars'



NASA

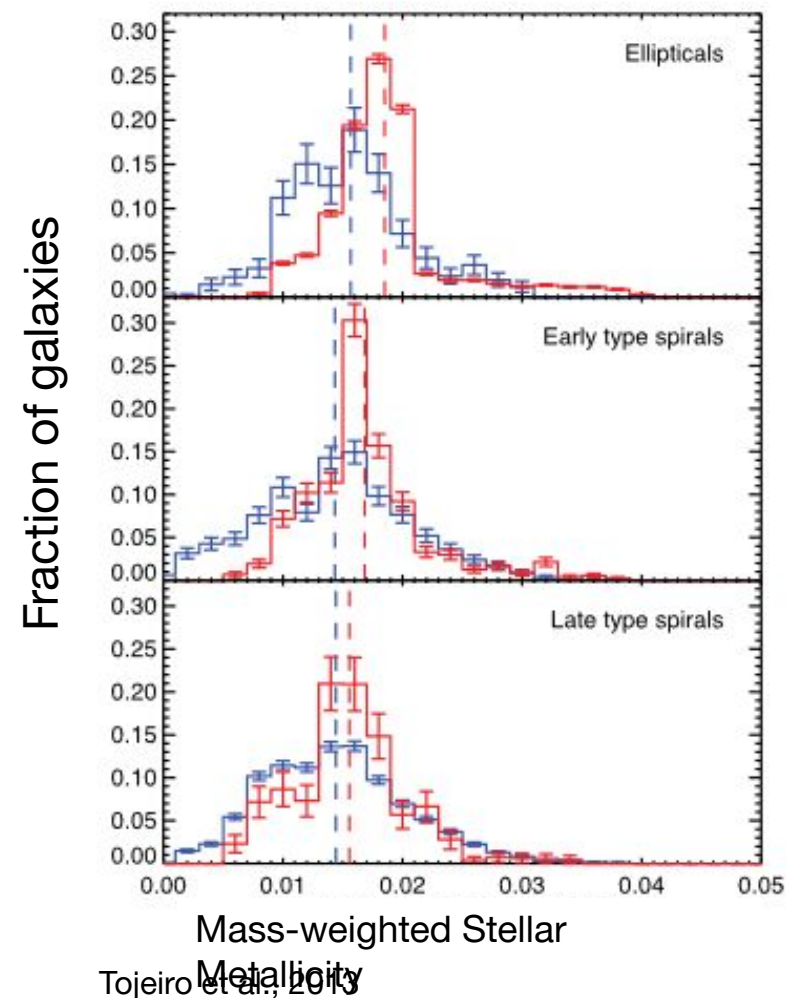
# Morphology vs Physical Properties

In general, there are well-defined connections between the morphological properties of a galaxy and its physical property. \*

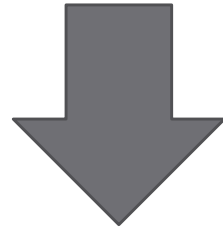
- Ellipticals tends to be older.\*
- Ellipticals tend to have higher metallicity
- Spirals tend to have higher SFR than Ellipticals\*\*
- Spirals' bulges tend to be older than their disc components

**Can we predict a galaxy's physical properties just from the features in its image?**

Fraction of galaxies with different metallicities for Ellipticals and Blue/Red Spirals



**Why do we want to predict a galaxy's physical properties just from the features in its image?**



**Direct calculation of physical properties is:**

- 1. Calculation/time intensive.**
- 2. Requires us to collect large amounts of other data.**

# Method

# What needs to be done?

We need a way to extract features from the images autonomously.

## Past works

*“Eigengalaxies: describing galaxy morphology using principal components in image space” – 2020*

- Demonstrated that galactic images could be decomposed into a linear sum of ‘eigen-galaxies’ via PCA.

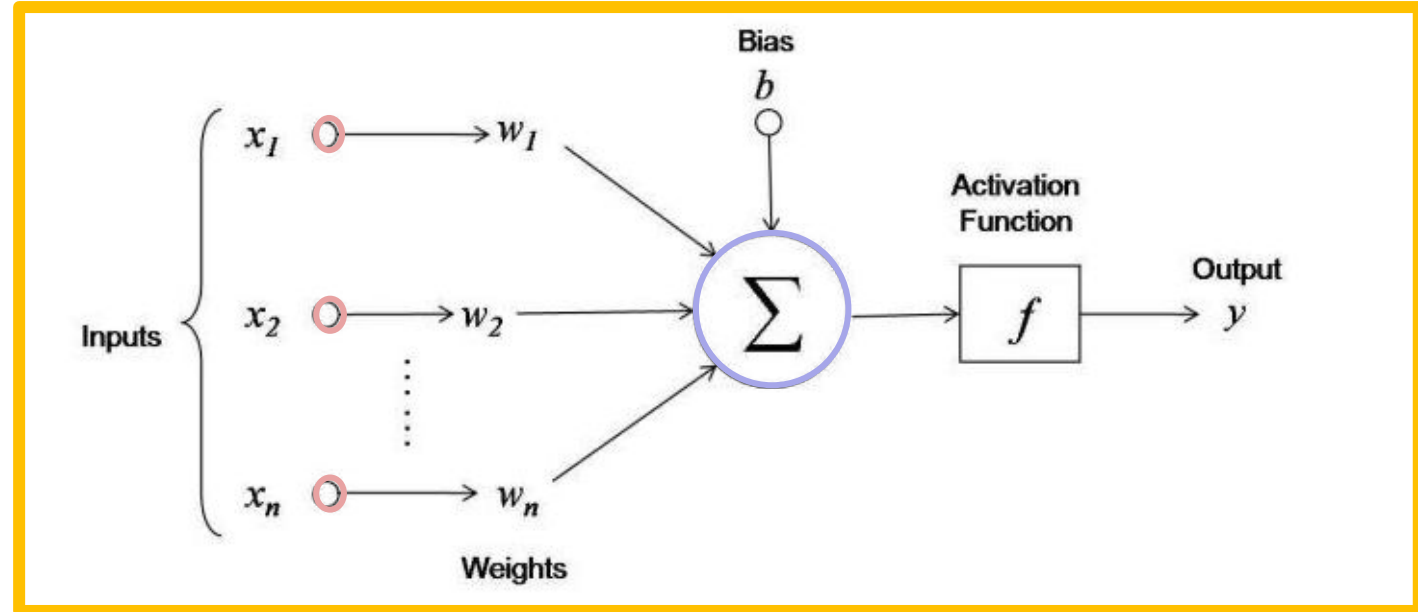
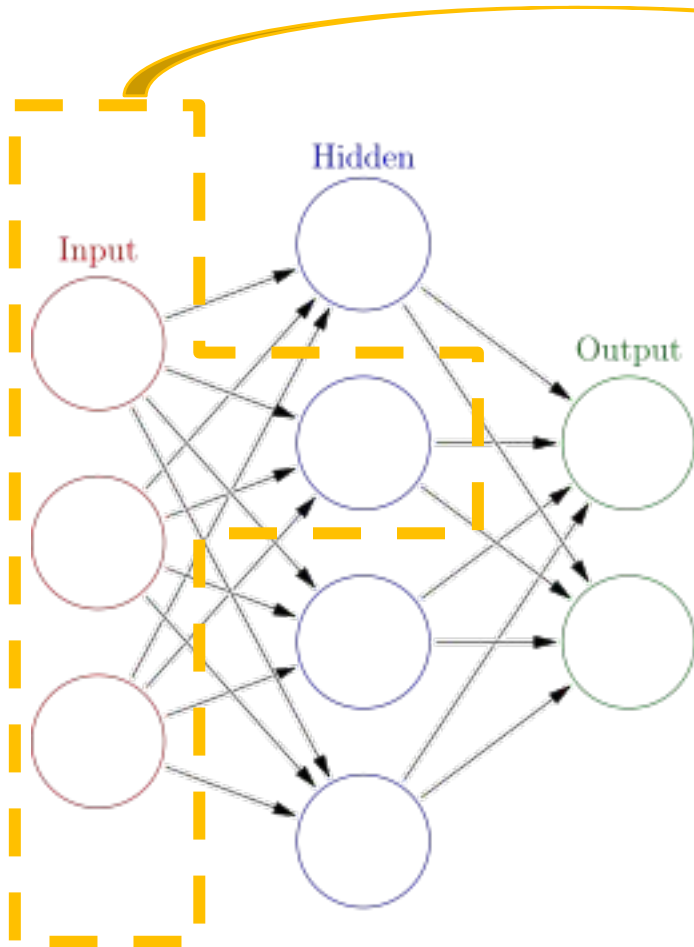
## This work

We use unsupervised deep learning (autoencoders).



