# AstroStatistics and Cosmology Homework

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## **Contents**

#### 1 November exercises

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Exercises 1–3 and  $7^1$  are in Jupyter notebooks in the folder astrostat\_homework. They can be most easily accessed through the following links:

- https://nbviewer.jupyter.org/github/jacopok/notes/blob/master/ap\_third\_semester/astrostat\_homework/exercises\_123.ipynb
- 2. https://nbviewer.jupyter.org/github/jacopok/notes/blob/master/ap\_third\_semester/astrostat\_homework/exercise\_7.ipynb.

### 1 November exercises

#### Exercise 4

After being given a probability distribution  $\mathbb{P}(x)$ , we define the *characteristic function*  $\phi$  as its Fourier transform, which can also be expressed as the expectation value of  $\exp(-i\vec{k}\cdot\vec{x})$ :

$$\phi(\vec{k}) = \int d^n x \exp\left(-i\vec{k} \cdot \vec{x}\right) \mathbb{P}(x) = \mathbb{E}\left[\exp\left(-i\vec{k} \cdot \vec{x}\right)\right]. \tag{1.1}$$

Claim 1.1. A multivariate normal distribution

$$\mathcal{N}(\vec{x}|\vec{\mu},C) = \frac{1}{(2\pi)^{n/2}\sqrt{\det C}} \exp\left(-\frac{1}{2}\vec{y}^{\top}C^{-1}\vec{y}\right)\Big|_{\vec{y}=\vec{x}-\vec{\mu}},$$
(1.2)

has a characteristic function equal to

$$\phi(\vec{k}) = \exp\left(-i\vec{\mu}\cdot\vec{k} - \frac{1}{2}\vec{k}^{\top}C\vec{k}\right). \tag{1.3}$$

<sup>&</sup>lt;sup>1</sup> It is not finished yet.

*Proof: completing the square.* The integral we need to compute is given, absorbing the normalization into a factor *N*, by

$$\phi(\vec{k}) = N \int d^n x \, \exp\left(-i\vec{k} \cdot \vec{x} - \frac{1}{2} \vec{y}^\top C^{-1} \vec{y}\right) \bigg|_{\vec{y} = \vec{x} - \vec{u}} \,. \tag{1.4}$$

The only integrals we really know how to do are Gaussian ones, so we want to rewrite the argument of the exponential so that it is a quadratic form. The manipulation goes as follows, considering the opposite of the argument the exponential in order to have less minus signs and defining the symmetric matrix  $V = C^{-1}$ :

$$i\vec{k} \cdot \vec{x} + \frac{1}{2}\vec{y}^{\top}V\vec{y} = i\vec{k} \cdot \vec{x} + \frac{1}{2}\vec{x}^{\top}V\vec{x} - \vec{x}^{\top}V\vec{\mu} + \frac{1}{2}\vec{\mu}^{\top}V\vec{\mu}$$

$$= \frac{1}{2}\vec{x}^{\top}V\vec{x} + \vec{x}^{\top}\left(i\vec{k} - V\vec{\mu}\right) + \frac{1}{2}\vec{\mu}^{\top}V\vec{\mu}$$

$$= \underbrace{\frac{1}{2}\left(\vec{x} + V^{-1}(i\vec{k} - V\vec{\mu})\right)^{\top}V\left(\vec{x} + V^{-1}(i\vec{k} - V\vec{\mu})\right)}_{(1)} + \underbrace{\frac{1}{2}\left(i\vec{k} - V\vec{\mu}\right)^{\top}V^{-1}\left(i\vec{k} - V\vec{\mu}\right) + \frac{1}{2}\vec{\mu}^{\top}V\vec{\mu}}_{(2)},$$

$$(1.5)$$

which we can now integrate, since it is now a quadratic form in terms of a shifted variable,  $\vec{x} + \vec{p}$ , where the constant (with respect to  $\vec{x}$ ) vector  $\vec{p}$  is given by  $V^{-1}(i\vec{k} - V\vec{\mu})$ .<sup>2</sup>

