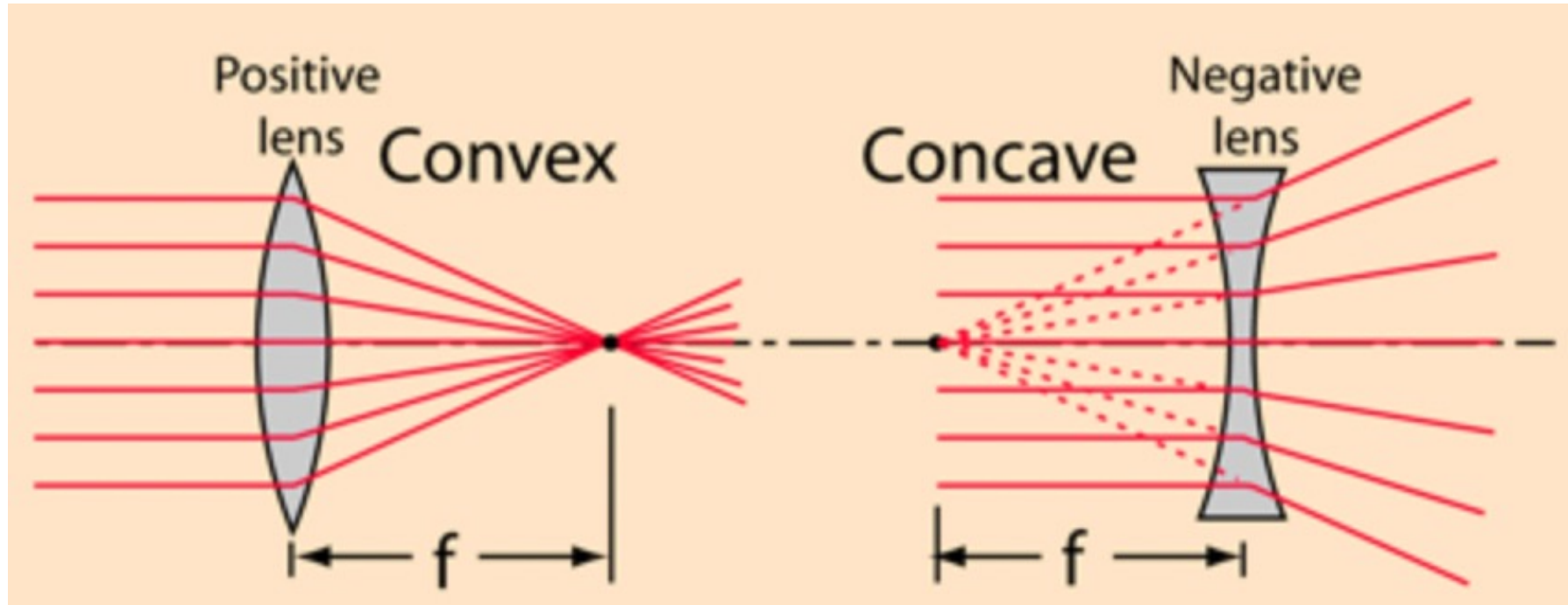


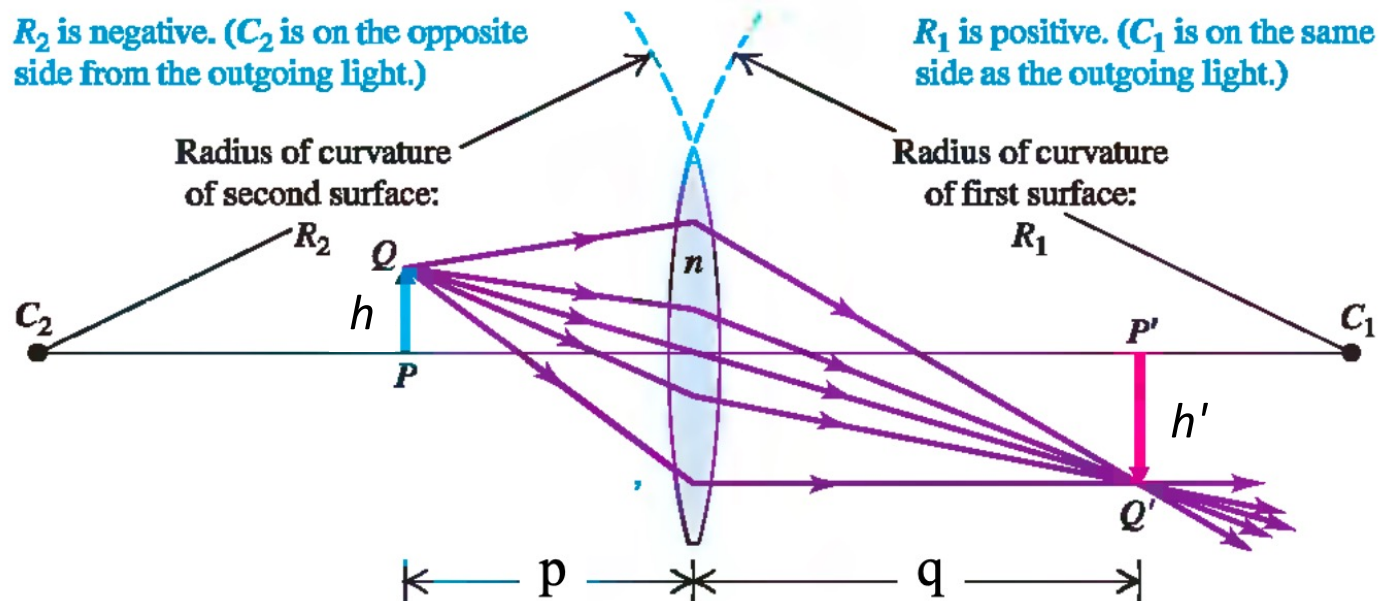
# 1. Thin lens



$$\frac{1}{f} = (n - 1) \left( \frac{1}{R_1} - \frac{1}{R_2} \right) \quad \text{(lens makers' equation)}$$

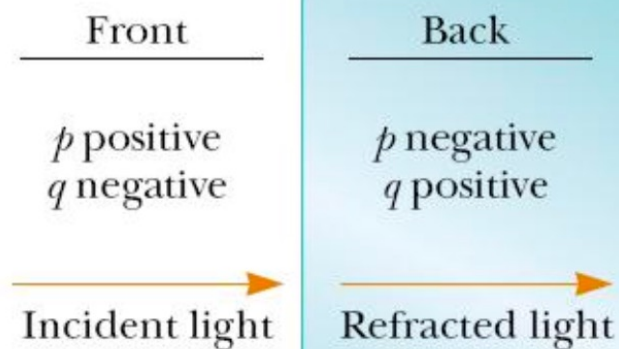
$f$  is **positive** if the lens is **converging**.

$f$  is **negative** if the lens is **diverging**.