(Uzeirbegovic et al.,

# Neural Networks

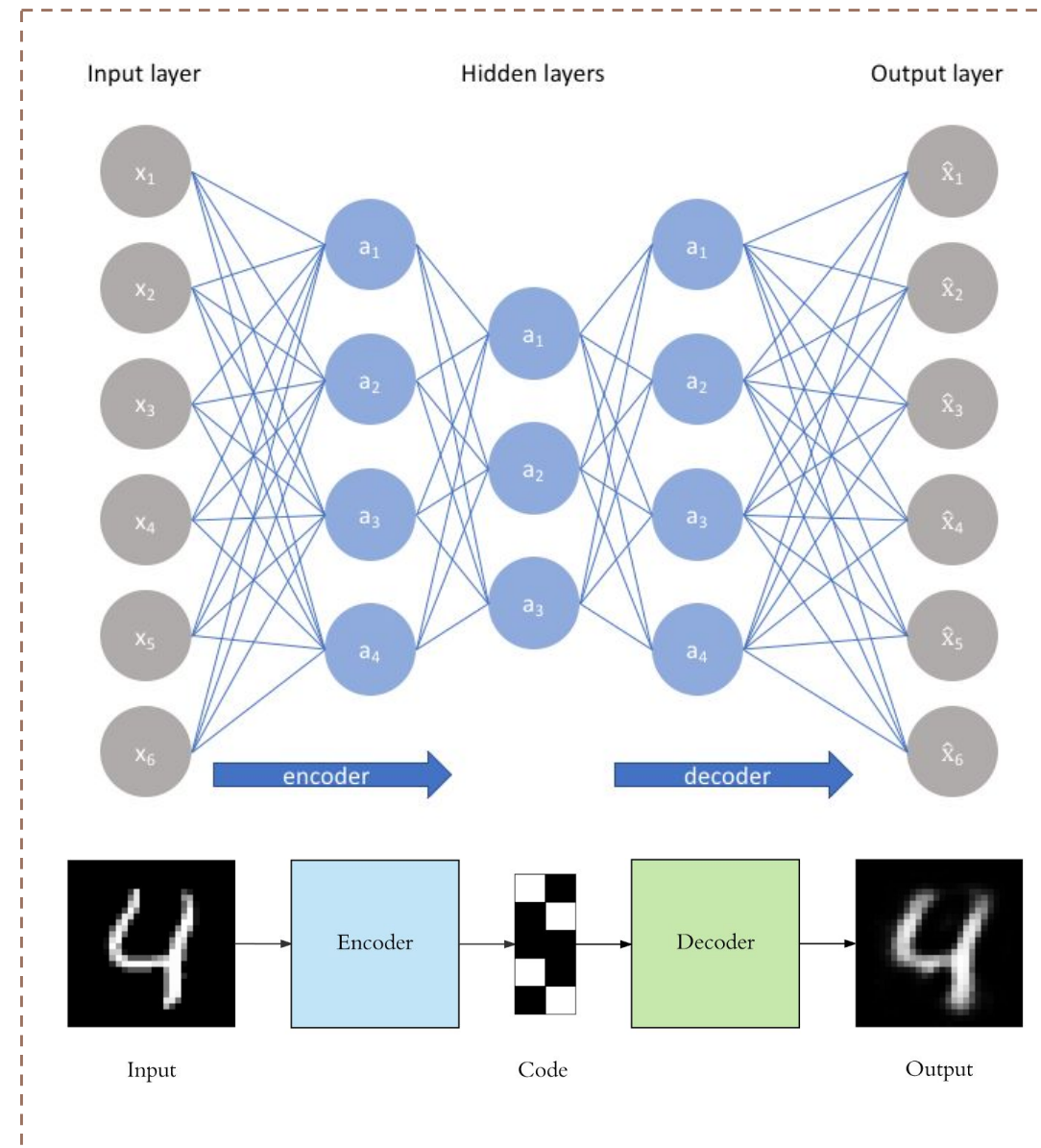


- Each **node** sums the (**inputs** \* **weight**) from ALL the nodes of the previous layer.
- The sum is passed through an activation function.
- Activation function output is passed as inputs to next layer's nodes.
- "Training" the NN involves adjusting the **weight** until it gives "good" outputs from the **final layer**.



# Autoencoders

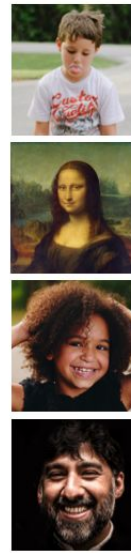
- Autoencoders are a type of feed forward neural network.
- Consists of a 'encoder' and 'decoder'.
- Reconstruction-loss function helps check and guide training progress
- The encoder is trained to extract features from the image.
  - Each node in the bottleneck layer is a 'feature'.
  - When an image is passed into the network, it outputs one fixed value from each node.
  - These values are the 'amounts' of each feature in the image.





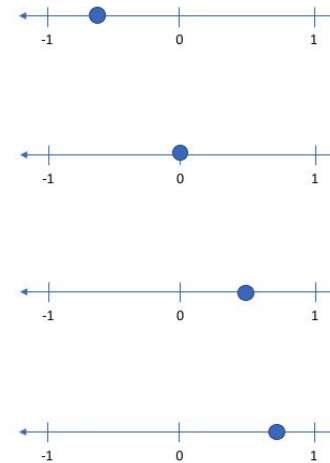
# Variational Autoencoders

- Unlike a basic autoencoder, a variational autoencoder (VAE) represents each latent variable as a **probability distribution**.
- We can 'shape' our distribution to make it smooth.
- Typically, we use the very simple Gaussian distribution.
- To encourage our latent variables to resemble a Gaussian, we add a loss term that measures the difference between the latent and Gaussian distributions
  - Kullback-Leibler Divergence Term



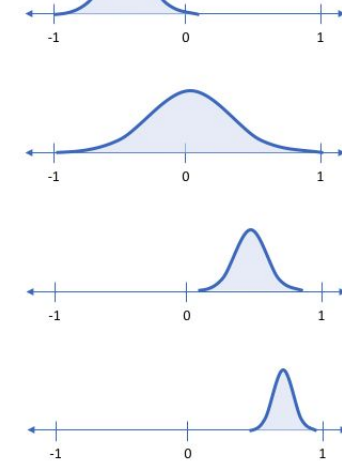
## Autoencoder

Smile (discrete value)



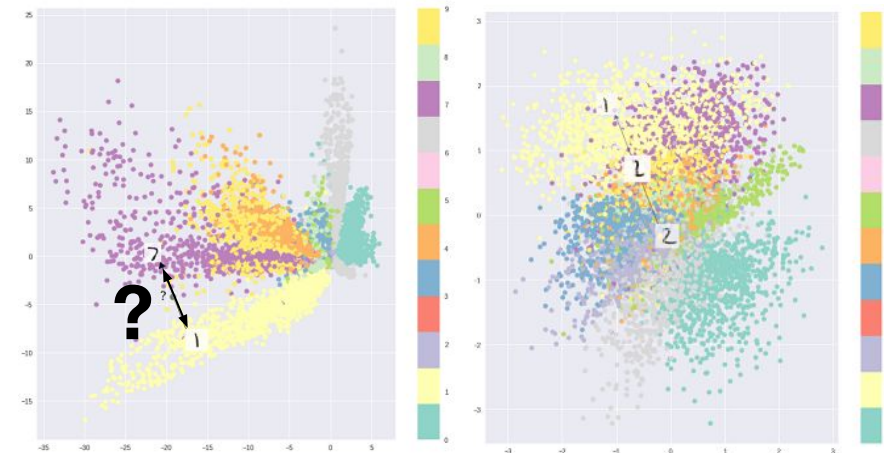
## VAE

Smile (probability distribution)



vs.

