

“Experimental Data Processing”

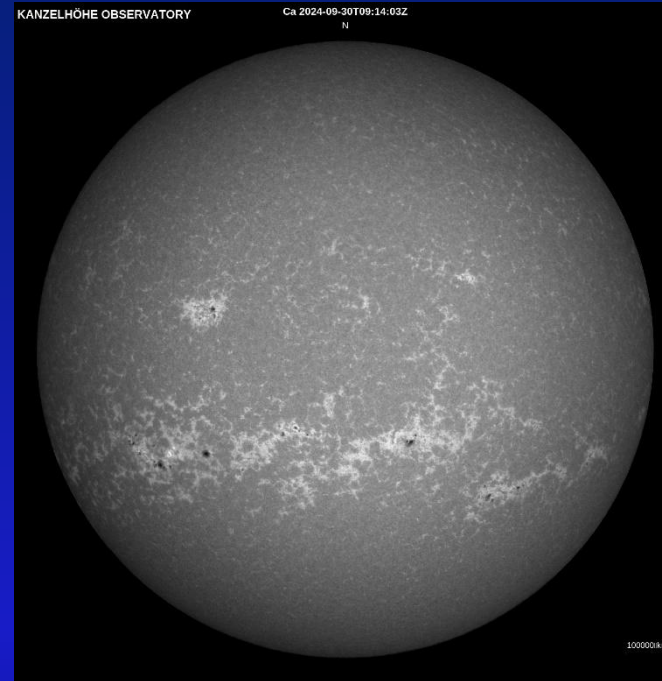
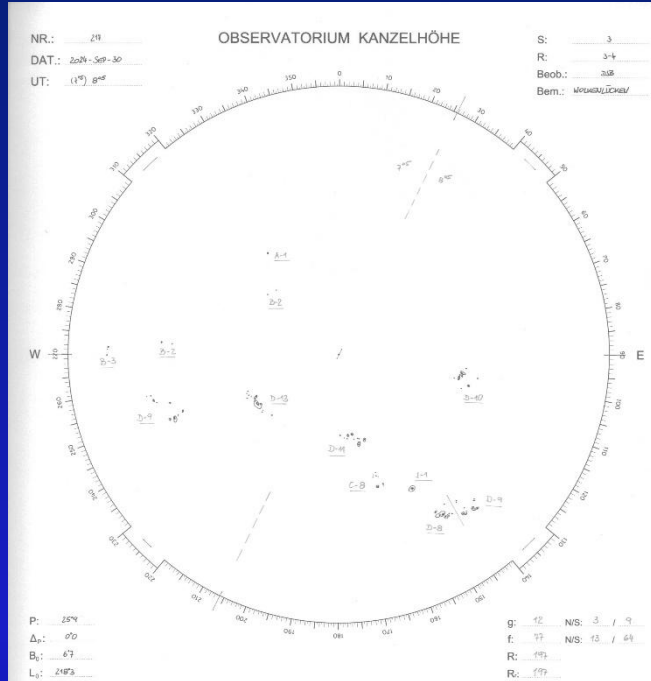
Assignment 1

Reconstructing the past:

Estimating solar radio flux F10.7 from sunspot records.