## Thin-lens equation :

$$\frac{1}{p} + \frac{1}{q} = \frac{1}{f}$$

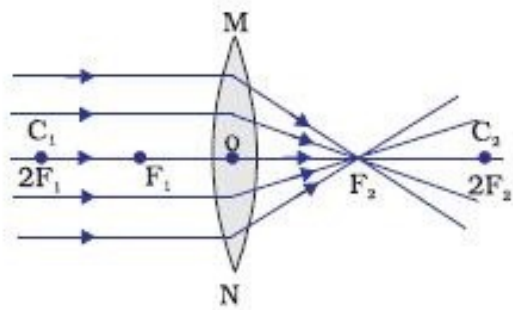


## Lateral magnification :

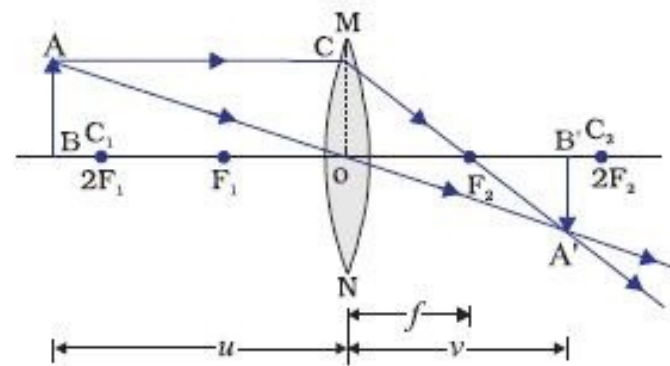
$$M = \frac{h'}{h} = -\frac{q}{p}$$

Real object  $\leftrightarrow p > 0$     Virtual object  $\leftrightarrow p < 0$

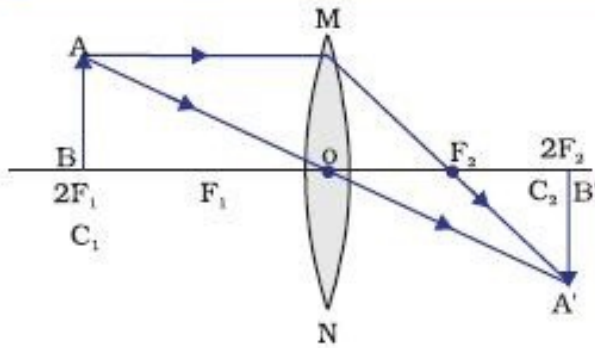
Real image  $\leftrightarrow q > 0$     Virtual image  $\leftrightarrow q < 0$



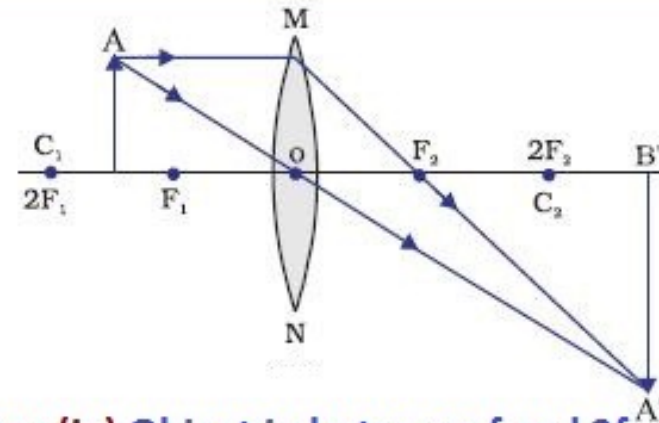
**Case (i) Object at infinity**



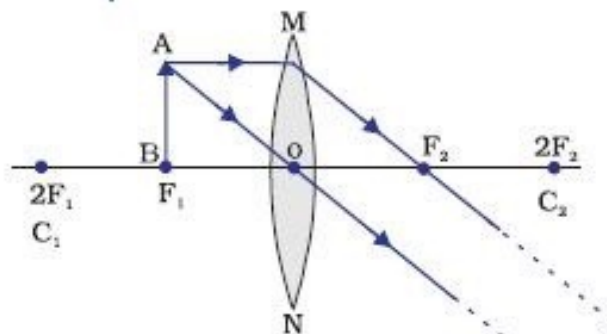
**Case (ii) Object at beyond  $2f$**



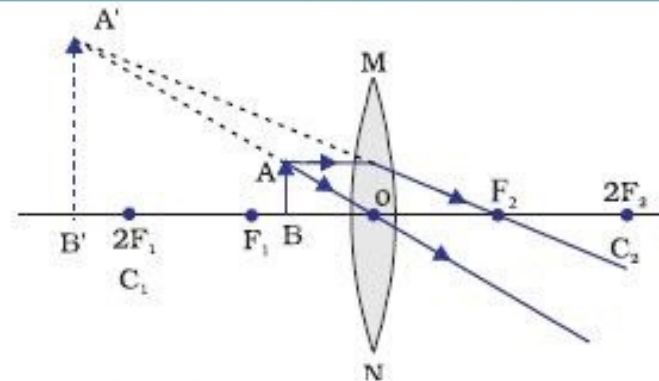
**Case (iii) Object at  $2f$**



**Case (iv) Object in between  $f$  and  $2f$**



**Case (v) Object at  $f$**



**Case (vi) Object distance  $< f$**

## 2. Matter-wave principle

### De Broglie's relationships

We recall that

for a photon ( $E, p$ ) associated to an electromagnetic wave ( $f, \lambda$ ):