<https://www.jeremyjordan.me/variational-autoencoders/>

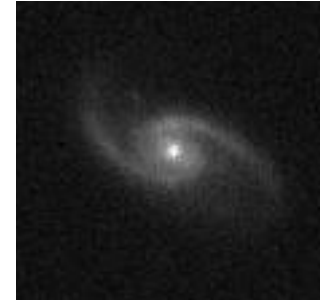


# Data

# Data used

- Sloan Digital Sky Survey (SDSS) is a multi-spectral imaging and redshift survey.
  - Data Release 17 optical images are downloaded via SkyView\*.
- SDSS *i* Band Images Used
  - Near-infrared wavelengths sensitive to long-living stars' emission
  - Representative of stellar mass distribution
- 150x150 pixel images

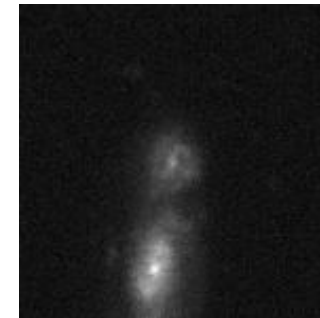
## Example Images



Example Image



Relatively small object



Center object  
disturbed by extra  
objects



Presence of extra  
objects that have  
negligible disturbance

<https://skyview.gsfc.nasa.gov/current/cgi/titlepage.pl>

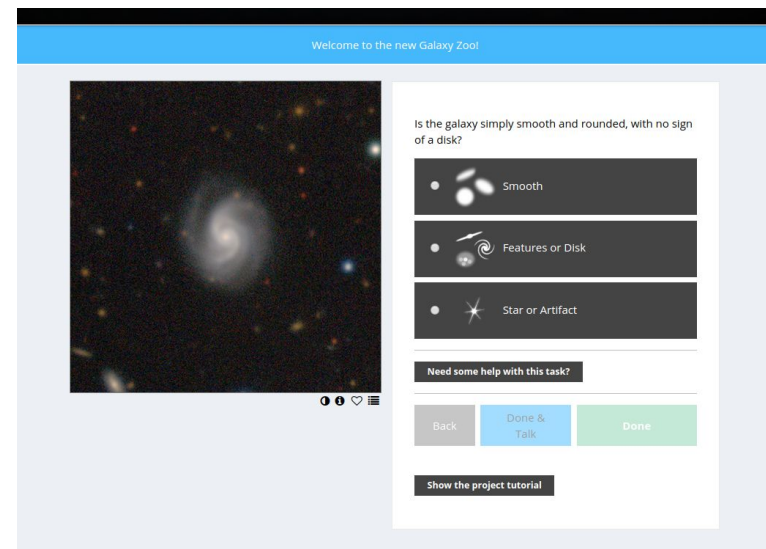
# Filtering by Galaxy Morphology with Galaxy Zoo 2

Galaxy Zoo is a project **where humans assist in the manual identification of morphological features.**

Using this data, we filtered out the following:

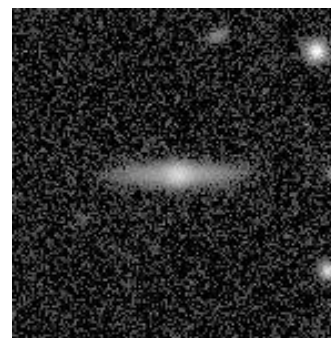
- Edge-on galaxies
  - Difficult to observe morphological features.
- Galaxies with 'oddities'
  - Galactic merging, overlap, lensing, dust lane, etc.

