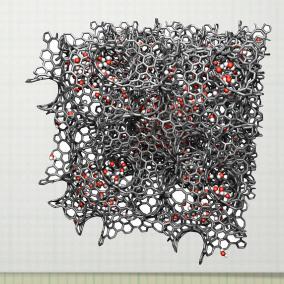
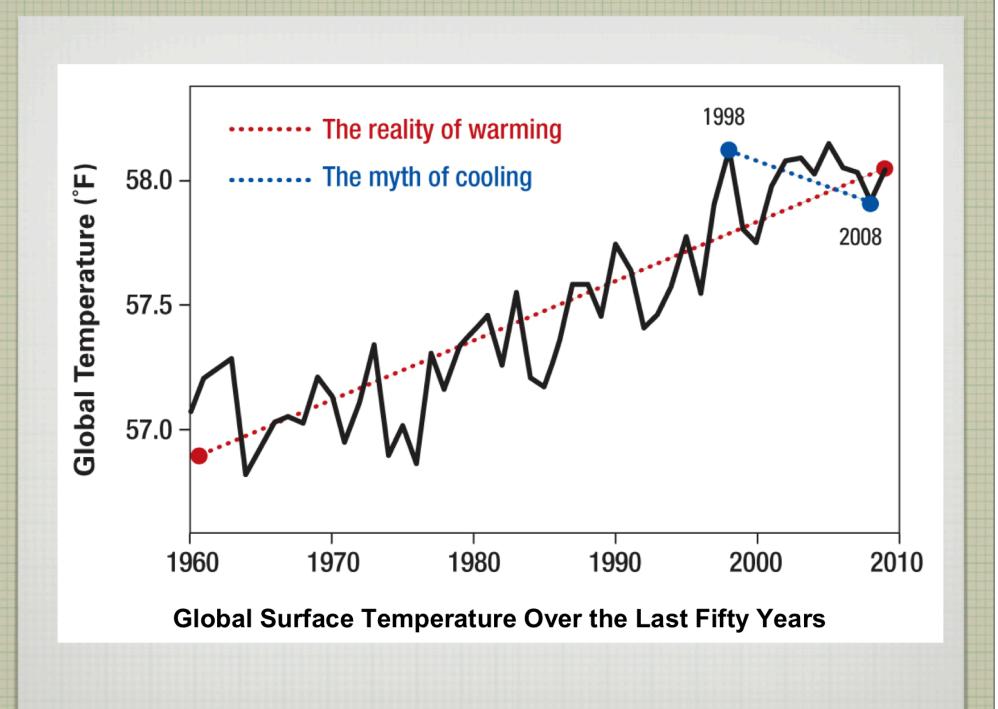
PHY-4810 COMPUTATIONAL PHYSICS

LECTURE 7: DATA FITTING





HTTP://DATA.GISS.NASA.GOV/GISTEMP/TABLEDATA/GLB.TS+DSST.TXT

GLOBAL Land-Ocean Temperature Index in 0.01 degrees Celsius - base period: 1951-1980

sources: GRDN 1880-01/2011 + 33T: 1880-11/1981 HadISSTI 12/1981-01/2011 Reynolds v2 using elimination of outliers and homogeneity adjustment Notes: 1950 DJT = Dea 1949 - Ieb 1950 ; **** = missing

The property of the property o

Mar 10 12 -13 -26 -10 11 7 24 -2 9 -21 -2 26 -5 11 -28 15 11 8 23 Year 1961 1962 1963 1964 1965 1966 1967 1970 1971 1972 1973 1974 1975 1976 1977 1978 1979 1979 Jon Feb 5 21 5 18 1 19 -5 -6 -10 -18 -16 -1 -7 -23 -22 -14 -26 -26 -20 -26 -28 -14 -27 -18 10 15 3 8 6 -18 21 28 8px 11 11 -7 -33 -19 -11 -2 -5 20 4 -11 -1 24 -12 -3 -16 19 9 May 22 -10 -2 -28 -6 -7 11 -9 13 -5 -12 -2 22 -5 18 -29 29 1 -6 29 Jun
12
7
6
-2
-11
-6
0
11
-2
-22
6
16
-5
-3
-15
22
-8
4
9 Jul 1 -3 14 -5 -20 11 6 -8 -4 -1 10 9 -2 -5 -13 19 4 -7 22 Aug 3 -5 26 -24 -6 -5 1 -5 0 -12 -2 20 1 10 -21 -20 17 -21 Sep 8 2 25 -37 -16 1 0 -14 9 13 -2 5 7 -9 -6 -12 -5 5 18 12 0ct 7 1 9 -30 -6 -15 10 13 14 3 -5 4 13 -7 -11 -29 -5 -4 18 8 Nov 5 8 14 -20 -6 0 0 -2 -3 16 4 -7 -2 4 -10 -17 -11 12 6 19 22 Dec -15 0 3 -30 -5 -4 -1 -12 30 -12 -11 18 -8 -11 -22 0 2 1 39 8 Year 1961 1962 1963 1964 1965 1966 1967 1970 1971 1972 1973 1974 1975 1976 1977 1978 -10 -16 19 -9 2 16 J-D D-H 26 24 4 4 4 25 27 9 11 4 2 12 12 27 24 31 33 19 19 36 36 Feb Mor 35 44 7 -12 34 36 7 22 -14 11 37 24 38 11 35 45 28 27 28 44 0ct Hov 6 18 2 3 9 23 4 -5 4 -1 5 1 26 21 29 -3 25 8 30 42 DJT MAM 30 27 12 -2 37 31 13 18 -5 7 21 20 23 15 43 39 17 16 28 50 Jul 29 11 11 11 -8 5 41 26 28 45 Dec 28 34 10 -15 7 6 42 19 26 36 Year 1981 1982 1983 1984 1985 1986 1987 1988 1989 Jan 47 0 42 22 15 19 25 51 3 Apr 23 -3 27 3 5 18 17 35 17 May 14 10 30 29 5 16 17 38 4 Jun 21 14 -2 12 4 32 40 5 26 Aug 31 -3 27 11 10 7 19 31 26 25 Sep 11 1 34 16 6 0 30 30 31 JJA 27 3 18 7 5 6 31 33 20 50N 12 2 22 5 3 2 26 19 22

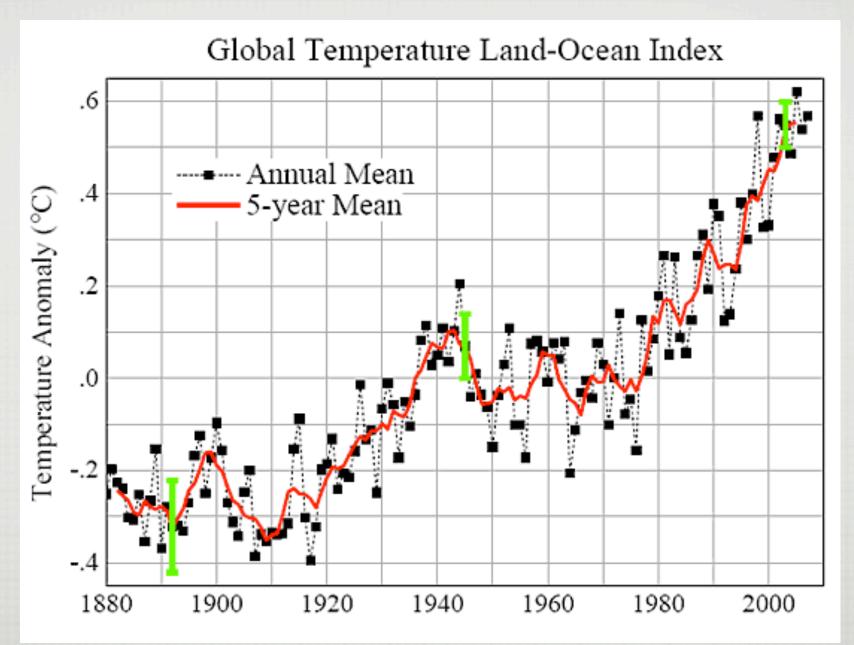
GLOBAL LAND-OCEAN TEMPERATURE INDEX IN 0.01 DEGREES CELSIUS

2006 46 61 58 42 38 55 42 65 57 59 64 71 55 54 56 46 54 60 2006 2007 89 63 64 68 60 52 55 54 50 54 47 40 58 61 74 64 54 50 2007 2008 17 26 66 43 41 34 52 36 55 56 45 50 54 47 40 58 61 74 64 54 50 2007 2008 17 26 66 43 41 34 52 36 55 56 45 50 54 47 44 3 28 50 40 55 2008 2009 55 46 47 50 54 61 66 57 65 61 68 60 58 56 49 50 61 64 2009 2010 70 75 65 75 64 55 50 64 54 56 61 68 60 58 56 49 50 61 64 2009 2011 70 75 65 75 64 55 50 64 54 62 73 40 63 65 69 75 53 63 2010 2011 464***

David by 100 to get changes in degrees Calsius (4gg-C).

