

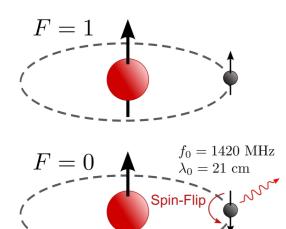
An Update on the REACH experiment and its Bayesian Data Analysis Pipeline

**Dominic Anstey** 

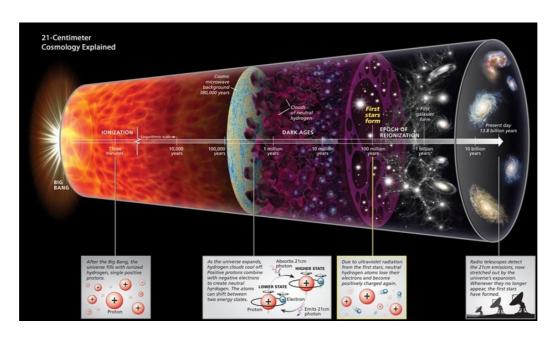
On behalf of the REACH collaboration



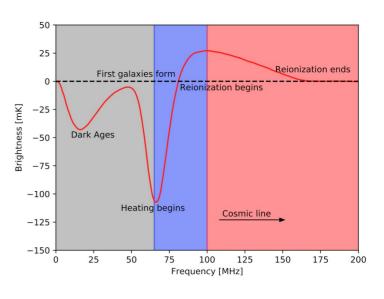
### Global 21cm Experiments



https://en.wikipedia.org/wiki/Hydrogen\_line



Roen Kelly, Discover Magazine



https://www.astro.phy.cam.ac.uk/research/research-projects/reach/overview







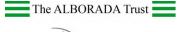




**Kavli Institute** for Cosmology Cambridge<sub>in</sub> Cambridge University

The Kavli Cavendish tellenbosch
Foundation Astrophysics Iniversity

Astrophysics Iniversity South African Radio **Astronomy Observatory** 





eibniz-Institut für strophysik Potsdam



**Tata Institute** for **Fundamental Research University** 









SCIENTA

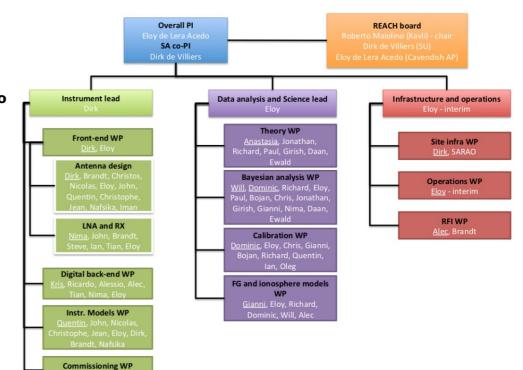
**Imperial** 

College











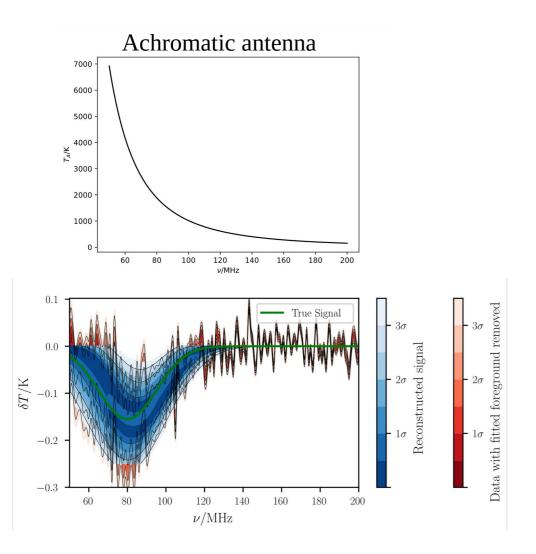
REACH

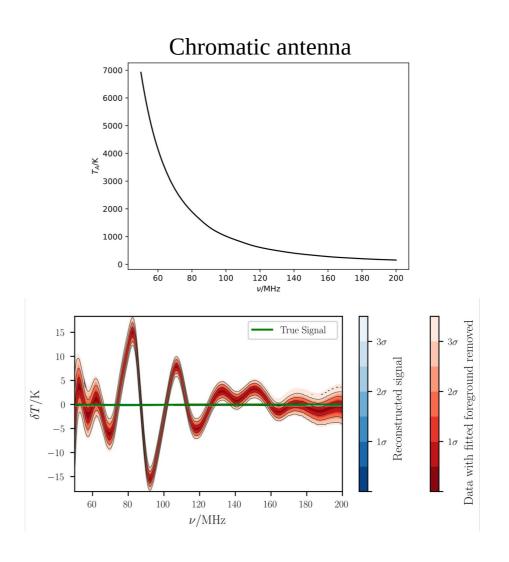
- Designed with Bayesian calibration and data analysis in mind
- Use physics-rooted, physically interpretable models to understand and account for systematics
- Simultaneous observations with two different antennae