Final Dataset: 12539 galaxies

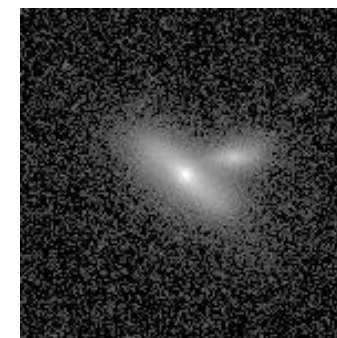


<https://www.zooniverse.org/projects/zookeeper/galaxy-zoo/>

Edge-on



Oddities

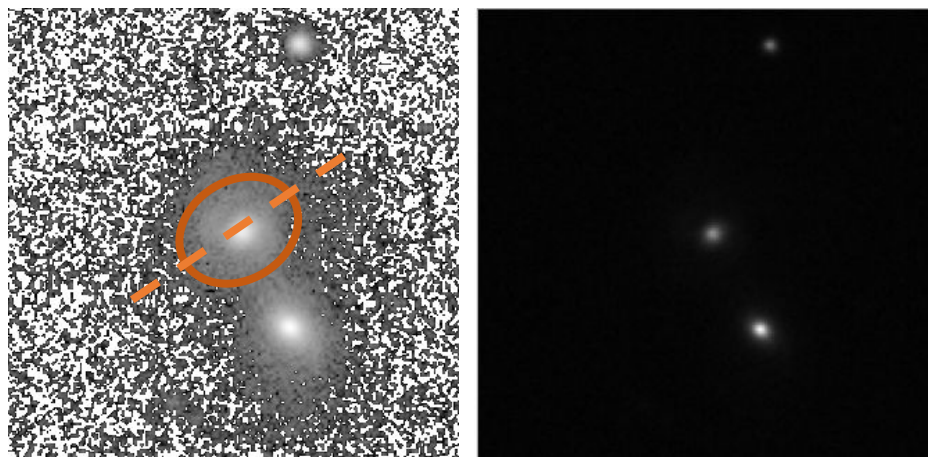


<https://skyview.gsfc.nasa.gov/current/cgi/titlepage.pl>

# Data Preprocessing

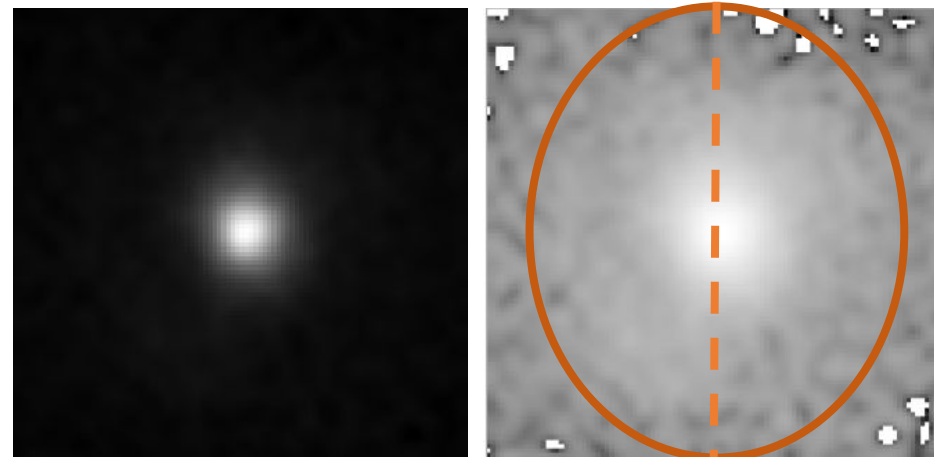


Original Image



Log-scaled

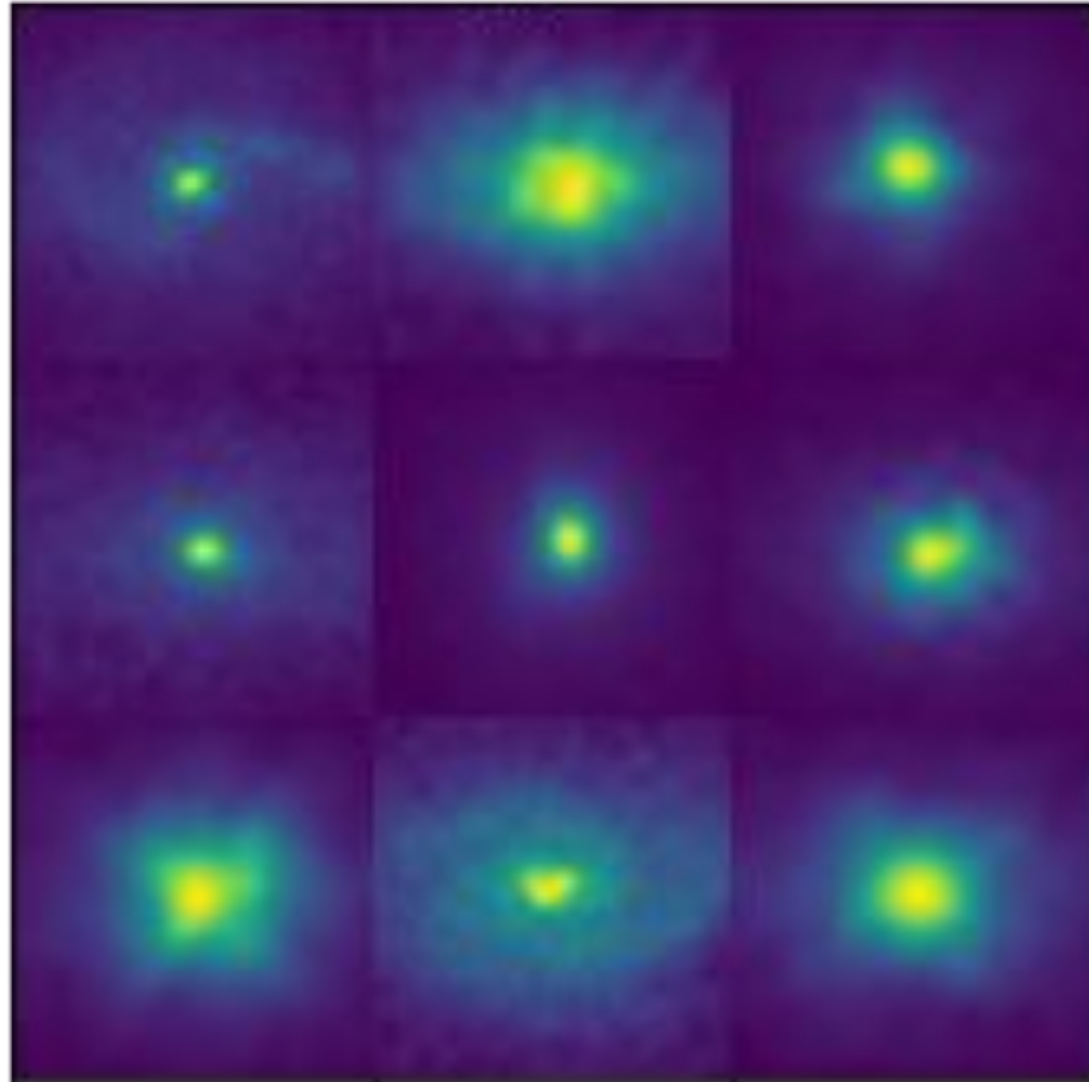
Final Image



Log-scaled



# Data Examples



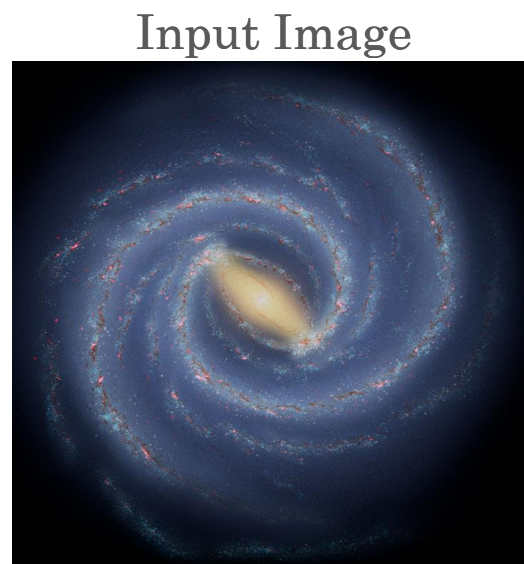
# **Astrophysics Applications**



# How did we use our autoencoders?

- We trained our autoencoder to find a set of the most common features in all our galactic images
- If we pass a galactic image into this autoencoder, it will give us a probability distribution of values for each feature.
  - If we passed it into a normal AE, it would just give a single value for each feature.
- The values represent ‘how much’ of that feature is present in the input galactic image.

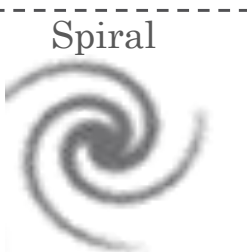
## EXAMPLE



Encoding  
→

### Features

Spiral



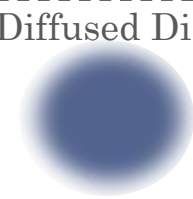
Bar



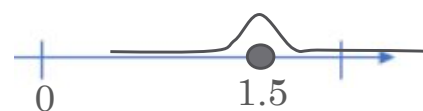
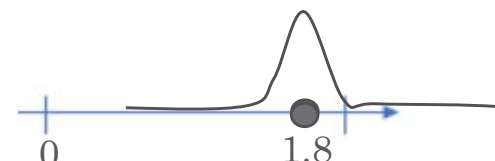
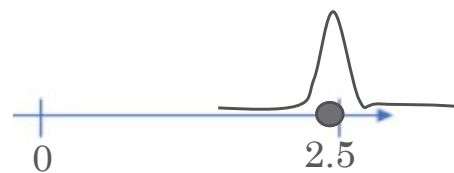
Center Bulge



Diffused Disc



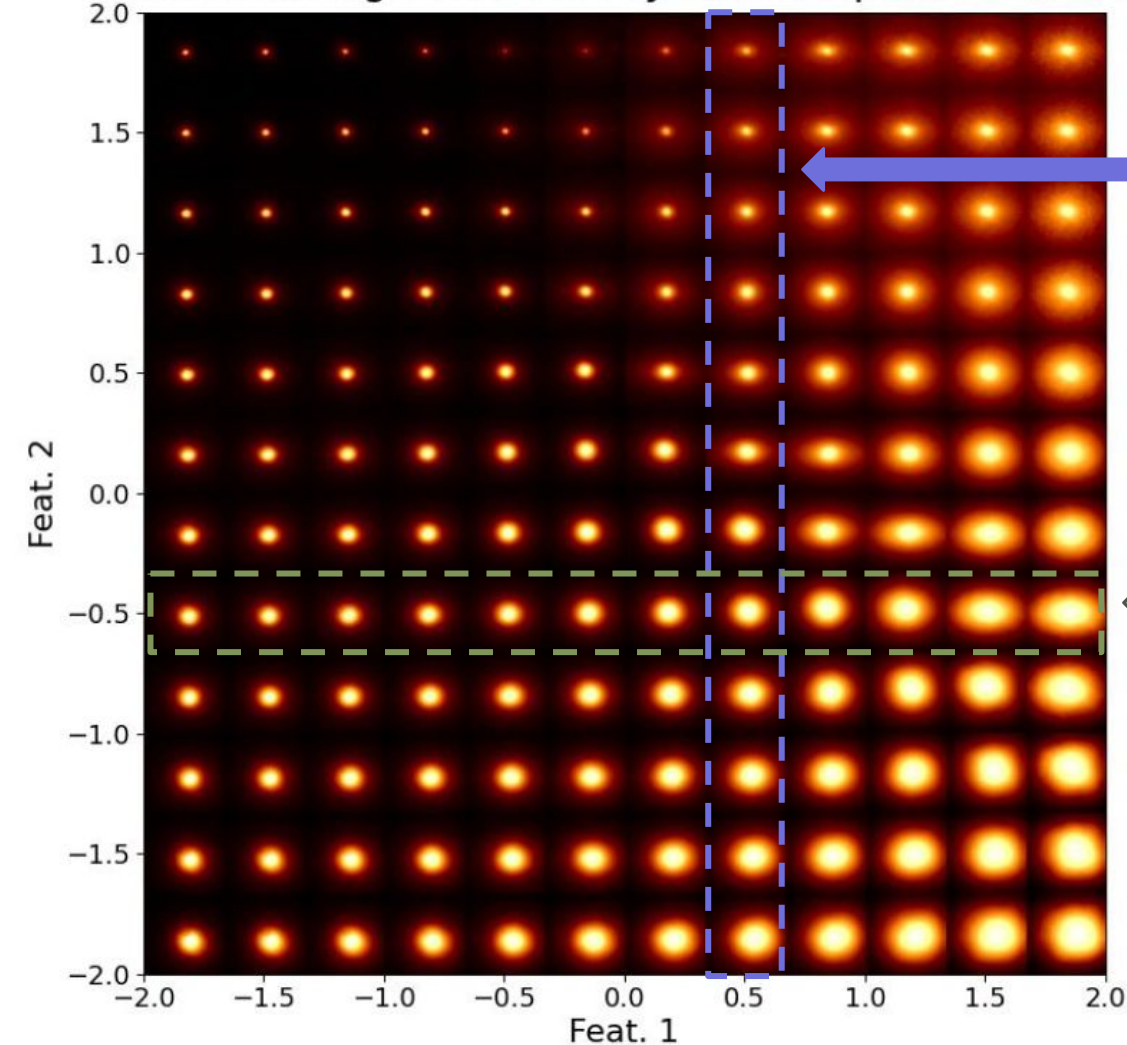
### “Amount” of features



# Extraction of Morphological Features

We set our VAE to find only 2 latent features that best describes the images.