SUNSPOT NUMBER OBSERVATIONS

30 September 2024



$$R = k(n + 10g)$$

n – number of observed sunspots

g – number of observed sunspot groups

k – coefficient of a telescope

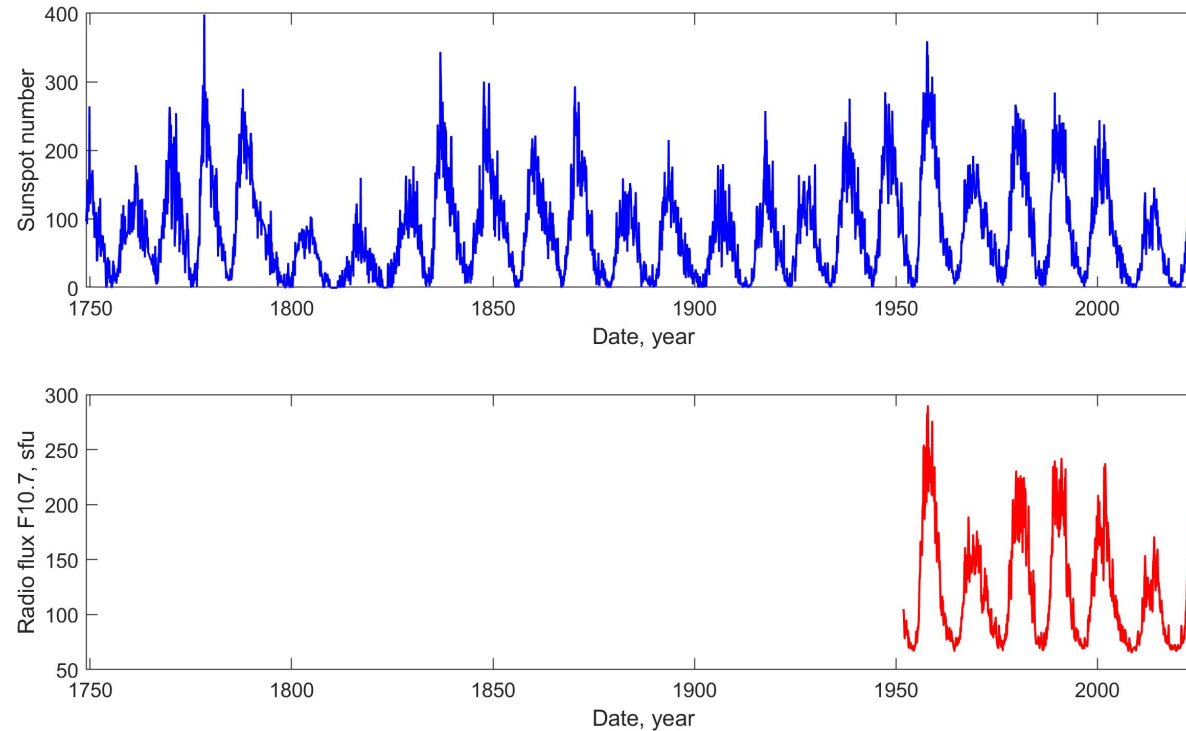
Around 80 cooperating stations around the globe observe sunspot numbers daily

Main indicators of solar activity

Sunspot number



From ground-based
observatories



Solar radio flux at 10.7 cm (in sfu)