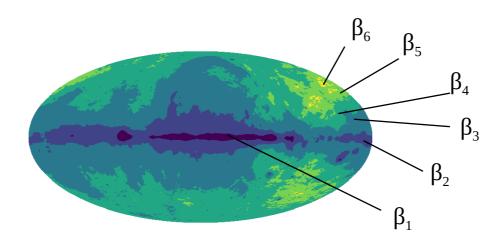
# Radio Foregrounds and Chromaticity







### Summary of REACH Pipeline



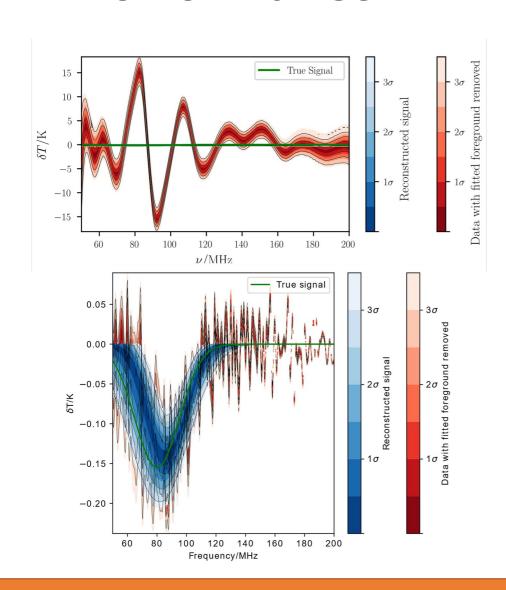
$$T_{\rm F}\left(\nu,\theta_{\rm F}\right) = \frac{1}{4\pi} \int_{0}^{4\pi} D\left(\theta,\phi,\nu\right) \times \int_{t_{\rm start}}^{t_{\rm end}} \sum_{i=1}^{N} M_{i}\left(\theta,\phi\right) \left(T_{230}\left(\theta,\phi\right) - T_{\rm CMB}\right) \left(\frac{\nu}{230}\right)^{-\beta_{i}} dt d\Omega + T_{\rm CMB}$$

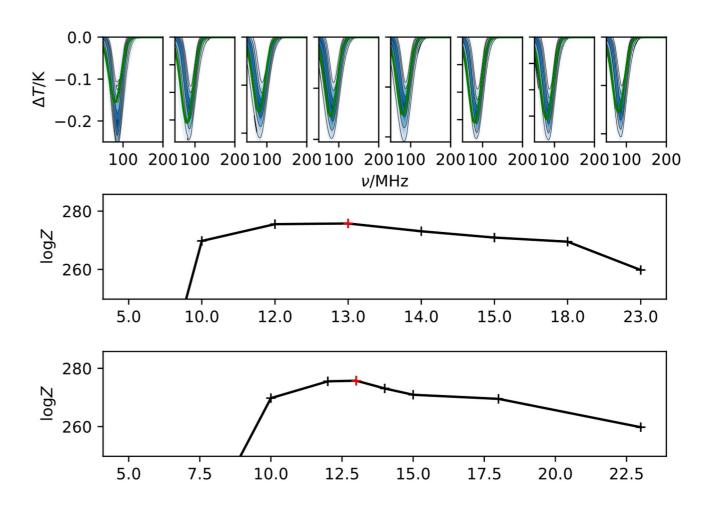
$$\log \mathcal{L} = \sum_{i} -\frac{1}{2} \log \left(2\pi\sigma_{\mathrm{n}}^{2}\right) - \frac{1}{2} \left(\frac{T_{\mathrm{data}}\left(\nu_{i}\right) - \left(T_{\mathrm{F}}\left(\nu_{i}, \theta_{\mathrm{F}}\right) + T_{\mathrm{S}}\left(\nu_{i}, \theta_{\mathrm{S}}\right)\right)}{\sigma_{\mathrm{n}}}\right)^{2}$$