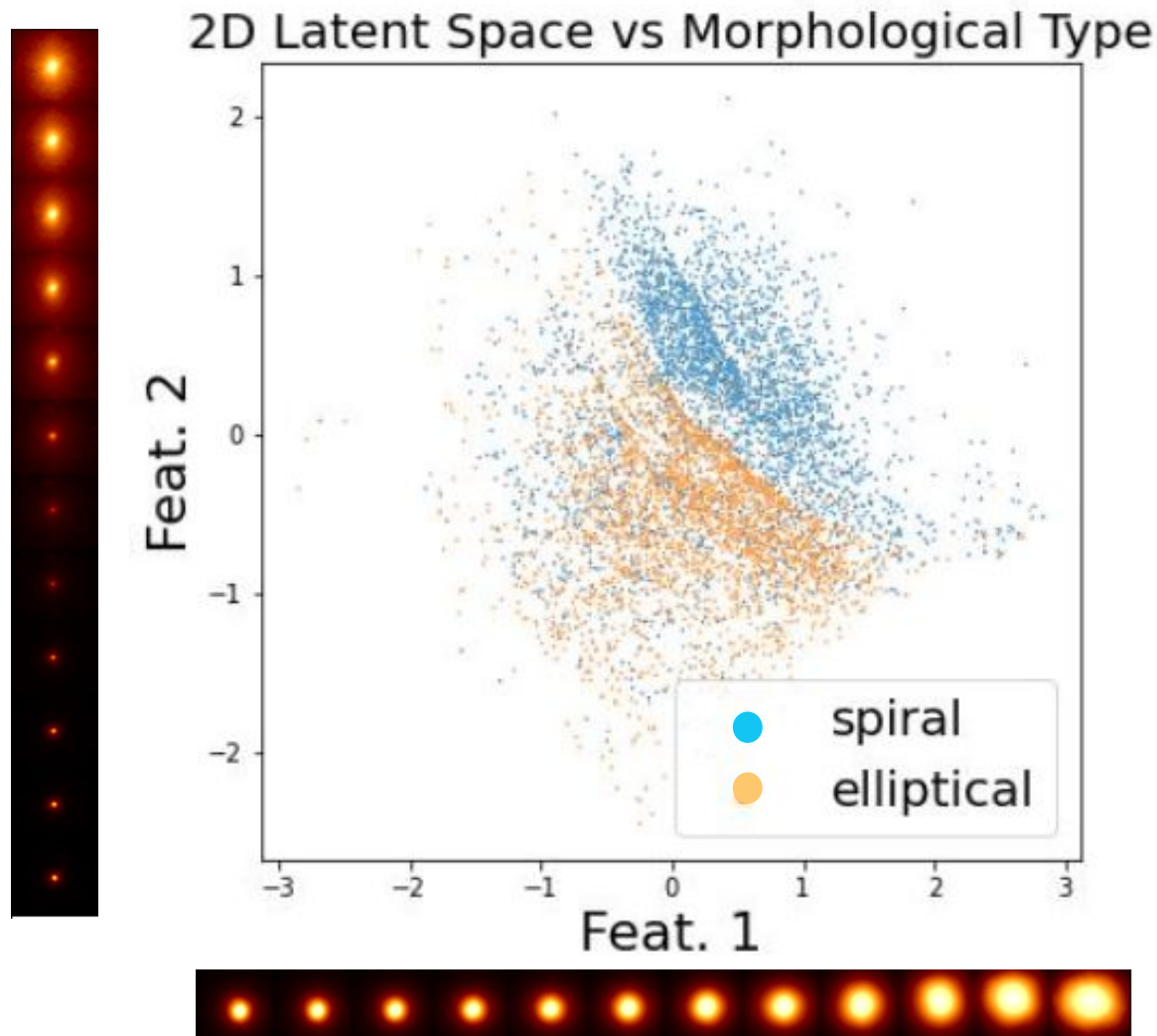
Galactic Images created by Latent Space Features



Higher Feat. 2 corresponds to bigger diffused disc-like structures.

Higher Feat. 1 corresponds to bigger center bulges

# Morphological Class Distribution in Latent Space



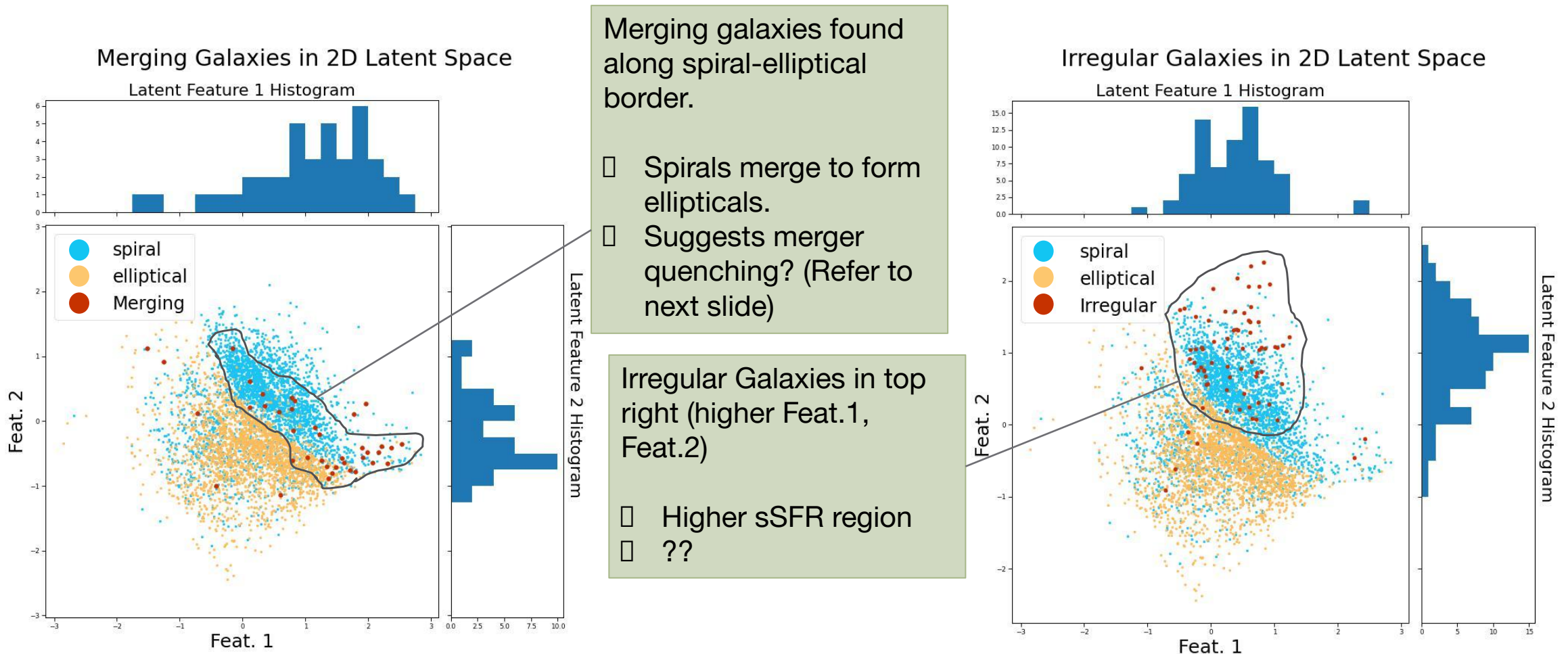
	Feat. 1	Feat. 2
<b>spiral</b>	0.322450	0.198545
<b>elliptical</b>	0.137961	-0.344143