$1 \text{ sfu} = 10^{-22} \text{ W m}^{-2} \text{ Hz}^{-1}$



Ottawa and
Pentinton
Observatory, Canada

$$R = k(n + 10g)$$

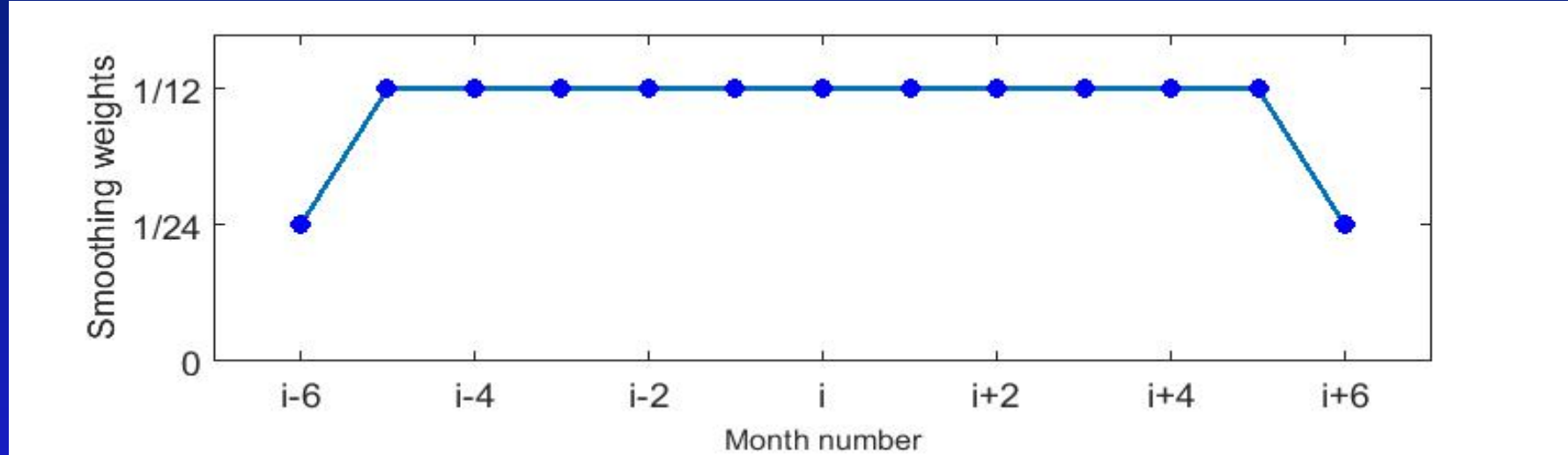
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sunspot groups

k – coefficient of a telescope

