

Problem 1.

- (a) Assuming the ansatz
- $E(x, t) = X(x)T(t)$
- , the wave equation becomes

$$\begin{aligned}\frac{\partial^2 E}{\partial x^2} &= \frac{1}{c^2} \frac{\partial^2 E}{\partial t^2} \\ \implies X''(x)T(t) &= \frac{1}{c^2} X(x)T''(t) \\ \implies \frac{X''}{X} &= \frac{1}{c^2} \frac{T''}{T}.\end{aligned}$$

As the LHS is a function of only x and the RHS is a function of only t , both sides must be equal to some separation constant, which we define as $-k^2$. Thus

$$\begin{aligned}X'' &= -k^2 X \\ T'' &= -c^2 k^2 T\end{aligned}$$

The solutions to each of these equations are

$$\begin{aligned}X(x) &= A \sin(kx) + B \cos(kx) \\ T(t) &= C \sin(ckt) + D \cos(ckt).\end{aligned}$$

As we have the initial condition $E(x, 0) = X(x)T(0) = 0$, a nontrivial solution must have that $T(0) = 0$, hence $D = 0$ by substitution into the above equation. Thus

$$T(t) = C \sin(ckt).$$

- (b) We thus have that
- E
- is of the form (absorbing
- C
- into the other constants):

$$E(x, t) = (A \sin(kx) + B \cos(kx)) \sin(ckt).$$

Applying the boundary condition $E(0, t) = 0$:

$$\begin{aligned}0 &= B \sin(ckt) \\ \implies 0 &= B.\end{aligned}$$

Applying the boundary condition $E(L, t) = 0$:

$$\begin{aligned}0 &= A \sin(kL) \sin(ckt) \\ \implies 0 &= \sin(kL) \\ \implies kL &= n\pi, \quad n = 1, 2, 3, \dots \\ \implies k &= \frac{n\pi}{L}, \quad n = 1, 2, 3, \dots\end{aligned}\tag{1}$$

We thus have a family of solutions indexed by n :

$$E_n(x, t) = C_n \sin\left(\frac{n\pi}{L}x\right) \sin\left(\frac{cn\pi}{L}t\right).$$

- (c) Defining
- $\omega = ck$
- , the allowed angular frequencies are

$$\omega_n = \frac{cn\pi}{L}, \quad n = 1, 2, 3, \dots$$

For each mode of the wave, we define N as its index in order to count the number of possible modes. We want to know how a change in N leads to a change in the wavenumber k . As the number of modes

becomes very large, a small change is comparatively infinitesimal, and so we can approximately use the language of differentials. In 1D, from (1):

$$dN = \frac{L}{\pi} dk.$$

In 3D, we can choose modes along each of the three axes, yielding an $(L/\pi)^3$ term. However, we also must account for polarization, thus

$$dN = 2 \left(\frac{L}{\pi} \right)^3 d^3k.$$

In abstract \mathbf{k} -space, we treat the volume element d^3k as a spherical shell of radius k and width dk centered at the origin. We are only concerned with positive k values, so we pick up a factor of $1/8$ as well:

$$d^3k = \frac{2\pi k^2}{8} dk.$$

As we are concerned with angular frequencies, we substitute $k = \omega/c$:

$$dN = 2 \left(\frac{L}{\pi} \right)^3 \cdot \frac{4\pi(\omega/c)^2}{8} \cdot \frac{d\omega}{c} = \frac{L^3 \omega^2}{\pi^2 c^3} d\omega. \quad (2)$$

(d) We will use Equation 2 to compute the energy density per unit ω , given as

$$\mathcal{U}(\omega) = \frac{1}{L^3} \frac{\partial E}{\partial \omega}.$$

We can compute $d\omega$ from above, now we need dE . From the Equipartition Theorem, each degree of freedom in a statistical system has the same amount of energy:

$$E = \underbrace{\frac{1}{2}k_B T}_{\text{K.E.}} + \underbrace{\frac{1}{2}k_B T}_{\text{P.E.}} = k_B T,$$

where k_B is Boltzmann's constant. Thus

$$\begin{aligned} dE &= k_B dT \\ &= k_B T dN \\ &= \frac{L^3 \omega^2}{\pi^2 c^3} d\omega \\ \implies \frac{\partial E}{\partial \omega} &= \frac{L^3 \omega^2}{\pi^2 c^3}. \end{aligned}$$