Now, shifting the integral from one in  $d^n x$  to one in  $d^n (x + p)$  does not change the measure, since the Jacobian of a shift is the identity. Then, we have

$$\phi(\vec{k}) = N \int d^n(x+p) \exp\left(-(1)-(2)\right)$$
(1.12)

$$= N\sqrt{\frac{(2\pi)^n}{\det V}}\exp\left(-2\right) \tag{1.13}$$

$$= \underbrace{\frac{1}{\sqrt{\det V \det C}}}_{=1} \exp\left(-2\right), \tag{1.14}$$

since the determinant of the inverse is the inverse of the determinant.

$$\frac{1}{2} \left( \vec{x} + A^{-1} \vec{b} \right)^{\top} A \left( \vec{x} + A^{-1} \vec{b} \right) - \frac{1}{2} \vec{b}^{\top} A^{-1} \vec{b} = \tag{1.8}$$

$$= \frac{1}{2} \left[ \vec{x}^{\top} A \vec{x} + \vec{x}^{\top} A A^{-1} \vec{b} + \left( A^{-1} \vec{b} \right)^{\top} A \vec{x} + \left( A^{-1} \vec{b} \right)^{\top} A A^{-1} \vec{b} - \vec{b}^{\top} A^{-1} \vec{b} \right]$$
(1.9)

$$= \frac{1}{2} \left[ \vec{x}^{\top} A \vec{x} + \vec{x}^{\top} \vec{b} + \vec{b}^{\top} (A^{-1})^{\top} A \vec{x} + \vec{b}^{\top} (A^{-1})^{\top} \vec{b} - \vec{b}^{\top} A^{-1} \vec{b} \right]$$
(1.10)

$$= \frac{1}{2}\vec{x}^{\top}A\vec{x} + \vec{b}^{\top}\vec{x}, \tag{1.11}$$

which we used with  $\vec{b} = i\vec{k} - V\vec{\mu}$ .

In the last step we applied the matrix square completion formula: for a symmetric matrix A and vectors  $\vec{x}$ ,  $\vec{b}$  we have

Now, we only need to simplify (2):

$$=\frac{1}{2}\vec{k}^{\top}C\vec{k}+i\vec{\mu}^{\top}\vec{k}\,,\tag{1.16}$$

inserting which into the exponent yields the desired result.

*Proof: by diagonalization.* We now follow a different approach: the covariance matrix C is symmetric, so we will always be able to find an orthogonal matrix O (satisfying  $O^{\top} = O^{-1}$ ) such that  $C = O^{\top}DO$ , where D is diagonal. We will then also have  $V = C^{-1} = O^{\top}D^{-1}O$ . Let us denote the eigenvalues of D as  $\lambda_i$ , and the eigenvalues of  $D^{-1}$  as  $d_i = \lambda_i^{-1}$ .

Defining  $\vec{z} = O\vec{x}$ ,  $\vec{m} = O\vec{\mu}$ ,  $\vec{u} = O\vec{k}$  the negative of the argument of the integral becomes:

$$i\vec{k} \cdot \vec{x} + \frac{1}{2} (\vec{x} - \vec{\mu})^{\top} C^{-1} (\vec{x} - \vec{\mu}) = i\vec{u} \cdot \vec{z} + \frac{1}{2} (\vec{z} - \vec{m})^{\top} D^{-1} (\vec{z} - \vec{m})$$
(1.17)

$$= i\vec{u} \cdot \vec{z} + \frac{1}{2} \sum_{i} d_{i} (z_{i} - m_{i})^{2}$$
 (1.18)

$$= \sum_{i} \left[ iu_{i}z_{i} + \frac{d_{i}}{2} \left( z_{i}^{2} + m_{i}^{2} - 2m_{i}z_{i} \right) \right]$$
 (1.19)

$$= \sum_{i} \left[ z_i^2 \frac{d_i}{2} + z_i (iu_i - m_i d_i) + \frac{d_i}{2} m_i^2 \right]. \tag{1.20}$$

