

# General Relativity notes

Jacopo Tissino, Giorgio Mentasti

October 8, 2019

3 October 2019

Marco Peloso, [marco.peloso@pd.infn.it](mailto:marco.peloso@pd.infn.it)

## 1 Special relativity

**Definition 1.1.** *An inertial frame is one in which Newton's laws hold: a free body moves with acceleration  $a^i = 0$ .*

Newton's first law establishes the *existence* of inertial frames.

**Proposition 1.1.** *The frames  $O$  and  $O'$  are both inertial frames iff  $O'$  moves with constant velocity wrt  $O$ .*

**Proposition 1.2.** *Coordinate transformations between inertial frames are Lorentz boosts, which in some coordinate frame can be written as*

$$t' = \gamma_v \left( t - \frac{vx}{c^2} \right) \quad (1a)$$

$$x' = \gamma_v (x - vt) \quad (1b)$$

$$y' = y \quad (1c)$$

$$z' = z, \quad (1d)$$

where  $\gamma_v = 1/\sqrt{1 - v^2/c^2}$ .

If  $v \ll c$ , so  $v/c \sim 0$ , they simplify to the identity for  $t, y, z$  and  $x' = x - vt$ : these are Galilean transformations.

If we have two events,  $x^\mu$  and  $y^\mu$ , they occur with some time and space separation  $\Delta x^\mu = x^\mu - y^\mu$ . We can compute  $\Delta s^2 = \eta_{\mu\nu} \Delta x^\mu \Delta x^\nu$ , where

$$\eta_{\mu\nu} = \text{diag}(-c^2, 1, 1, 1). \quad (2)$$

**Proposition 1.3.** Under Lorentz transformations  $\Delta s^2$  is invariant.

We can classify separations between events as

- time-like when  $\Delta s^2 < 0$ ;
- null-like when  $\Delta s^2 = 0$ ;
- space-like when  $\Delta s^2 > 0$ .

We can draw spacetime diagrams. A light cone is the set of points which are null-like separated from a select point. Things can be only causally related to events inside the light-cone (with  $\Delta s^2 \geq 0$ ).

## 1.1 Time dilation

Take two events which occur at the same location for  $O'$ . In the primed frame they will have coordinates  $x^\mu = (t_0, x_0)$  and  $y^\mu = (t_1, x_0)$ .

**Definition 1.2.** The proper time between these two events is  $t_1 - t_0 \stackrel{\text{def}}{=} \Delta\tau$ .

We now see that  $\Delta s'^2 = -c^2 \Delta\tau^2$ . Then, any other observer will see the same  $\Delta s^2 = -c^2 \Delta t^2 + \Delta x^2 = \Delta s'^2$ .

This directly implies that  $\Delta\tau \leq \Delta t$  for any observer, since  $\Delta\tau^2 = \Delta t^2 - \Delta x^2/c^2$ . This effect is called *time dilation*.

By how much exactly is time dilated? Of course  $\Delta x = v\Delta t$ , therefore  $\Delta t = \gamma_v \Delta\tau$ .  
-> Muon problem.

Inverse Lorentz transformation have the same expression, but with  $v \rightarrow -v$ . This can be proved both mathematically by solving the equations and phisically by reasoning about their meaning. There is no preferential inertial frame.

A Lorentz transformation can be written in matrix form in the  $(ct, x)$  plane as:

$$\Lambda = \begin{bmatrix} \gamma & -\gamma\beta \\ -\gamma\beta & \gamma \end{bmatrix} = \begin{bmatrix} \cosh \theta & -\sinh \theta \\ -\sinh \theta & \cosh \theta \end{bmatrix} \quad (3)$$

since there is an angle  $\theta$  such that  $\gamma = \cosh \theta$  and  $\gamma\beta = \sinh \theta$ : the angle  $\theta$  will be  $\theta = \tanh^{-1}(v/c)$ . This is true because  $\gamma^2 - \beta^2\gamma^2 = 1$ .

After a boost the  $ct'$  and  $x'$  axes are respectively the lines  $ct = x/\beta$  and  $ct = \beta x$ .

## 4 October 2019

Last lecture we saw the fact that the  $ct'$  and  $x'$  axes are rotated by equal angles from the  $ct$  and  $x$  axes towards the  $ct = x$  axis.

## 1.2 Relativity of simultaneity

Consider two events which are simultaneous in the  $O'$  frame. Their times in this frame are  $t'_A = t'_B$ .

In the  $O$  frame, instead, we have

$$ct_{A,B} = \frac{v}{c}x_{A,B} + \underbrace{\sqrt{1 - \frac{v^2}{c^2}}}_{\text{a constant}} ct'_{A,B}, \quad (4)$$

so the events are not simultaneous in the  $O$  frame.