- Spirals tend to have higher Feat. 2 while Ellipticals have lower.
- Ellipticals tend to have higher amounts of Feat. 1
- Surprisingly, distribution of spirals is skewed towards greater Feat. 1

# What about irregular galaxies?

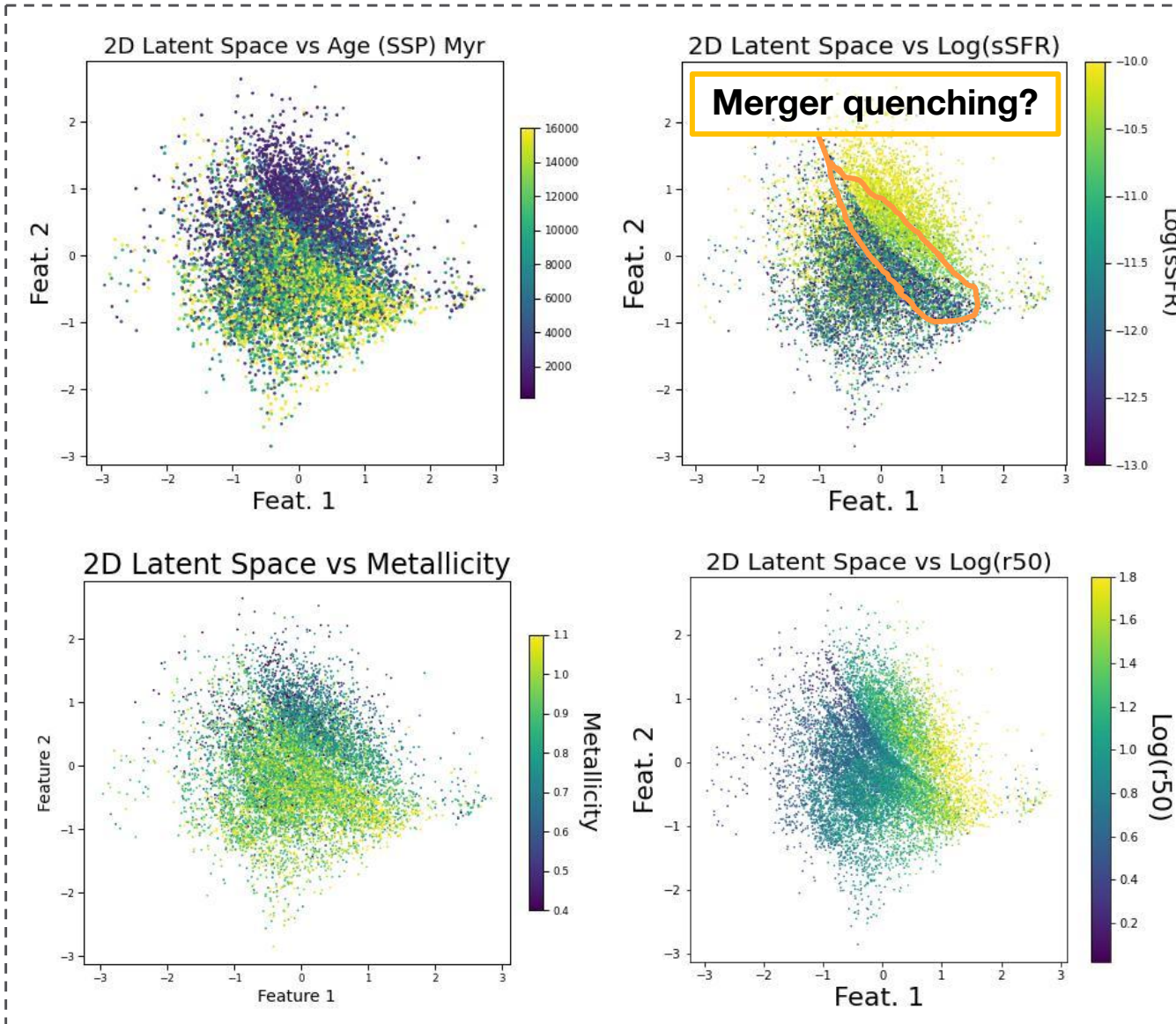
What if we now encoded images of irregular/merging galaxies?

- Recall that we trained only with 'regular' galaxies (Spirals/Ellipticals)





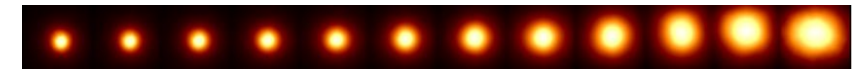
# Latent Space vs Physical Properties



- We see that there are gradients in the physical properties across Latent Space.
- Some of these make intuitive sense based on our knowledge of galactic structures.



For example



Feat. 1 ↑

- Since stars in center bulges tend to be older, a larger bulge implies an overall older galaxy (assuming equally sized galaxies)



Feat. 2 ↑

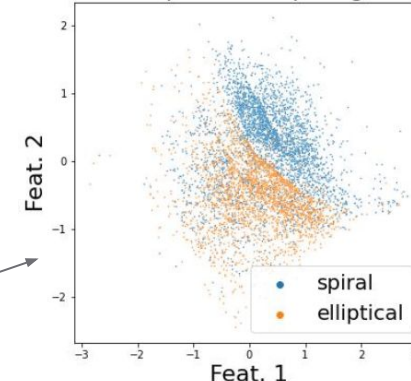
- Since SFR in spiral arms/discs tend to be higher, larger proportion of them in galaxy => higher sSFR.

# Conclusion/Discussion

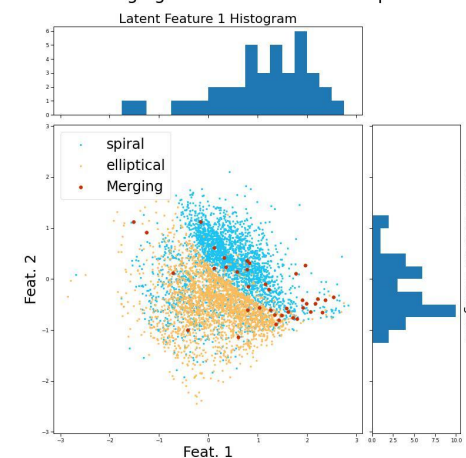
# Conclusion

- Our VAE was able to find a set of just 2 features that distinctly defines the morphology of spiral and elliptical galaxies.
- Even though we only trained with ‘regular’ galaxies, our VAE could still be applied to ‘irregular’ galaxies with physically interpretable results.
- The image latent features our VAE found also had monotonic relationships with actual physical properties.

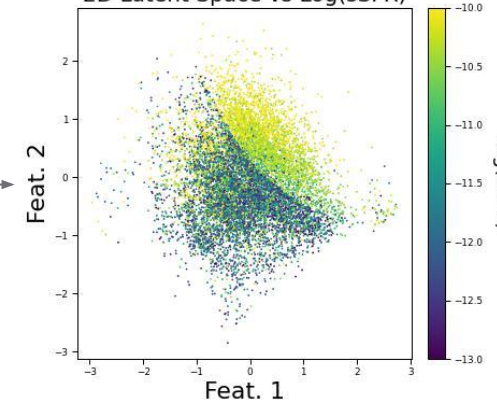
2D Latent Space vs Morphological Type



Merging Galaxies in 2D Latent Space



2D Latent Space vs Log(sSFR)



# Future Work

- Use multiple band images to better capture other sources of radiation/physical properties.
- Use maps of other physical properties such as emission line maps or stellar velocity dispersion maps.
- Build parametric models based on these latent feature/physical property relationships.

I would greatly appreciate further suggestions to improve on this work!