Anstey et al. 2021



### Performance

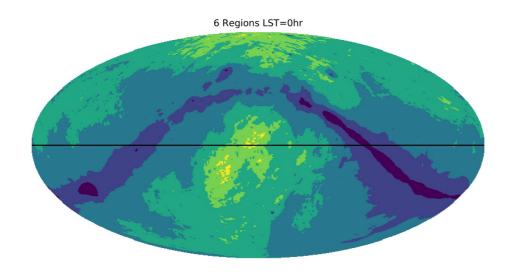


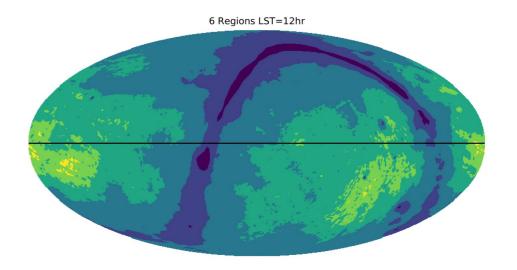




### Time-Separated Fitting

$$\log \mathcal{L} = \sum_{i} -\frac{1}{2} \log \left(2\pi\sigma_{\mathrm{n}}^{2}\right) - \frac{1}{2} \left(\frac{\frac{1}{n_{j}} \sum_{j} \left[T_{\mathrm{data}}\left(\nu_{i}, t_{j}\right)\right] - \left(\frac{1}{n_{j}} \sum_{j} \left[T_{\mathrm{F}}\left(\nu_{i}, t_{j}, \theta_{\mathrm{F}}\right)\right] + T_{\mathrm{S}}\left(\nu_{i}, \theta_{\mathrm{S}}\right)\right)}{\sigma_{\mathrm{n}}}\right)^{2}$$



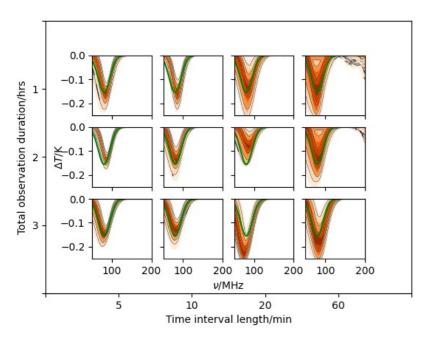


$$\log \mathcal{L} = \sum_{i} \sum_{j} -\frac{1}{2} \log \left(2\pi\sigma_{\mathrm{n}}^{2}\right) = \frac{1}{2} \left(\frac{T_{\mathrm{data}}\left(\nu_{i}, t_{j}\right) - \left(T_{\mathrm{F}}\left(\nu_{i}, t_{j}, \theta_{\mathrm{F}}\right) + T_{\mathrm{S}}\left(\nu_{i}, \theta_{\mathrm{S}}\right)\right)}{\sigma_{\mathrm{n}}}\right)^{2}$$