Nultiply that result by 1.8(=9/5) to get changes in degrees Tahrenheit (4gg-T).

Best estimate for absolute global mean for 1951-1960 is 14.0 deg-C or 57.2 deg-T, so add that to the temperature change if you want to use an absolute scale (this note applies to global annual means only, J-3 and J-8 H)



HTTP://WWW.ECY.WA.GOV/CLIMATECHANGE/WHATIS.HTM

ON THE MENU

- Lagrange multipliers
- □ Spline
- ☐ Fit to theory
- ☐ How to use Numerical Recipes

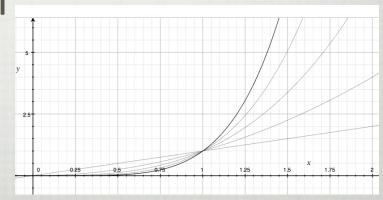
PART I: LAGRANGE INTERPOLATION

METHOD AND ALGORITHM

Any function g(x) can be approximated as a polynomial of degree (n-1) in each interval i for a sufficiently high n:

$$g_i(x) \simeq a_0 + a_1 x + a_2 x^2 + \dots + a_{n-1} x^{n-1}$$

Because our fit is <u>local</u>, we do not assume that one g(x) can fit all the data in a list, but instead will use a different polynomial, that is, a different set of <u>ai</u> values, for each region of the list.



LAGRANGE INTERPOLATION FORMULA

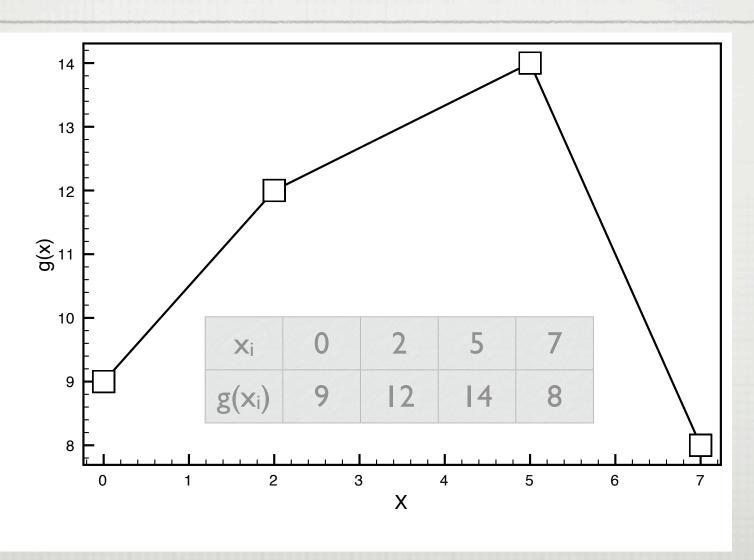
Lagrange figured out a closed-form approach that directly fits the (n - 1)-order polynomial to n values of the function g(x) evaluated at the points x_i . The formula is written as the sum of polynomials:

$$g(x) \simeq g_1 \lambda_1(x) + g_2 \lambda_2(x) + \dots + g_n \lambda_n(x)$$

$$\lambda_i(x) = \prod_{j(\neq i)=1}^n \frac{x - x_j}{x_i - x_j} = \frac{x - x_1}{x_i - x_1} \frac{x - x_2}{x_i - x_2} \cdots \frac{x - x_n}{x_i - x_n}$$

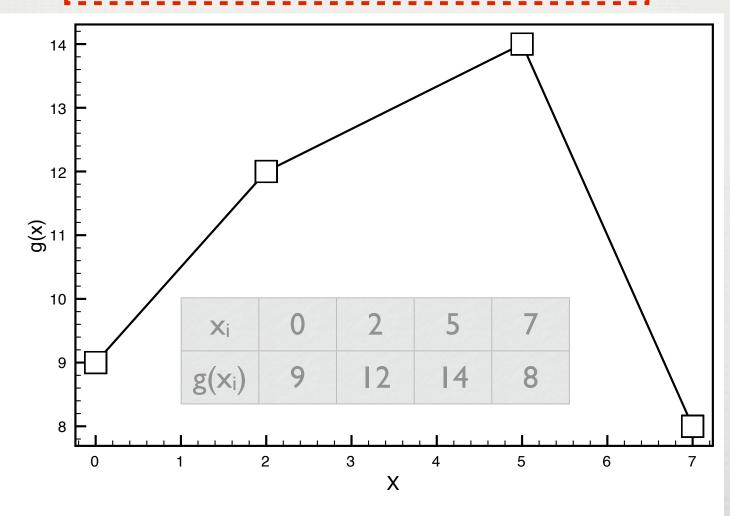
For <u>three</u> points, this formula provides a <u>second-degree polynomial</u>, while for <u>eight</u> points it gives a <u>seventh-degree</u> polynomial.

EXAMPLE: POLYNOMIAL OF ORDER 3



$$g(x) \simeq g_1 \lambda_1(x) + g_2 \lambda_2(x) + \dots + g_n \lambda_n(x)$$

$$\lambda_i(x) = \prod_{j(\neq i)=1}^n \frac{x - x_j}{x_i - x_j} = \frac{x - x_1}{x_i - x_1} \frac{x - x_2}{x_i - x_2} \cdots \frac{x - x_n}{x_i - x_n}$$



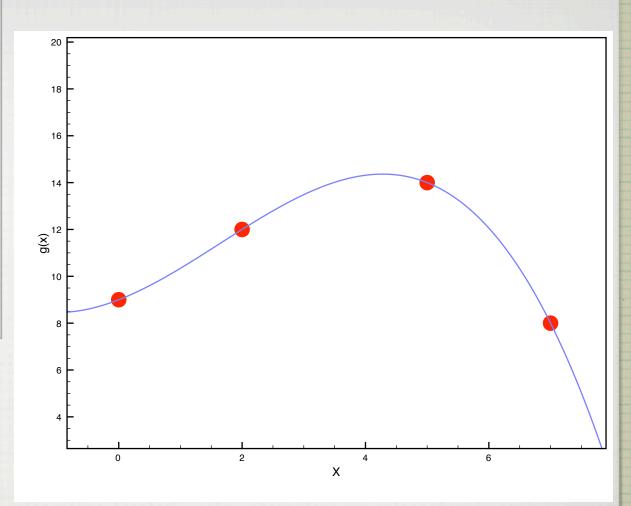
$$g(x) = g_1 \lambda_1(x) + g_1 \lambda_2(x) + g_1 \lambda_3(x)$$

$$\lambda_1(x) = \frac{(x-2)(x-5)(x-7)}{-2 \times -5 \times -7}$$

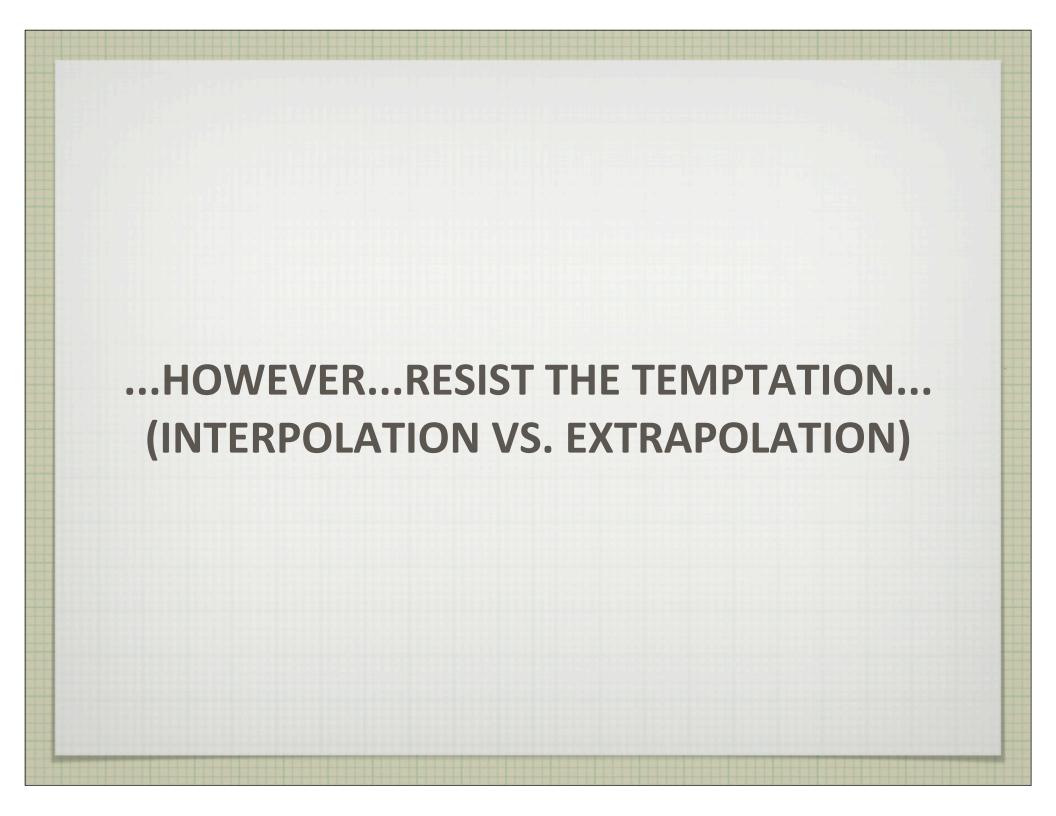
$$\lambda_2(x) = \frac{x(x-5)(x-7)}{2 \times -3 \times -5}$$