Therefore, we obtain the Rayleigh-Jeans law:

$$\mathcal{U}(\omega) = \frac{1}{L^3} \frac{\partial E}{\partial \omega} = \frac{k_B \omega^2 T}{\pi^2 c^3}.$$

(e) Suppose our black body is made of atoms with two energy levels, a ground state $|g\rangle$ and an excited state $|e\rangle$. Classically, the ratio between the number of systems in $|e\rangle$ vs the number in $|g\rangle$ is given by the Boltzmann distribution:

$$\frac{N_e}{N_g} = e^{-E/(k_B T)}.$$

Let \bar{n} count the average number of photons in the black box and $p_{e \leftrightarrow g}$ the probability of transitioning between either state. Then the absorption rate is $N_g \bar{n} p_{e \leftrightarrow g}$, and the emission rate is $N_e (\bar{n} + 1) p_{e \leftrightarrow g}$.

Assuming thermal equilibrium, the system must be in detailed balance, and so these rates must be the same:

$$\begin{aligned} N_g \bar{n} p_{e \leftrightarrow g} &= N_e (\bar{n} + 1) p_{e \leftrightarrow g} \\ \implies \frac{\bar{n}}{\bar{n} + 1} &= \frac{N_e}{N_g} = e^{-E/(k_B T)}. \end{aligned}$$

Taking Planck's assumption that the atoms can only absorb or emit photons of discrete chunks $E = \hbar\omega$, we get

$$\begin{aligned} \frac{\bar{n}}{\bar{n} + 1} &= e^{-E/(k_B T)} \\ \implies \bar{n} &= \frac{1}{e^{\hbar\omega/(k_B T)} - 1}. \end{aligned}$$

The total average energy of photons in the box is $\bar{E} = \bar{n}\hbar\omega$, thus

$$\bar{E} = \frac{\hbar\omega}{e^{\hbar\omega/(k_B T)} - 1}.$$

Counting photon states, we have that

$$\begin{aligned} dE &= \bar{n}\hbar\omega dN \\ &= \bar{n}\hbar\omega \frac{L^3\omega^2}{\pi^2 c^3} d\omega \\ \implies \frac{\partial E}{\partial \omega} &= \frac{L^3\omega^2}{\pi^2 c^3} \frac{\hbar\omega}{e^{\hbar\omega/(k_B T)} - 1}. \end{aligned}$$

Calculating the energy density as before, we obtain Planck's law:

$$\mathcal{U}(\omega) = \frac{\hbar\omega^3}{\pi^2 c^3} \frac{1}{e^{\hbar\omega/(k_B T)} - 1}.$$

Problem 2.

(a) We consider the 1D infinite square well and solve the 1D Schrödinger equation

$$\left[-\frac{\hbar}{2m} \frac{\partial^2}{\partial x^2} + V(x) \right] \psi(x, t) = i\hbar \frac{\partial}{\partial t} \psi(x, t),$$

where the potential energy function is

$$V(x) = \begin{cases} 0, & x \in [0, L] \\ \infty, & \text{otherwise.} \end{cases}$$

In effect, the form of V enforces that $\psi = 0$ outside of $[0, L]$ and the boundary conditions $\psi(0, t) = \psi(L, t) = 0$, thus we only need to solve for the wavefunction within the well. Taking the ansatz $\psi(x, t) = X(x)T(t)$, we get

$$\begin{aligned} -\frac{\hbar^2}{2m} X''(x) T(t) &= i\hbar X(x) T'(t) \\ \implies -\frac{\hbar^2}{2m} \frac{X''(x)}{X(x)} &= i\hbar \frac{T'(t)}{T(t)} := E, \end{aligned}$$

where E is a constant. Thus

$$X''(x) = -\frac{2mE}{\hbar^2} X(x),$$

$$T'(t) = -\frac{iE}{\hbar}T(t).$$

Solving the time component:

$$T(t) = C \exp\left(-\frac{iE}{\hbar}t\right).$$

Solving the spatial component:

$$X(x) = A \sin\left(\frac{\sqrt{2mE}}{\hbar}x\right) + B \cos\left(\frac{\sqrt{2mE}}{\hbar}x\right).$$