A measurement of radio
emission at a wavelength
of 10.7 cm (2800 MHz) from all sources
present on the solar disk

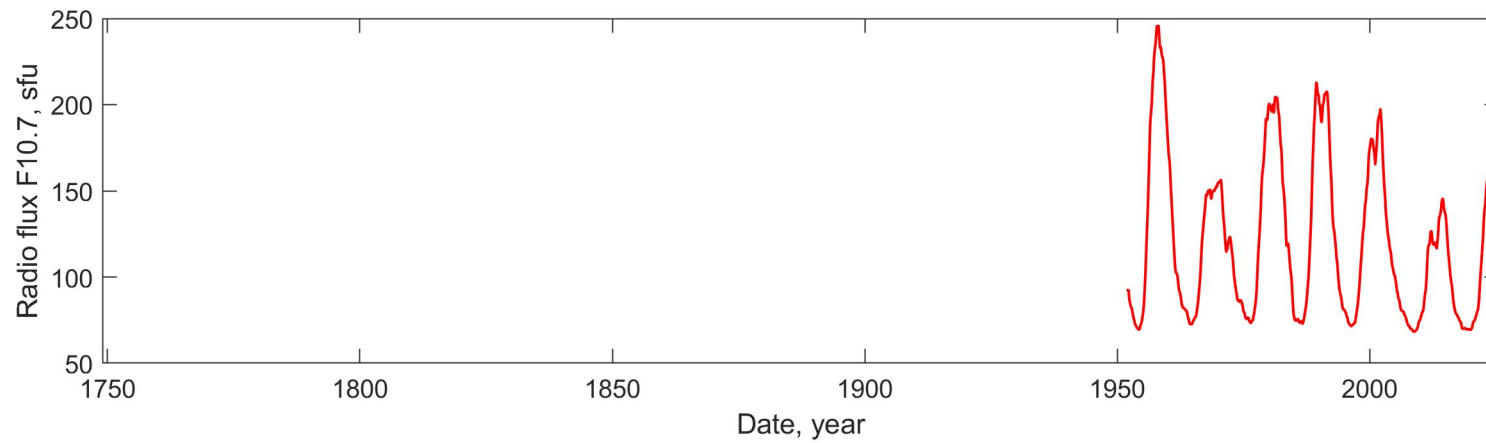
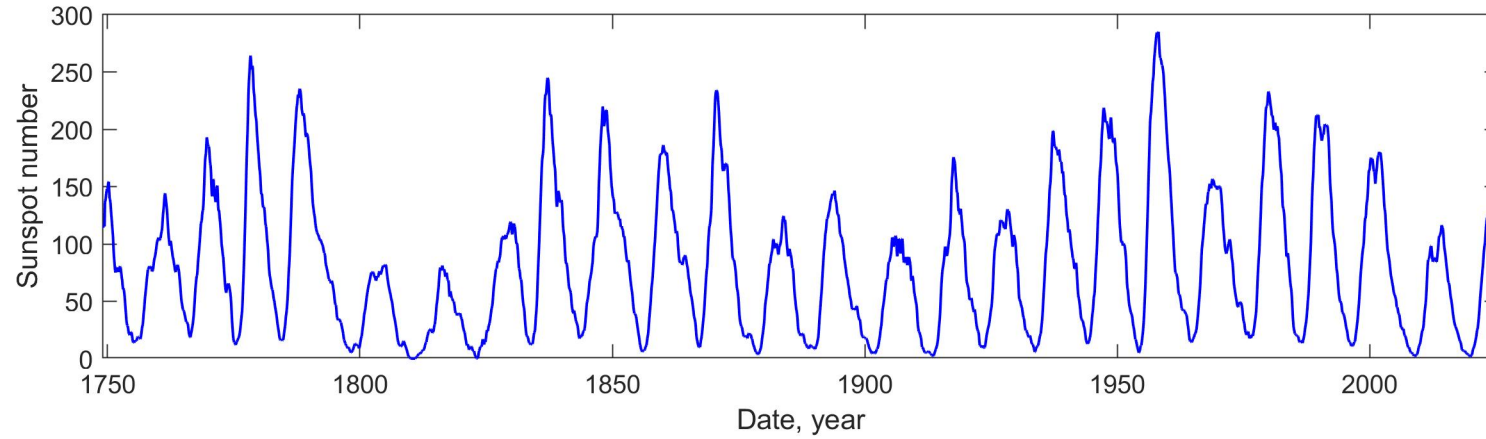
13-month sequent monthly mean sunspot numbers