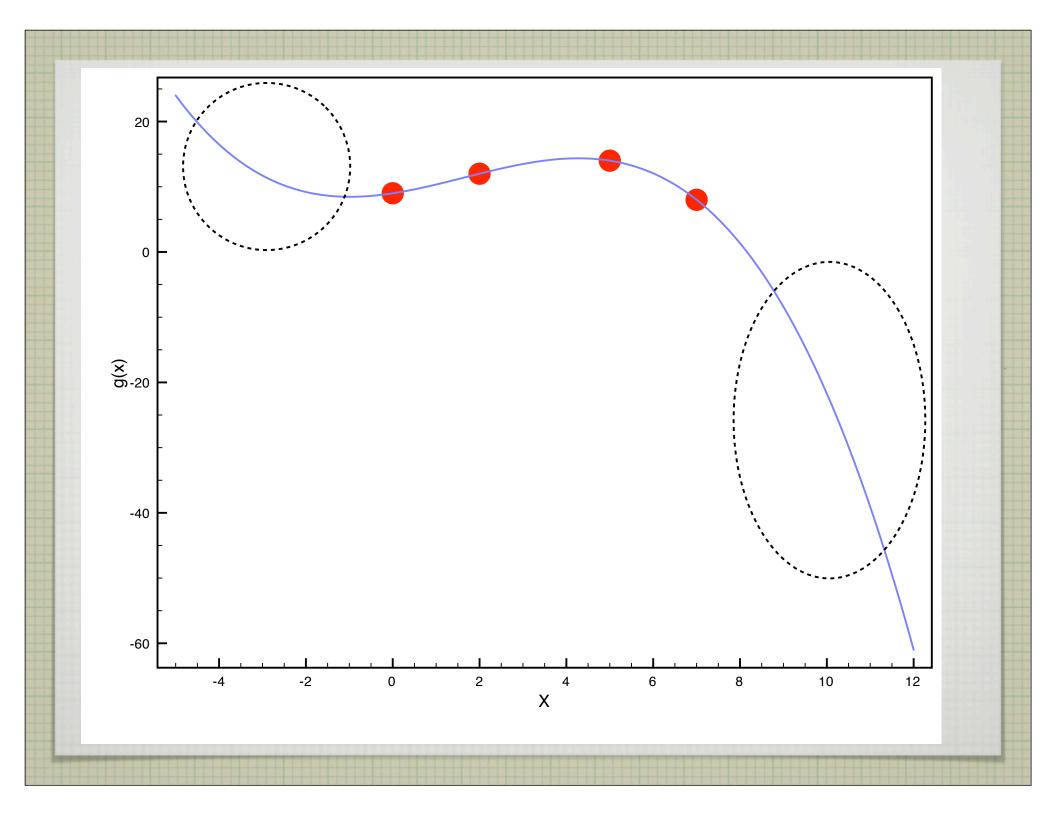
$$\lambda_3(x) = \frac{x(x-2)(x-7)}{5 \times 3 \times -2}$$

$$\lambda_4(x) = \frac{x(x-2)(x-5)}{7 \times 5 \times 2}$$

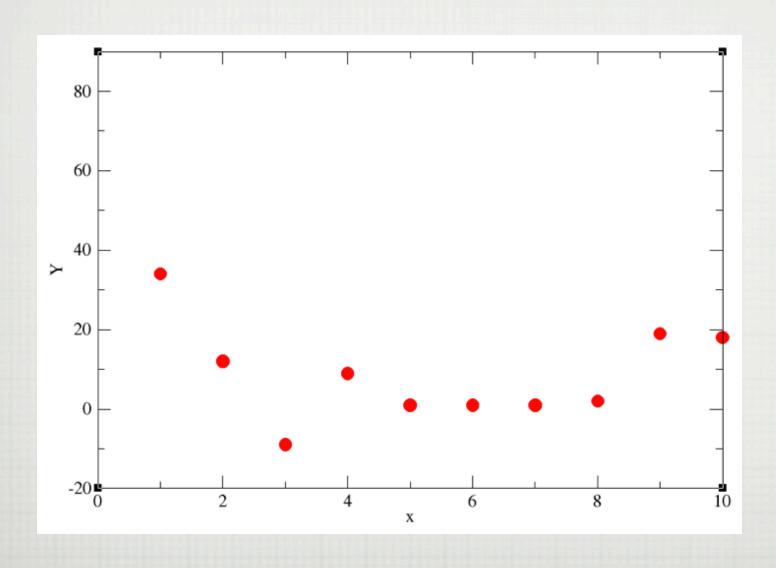


$$g(x) = \frac{-9}{70}(x-2)(x-5)(x-7) + \frac{2}{5}x(x-5)(x-7) + \frac{-7}{15}x(x-2)(x-7) + \frac{4}{35}x(x-2)(x-5)$$