With this, and since by  $\det O = 1$  we have  $d^n z = d^n x$ , we can decompose our Gaussian integral into a product of Gaussian integrals:

$$\phi(\vec{k}) = N \int d^n x \exp\left(-i\vec{k} \cdot \vec{x} - \frac{1}{2}(\vec{x} - \vec{\mu})^{\top} C^{-1}(\vec{x} - \vec{\mu})\right)$$
(1.21)

$$= N \int d^{n}z \exp\left(-\sum_{i} \left[z_{i}^{2} \frac{d_{i}}{2} + z_{i}(iu_{i} - m_{i}d_{i}) + \frac{d_{i}}{2}m_{i}^{2}\right]\right)$$
(1.22)

$$= N \prod_{i} \int dz_{i} \exp\left(-z_{i}^{2} \frac{d_{i}}{2} - z_{i} (iu_{i} - m_{i}d_{i}) - \frac{d_{i}}{2} m_{i}^{2}\right)$$
(1.23)

$$= N \prod_{i} \sqrt{\frac{2\pi}{d_{i}}} \exp\left(\frac{(iu_{i} - m_{i}d_{i})^{2}}{2d_{i}} - \frac{d_{i}m_{i}^{2}}{2}\right)$$
(1.24)

$$= \frac{1}{\sqrt{\det C \det V}} \prod_{i} \exp\left(\frac{-u_i^2 + m_i^2 d_i^2 - 2iu_i m_i d_i}{2d_i} - \frac{d_i m_i^2}{2}\right)$$
(1.25)

$$= \exp\left(\sum_{i} \left[ -\frac{u_i^2}{2d_i} - iu_i m_i \right] \right) \tag{1.26}$$

$$= \exp\left(-\frac{1}{2}\vec{u}^{\top}C\vec{u} - i\vec{u}\cdot\vec{m}\right) \tag{1.27}$$

$$= \exp\left(-\frac{1}{2}\vec{k}^{\top}C\vec{k} - i\vec{k}\cdot\vec{\mu}\right),\tag{1.28}$$

where we have used the expression for the single-variable Gaussian integral:

$$\int dz \exp\left(-az^2 + bz + c\right) = \sqrt{\frac{\pi}{a}} \exp\left(\frac{b^2}{4a} + c\right), \tag{1.29}$$

which comes from the one-variable completion of the square:

$$-az^{2} + bz + c = -a\left(z - \frac{b}{2a}\right)^{2} + \frac{b^{2}}{4a} + c.$$
 (1.30)

Also, we used the fact that orthogonal transformation do not change fully-contracted objects, such as scalar products or bilinear forms.

#### Exercise 5

We can calculate the moments of a distribution through its characteristic function:

$$\mathbb{E}\left[x_{\alpha}^{n_{\alpha}} \dots x_{\beta}^{n_{\beta}}\right] = \left. \frac{\partial^{n_{\alpha} \dots n_{\beta}} \phi(\vec{k})}{\partial (-ik_{\alpha})^{n_{\alpha}} \dots \partial (-ik_{\beta})^{n_{\beta}}} \right|_{\vec{k}=0}.$$
 (1.31)

In the multivariate Gaussian case we can then calculate the mean (component by component) as

$$\mathbb{E}(x_{\alpha}) = \left. \frac{\partial \phi(\vec{k})}{\partial (-ik_{\alpha})} \right|_{\vec{k}=0} \tag{1.32}$$

$$= \frac{\partial}{\partial (-ik_{\alpha})} \bigg|_{\vec{k}=0} \exp\left(-\frac{1}{2}\vec{k}^{\top}C\vec{k} - i\vec{k} \cdot \vec{\mu}\right)$$
(1.33)

$$= \left[ -i \sum_{\beta} k_{\beta} C_{\beta\alpha} + \mu_{\alpha} \right] \exp \left( -\frac{1}{2} \vec{k}^{\top} C \vec{k} - i \vec{k} \cdot \vec{\mu} \right) \bigg|_{\vec{k} = 0}$$
(1.34)

$$=\mu_{\alpha}\,,\tag{1.35}$$

where we used the fact that the differentiation of a symmetric bilinear form is as follows:

$$\frac{\partial}{\partial k_{\alpha}} \left( \sum_{\beta \gamma} k_{\beta} k_{\gamma} C_{\beta \gamma} \right) = 2 \sum_{\beta \gamma} \delta_{\beta \alpha} k_{\gamma} C_{\beta \gamma} = 2 \sum_{\gamma} k_{\gamma} C_{\alpha \gamma}. \tag{1.36}$$