$$\begin{array}{l} \boxed{E = h f} \\ \boxed{p = h \frac{1}{\lambda}} \end{array}$$

$\underbrace{\hspace{1.5cm}}_{\text{particle}} \quad \underbrace{\hspace{1.5cm}}_{\text{wave}}$

Planck constant:  
 $h = 6.63 \times 10^{-34} \text{ m}^2 \text{ kg/s}$

### De Broglie's hypothesis:

To a **particle** ( $E, p$ ) is associated a **matter wave**,  
which has a frequency  $f$  and a wavelength  $\lambda$

$$\boxed{f = \frac{E}{h}}$$

$$\boxed{\lambda = \frac{h}{p}}$$

- From  $f = \frac{E}{h}$  and  $\lambda = \frac{h}{p}$

if we put:  $h = 2\pi \hbar \Rightarrow E = 2\pi f \hbar$

$\Rightarrow p = \frac{2\pi}{\lambda} \hbar$

$$\left. \begin{array}{l} E = \hbar \omega \\ \vec{p} = \hbar \vec{K} \end{array} \right\} \text{Planck-Einstein's relationship}$$

$\lambda$  is called **de Broglie wavelength**

$\vec{K}$  is **the angular wave vector** (describing how many oscillations it completes per unit of distance)

$$\vec{K} = 2\pi/\lambda$$



## The Schrödinger's equation

For the case of one-dimensional motion,  
when a particle with the mass  $m$  has a potential energy  $U(x)$   
**Schrödinger's equation** is

$$\frac{d^2 \psi}{dx^2} + \frac{8\pi^2 m}{h^2} [E - U(x)] \psi = 0$$

### Probability density:

Assume that the free particle travels only in the positive direction

Relabel the constant  $A$  as  $\psi_0$  :  $\psi(x) = \psi_0 e^{iKx}$

The probability density is:

$$|\psi|^2 = |\psi_0 e^{iKx}|^2 = (\psi_0)^2 |e^{iKx}|^2$$

Because:

$$|e^{iKx}|^2 = (e^{iKx})(e^{iKx})^* = (e^{iKx})(e^{-iKx}) = e^0 = 1$$

we have:

$$|\psi|^2 = (\psi_0)^2 = \text{const}$$

### 3. A particle in an infinite well

With  $n$  is an integer:  $n = 1; 2; 3; \dots$

the energy can only have the discrete values:  $E_n = \left( \frac{\pi^2 \hbar^2}{2ma^2} \right) n^2$

$$E_n = \left( \frac{\pi^2 (h/2\pi)^2}{2ma^2} \right) n^2 = \frac{h^2}{8ma^2} n^2 \rightarrow \boxed{E_n = \frac{h^2}{8ma^2} n^2}$$

3rd excited  $E_4$

We say that the energy is quantized

these values of energy are called energy levels

$n = 1 \rightarrow$  ground state ( $E_1$ )

$n = 2 \rightarrow$  first excited state ( $E_2$ )

$n = 3 \rightarrow$  second excited state ( $E_3$ )

.

.

.