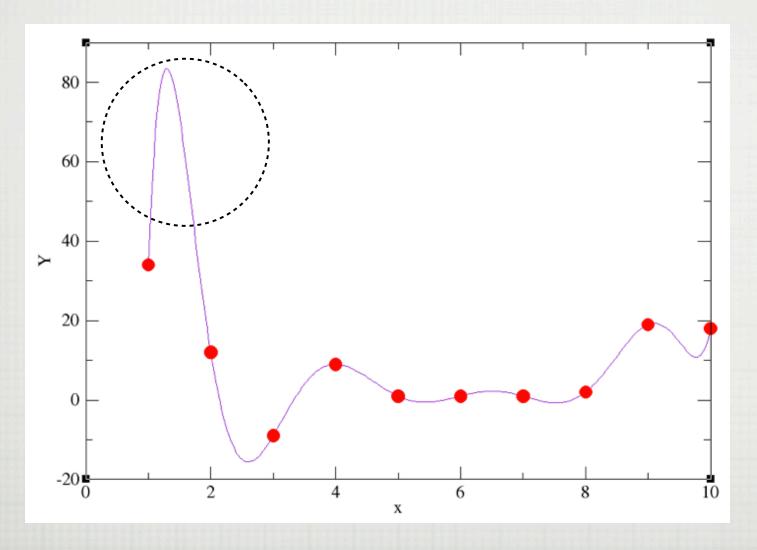




10 DATA POINTS: POLYNOMIAL OF ORDER 9



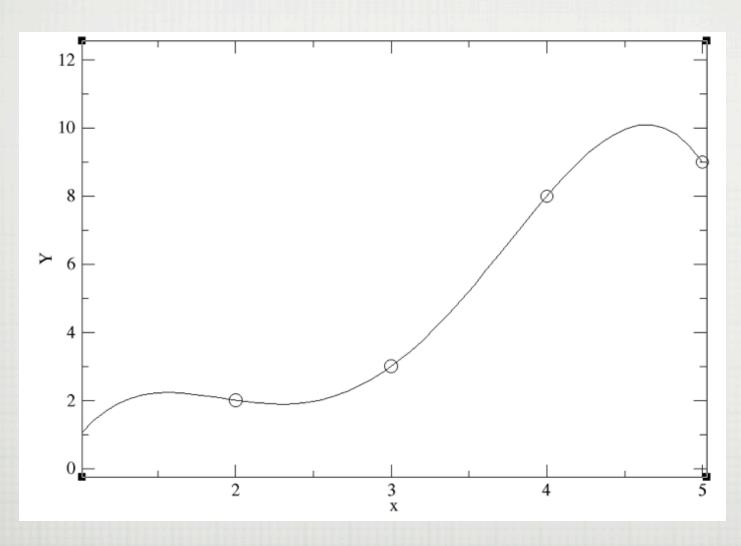
POLYNOMIAL FIT

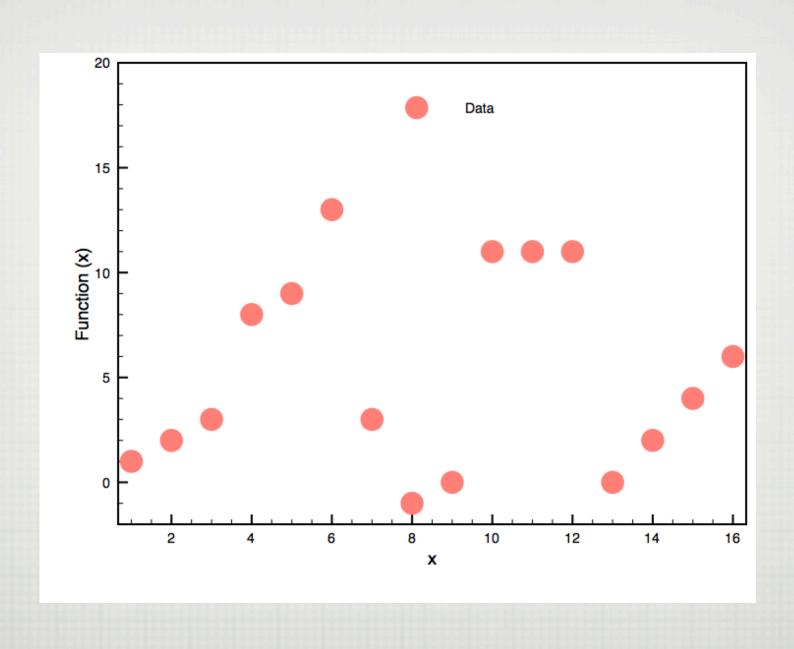


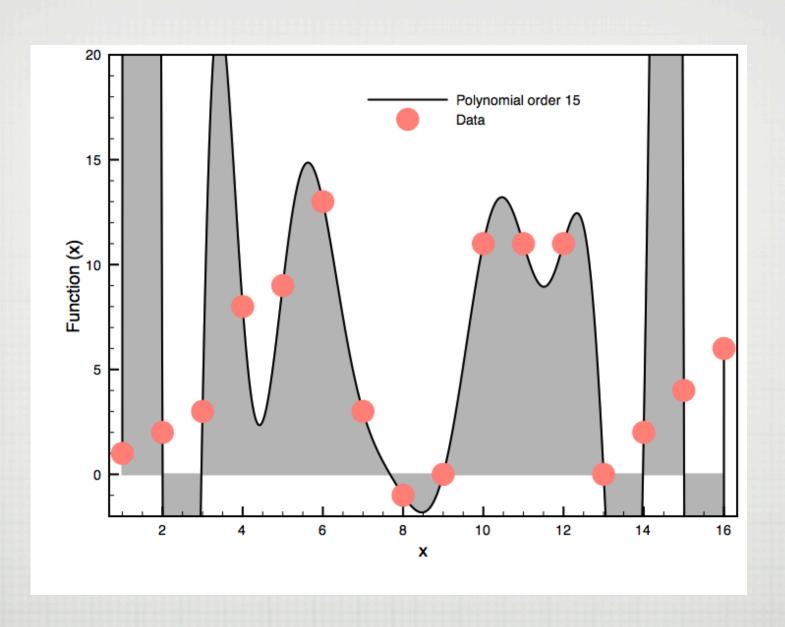
LET'S HAVE A LOOK AT THE C++ CODE

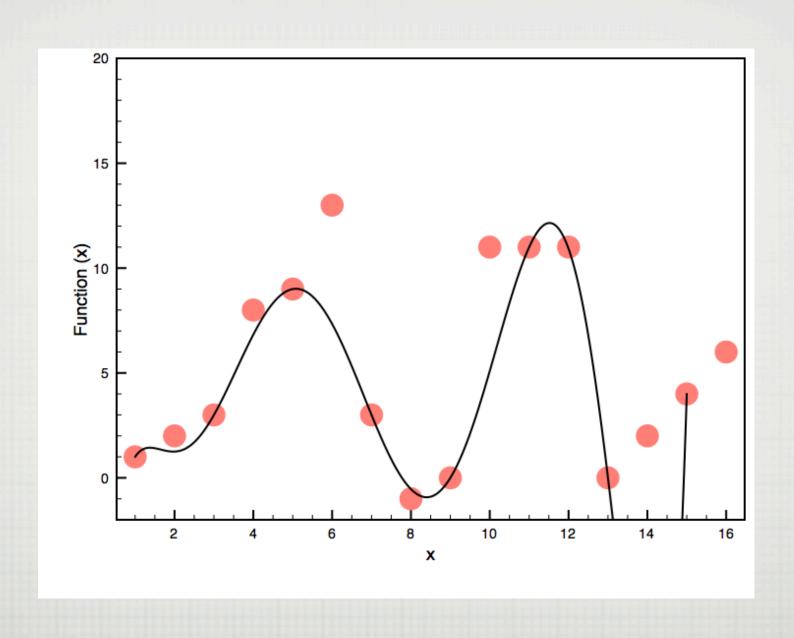
```
#INCLUDE <IOSTREAM>
USING NAMESPACE STD:
INT MAIN () {
    // INSERT CODE HERE...
    INT I, J, K;
    INT KMAX=393;
    DOUBLE XX=0, GXX=0;
    DOUBLE XXMIN, XXMAX;
    INT NPOINTS;
    CIN >> NPOINTS;
    DOUBLE X[NPOINTS], G[NPOINTS];
    DOUBLE LAMBDA[NPOINTS];
    FOR (I=0; I<NPOINTS; I++) {</pre>
           CIN >> X[I] >> G[I];
    //
    XXMIN=X[0];
    XXMAX=X[NPOINTS-1];
    FOR (K=0; K<KMAX; K++) {
           XX=XXMIN+(DOUBLE)K*(XXMAX-XXMIN)/((DOUBLE)KMAX-1);
           GXX=0.0;
           FOR (I=0;I<NPOINTS;I++){</pre>
                 LAMBDA[I]=1.;
                 FOR (J=0; J<NPOINTS; J++) {</pre>
                       IF(J==I) CONTINUE;
                       LAMBDA[I]*=(XX-X[J])/(X[I]-X[J]);
                 GXX+=LAMBDA[I]*G[I];
           COUT << XX << " " << GXX << ENDL;
    RETURN 0:
```

EXAMPLE OF POLYNOMIAL OBTAINED USING NEVILLE'S ALGORITHM









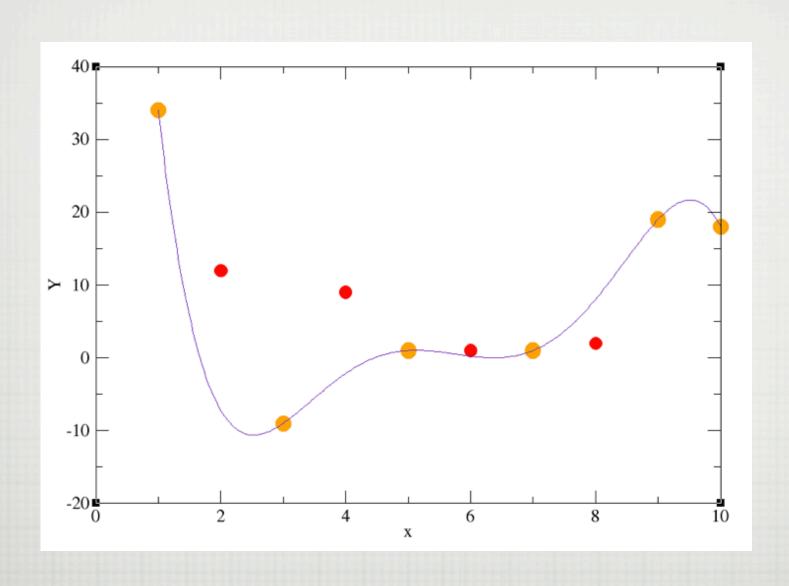
NEVILLE'S ALGORITHM

A much better algorithm to construct the (unique) polynomial is called Neville's algorithm

```
//NEVILLE's ALGORITHM
#include <iostream>
using namespace std;
int main () {
    int i,k,m;
    int kmax=50;
    double xx=0, gxx=0;
    double xxmin, xxmax;
    double sum;
    //
    int npoints;
    cin >> npoints;
    11
    double x[npoints], g[npoints];
    double P0[npoints], P[npoints];
    11
    for (i=0;i<npoints;i++){</pre>
           cin >> x[i] >> g[i];
    11
    xxmin=x[0];
    xxmax=x[npoints-1];
    //
    for(k=0; k<kmax; k++){</pre>
           xx=xxmin+(double)k*(xxmax-xxmin)/((double)kmax-1);
           qxx=0.0;
           sum=0;
           for(m=0;m<npoints;m++){</pre>
                 if(m==0){
                       for (i=0;i<npoints;i++){</pre>
                            P[i]=g[i];
                 }
                 else {
                       for (i=0;i<npoints-m;i++){</pre>
                             P[i]=((xx-x[i+m])*P0[i]+(x[i]-xx)*P0[i+1])/(x[i]-x[i+m]);
                 for (i=0;i<npoints-m;i++){</pre>
                       P0[i]=P[i];
           cout << xx << " " << P[0] << endl;</pre>
    return 0;
```