## 1.3 Length contraction

If in the  $O$  frame,  $A$  occurs at  $t, x = 0$  while  $B$  occurs at  $t = 0, x = L$ , then  $L$  is the measured length of their spatial interval by  $O$ . We assume that this is the frame in which the object is moving, and we transform into a frame in which it is stationary:  $O'$ .

In the primed frame their coordinates will be:

$$x'_A = \gamma_v \left( x_A - \frac{v}{c} ct_A \right) \quad (5a)$$

$$x'_B = \gamma_v \left( x_B - \frac{v}{c} ct_B \right), \quad (5b)$$

therefore  $x'_B - x'_A = \gamma_v(x_B - x_A)$ : the length is contracted in the  $O$  frame, since  $\gamma \geq 1$ .

## 1.4 Addition of velocities

Two observers see an object moving with  $v' = dx'/dt'$  and  $v = dx/dt$  respectively. Their relative velocity is  $u$ . Differentiating we get:

$$v' = \frac{\gamma(dx - v dt)}{\gamma\left(dt - \frac{u dx}{c^2}\right)} = \frac{v - u}{1 - \frac{uv}{c^2}}. \quad (6)$$

Two interesting limits of this formula are:  $v' = v - u$  if  $u \ll c$  or  $v \ll c$ ; and  $v' = c$  if  $v = c$  for whatever  $u$ .

## 1.5 Tensor notation

The position four-vector is  $x^\mu = (ct, x, y, z)$ . The Euclidean scalar product is given by  $x \cdot y = \delta_{\mu\nu} x^\mu x^\nu$ . If we substitute the identity  $\delta_{\mu\nu}$  with another metric we can find a more general metric space.

The Minkowski metric is  $\eta_{\mu\nu} = \text{diag}(-1, 1, 1, 1)$ . The separation 4-vector is  $dx^\mu = (c dt, dx, dy, dz)$ .

Using Einstein summation notation, we can write the spacetime interval as  $ds^2 = \eta_{\mu\nu} dx^\mu dx^\nu$ .

Specifically for the Minkowski metric we have the relation  $\eta_{\mu\nu} = \eta^{\mu\nu}$ : it is its own inverse. For a general metric  $g_{\mu\nu}$  this will not hold.

How do we express the Lorentz boosts? They preserve  $ds^2$ , therefore they look like  $x'^\mu = \Lambda^\mu_\nu x^\nu$ , with the  $(1, 1)$  tensors  $\Lambda_\mu^\nu$  satisfying  $\Lambda_\mu^\nu \Lambda_\rho^\sigma \eta_{\nu\sigma} = \eta_{\mu\rho}$ . This is called the *pseudo-orthogonality* relation.

The metric allows us to raise and lower indices. Raising an index in the pseudo-orthogonality relation gives us:  $\Lambda^\mu_\alpha \eta_{\mu\nu} \Lambda^\nu_\beta \eta^{\beta\sigma} = \delta_\alpha^\sigma$ , therefore  $\eta_{\mu\nu} \Lambda^\nu_\beta \eta^{\beta\sigma}$  is the inverse of a Lorentz transformation.

Four-vectors can also have their indices down, and they will transform according to the inverse of Lorentz transformations:

$$(\eta_{\alpha\mu} x^\mu)' = \eta_{\alpha\mu} \Lambda^\mu_\nu x^\nu \quad (7a)$$

$$= \Lambda_{\alpha\sigma} \delta^\sigma_\nu x^\nu \quad (7b)$$

$$= \Lambda_{\alpha\sigma} \eta^{\sigma\beta} \eta_{\beta\nu} x^\nu \quad (7c)$$

$$= \Lambda_\alpha^\beta x_\beta. \quad (7d)$$

We will write our laws as tensorial equations, which are covariant.

By pseudo-orthogonality, the scalar product  $A_\mu B^\mu$  is a covariant (that is, invariant) scalar. Of course it is equal to  $A^\mu B_\mu$ .

**Definition 1.3 (Tensor).** A  $(p, q)$  tensor is an object  $M_{\mu_1 \dots \mu_p}^{\nu_1 \dots \nu_q}$  with many components indexed by several indices, which transforms as:

$$M_{\mu_1 \dots \mu_p}^{\nu_1 \dots \nu_q} \rightarrow \Lambda_{\mu_1}^{\mu'_1} \dots \Lambda_{\mu_p}^{\mu'_p} \Lambda^{\nu_1}_{\nu'_1} \dots \Lambda^{\nu_q}_{\nu'_q} M_{\mu'_1 \dots \mu'_p}^{\nu'_1 \dots \nu'_q} \quad (8)$$

under Lorentz transformations  $\Lambda_\mu^\nu$ .