The integer  $n$  is called the quantum number

2nd excited  $E_3$

1st excited  $E_2$

ground  $E_1$

energy-level diagram

## 4. Spectral emission lines in Hydrogen atom

When the electron jumps down from an energy level  $E_m$  to a lower one  $E_n$ , the hydrogen atom emits a photon of energy:

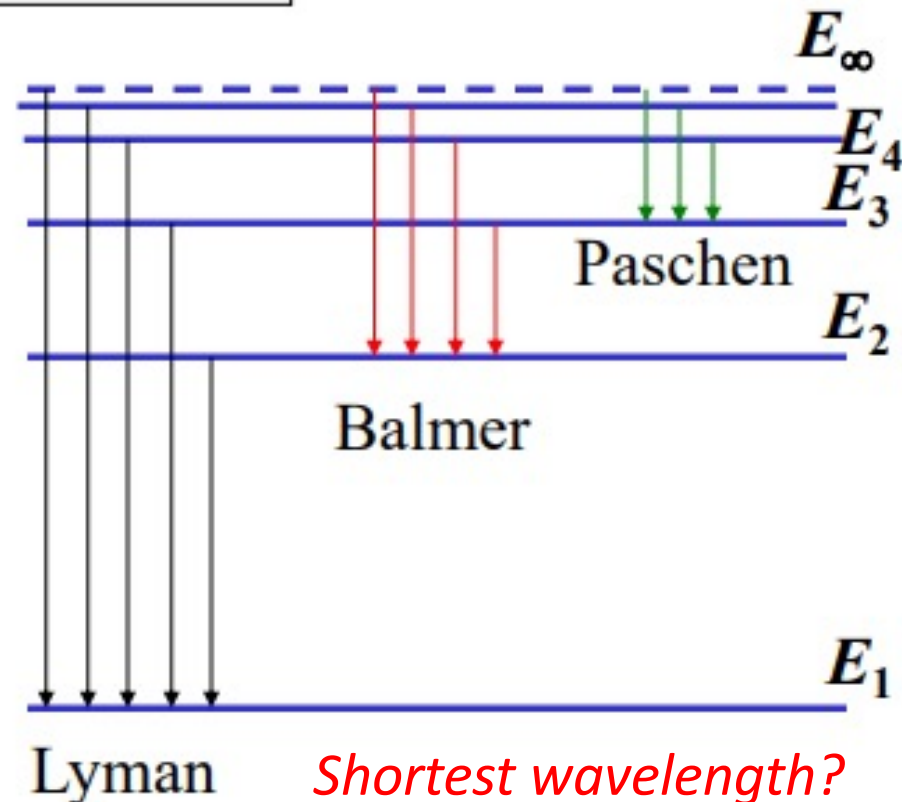
$$\varepsilon = hf_{mn} = \frac{hc}{\lambda_{mn}} = E_m - E_n \quad \text{with} \quad E_n = -\frac{13.6\text{eV}}{n^2}$$

- If  $E_n \equiv E_1$  : Lyman series
- If  $E_n \equiv E_2$  : Balmer series
- If  $E_n \equiv E_3$  : Paschen series
- If  $E_n \equiv E_4$  : Brackett series

$$\frac{1}{\lambda} = R_H \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right) \quad n_i, n_f \text{ are like } m, n \text{ above}$$

$$R_H = \frac{mk^2e^4}{4\pi c\hbar^3} = 1.097 \times 10^7 \text{ m}^{-1}$$

(Rydberg constant)



*Shortest wavelength?*  
*Longest wavelength?*



## 5. Schrodinger equation for hydrogen atom

After solving Schrodinger equation for the radial part we have:

$$E_n = -\left(\frac{me^4}{8\varepsilon_0^2 h^2}\right) \frac{1}{n^2}$$

- *Wave function for the ground state*

The wave function for the ground state of the hydrogen atom:

$$\psi(r) = \frac{1}{a^{3/2}\sqrt{\pi}} e^{-r/a}$$

Where  $a$  is the Bohr radius:  $a = 0.529 \times 10^{-10} m = 52.9 nm$

- *Wave function for the first excited state*

$$\psi(r) = Ce^{-ar/2}(2 - ar)$$

- *The radial probability density*

$$P(r) = \frac{4}{a^3} e^{-2r/a} r^2$$

## 6. Special relativity

$$\Delta t = \frac{\Delta t_0}{\sqrt{1 - u^2 / c^2}}$$

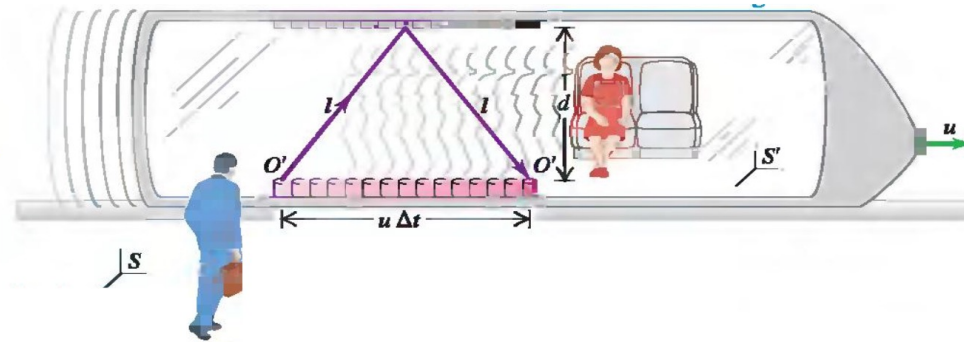
$$\Delta t = \gamma \Delta t_0 ;$$

$$\gamma = \frac{1}{\sqrt{1 - u^2 / c^2}}$$

Lorentz  
constant

$\Delta t > \Delta t_0$  : time dilation

Observers measure any clock to run slow if it moves relative to them



$$L = L_0 \sqrt{1 - \frac{v^2}{c^2}}$$

$$L = \frac{L_0}{\gamma} ;$$

$$\gamma = \frac{1}{\sqrt{1 - u^2 / c^2}}$$

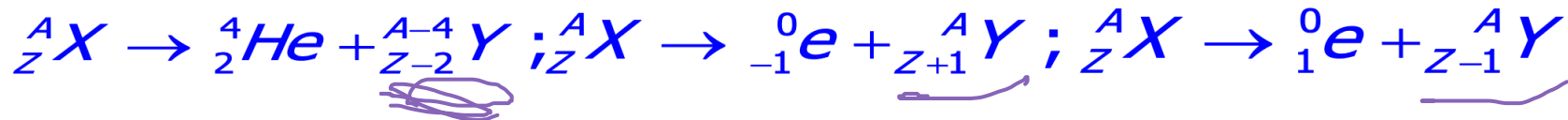
(  $L < L_0$  ) (Length contraction)

The length  $L$  measured in  $S$ , in which the ruler is moving, is shorter than the length  $L_0$  measured in its rest frame  $S'$ .

## 7. Nuclear Radioactivity

- When unstable nuclides decay into different nuclides, they usually emit **alpha** ( $\alpha$ ) or **beta** ( $\beta$ ) particles:

Alpha particle is a  ${}^4\text{He}$  nucleus, a beta-minus particle ( $\beta^-$ ) is an electron, beta-plus particle ( $\beta^+$ ) is a positron (antiparticle of electron)



- $dN(t)/dt$  is called the **decay rate or the activity** of the specimen.

**Activity:** becquerel (Bq) in SI or curie (Ci)

$$1\text{Ci} = 3.70 \times 10^{10} \text{Bq} = 3.70 \times 10^{10} \text{decays / s}$$

$$-\frac{dN(t)}{dt} = \lambda N(t) \quad \boxed{N(t) = N_0 e^{-\lambda t}} \quad \lambda : \text{decay constant}$$

$N(t)$  : the (very large) number of radioactive nuclei in a sample at time  $t$ .

- **The half-life  $T_{1/2}$**  is the time required for the number of radioactive nuclei to decrease to one-half the original number  $N_0$

$$T_{1/2} = \frac{\ln 2}{\lambda} = \frac{0.693}{\lambda}$$

- **The mean lifetime  $T_{\text{mean}}$**  (generally called the lifetime):

$$T_{\text{mean}} = \frac{1}{\lambda} = \frac{T_{1/2}}{\ln 2}$$