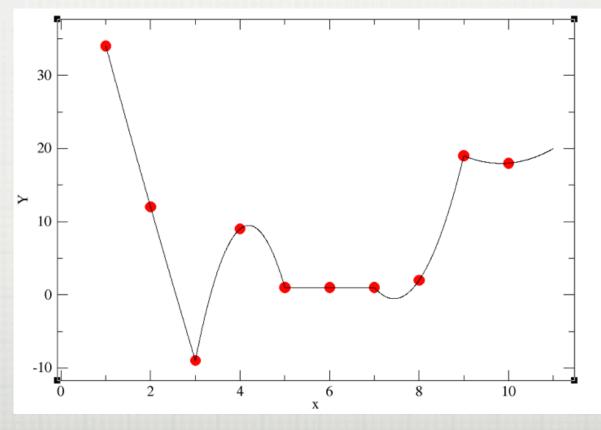
PRACTICAL CONSIDERATIONS

- If the data contain little noise, this polynomial can be used with some confidence within the range of data, but with risk beyond the range of data.
- □ Notice that Lagrange interpolation makes no restriction that the points in the table be evenly spaced.
- \square As a check, it is also worth noting that the sum of the Lagrange multipliers equals one, $\sum \lambda_i = 1$.
- ☐ Usually the Lagrange fit is made to only a small region of the table with a small value of n, even though the formula works perfectly well for fitting a high-degree polynomial to the entire table.

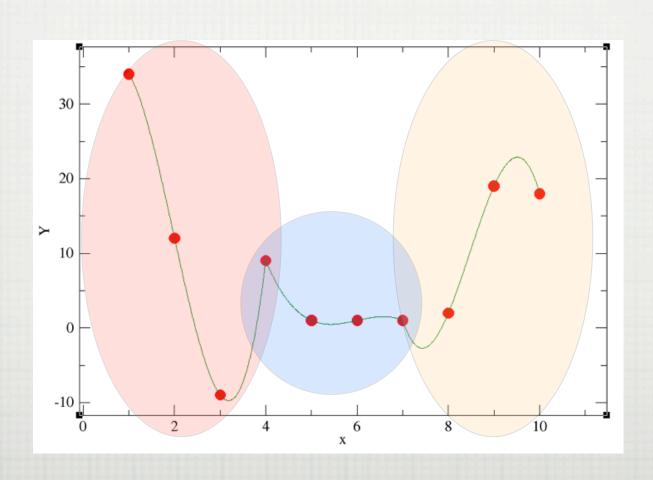


LOCAL INTERPOLATION

Here, we interpolate the function, small interval by small interval,
 with a low order polynomial (here: 3-points or parabola)

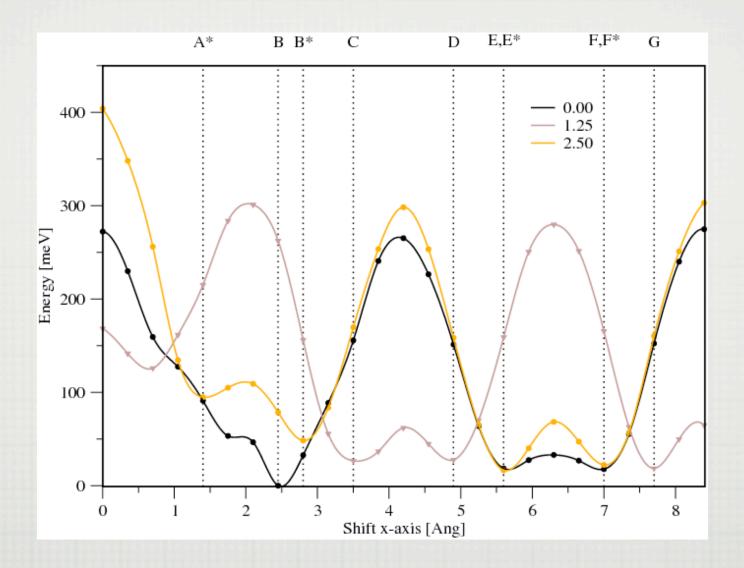


4-POINT SEGMENTATION (THIRD ORDER POLYNOMIAL)



PROBLEMS: DISCONTINUITY IN DERIVATIVE!

PART2: SPLINE



SPLINE: INTRODUCTION

- ☐ Fitting parabolas (three-point interpolation) within a table may avoid the erroneous and possibly catastrophic deviations of a high-order formula.
- ☐ (Two-point interpolation, which connects the points with straight lines, may not lead you far astray, but it is rarely pleasing to the eye or precise.)
- □ A sophisticated variation of n = 4 interpolation, known as cubic splines, often leads to surprisingly eye-pleasing fits.

SPLINE: INTRODUCTION (II)

In this approach, cubic polynomials are fit to the function in each interval, with the additional constraint that the first and second derivatives of the polynomials must be continuous from one interval to the next. This continuity of slope and curvature is what makes the spline fit particularly eye-pleasing.

SPLINE: C'TD

- The series of cubic polynomials obtained by splinefitting a table can be integrated and differentiated, and is guaranteed to have well-behaved derivatives.
- The complexity of simultaneously matching polynomials and their derivatives over all the interpolation points leads to many simultaneous, linear equations to be solved.
- This makes splines unattractive for hand calculations, yet easy for computers.

* The basic approximation of splines is the representation of the function g(x) in the subinterval $[x_i, x_{i+1}]$ with a cubic polynomial:

$$g(x) \simeq g_i(x)$$
 for $x_i \le x \le x_{i+1}$
 $g_i(x) = g_i + g_i'(x - x_i) + \frac{1}{2}g_i''(x - x_i)^2 + \frac{1}{6}g_i'''(x - x_i)^3$

- This representation makes it clear that the coefficients in the polynomial equal the values of g(x) and to its first, second, and third derivatives at the tabulated points x_i. Derivatives beyond the third vanish.
 - * The computational chore is to determine these derivatives in terms of the N tabulated values g_i.
 - * The matching of g_i from one interval to the next (at the nodes) provides the equations

$$g_i(x_{i+1}) = g_{i+1}(x_{i+1})$$
 $i = 1, N-1$

 The matching of the first and second derivatives at each subinterval's boundary provides the equations

$$g'_{i-1}(x_i) = g'_i(x_i)$$
 $g''_{i-1}(x_i) = g''_i(x_i)$

To provide the additional equations needed to determine all constants, the third derivatives at adjacent nodes are matched. Values for the third derivatives are found by approximating them in terms of the second derivatives (our old FD approximation):

$$g_i''' \simeq \frac{g_{i+1}'' - g_i''}{x_{i+1} - x_i}$$

It is straightforward though complicated to solve for all the parameters.

BOUNDARIES

Matching at the boundaries of the intervals results in only N – 2 linear equations for N unknowns. Further input is required.

It usually is taken to be the boundary conditions at the endpoints $a = x_1$ and $b = x_N$, specifically, the second derivatives g''(a), and g''(b).

- <u>Natural spline</u>: Set g"(a) = g"(b) = 0, (This is "natural" because the derivative vanishes for the flexible spline drafting tool (its ends being free).)
- <u>Input values for g' at boundaries</u>: The computer uses g'(a) to approximate g"(a). If you do not know the first derivatives, you can calculate them numerically from the table of gi values.
- Input values for g" at boundaries: Knowing values is of course better than assuming values, but it requires more input. If the values of g" are not known, they can be approximated by applying a forward-difference approximation to the tabulated values