Therefore (absorbing C) into the other integration factors

$$\psi(x, t) = \exp\left(-\frac{iE}{\hbar}t\right) \left[A \sin\left(\frac{\sqrt{2mE}}{\hbar}x\right) + B \cos\left(\frac{\sqrt{2mE}}{\hbar}x\right) \right].$$

Applying $\psi(0, t) = 0$:

$$\begin{aligned} 0 &= \exp\left(-\frac{iE}{\hbar}t\right) \cdot B \\ \implies B &= 0. \end{aligned}$$

Then, applying $\psi(L, t) = 0$:

$$\begin{aligned} 0 &= A \exp\left(-\frac{iE}{\hbar}t\right) \sin\left(\frac{\sqrt{2mE}}{\hbar}L\right) \\ \implies 0 &= \sin\left(\frac{\sqrt{2mE}}{\hbar}L\right) \\ \implies \frac{\sqrt{2mE}}{\hbar}L &= n\pi, \quad n = 1, 2, 3, \dots \\ \implies E_n &= \frac{n^2\pi^2\hbar^2}{2mL^2}, \quad n = 1, 2, 3, \dots \end{aligned}$$

Thus the non-normalized solutions are of the form

$$\psi_n(x, t) = A_n \exp\left(-\frac{iE_n}{\hbar}t\right) \sin\left(\frac{n\pi}{L}x\right).$$

Normalization entails that $\|\psi_n\|_{L^2} = 1$, where $\|\cdot\|_{L^2}$ is the $L^2(\mathbb{R})$ -norm. Thus

$$\begin{aligned} 1 &= \|\psi_n\|_{L^2}^2 \\ &= \int_{\mathbb{R}} dx \psi_n^* \psi_n \\ &= \int_{[0, L]} dx \psi_n^* \psi_n \quad (\text{supp}(\psi_n) = [0, L]) \\ &= \int_0^L dx A_n^* \exp\left(\frac{iE_n}{\hbar}t\right) \sin\left(\frac{n\pi}{L}x\right) A_n \exp\left(-\frac{iE_n}{\hbar}t\right) \sin\left(\frac{n\pi}{L}x\right) \\ &= |A_n|^2 \int_0^L dx \sin^2\left(\frac{n\pi}{L}x\right) \\ &= |A_n|^2 \int_0^L dx \frac{1 - \cos\left(\frac{2n\pi}{L}x\right)}{2} \end{aligned}$$

$$= \frac{L}{2} |A_n|^2$$

$$\Rightarrow |A_n| = \sqrt{\frac{2}{L}}.$$

By linearity of the Schrödinger equation, a general solution ψ can be built via a series expansion

$$\psi = \sum_n c_n \psi_n.$$

Without loss of generality we can move any complex phase factor of A_n to the coefficient c_n and assume that A_n is real and nonnegative, hence $A_n := \sqrt{2/L}$. Thus the basis eigenfunctions and their corresponding energies are

$$\psi_n(x, t) = \sqrt{\frac{2}{L}} \exp\left(-\frac{iE_n}{\hbar}t\right) \sin\left(\frac{n\pi}{L}x\right),$$

$$E_n = \frac{n^2 \pi^2 \hbar^2}{2mL^2}, \quad n = 1, 2, 3, \dots$$

Orthogonality of the basis $\{\psi_n\}$ follows from the orthogonality of the subset $\{\sin(\pi nx/L)\}$ of the Fourier basis. Thus for the general solution to be normalized:

$$\begin{aligned} 1 &= \|\psi\|_{L^2}^2 \\ &= \langle \psi, \psi \rangle_{L^2} \\ &= \left\langle \sum_m c_m \psi_m, \sum_n c_n \psi_n \right\rangle_{L^2} \\ &= \sum_{m,n} c_m^* c_n \langle \psi_m, \psi_n \rangle \\ &= \sum_{m,n} c_m^* c_n \delta_{m,n} \\ &= \sum_m |c_m|^2 \\ &= \|c\|_{\ell^2}^2 \\ \Rightarrow \|c\|_{\ell^2} &= 1, \end{aligned}$$

where $\langle \cdot, \cdot \rangle_{L^2}$ is the L^2 -inner product, $\|\cdot\|_{\ell^2}$ is the ℓ^2 -norm, and c is the (possibly finite or infinite) coefficient sequence $c = (c_1, c_2, c_3, \dots)$.