Thank you

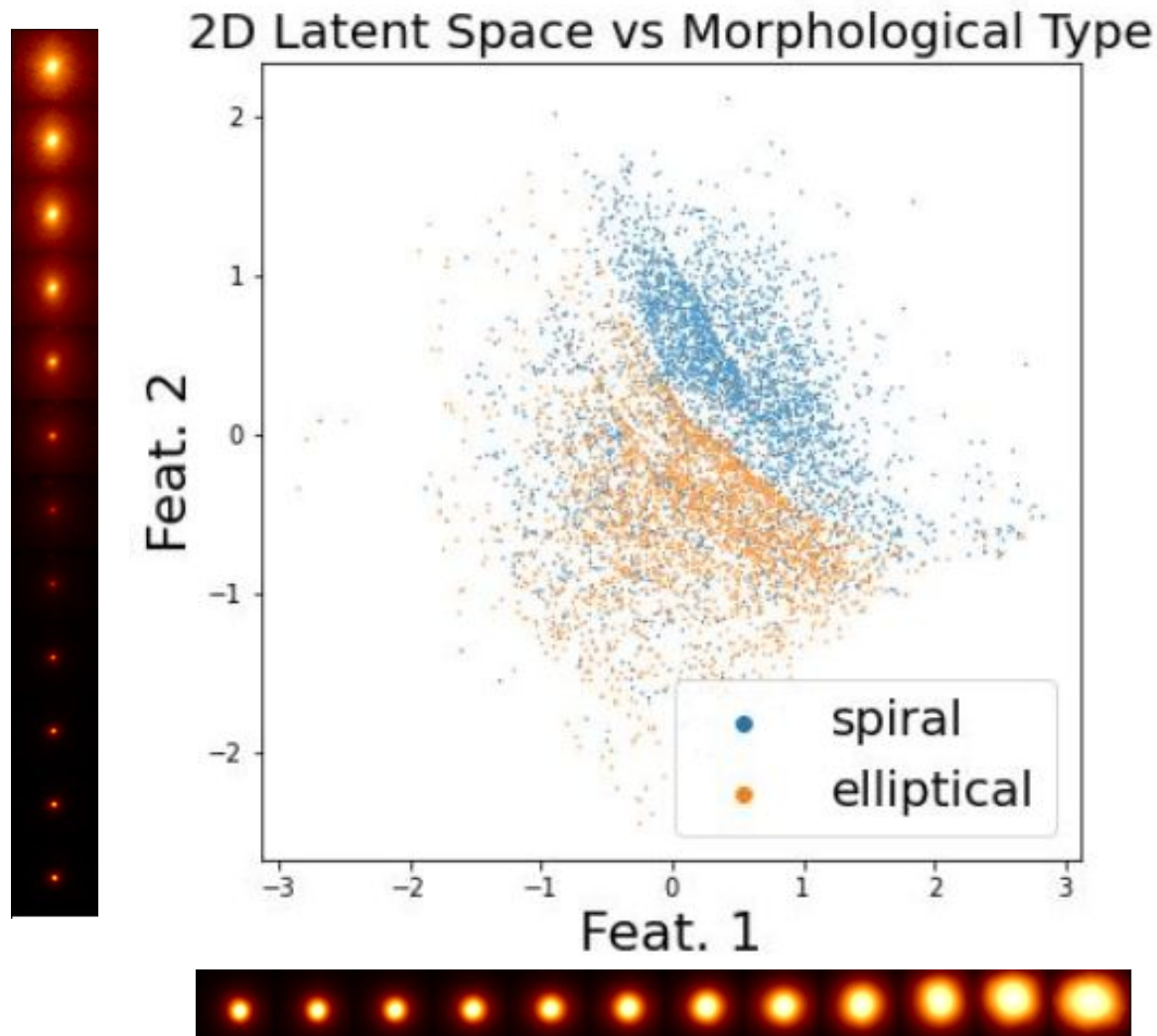
# Correlations Tables

	Feat. 1	Feat. 2
<b>Vel. Dispersion (SSP) km/s</b>	-0.051947	-0.465508
<b>Age (SSP) Myr</b>	0.034974	-0.417873
<b>Metallicity (SSP)</b>	0.019536	-0.455701
<b>Vel. Dispersion (exp SFH) km/s</b>	-0.058639	-0.472199
<b>Age (exp SFH) Myr</b>	-0.049727	0.466863
<b>Metallicity (exp SFH)</b>	0.057441	-0.445452
<b>Petrosian 50 Radius arcsec</b>	0.717488	-0.033900
<b>LOGSFRSED</b>	-0.058900	0.045080
<b>LOGMSTAR</b>	-0.069458	-0.008703
<b>LOGSSFR</b>	0.110948	0.334421

	Feat. 1	Feat. 2
<b>spiral</b>	0.018326	-0.390460
<b>elliptical</b>	0.282897	-0.045579
<b>uncertain</b>	-0.249939	0.376698

	Feat. 1	Feat. 2
<b>Has Smooth Profile</b>	-0.123796	0.563262
<b>Has Features or Disk</b>	0.126552	-0.561337
<b>Has Spiral Arms</b>	-0.022620	-0.425911
<b>Has Obvious Bulge</b>	0.292862	0.088417
<b>Is Completely Round</b>	0.172631	0.014948

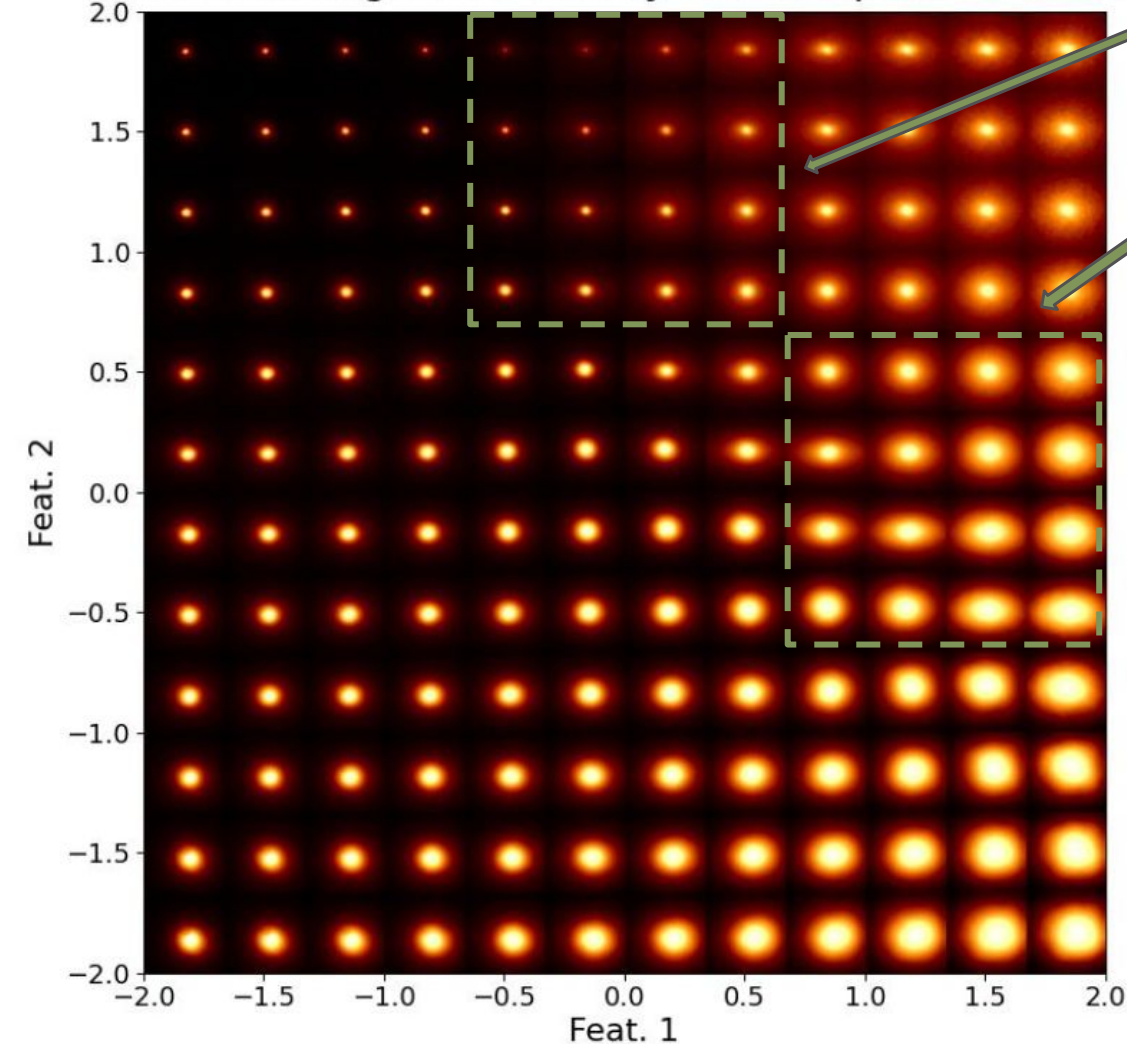
# Morphological Class Distribution in Latent Space