HARD WARK

NUMERICAL RECIPES

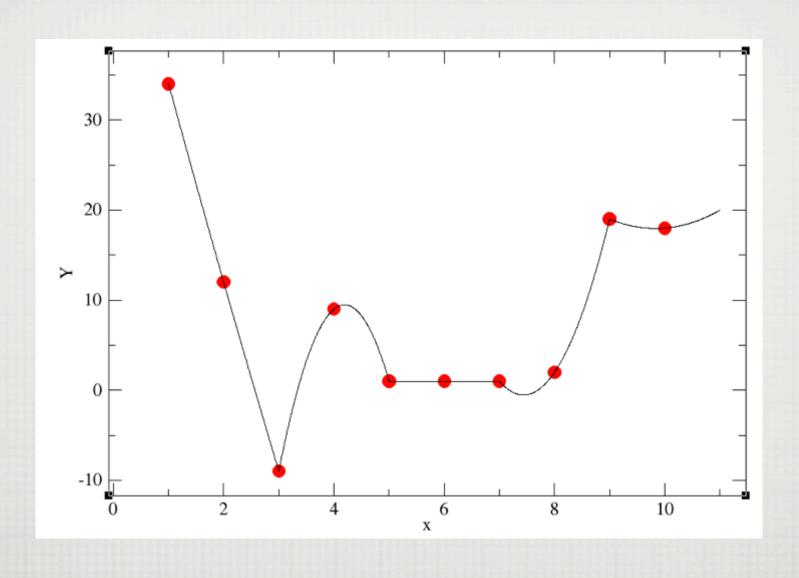
The Art of Scientific Computing

THIRD EDITION

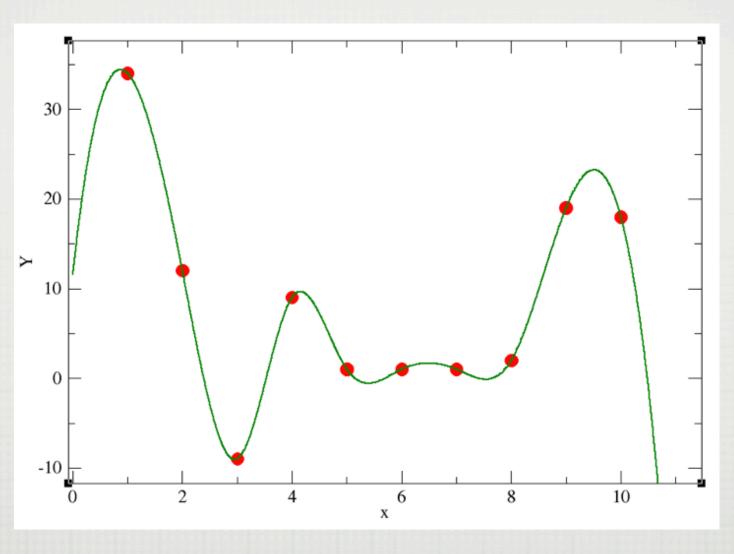
William H. Press Saul A. Teukolsky William T. Vetterling Brian P. Flannery



```
#include <iostream>
#include "nr3.h"
#include "interp 1d.h"
using namespace std;
int main () {
  int i,k;
  int kmax=50;
  double xx=0, gxx=0;
  double xxmin, xxmax;
                                          SPLINE WITH NUMREC
  //
  int npoints;
//GET INPUT
   cin >> npoints;
//allocate memory
  VecDoub x(npoints), g(npoints);
//read in data points
  for (i=0;i<npoints;i++){</pre>
      cin >> x[i] >> q[i];
//CREATE SPLINE OBJECT
  Spline interp myspline (x,q);
//CREATE INTERPOLATION
  xxmin=x[0];
  xxmax=x[npoints-1];
    11
  for(k=0; k<kmax; k++) {</pre>
      xx=xxmin+(double)k*(xxmax-xxmin)/((double)kmax-1);
      gxx= myspline.interp(xx);
       cout << xx << " " << qxx <<endl;</pre>
```







USING SPLINES FOR INTEGRATION

- A powerful integration scheme is to fit an integrand with splines, and then integrate the cubic polynomials analytically.
- If you have the ability to actually calculate the function for arbitrary x, Gaussian quadrature may be preferable. We know that the spline fit to g in each interval is the cubic (9.8),

$$g(x) \simeq g_i + g_i'(x - x_i) + \frac{1}{2}g_i''(x - x_i)^2 + \frac{1}{6}g_i'''(x - x_i)^3$$

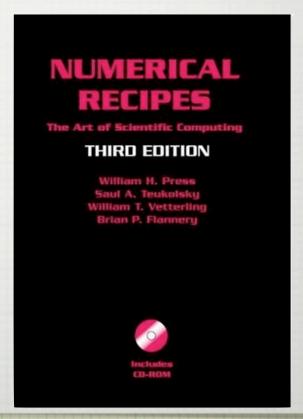
☐ It is easy to integrate this to obtain the integral of g for this interval, and then to sum over all intervals:

$$\int_{x_i}^{x_{i+1}} g(x)dx \simeq \left(g_i x + \frac{1}{2} g_i' x_i^2 + \frac{1}{6} g_i'' x^3 + \frac{1}{24} g_i''' x^4 \right) \Big|_{x_i}^{x_{i+1}}$$
$$\int_{x_j}^{x_k} g(x)dx = \sum_{i=j}^k \int_{x_i}^{x_{i+1}} g(x)dx$$

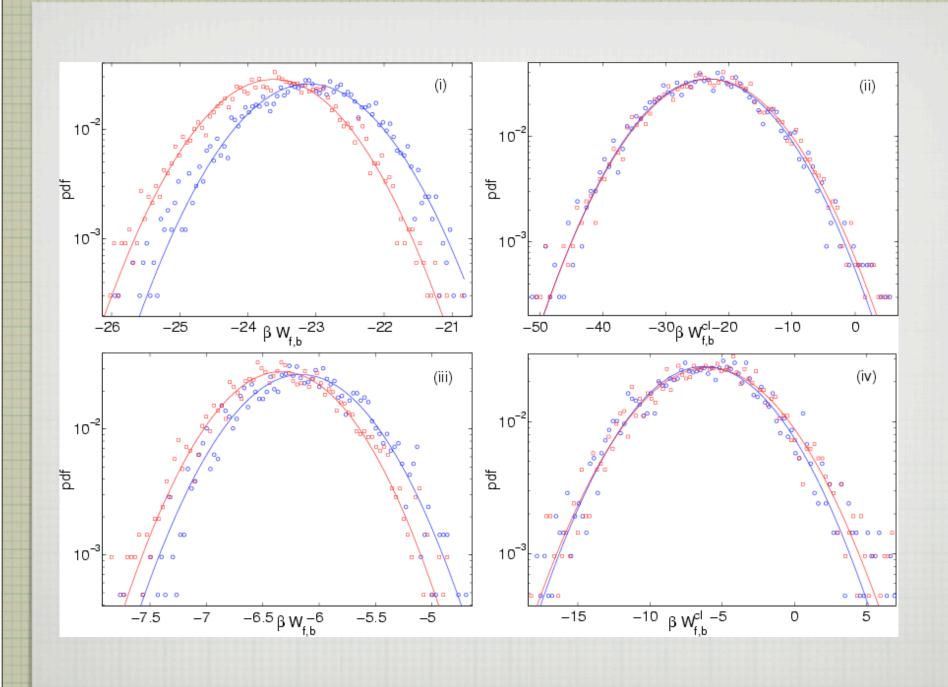
☐ Making the intervals smaller does not necessarily increase precision as subtractive cancellations may get large.

HOW DO WE USE THESE METHODS?

- Preferred method: using numerical libraries
- ☐ Go online to nr.com or look in a book...or download the routines...
- ☐ Example: spline
- http://www.nr.com/dependencies/index.php



PART 3: FIT TO A FUNCTION ("FIT TO THEORY")



LEAST SQUARE FITTING: "GOODNESS OF A FIT"

Imagine that you have measured N_D data values of the independent variable y as a function of the dependent variable x:

$$(x_i, y_i \pm \sigma_i)$$
 $i = 1, N_D$

- \square where $\pm \sigma_i$ is the uncertainty in the ith value of y.
- Our goal is to determine how well a mathematical function y = g(x) (also called theory or model) can describe these data.

$$g(x) = g(x; \{a_1, a_2, \dots, a_{M_P}\}) = g(x; \{a_m\})$$

- We assume that the model function g(x) contains, in addition to the functional dependence on x, an additional dependence upon MP parameters {a₁, a₂, . . . , a_{MP} }.
- * The parameters {a_m} are not variables, in the sense of numbers read from a meter, but rather are parts of the theoretical model such as the size of a box, the mass of a particle, or the depth of a potential well.

MINIMIZATION "CHI-SQUARE"

$$\chi^2 \stackrel{\text{def}}{=} \sum_{i=1}^{N_D} \left(\frac{y_i - g(x_i; \{a_m\})}{\sigma_i} \right)^2$$

* We use the chi-squared (χ^2) measure as a gauge of how well a theoretical function g reproduces data

$$\chi^2 \stackrel{\text{def}}{=} \sum_{i=1}^{N_D} \left(\frac{y_i - g(x_i; \{a_m\})}{\sigma_i} \right)^2$$

- * The sum is over the ND experimental points $(x_i, y_i \pm \sigma_i)$.
- * The definition is such that smaller values of χ^2 are better fits, with $\chi^2 = 0$ occurring if the theoretical curve went through the center of every data point.
- * Notice also that the $1/\sigma_i^2$ weighting means that measurements with larger errors contribute less to χ^2 .