(b) (A) As per the conditions above, we require $1 = \sqrt{A^2 + A^2}$, thus $A = 1/\sqrt{2}$, and so

$$\Psi(x, 0) = \frac{1}{\sqrt{2}} \psi_1(x) + \frac{1}{\sqrt{2}} \psi_2(x).$$

(B) The coefficients in the series expansion do not change with time, so

$$\begin{aligned} \Psi(x, t) &= \frac{1}{\sqrt{2}} \psi_1(x, t) + \frac{1}{\sqrt{2}} \psi_2(x, t) \\ &= \frac{1}{\sqrt{L}} \exp\left(-\frac{iE_1}{\hbar}t\right) \sin\left(\frac{\pi}{L}x\right) + \frac{1}{\sqrt{L}} \exp\left(-\frac{iE_2}{\hbar}t\right) \sin\left(\frac{2\pi}{L}x\right) \end{aligned}$$

and

$$|\Psi(x, t)|^2 = \frac{1}{L} \sin^2\left(\frac{\pi}{L}x\right)$$

$$\begin{aligned}
& + \frac{1}{L} \exp\left(\frac{i(E_1 - E_2)t}{\hbar}\right) \sin\left(\frac{\pi}{L}x\right) \sin\left(\frac{2\pi}{L}x\right) \\
& + \frac{1}{L} \exp\left(\frac{i(E_2 - E_1)t}{\hbar}\right) \sin\left(\frac{2\pi}{L}x\right) \sin\left(\frac{\pi}{L}x\right) \\
& + \frac{1}{L} \sin^2\left(\frac{2\pi}{L}x\right) \\
& = \frac{1}{L} \left[\sin^2\left(\frac{\pi}{L}x\right) + \sin^2\left(\frac{2\pi}{L}x\right) + 2 \cos\left(\frac{E_1 - E_2}{\hbar}t\right) \sin\left(\frac{\pi}{L}x\right) \sin\left(\frac{2\pi}{L}x\right) \right] \\
& = \frac{1}{L} \left[\sin^2\left(\frac{\pi}{L}x\right) + \sin^2\left(\frac{2\pi}{L}x\right) + 2 \cos((1^2\omega - 2^2\omega)t) \sin\left(\frac{\pi}{L}x\right) \sin\left(\frac{2\pi}{L}x\right) \right] \\
& = \frac{1}{L} \left[\sin^2\left(\frac{\pi}{L}x\right) + \sin^2\left(\frac{2\pi}{L}x\right) + 2 \cos(3\omega t) \sin\left(\frac{\pi}{L}x\right) \sin\left(\frac{2\pi}{L}x\right) \right].
\end{aligned}$$