The covariance matrix can be computed by linearity as

$$\widetilde{C}_{\alpha\beta} = \mathbb{E}\left[\left(x_{\alpha} - \mathbb{E}(x_{\alpha})\right)\left(x_{\beta} - \mathbb{E}(x_{\beta})\right)\right] = \mathbb{E}\left[x_{\alpha}x_{\beta}\right] - \mu_{\alpha}\mu_{\beta}, \tag{1.37}$$

the first term of which reads as follows:

$$\mathbb{E}[x_{\alpha}x_{\beta}] = \left. \frac{\partial^2 \phi(\vec{k})}{\partial (-ik_{\beta})\partial (-ik_{\alpha})} \right|_{\vec{k}=0}$$
(1.38)

$$= \frac{\partial}{\partial (-ik_{\beta})} \bigg|_{\vec{k}=0} \left[ -i\sum_{\beta} k_{\beta} C_{\beta\alpha} + \mu_{\alpha} \right] \exp\left( -\frac{1}{2} \vec{k}^{\top} C \vec{k} - i \vec{k} \cdot \vec{\mu} \right)$$
(1.39)

$$=C_{\alpha\beta}+\mu_{\alpha}\mu_{\beta}\,,\tag{1.40}$$

therefore, as expected,  $\widetilde{C}_{\alpha\beta}$  is indeed  $C_{\alpha\beta}$ .

#### Exercise 6

**Claim 1.2.** The characteristic function of a multivariate Gaussian is, up to normalization, a multivariate Gaussian.

*Proof.* The characteristic function is the exponential of (minus)

$$\frac{1}{2}\vec{k}^{\top}C\vec{k} + i\vec{k} \cdot \vec{\mu} = \frac{1}{2} \left( \vec{k} + iC^{-1}\vec{\mu} \right)^{\top} C \left( \vec{k} + iC^{-1}\vec{\mu} \right) + \frac{1}{2}\vec{\mu}^{\top}C^{-1}\vec{\mu} , \tag{1.41}$$

which means that the characteristic function is in the form

$$\phi(\vec{k}) = \text{const} \times \exp\left(-\frac{1}{2}(\vec{k} - \vec{m})^{\top}C(\vec{k} - \vec{m})\right), \tag{1.42}$$

a multivariate normal with mean  $\vec{m} = -iC^{-1}\vec{\mu}$  and covariance matrix  $C^{-1}$ , the inverse of the covariance matrix of the corresponding MVN.

#### Exercise 8

For clarity, we denote with Greek indices those ranging from 1 to N, the size of the vector of data; and with Latin indices those ranging from 1 to M, the number of templates.

We are assuming that the data have a Gaussian distribution with a covariance matrix C, and we are modelling their mean  $\mu_{\alpha}$  as a sum of templates  $t_{i\alpha}$  with coefficients  $A_i$ :

$$\mu_{\alpha} = t_{i\alpha} A_i \,, \tag{1.43}$$

where the Einstein summation convention has been used. Therefore, the likelihood is proportional to

$$\mathscr{L}(d_{\alpha}|A_{i}) \propto \exp\left(-\frac{1}{2}(d_{\alpha} - A_{i}t_{i\alpha})C_{\alpha\beta}^{-1}(d_{\beta} - A_{j}t_{j\beta})\right). \tag{1.44}$$

The normalization only depends on the covariance matrix  $C_{\alpha\beta}$ , which we assume is fixed. Therefore, maximizing the likelihood<sup>3</sup> is equivalent to minimizing the  $\chi^2$ , which reads

$$\chi^2 = (d_{\alpha} - A_i t_{i\alpha}) C_{\alpha\beta}^{-1} \left( d_{\beta} - A_j t_{j\beta} \right). \tag{1.45}$$

<sup>&</sup>lt;sup>3</sup> Which is equivalent to maximizing the posterior if we are using a flat prior.