- * Least-squares fitting refers to adjusting the theory until a minimum in χ^2 is found
- * The goal is to find a curve that produces the least value for the summed squares of the deviations of the data from the function g(x).
- * In general, this is the best fit possible or the best way to determine the parameters in a theory.
- * The MP parameters $\{a_m, m = 1, MP\}$ that make χ^2 an extremum are found by solving the MP equations:

$$\frac{\partial \chi^2}{\partial a_m} = 0 \quad \Rightarrow \quad \sum_{i=1}^{N_D} \frac{[y_i - g(x_i)]}{\sigma_i^2} \frac{\partial g(x_i)}{\partial a_m} = 0 \quad (m = 1, M_P)$$

GOODNESS OF A FIT

- ☐ When the deviations from theory are due to random errors and when these errors are described by a Gaussian distribution, there are some useful rules of thumb to remember.
- You know that your fit is good if the value of χ^2 calculated via the definition is approximately equal to the number of degrees of freedom $\chi^2 \approx ND MP$, where ND is the number of data points and MP the number of parameters in the theoretical function.
- If your χ^2 is much less than ND MP, it does not mean that you have a "great" theory or a really precise measurement; instead, you probably have too many parameters or have assigned errors (σ_i values) that are too large. In fact, too small a χ^2 may indicate that you are fitting the random scatter in the data rather than missing ~ 1 of the error bars, as expected for Gaussian statistics.

FITTING TO A STRAIGHT LINE: LINEAR REGRESSSION

The MP simultaneous equations simplify considerably if the functions g(x; {am}) depend linearly on the parameter values a:

$$g(x; \{a_1, a_2\}) = a_1 + a_2 x$$

- ☐ In this case there are MP = 2 parameters, the slope a_2 and the y intercept a_1 .
- A unique solution is not possible unless the number of data points is equal to or greater than the number of parameters.
- For this linear case, there are just two derivatives,

$$\frac{\partial g(x_i)}{\partial a_1} = 1 \qquad \frac{\partial g(x_i)}{\partial a_2} = x_i$$

 \square and after substitution, the χ^2 minimization equations can be solved

LINEAR REGRESSION (2)

$$a_{1} = \frac{S_{xx}S_{y} - S_{x}S_{xy}}{\Delta} \qquad a_{2} = \frac{SS_{xy} - S_{x}S_{y}}{\Delta}$$

$$S = \sum_{i=1}^{N_{D}} \frac{1}{\sigma_{i}^{2}} \qquad S_{x} = \sum_{i=1}^{N_{D}} \frac{x_{i}}{\sigma_{i}^{2}} \qquad S_{y} = \sum_{i=1}^{N_{D}} \frac{y_{i}}{\sigma_{i}^{2}}$$

$$S_{xx} = \sum_{i=1}^{N_{D}} \frac{x_{i}^{2}}{\sigma_{i}^{2}} \qquad S_{xy} = \sum_{i=1}^{N_{D}} \frac{x_{i}y_{i}}{\sigma_{i}^{2}} \qquad \Delta = SS_{xx} - S_{x}^{2}$$

☐ Statistics also gives you an expression for the variance or uncertainty in the deduced parameters:

$$\sigma_{a_1}^2 = \frac{S_{xx}}{\Delta} \qquad \sigma_{a_2}^2 = \frac{S}{\Delta}$$

```
#include <iostream>
#include "nr3.h"
#include "gamma.h"
#include "incgammabeta.h"
#include "fitab.h"
using namespace std;
int main () {
                                                    LINEAR REGRESSION
   int i,k;
   int kmax=500;
   double xx=0, qxx=0;
   double aa, bb, xxmin, xxmax;
   //
  int npoints;
   cin >> npoints;
   //
  VecDoub x(npoints), g(npoints);
   //
   for (i=0;i<npoints;i++){</pre>
       cin >> x[i] >> g[i];
   Fitab myreg(x,g);
11
  xxmin=x[0];
   xxmax=x[npoints-1];
   //
   aa=myreg.a;
   bb=myreg.b;
   cout << aa << " " << bb << " " << myreg.chi2<< endl;</pre>
   for(k=0; k<kmax; k++) {</pre>
       xx=xxmin+(double)k*(xxmax-xxmin)/((double)kmax-1);
       gxx=bb*xx+aa;
       cout << xx << " " << gxx <<endl;</pre>
```

sources: GMCH 1880-01/2011 + SST: 1880-11/1981 HodISSTI 12/1981-01/2011 Reymolds v2 using elimination of outliers and homogeneity adjustment Hotes: 1950 DJT = Dec 1949 - Teb 1950 ; ***** = missing

GLOBAL Land-Ocean Temperature Index in 0.01 degrees Celsius | base period: 1951-1980

J-B B-H -15 -13 -24 -23 -30 -31 -34 -36 -24 -24 -19 -19 -39 -37 -33 -33 DJF MAM JJA -9 -5 -24 -23 -12 -24 -23 -21 -27 -36 -47 -36 -33 -35 -26 -20 -25 -13 -15 -34 -40 -40 -36 -39 -27

Year 1961 1962 1963 1964 1965 1966 1967 1970 1971 1972 1973 1974 1975 1976 1977 1978 1979 1979 Jon Feb 5 21 5 18 1 19 -5 -6 -10 -18 -16 -1 -7 -23 -22 -14 -26 -26 -20 -26 -28 -14 -27 -18 10 15 3 8 6 -18 21 28 8px 11 11 -7 -33 -19 -11 -2 -5 20 4 -11 -1 24 -12 -3 -16 19 9 May 22 -10 -2 -28 -6 -7 11 -9 13 -5 -12 -2 22 -5 18 -29 29 1 -6 29 Jun
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GLOBAL LAND-OCEAN TEMPERATURE INDEX IN 0.01 DEGREES CELSIUS

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-6 1959
0 1960 Jun 7 5 2 12 0 -14 -7 4 -20 -5 -3 4 5 -12 -13 16 -8 7 -3 1 18 -52 | 18 - 52 | 18 - 52 | 18 - 52 | 18 - 52 | 18 - 52 | 18 - 52 | 18 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 52 | 19 - 8 -17 1
B15 114 5
3 15 5 0
3 7 -8 115
-3 -9 -8 115
-19 -12 -13
-8 -3 2 -11
-19 -12 -13
-11 6 0
-12 4 -4
-9 10 3
21 3 -2
-11 -15 -12
-19 -2 9 -2
-10 -2 9 -10
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-3 9 -10
-3 9 -10
-4 -24 16
-8 21 19
-4 7 -9 J-D D-N
7 10
4 3
8 8
-21 -18
-11 -13
-3 -3
-1 -1
-4 -4
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3 6
-10 -11
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-5 -4
-16 -18
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56 46 54 60 2006 74 64 54 50 2007 28 50 40 55 2008 49 50 61 64 2009 69 75 53 63 2010 2011 46 61 58 42 38 55 42 65 57 59 64 71 89 63 64 66 80 52 58 54 50 54 47 47 40 17 29 64 643 41 34 52 34 53 55 55 47 75 46 47 7 50 54 61 66 57 65 61 66 60 70 75 65 75 64 55 50 54 62 73 40 46 75 75 75 64 55 50 54 62 73 40 55 54 58 61 44 43 58 56 63 65

Divide by 100 to get changes in degrees Celsius (deg-C).
Multiply that result by 1.8(=9/5) to get changes in degrees Fahrenheit (deg-F)

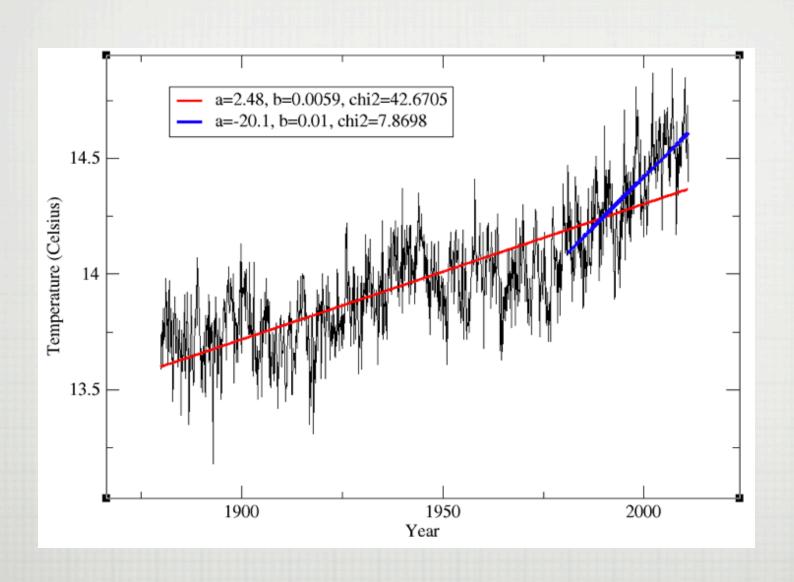
Best estimate for absolute alobal mean for 1951-1980 is 14.0 deg-C or 57.2 deg-F

so add that to the temperature change if you want to use an absolute scale (this note applies to global annual means only, J-D and D-N !)