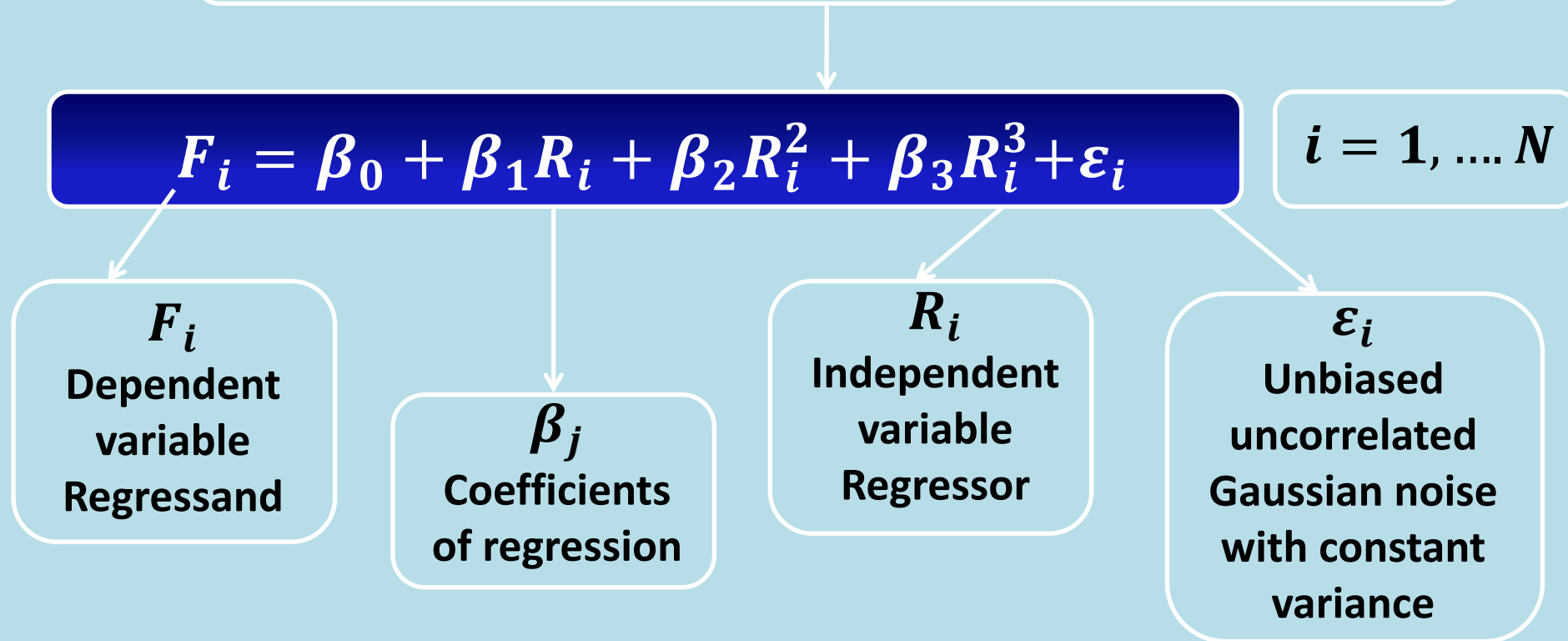
13-month running mean \bar{R}

$$\frac{1}{24}R_{i-6} + \frac{1}{12}(R_{i-5} + R_{i-4} + \dots + R_{i-1} + R_i + R_{i+1} + \dots + R_{i+5}) + \frac{1}{24}R_{i+6}$$