We want to maximize this as the amplitudes vary: therefore, we set the derivative with respect to  $A_k$  to zero,

$$\frac{\partial \chi^2}{\partial A_k} = -2t_{k\alpha} C_{\alpha\beta}^{-1} \left( d_{\beta} - A_j t_{j\beta} \right) = 0, \qquad (1.46)$$

which means that

$$t_{k\alpha}C_{\alpha\beta}^{-1}d_{\beta} = (t_{k\alpha}C_{\alpha\beta}^{-1}t_{j\beta})A_{j}, \qquad (1.47)$$

a linear system of M equations (indexed by k) in the M variables  $A_j$ . If we denote the evaluations of bilinear forms in the data (N-dimensional) space with brackets, as  $a_{\alpha}C_{\alpha\beta}b_{\beta}\stackrel{\text{def}}{=}(a|C|b)$ , this reads

$$(t|C^{-1}|d)_k = (t|C^{-1}|t)_{kj}A_j (1.48)$$

$$\left[ (t|C^{-1}|t)^{-1} \right]_{mk} (t|C^{-1}|d)_k = \underbrace{\left[ (t|C^{-1}|t)^{-1} \right]_{mk} (t|C^{-1}|t)_{kj}}_{=\delta_{mi}} A_j = A_m \tag{1.49}$$

$$A_m = \left[ (t|C^{-1}|t)^{-1} \right]_{mk} (t|C^{-1}|d)_k, \qquad (1.50)$$

where the inverse of  $(t|C^{-1}|t)$  is to be computed in the *M*-dimensional vector space.

### Exercise 9

Our model for the mean value is in the form  $\mu(\Theta, A) = A\overline{x}(\Theta)$ , where  $\overline{x}$  is a generic function of  $\Theta$ , while A is our scale parameter.<sup>4</sup> Our likelihood then reads

$$\mathscr{L}(x|\Theta,A) = \underbrace{\frac{1}{(2\pi)^{N/2}\sqrt{\det C}}}_{B_1} \exp\left(-\frac{1}{2}(x-A\overline{x}(\Theta))^{\top}C^{-1}(x-A\overline{x}(\Theta))\right). \tag{1.51}$$

If the priors for both A and  $\Theta$  are flat, this corresponds to the joint posterior  $P(\Theta, A|x)$ . We want to marginalize over A, which amounts to integrating over it: dropping the dependence on  $\Theta$  of  $\overline{x}$  and defining  $V = C^{-1}$  we find

$$P(\Theta|x) = B_1 \int \exp\left(-\frac{1}{2}(x - A\overline{x})^{\top}V(x - A\overline{x})\right) dA$$
 (1.52)

$$= B_1 \int \exp\left(-\frac{1}{2}\left(x^\top V x - 2A\overline{x}^\top V x + A^2\overline{x}^\top V \overline{x}\right)\right) dA .$$
 Used the symmetry of  $V$ .

The amplitude being negative makes little sense in a typical physical context, however the Gaussian integral can be done analytically only over the whole of  $\mathbb{R}$ .

In order to get analytical results, here we will marginalize by integrating over negative amplitudes as well  $(A \in \mathbb{R})$ ; the last figure 1 will show how only integrating over positive

<sup>&</sup>lt;sup>4</sup> This is not specified in the problem, but it seems natural to think that  $|\overline{x}(\Theta)|$  is a constant for varying  $\Theta$ .

amplitudes only would have looked (by numerical calculation) in a simple case. In general if one wishes to perform the integral over  $A \in (0, +\infty)$  the tabulated values of the error function may be used.

Applying the formula for the single-variable Gaussian integral (1.29) (the bilinear forms are all evaluated to yield scalars, we are only integrating over the scalar A!) we then get

$$P(\Theta|x) = \underbrace{B_1 \exp\left(-\frac{1}{2}x^\top V x\right)}_{B_1} \exp\left(\frac{(\overline{x}^\top V x)^2}{(\overline{x}^\top V \overline{x})}\right) \sqrt{\frac{\pi}{\overline{x}^\top V \overline{x}}}$$
(1.54)

$$= B_2 \sqrt{\frac{\pi}{\overline{x}^{\top} V \overline{x}}} \exp\left(\frac{\overline{x}^{\top} \Omega \overline{x}}{\overline{x}^{\top} V \overline{x}}\right), \tag{1.55}$$

where we defined the bilinear form  $\Omega = Vxx^{\top}V^{\top}.^{5}$ 

Let us consider a simple example of this as a sanity check: suppose that x is two-dimensional, and  $\overline{x}(\Theta) = (\cos \Theta, \sin \Theta)^{\top}$ ; further, suppose that V is diagonal, so that

$$V = \begin{bmatrix} \sigma_x^{-2} & 0\\ 0 & \sigma_y^{-2} \end{bmatrix}. \tag{1.56}$$