 Example
 - Table Value
 40

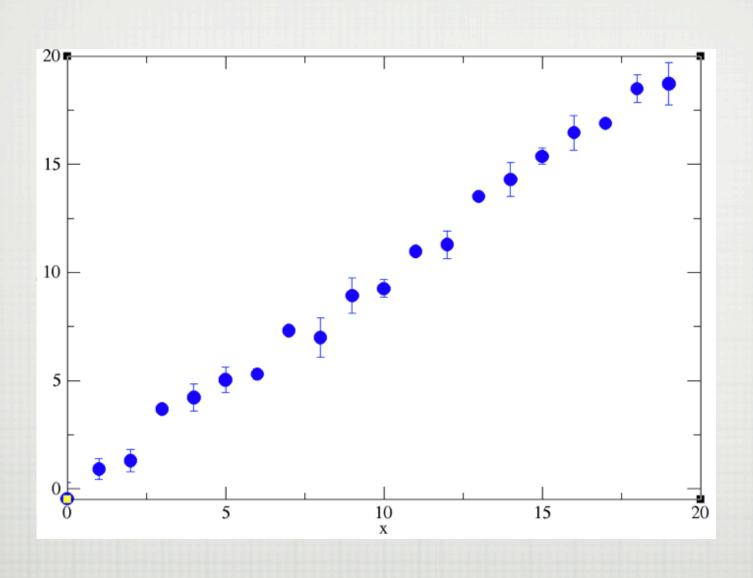
 change
 0.40 deg-C
 or 0.72 deg-T

 abs. scale if global annual mean
 14.40 deg-C
 or 57.92 deg-T

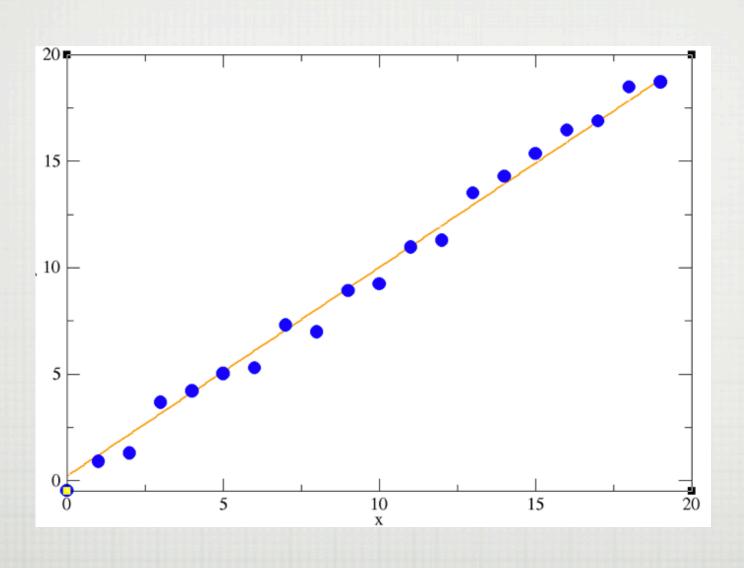


```
#include <iostream>
#include "nr3.h"
#include "gamma.h"
#include "incgammabeta.h"
                                                   WITH EXPERIMENTAL ERROR!
#include "fitab.h"
using namespace std;
double randDouble(double low, double high)
   double temp;
   temp = ((double) rand() / (static_cast<double>(RAND_MAX) + 1.0))* (high - low) + low;
   return temp;
int main () {
   int i,k;
   int kmax=500;
   double xx=0, qxx=0;
   double aa, bb, xxmin, xxmax;
   int npoints=20;
   srand((unsigned) time(0));
   VecDoub x(npoints), g(npoints), sigma(npoints);
   for (i=0;i<npoints;i++){</pre>
        x[i]=double(i);
        g[i] = randDouble(-1.0, 1.0) + x[i];
        sigma[i]=randDouble(-1.,1.0);
        cout << x[i] << " " << q[i] << " " << sigma[i] << endl;</pre>
   Fitab myreg(x,g,sigma);
   //
   xxmin=x[0];
   xxmax=x[npoints-1];
   //
   aa=myreq.a;
   bb=myreq.b;
   cout << aa << " " << bb << " " << myreg.chi2<< endl;</pre>
   for(k=0; k<kmax; k++){</pre>
        xx=xxmin+(double)k*(xxmax-xxmin)/((double)kmax-1);
        gxx=bb*xx+aa;
        cout << xx << " " << qxx <<endl;</pre>
```

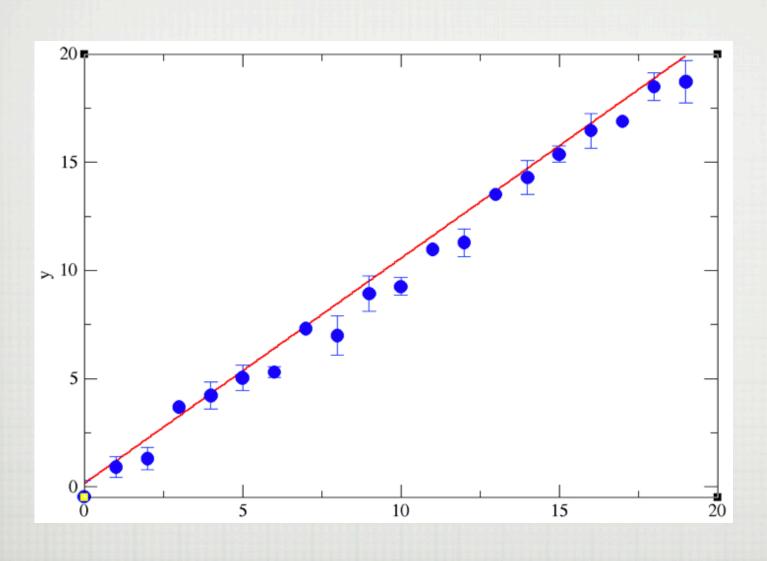




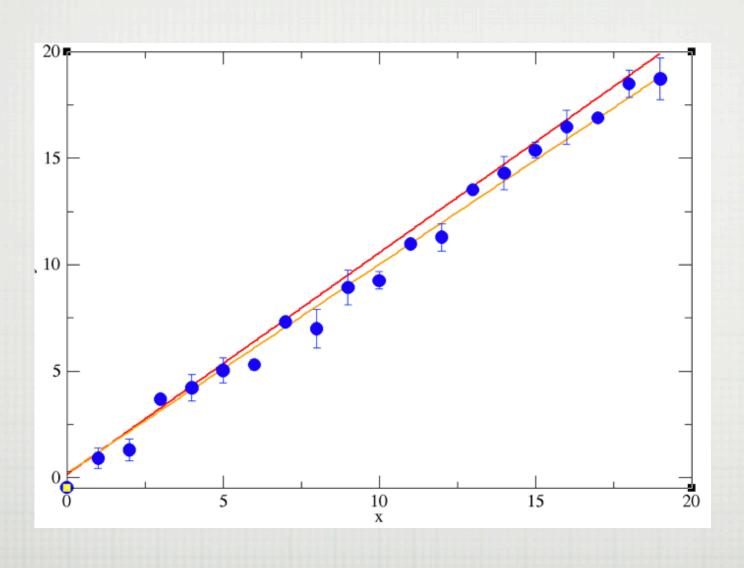




REGRESSION WITH ERRORS



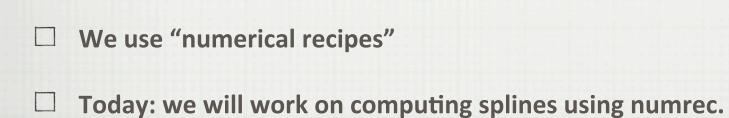
COMPARISON



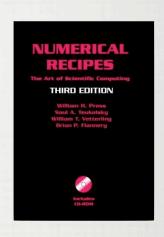
SUMMARY

Pol	Polynomial fits	
	Lagrange formula and Neville's algorithm	
	Can fit polynomial of order N-1 passing exactly through N points	
Spl	Spline fits	
	More pleasing for the eye	
	3rd order polynomial with continuity of slope and curvature	
Lea	Least Xi ² fits	
	best fit does not necessarily goes though all the points	
	very powerful when we know the analytical form of the solution	

IN PRACTICE



☐ See LMS



```
#include <iostream>
#include "nr3.h"
#include "interp 1d.h"
using namespace std;
int main () {
  int i,k;
  int kmax=50;
  double xx=0, gxx=0;
  double xxmin, xxmax;
                                          SPLINE WITH NUMREC
  //
  int npoints;
//GET INPUT
   cin >> npoints;
//allocate memory
  VecDoub x(npoints), g(npoints);
//read in data points
  for (i=0;i<npoints;i++){</pre>
      cin >> x[i] >> q[i];
//CREATE SPLINE OBJECT
  Spline interp myspline (x,q);
//CREATE INTERPOLATION
  xxmin=x[0];
  xxmax=x[npoints-1];
    11
  for(k=0; k<kmax; k++) {</pre>
      xx=xxmin+(double)k*(xxmax-xxmin)/((double)kmax-1);
      gxx= myspline.interp(xx);
       cout << xx << " " << qxx <<endl;</pre>
```