- Spirals tend to have higher Feat. 2 while Ellipticals have lower.
- Ellipticals tend to have higher amounts of Feat. 1
- Surprisingly, distribution of spirals is skewed towards greater Feat. 1

# Morphological Class Distribution in Latent Space

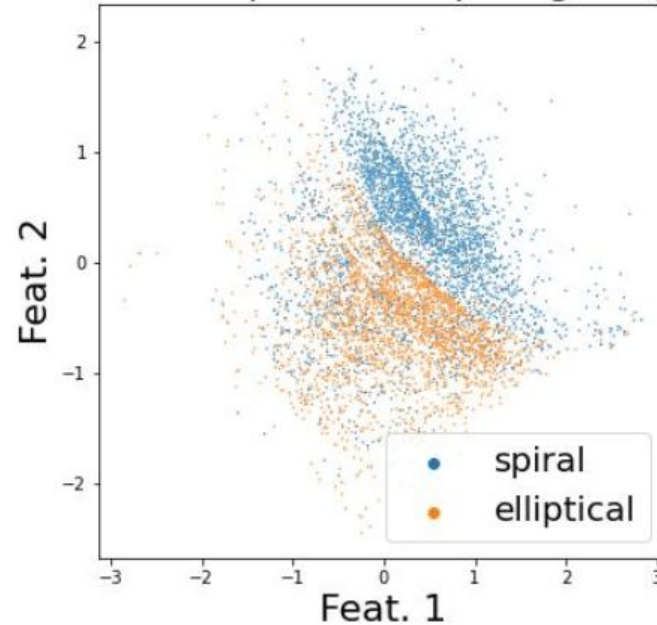
Galactic Images created by Latent Space Features



Higher Feat. 2 corresponds to more diffused disc-like structures.

Higher Feat. 1 corresponds to bigger center bulges

2D Latent Space vs Morphological Type

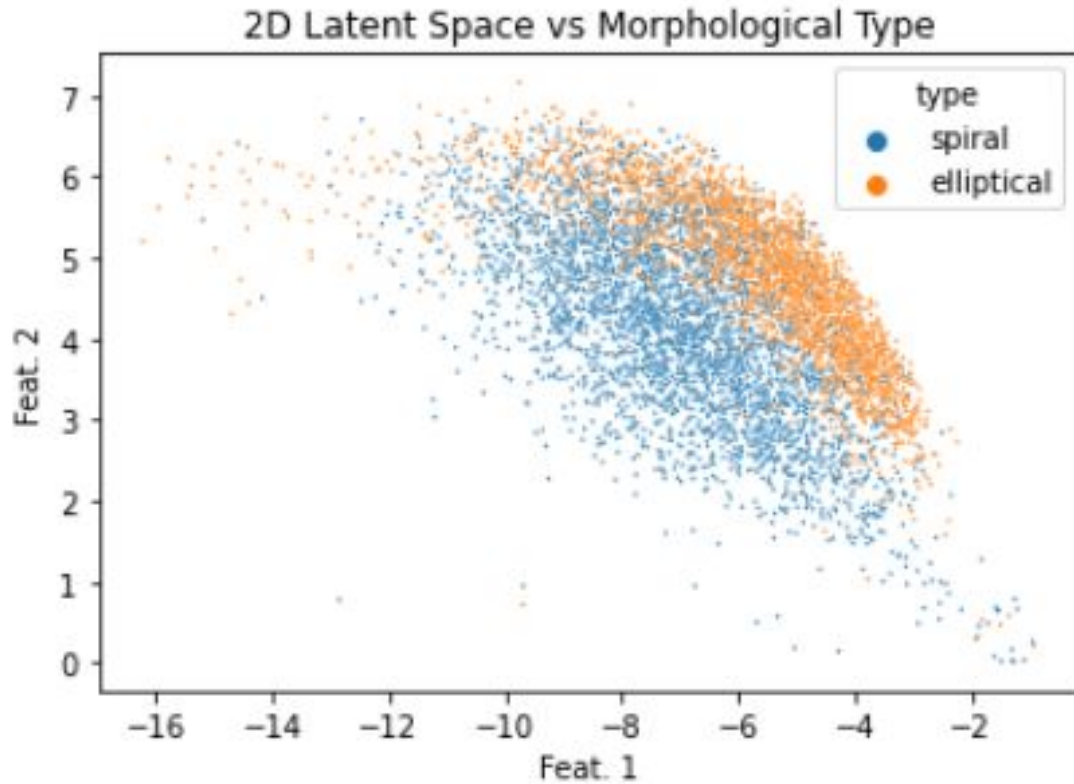


Elliptical galaxies have more less Feat. 2 while Spirals have more.

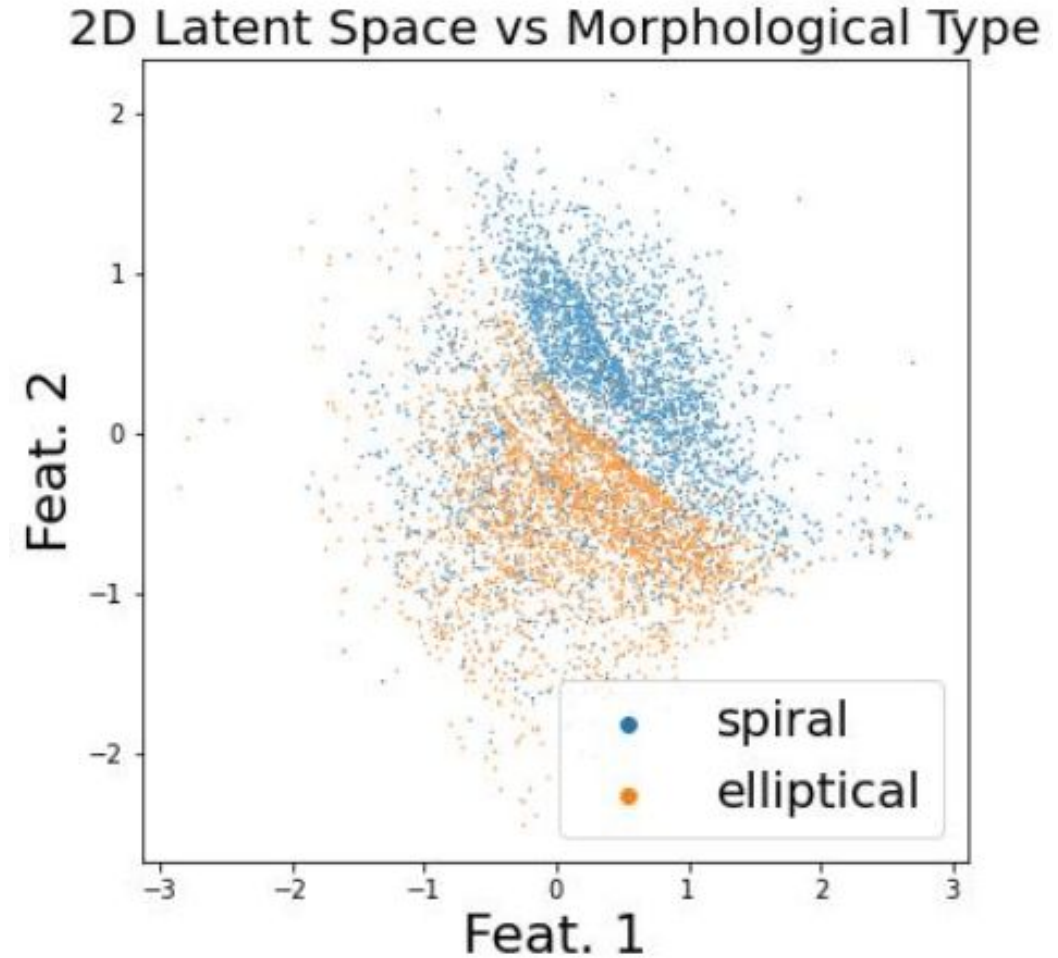
Both classes are similarly distributed along Feat. 1



# Morphological Class Distribution in Latent Space



Autoencoder



Variational Autoencoder

