## Astronomy 211 Formulas

**Distance** d[pc] = 1/parallax[arcsec]

**magnitudes**  $m = C - 2.5 \log_{10} F$ ;  $m_1 - m_2 = -2.5 \log_{10} (F_1/F_2)$ ,  $M - m = 5 - 5 \log_{10} (d/\text{pc})$ 

Doppler shift  $\Delta \lambda / \lambda = -\Delta \nu / \nu = v_{\rm rad} / c$ 

Gravity and Tides accel= $GM/r^2$ , accel<sub>tide</sub>  $\approx (GM/r^2)(2R/r)$ ; escape speed  $v_{\rm esc} = \sqrt{2GM/R}$ 

**Kepler**  $GM = \Omega^2 a^3 = \left(\frac{2\pi}{P}\right)^2 a^3$ ;  $\frac{a_1}{a} = \frac{v_1}{v} = \frac{m_2}{M}$ ; reduced mass  $\mu = m_1 m_2/(m_1 + m_2)$ , total mass  $M = m_1 + m_2$ ; for circular orbit  $v = \sqrt{GM/a}$ ,  $L = \mu\sqrt{GMa}$ ; for Solar units (years, AU,  $M_{\odot}$ ):  $P^2 = a^3/M$ 

Virial Theorem  $E_{\text{kin}} = -\frac{1}{2}E_{\text{pot}}$ ;  $E_{\text{tot}} = E_{\text{kin}} + E_{\text{pot}} = \frac{1}{2}E_{\text{pot}}$ , [where  $E_{\text{pot,binary}} = -\frac{Gm_1m_2}{a}$  and  $E_{\text{pot,star}} \approx -\frac{GM^2}{R}$ ]

**Ideal Gas**  $P = nk_BT = \frac{\rho}{\mu m_H}k_BT$ ; typical KE per particle is  $\frac{3}{2}k_BT$ ; energy density  $u = \frac{3}{2}nk_BT = \frac{3}{2}P$ 

Degenerate Gas  $\Delta x \Delta p \sim \hbar$ ;  $E_{\rm F} = \frac{1}{2} \frac{p_{\rm F}^2}{m_e} \propto n_e^{2/3}$ ;  $P \propto n_e E_{\rm F} \propto n_e^{5/3} \propto (\rho/\mu)^{5/3} \rightarrow R \propto M^{-1/3}$  [non-relativistic]

Photon Propagation  $l_{\mathrm{mfp}} = \frac{1}{n\sigma} = \frac{1}{\kappa\rho}; \, t_{\mathrm{randomwalk}} = \frac{R}{l_{\mathrm{mfp}}} \frac{R}{c}$ 

**Blackbody**  $L=4\pi R^2\sigma T_{\mathrm{eff}}^4,\,F=\sigma T_{\mathrm{eff}}^4;\,\lambda_{\mathrm{peak}}=0.29\,\mathrm{cm}/T$ 

**Light**  $c = \lambda \nu$ ,  $E = h\nu = hc/\lambda$ , p = E/c, energy density  $u = aT^4$ , pressure  $P = (a/3)T^4$ 

Hydrostatic Equilibrium  $\frac{\Delta P}{\Delta r}=\rho\frac{GM}{r^2}=-g\rho\rightarrow P\propto M^2/R^4$ 

Stars  $T_c \propto M/R$ ,  $\rho_c \propto M/R^3$ ,  $P_c \propto M^2/R^4$ 

Timescales  $\tau_{\rm free-fall} \sim \sqrt{1/G\rho}$ ;  $\tau_{\rm Kelvin-Helmholtz} \sim \frac{GM^2/R}{L}$ 

Hydrogen Fusion  $E = \Delta mc^2 \approx 0.7\%c^2$ 

Hydrogen Atom  $E_n = -13.6 \,\mathrm{eV}/n^2$ 

Jeans mass  $M_J = 3k_BTR/2G\mu m_{\rm H}; \ \rho_J = (3/4\pi M^2)(3k_BT/2G\mu m_{\rm H})^3$ 

Equilibrium Temperature  $T_p = T_s(1-A)^{1/4}(R_s/2d)^{1/2}$ 

Diffraction Limit  $\theta = \lambda/D$ 

**Poisson noise** uncertainty on n counts is  $\sqrt{n}$