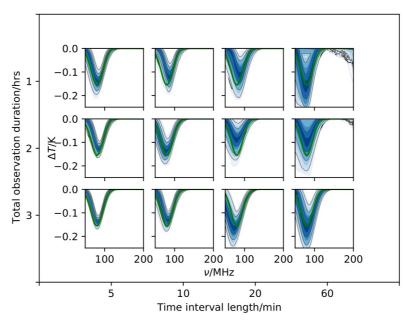


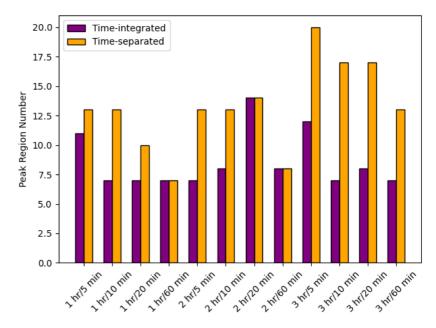
### Time-Separated Fitting Results

#### Time-averaged



#### Time-separated

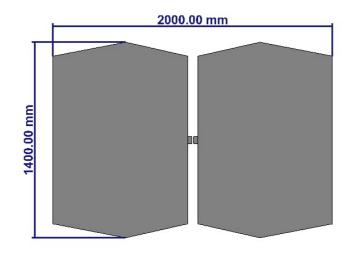


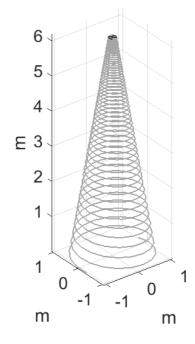




### Multiple Antennae

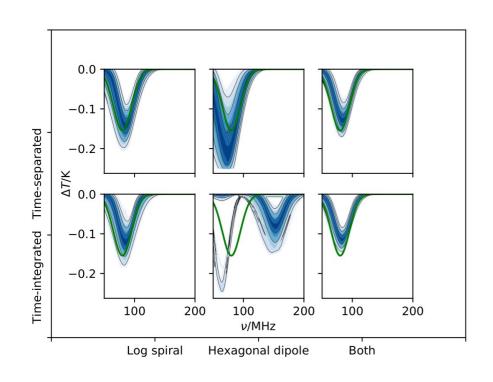
$$\log \mathcal{L} = \sum_{i} \sum_{j} \sum_{k} -\frac{1}{2} \log \left(2\pi\sigma_{\mathrm{n}}^{2}\right) - \frac{1}{2} \left(\frac{T_{\mathrm{data}\,k}\left(\nu_{i},t_{j}\right) - \left(T_{\mathrm{F}\,k}\left(\nu_{i},t_{j},\theta_{\mathrm{F}}\right) + T_{\mathrm{S}}\left(\nu_{i},\theta_{\mathrm{S}}\right)\right)}{\sigma_{\mathrm{n}}}\right)^{2}$$

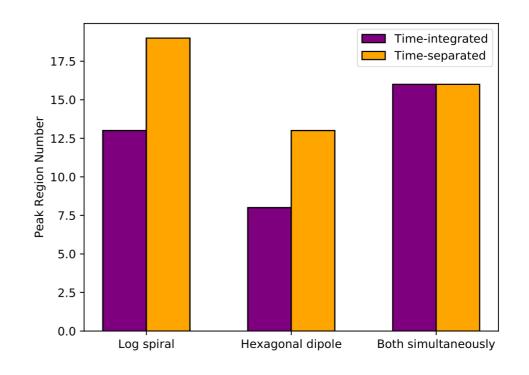






### Multiple Antennae Results







## Antenna Beam Modelling

The antenna beam pattern may not be known exactly in practice:

- Soil permittivity
- Horizon effects
- Imperfections in construction
- Etc.

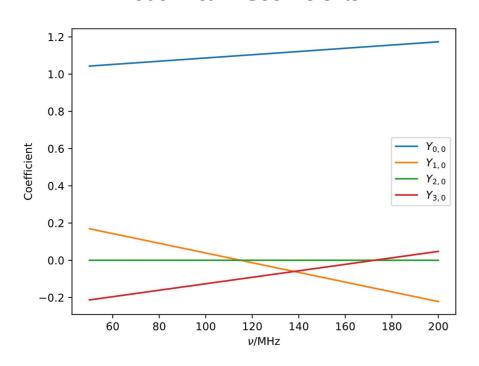
$$D\left(\Omega, \nu, \bar{\theta}\right) = \sum_{j=1}^{M} \theta_j X_j\left(\Omega, \nu\right)$$

$$D\left(\Omega, \nu, \bar{\theta}\right) = \sum_{j=1}^{M} \Gamma\left(\nu, \bar{\theta}_{j}\right) Y_{j}\left(\Omega\right)$$

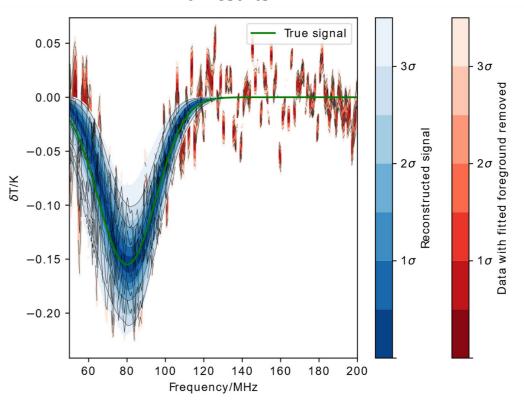


## Fitting for Antenna Beam

#### Model Beam Coefficients



#### Fit Results





- REACH is a global 21 cm experiment with a focus on physicallyinterpretable models, particularly of systematics
- The REACH data analysis pipeline allows modelling of foregrounds and chromaticity based on spectral index
- Using a constant physical property as the parameter allows data from multiple observing times and antennae to be fit simultaneously, giving improvements in signal and particularly foreground reconstruction
- Ongoing work is underway to also model antenna beams











Antenna EM simulations provided by John Cumner and Quentin Gueuning

Plots produced using fgivenx tool: Handley, 2018