(C) We have that

$$\begin{aligned}
\langle x \rangle &= \int_{\mathbb{R}} dx \psi^* x \psi \\
&= \frac{1}{L} \int_0^L dx x \left[\sin^2\left(\frac{\pi}{L}x\right) + \sin^2\left(\frac{2\pi}{L}x\right) + 2 \cos(3\omega t) \sin\left(\frac{\pi}{L}x\right) \sin\left(\frac{2\pi}{L}x\right) \right] \\
&= \frac{1}{L} \int_0^L dx x \left[\frac{1 - \cos\left(\frac{2\pi}{L}x\right)}{2} + \frac{1 - \cos\left(\frac{4\pi}{L}x\right)}{2} + \cos(3\omega t) \left(\cos\left(\frac{\pi}{L}x\right) - \cos\left(\frac{3\pi}{L}x\right) \right) \right] \\
&= 1 - \frac{1}{2L} \int_0^L dx x \left[\cos\left(\frac{2\pi}{L}x\right) + \cos\left(\frac{4\pi}{L}x\right) \right] \\
&\quad + \frac{\cos(3\omega t)}{L} \int_0^L dx x \left(\cos\left(\frac{\pi}{L}x\right) - \cos\left(\frac{3\pi}{L}x\right) \right) \\
&= 1 - \frac{1}{2L} \left[\frac{\cos\left(\frac{2\pi}{L}x\right)}{(2\pi/L)^2} + \frac{x \sin\left(\frac{2\pi}{L}x\right)}{(2\pi/L)} + \frac{\cos\left(\frac{4\pi}{L}x\right)}{(4\pi/L)^2} + \frac{x \sin\left(\frac{4\pi}{L}x\right)}{(4\pi/L)} \right] \Big|_0^a \\
&\quad + \frac{\cos(3\omega t)}{L} \left[\frac{\cos\left(\frac{\pi}{L}x\right)}{(\pi/L)^2} + \frac{x \sin\left(\frac{\pi}{L}x\right)}{(\pi/L)} + \frac{\cos\left(\frac{3\pi}{L}x\right)}{(3\pi/L)^2} + \frac{x \sin\left(\frac{3\pi}{L}x\right)}{(3\pi/L)} \right] \Big|_{x=0}^a \\
&= 1 - \frac{1}{2L} \left[\frac{\cos\left(\frac{2\pi a}{L}\right) - 1}{(2\pi/L)^2} + \frac{a \sin\left(\frac{2\pi a}{L}\right)}{(2\pi/L)} + \frac{\cos\left(\frac{4\pi a}{L}\right) - 1}{(4\pi/L)^2} + \frac{a \sin\left(\frac{4\pi a}{L}\right)}{(4\pi/L)} \right] \\
&\quad + \frac{\cos(3\omega t)}{L} \left[\frac{\cos\left(\frac{\pi a}{L}\right) - 1}{(\pi/L)^2} + \frac{a \sin\left(\frac{\pi a}{L}\right)}{(\pi/L)} + \frac{\cos\left(\frac{3\pi a}{L}\right) - 1}{(3\pi/L)^2} + \frac{a \sin\left(\frac{3\pi a}{L}\right)}{(3\pi/L)} \right]
\end{aligned}$$

The frequency of oscillation is 3ω . The amplitude of oscillation is

$$\frac{1}{L} \left[\frac{\cos\left(\frac{\pi a}{L}\right) - 1}{(\pi/L)^2} + \frac{a \sin\left(\frac{\pi a}{L}\right)}{(\pi/L)} + \frac{\cos\left(\frac{3\pi a}{L}\right) - 1}{(3\pi/L)^2} + \frac{a \sin\left(\frac{3\pi a}{L}\right)}{(3\pi/L)} \right] \quad (???)$$

(D) The expected value of momentum can be computed as

$$\begin{aligned}
\langle p \rangle &= \frac{d}{dt} \langle x \rangle \\
&= \frac{-3\omega \sin(3\omega t)}{L} \left[\frac{\cos\left(\frac{\pi a}{L}\right) - 1}{(\pi/L)^2} + \frac{a \sin\left(\frac{\pi a}{L}\right)}{(\pi/L)} + \frac{\cos\left(\frac{3\pi a}{L}\right) - 1}{(3\pi/L)^2} + \frac{a \sin\left(\frac{3\pi a}{L}\right)}{(3\pi/L)} \right].
\end{aligned}$$

(E) You might get E_1 or E_2 . The expectation value is

$$\langle H \rangle = \left(\frac{1}{\sqrt{2}} \right)^2 E_1 + \left(\frac{1}{\sqrt{2}} \right)^2 E_2$$

$$= \frac{E_1 + E_2}{2},$$

which is their mean.

(c) The probability of finding the particle in $[0, L/a]$ is

$$\begin{aligned} P_n([0, L/a]) &= \int_0^a dx \psi_n^* \psi_n \\ &= \frac{2}{L} \int_0^a \sin^2\left(\frac{n\pi}{L}x\right) \\ &= \frac{2}{L} \int_0^a \frac{1 - \cos\left(\frac{2n\pi}{L}x\right)}{2} \\ &= \frac{1}{L} \left[x - \frac{L}{2n\pi} \sin\left(\frac{2n\pi}{L}x\right) \right] \Big|_0^a \\ &= \frac{1}{L} \left[a - \frac{L}{2n\pi} \sin\left(\frac{2n\pi a}{L}\right) \right]. \end{aligned}$$

(d) As $n \rightarrow \infty$, the second term vanishes, leaving $P_n = a/L$. If the particle were a classical one, bouncing around randomly in the square well, one would also expect it to be in $[0, a]$ around a proportion a/L of the time. This suggests that quantum mechanical objects begin to behave classically when highly excited.