Franck-Hertz Experiment; Measuring Energy Levels of Neon and Mercury

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Abstract

I. INTRODUCTION

II. METHODS

A. Neon

B. Mercury

III. RESULTS

A. Neon

We did four data runs using neon. For each run we recorded the accelerating voltage (x data) and the electron current measured by the anode???, this current was recorded as a voltage measured across an internal resister. Each of our runs showed three discernible dips in the voltage corresponding to electron current, see the sample data run Figure **??**. These dips are the voltages just before the electrons have enough energy (from the accelerating voltage) to reach the anode even after an inelastic collision with a neon atom. The apparent double minima, see the second and third dips in Figure **??**, is most likely caused by by energy levels with very similar excitation energies.[?] Also of interest is the voltage corresponding to the steepest negative slope which is when the majority of electrons have enough energy to cite the neon atoms. However the location of the steepest negative slope was difficult to determine from our data, again see Figure **??**.

B. Mercury

Since the number of discernible dips for mercury depends on temperature, we collected data for 10°C increments starting at 150°C and ranging to 210°C. From this data, see Figure **??**, it can be seen that there is an optimum temperature at around 200°C at which the most dips can be observed.

Using the information that we obtained about the optimum range of temperature, we took a more data using a lock-in. With the lock-in we took data from 185°C to 215°C in increments of 5°C. The output from the lock-in, see Figure **??**, is the derivative of the output without the lock-in. Therefore, the dips in the lock-in output correspond to the
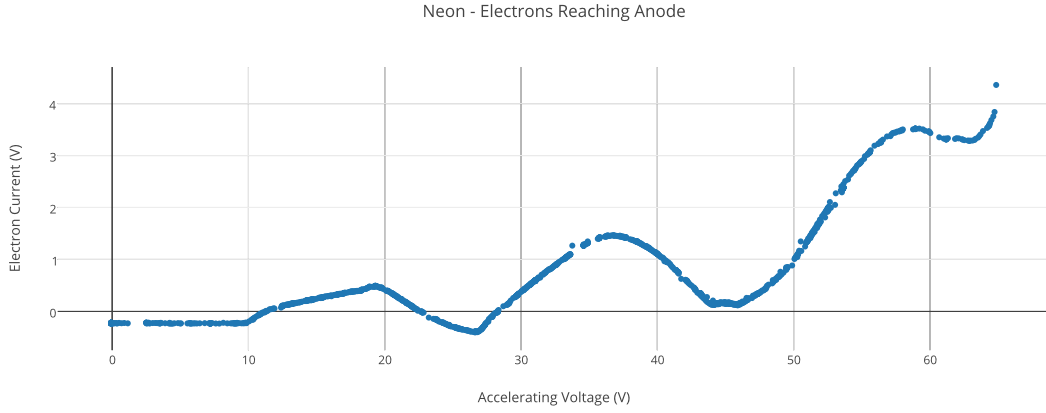
FIG. 1. Neon data with electron current as a voltage plotted against accelerating voltage. The three dips correspond to electrons exciting three neon atoms before reaching the anode. The double minima in the second and third dips is due to multiple excitations with similar energies.
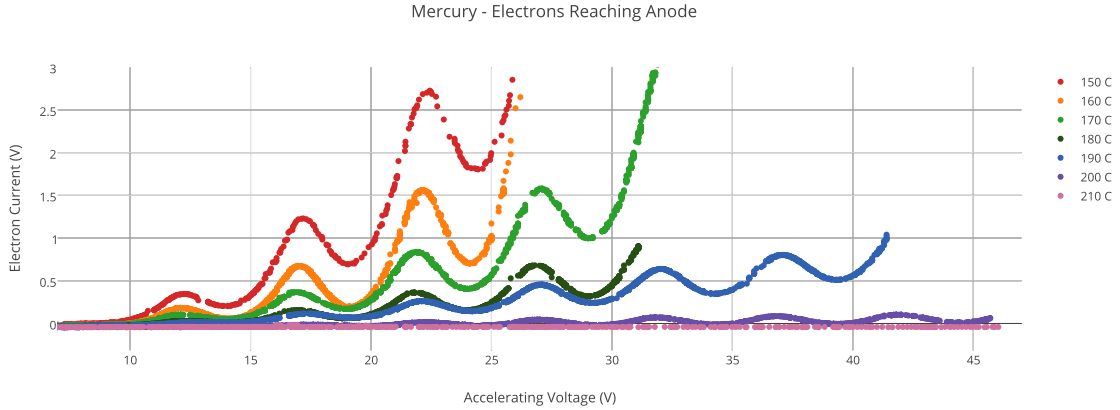

FIG. 2. Mercury data for temperatures ranging from 150°C to 210°C. The range of temperatures shows the optimum temperature to be approximately 190 - 200°C. For higher temperatures, such as 210°C, the dips in electron current are not discernible. For lower temperatures, for example 150°C and 160°C, there were fewer dips before the mercury atoms ionized.

steepest negative slope of the direct output and the places the lock-in data passes through zero correspond to the dips and peaks of the direct output. The lock-in detector was highly sensitive to ionization of the mercury atoms. Our 185°C data already showed significant reduction in the number of dips because of ionization.
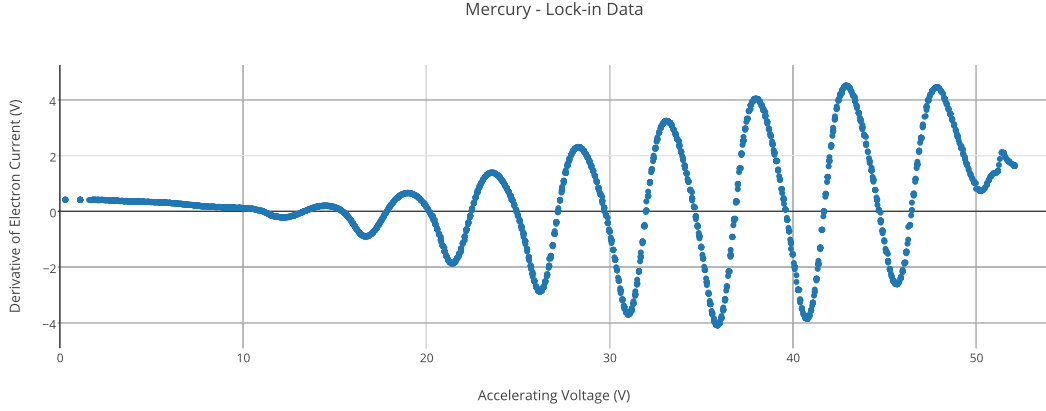
FIG. 3. Sample mercury data using lock-in detector at 205°C. The lock-in output is the derivative of the direct output in Figure **??**. Dips in the voltage correspond to the steepest negative slope of the electron current and the lock-in x intercepts with positive slopes correspond to the dips in the direct output.

IV. ANALYSIS

A. Neon

To determine the energy need to excite the neon atoms we needed to find the spacing of the dips in electron current (which we measured as a voltage.) We visually determined the locations of the dips and the steepest negative slopes with uncertainty for each of our data runs, see sample data in Figure **??**. We then calculated the change in voltage between adjacent dips and steepest slopes. These values can be seen in Table **??**.

We then averaged these values and got $17.9 \pm .3$ eV from the difference between the dips where the uncertainty is the standard deviation of the mean. From the difference between the steepest slopes we got a value of $19.19 \pm .19$ eV. If we then average these two values, we get a final value of $18.5 \pm .4$ eV where uncertainty is again the standard deviation of the mean. This value agrees with our expectations because neon has a large number of transitions with energies ranging from 18.3 eV to 18.9 eV.

TABLE I.			
ΔV Dips	Error ΔV Dips	ΔV Steepest Slope	Error ΔV Steepest Slope
19.0	2.0	19.2	1.5
16.4	2.0	19.5	1.5
18.5	1.3	20.0	2.5
17.6	1.6	18.5	2.5
18.4	1.2	19.5	2.0
17.0	1.5	18.8	2.0
18.5	1.5	18.5	2.5
17.5	2.5	19.5	2.0

B. Mercury

V. DISCUSSION

VI. CONCLUSION

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[1] Gerald Rapior, Klaus Sengstock, and Valery Baev, "New Features of the Franck-Hertz Experiment," Am. J. Phys. **74**, 423–428 (2006).

[2] Adrian C. Melissanos and Jim Napoitano, *Experiments in Modern Physics* 2nd edition (Academic Press, Boston, 2003).

[3] Hyper-Physics: Franck-Hertz Experiment Website <http://hyperphysics.phy-astr.gsu.edu/hbase/frhz.html>