Smoothing: 13-month running mean



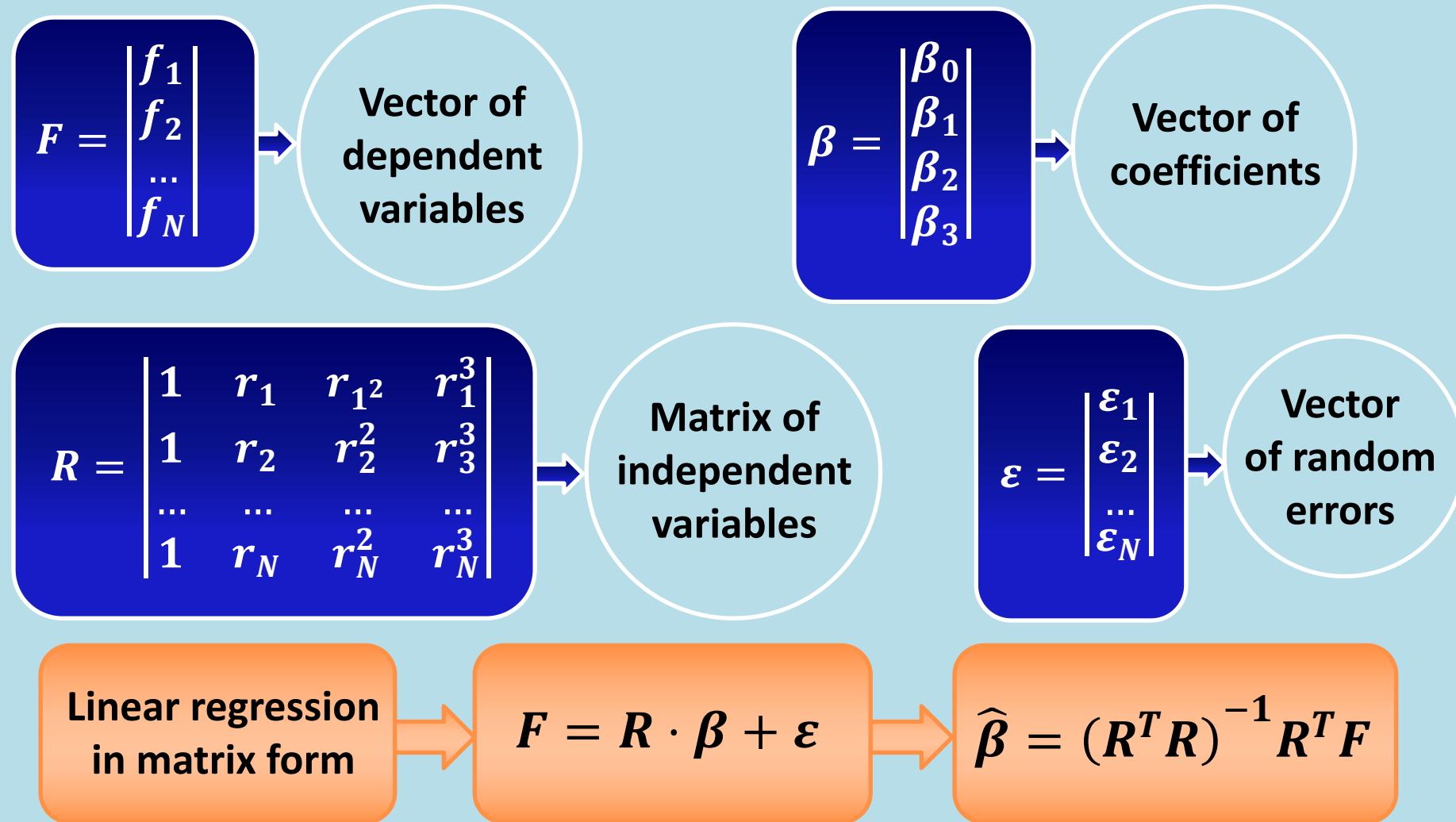
Multi-dimensional linear regression



Coefficients β_j are determined by LSM

$$\sum_{i=1}^N \varepsilon_i^2 \rightarrow \min$$

Multi-dimensional linear regression

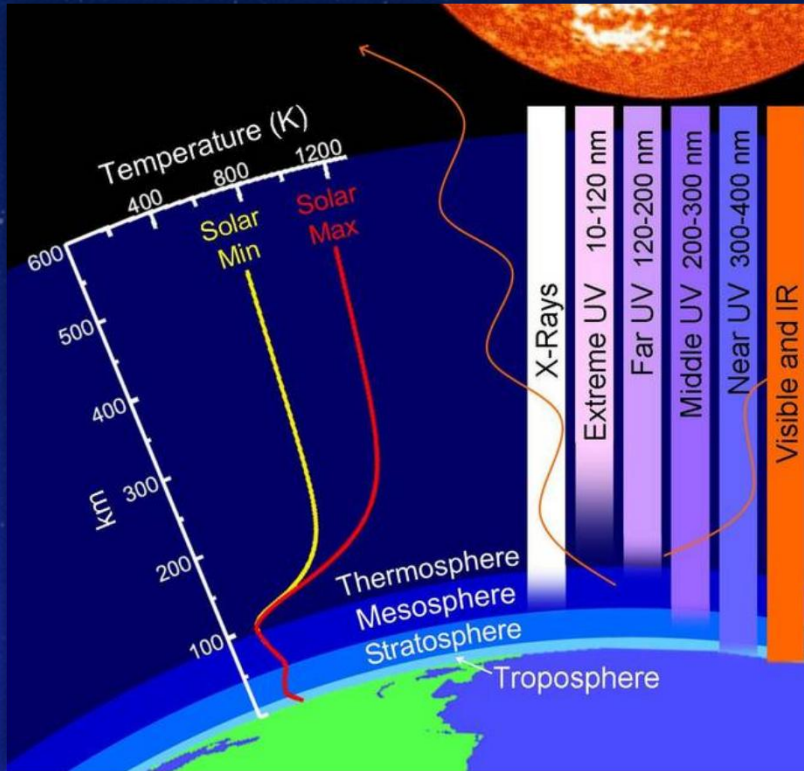


Estimation error of solar radio flux F10.7

Covariance
matrix of
estimation
error



$$\sigma^2 = \frac{1}{N-1} \sum_{i=1}^N (f_i - \hat{f}_i)^2$$



Properties of Earth's atmosphere, NASA