Also, suppose that the observed data parameter is

$$x = A_x \begin{bmatrix} \cos \varphi \\ \sin \varphi \end{bmatrix}. \tag{1.57}$$

Then, the multiplicative constant in front of the marginalized posterior reads

$$B_2 = B_1 \exp\left(-\frac{1}{2}A_x^2 \left(\frac{\cos^2 \varphi}{\sigma_x^2} + \frac{\sin^2 \varphi}{\sigma_y^2}\right)\right); \tag{1.58}$$

while the bilinear form  $\Omega$  is

$$\Omega = A_x^2 \begin{bmatrix} \sigma_x^{-2} & 0 \\ 0 & \sigma_y^{-2} \end{bmatrix} \begin{bmatrix} \cos^2 \varphi & \cos \varphi \sin \varphi \\ \cos \varphi \sin \varphi & \sin^2 \varphi \end{bmatrix} \begin{bmatrix} \sigma_x^{-2} & 0 \\ 0 & \sigma_y^{-2} \end{bmatrix}$$
(1.59)

$$= A_x^2 \begin{bmatrix} \cos^2 \varphi / \sigma_x^4 & \cos \varphi \sin \varphi / \sigma_x^2 \sigma_y^2 \\ \cos \varphi \sin \varphi / \sigma_x^2 \sigma_y^2 & \sin^2 \varphi / \sigma_y^4 \end{bmatrix}. \tag{1.60}$$

Then, when we evaluate the marginalized posterior we will find something in the form

$$P(\Theta|x) = B_1 \sqrt{\pi} \left( \frac{\cos^2 \Theta}{\sigma_x^2} + \frac{\sin^2 \Theta}{\sigma_y^2} \right)^{-1/2} \exp\left( A_x^2 F(\Theta, \varphi) \right), \tag{1.61}$$

<sup>&</sup>lt;sup>5</sup> With explicit indices,  $\Omega_{im} = V_{ij}x_jx_kV_{km}$ .

where  $F(\Theta, \varphi)$  is some function whose specific form does not really matter;<sup>6</sup> the point is that the amplitude of the observed data vector,  $A_x$ , appears only as a multiplicative prefactor: its exact value will be taken care of by the evidence, and it cannot affect the shape of the distribution. Therefore, we see that by marginalizing over A we have "forgotten" any scaling information about x.

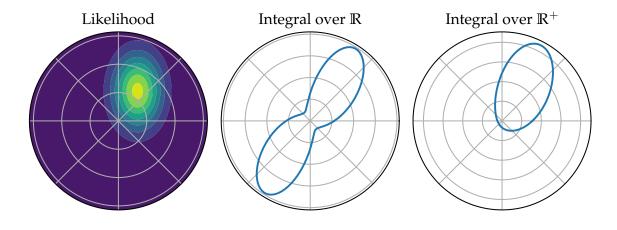


Figure 1: Marginalization: the left plot shows the full likelihood in terms of A and  $\Theta$ ; the middle plot shows the result of marginalization as shown in the previous calculation (the posterior as a function of  $\Theta$ ); the right plot shows the result of the more physically meaningful marginalization over  $A \in (0, +\infty)$  only. Here the likelihood is a diagonal Gaussian with  $\sigma_x = 1.2$  and  $\sigma_y = 1.8$ , centered in  $A_x = 2.5$  and  $\varphi = 1$  rad.

$$F(\Theta, \varphi) = -\frac{1}{2} \left( \frac{\cos^2 \varphi}{\sigma_x^2} + \frac{\sin^2 \varphi}{\sigma_y^2} \right) + \left( \frac{\cos^2 \Theta}{\sigma_x^2} + \frac{\sin^2 \Theta}{\sigma_y^2} \right)^{-1} \left[ \frac{\cos^2 \Theta \cos^2 \varphi}{\sigma_x^4} + 2 \frac{\cos \Theta \sin \Theta \cos \varphi \sin \varphi}{\sigma_x^2 \sigma_y^2} + \frac{\sin^2 \Theta \sin^2 \varphi}{\sigma_y^4} \right].$$

$$(1.62)$$

<sup>&</sup>lt;sup>6</sup> For completeness, here is the full expression: