

# Astronomical Coordinates & Telescopes

James Graham

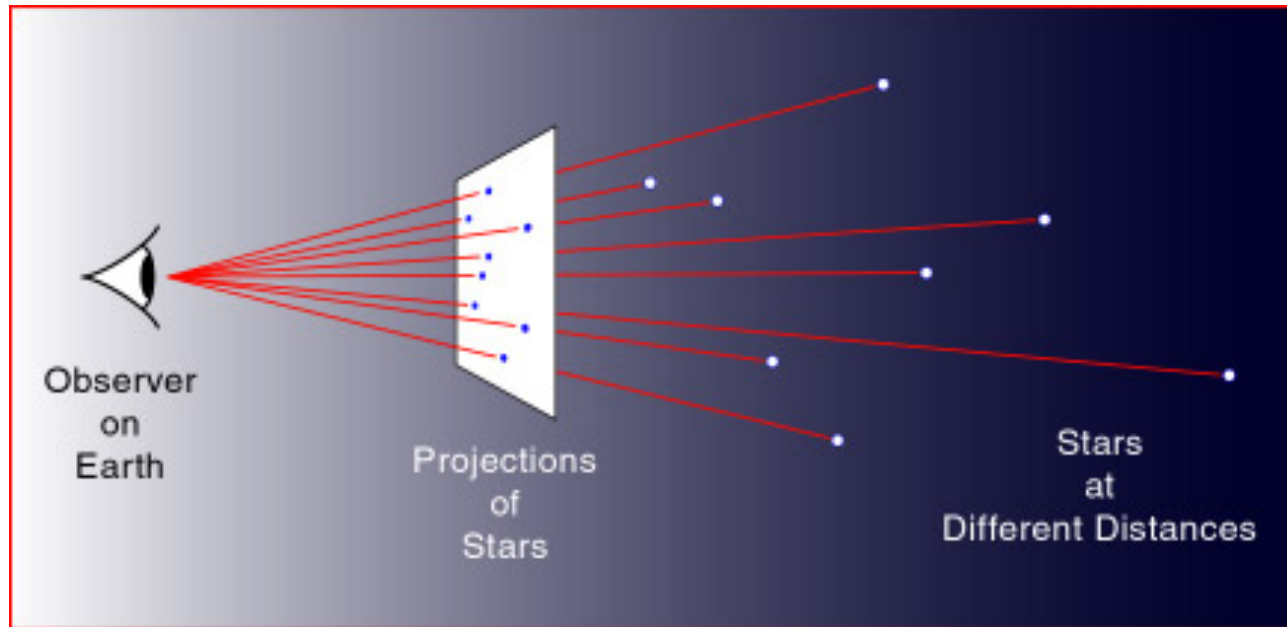
2014/10/14

# Observing Basics

- Astronomical coordinates
- The telescope
- The infrared camera

# Astronomical Coordinates

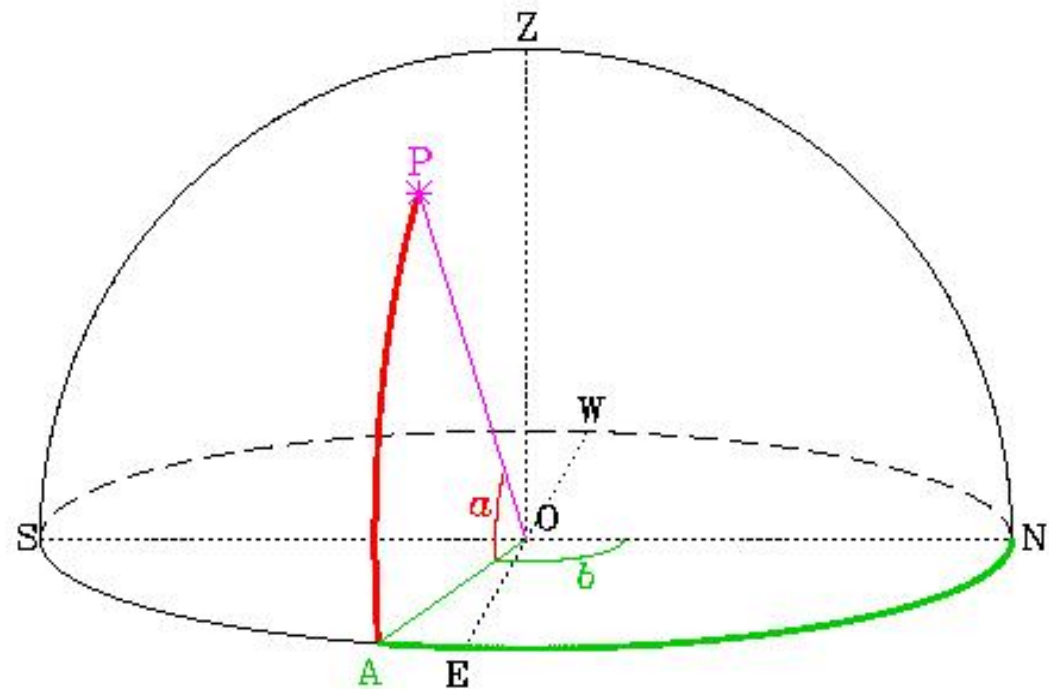
- Celestial sphere
  - Stars and other astronomical objects occupy 3-d space
  - As stars very distant, it is convenient to describe their location as the projection of their position onto a sphere centered at the observer



# Observer's Coordinates

Coordinates of star P  
relative to observer O  
are

- **Altitude** ( $a$ ) is the angle AOP
  - Altitude is the angle above the horizon
- **Azimuth** ( $b$ ) is the angle NOA
  - Degrees E from N
- Observer's coordinates (horizon coordinates) for stars change!

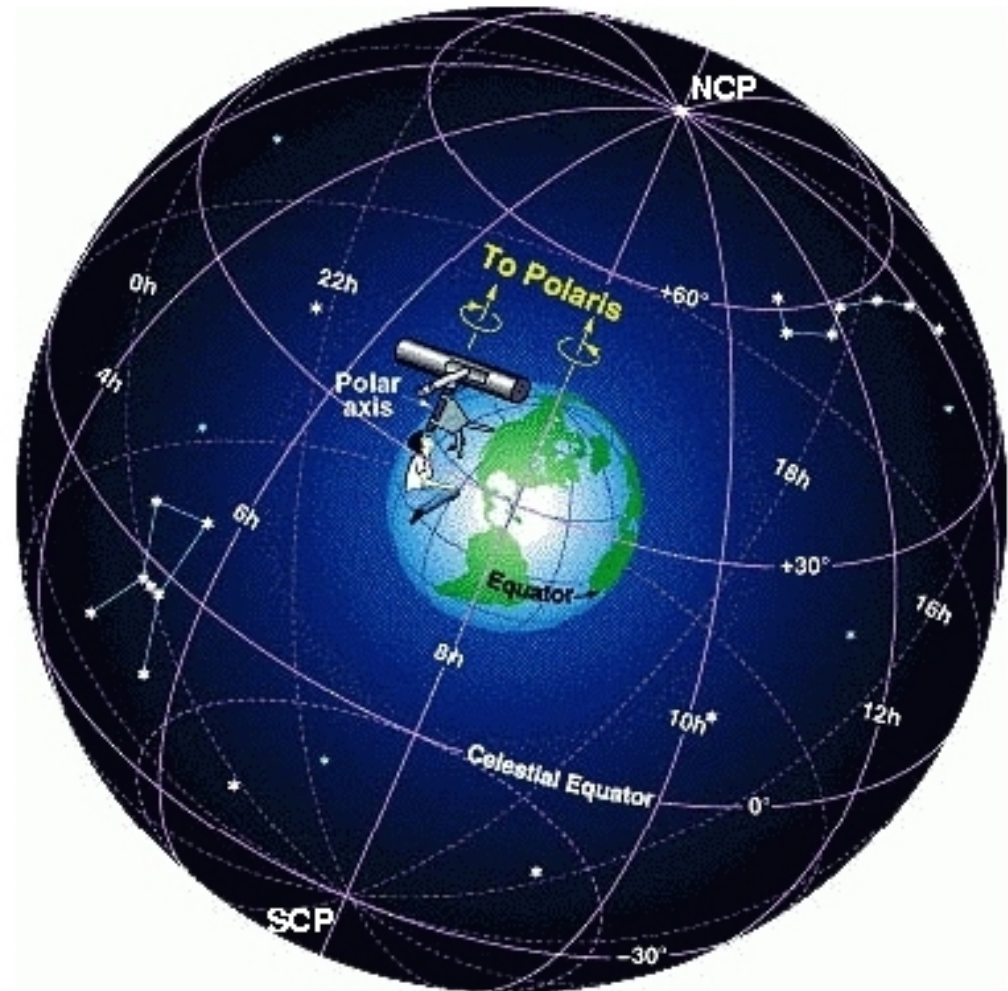


# Sunrise, Sunset & Star Trails



# Celestial Sphere

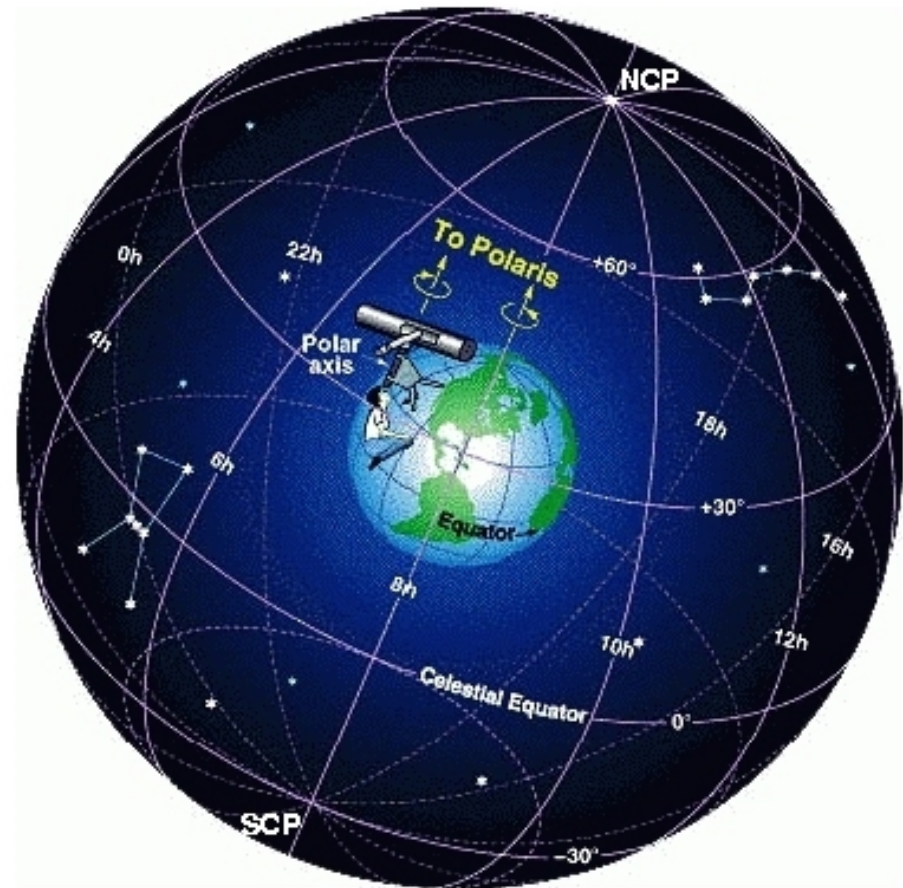
- Coordinates on the celestial sphere are analogous to latitude & longitude
  - *Declination*
  - *Right ascension*





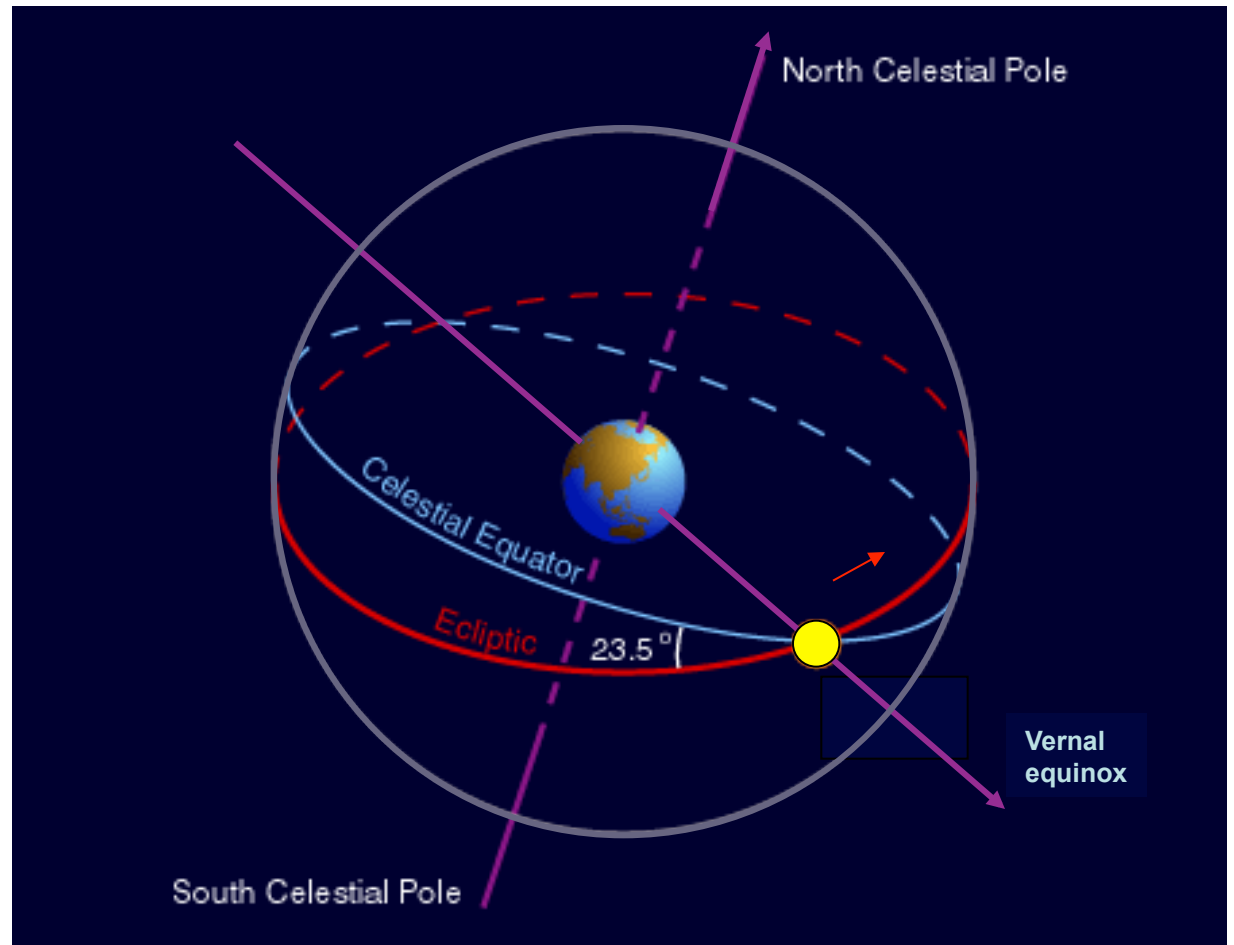
# Celestial Poles & Equator

- N & S *celestial poles* are defined by the projection of the earth's spin axis onto the celestial sphere
- *Celestial equator* is the projection of the earth's equator onto the celestial sphere
  - Objects on the celestial equator have a declination of zero degrees
  - The N & S poles are declination  $\pm 90$  degrees

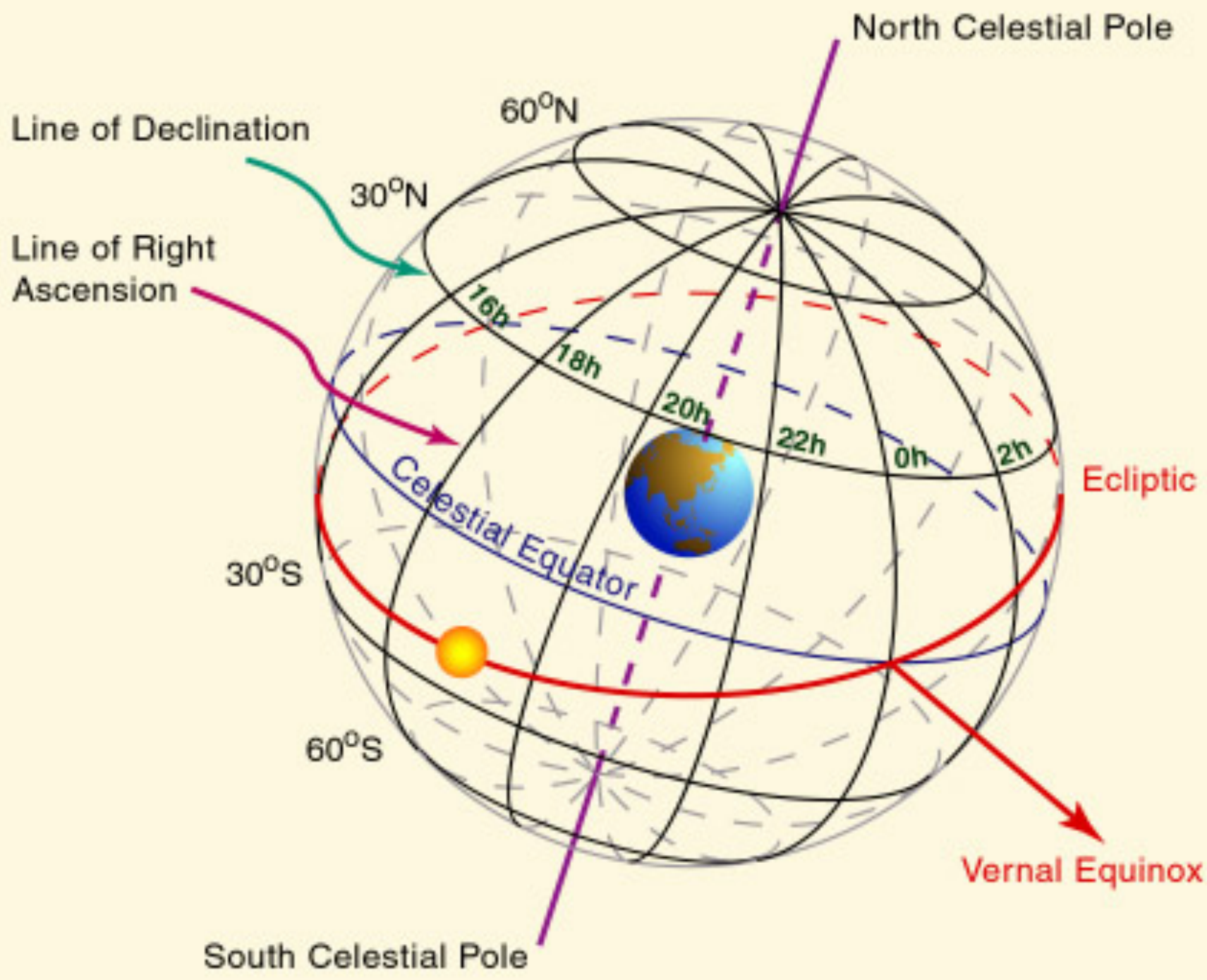


# Defining Celestial Zero Points

- The spin axis of the earth provides a natural definition of one direction
- The location of the Sun, when it crosses the celestial equator in the spring defines the other
  - Defines  $RA = 0$

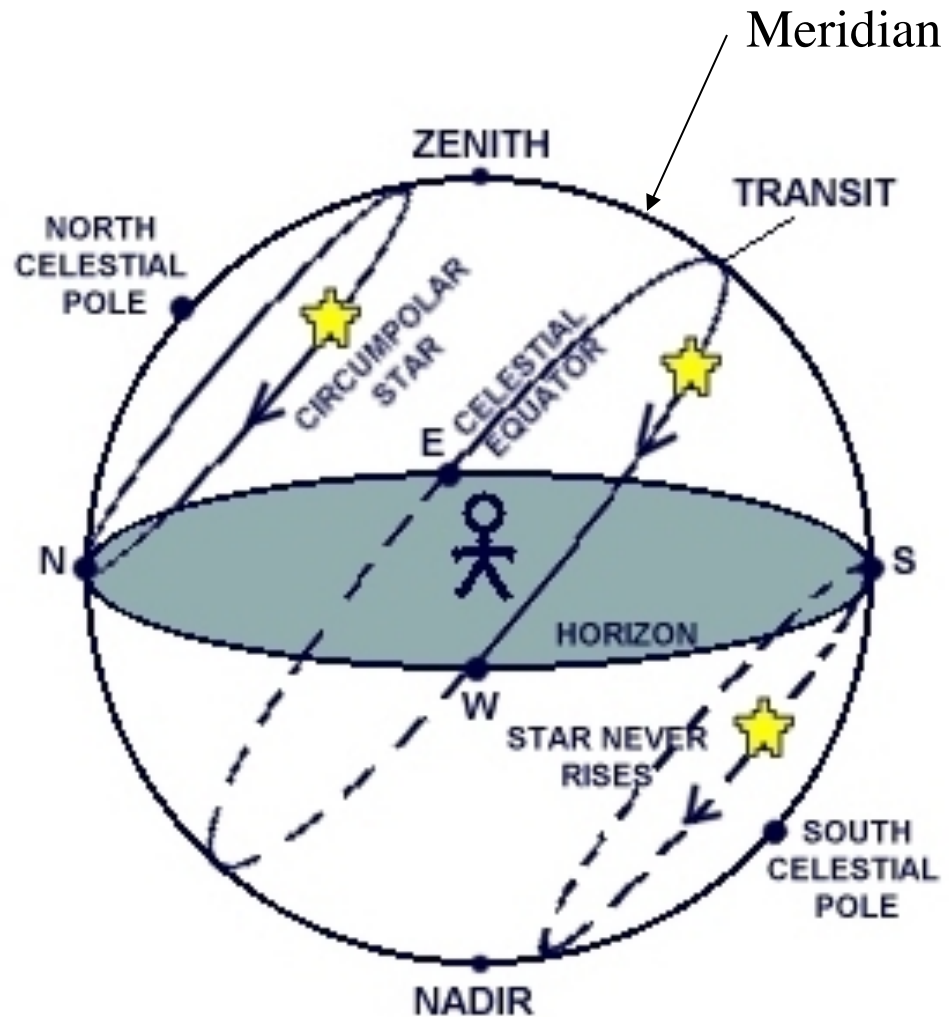






# Rising & Setting

- The observer's *meridian* is the great circle drawn through the N pole, which passes directly overhead
- An star *transits* when it crosses the meridian
- The elevation of a star is greatest when it transits



# Sidereal Time

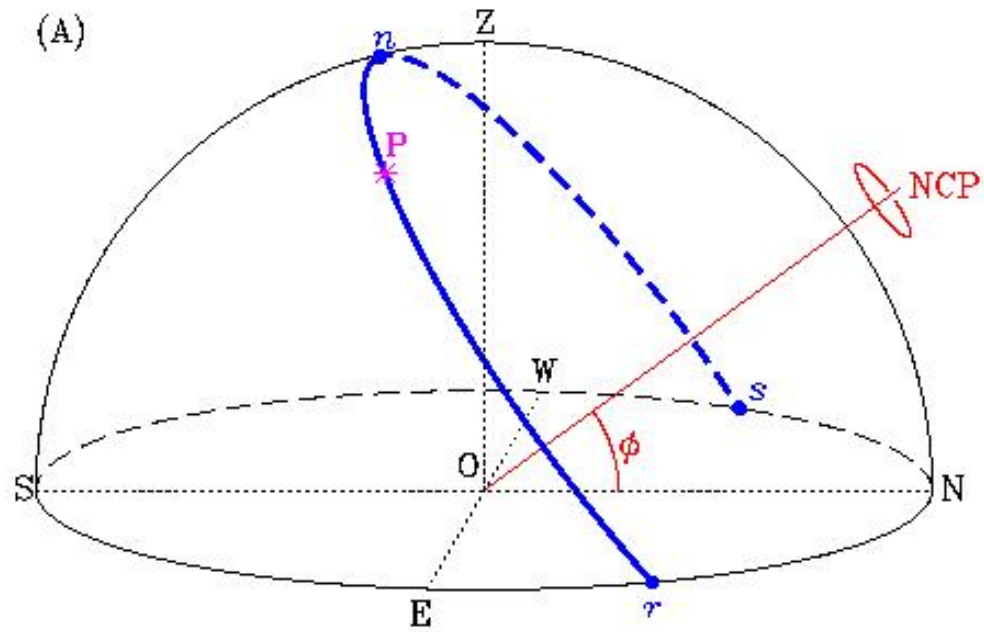
- The *sidereal time* is measured by the rotation of the Earth, with respect to the stars (rather than relative to the Sun)
  - Local sidereal time is the right ascension of a star on the observer's meridian
  - One sidereal day corresponds to the time taken for the earth to rotate once with respect to the stars and lasts approximately 23 h 56 min

# Hour Angle and Right Ascension

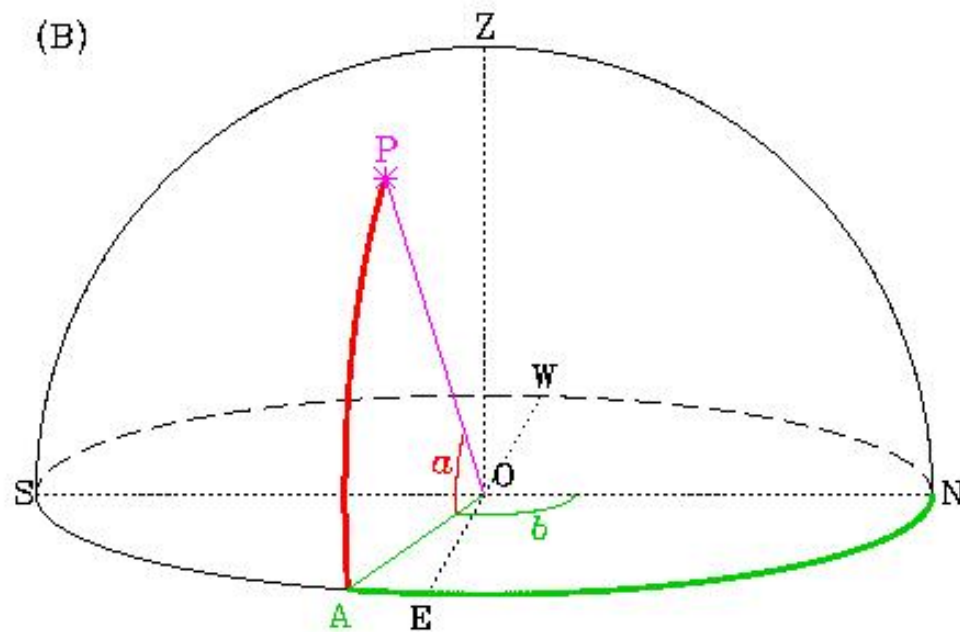
- *Hour angle* (HA) is the angle between an observer's meridian and the hour circle on which some celestial body lies
- Expressed in hours, minutes & seconds, HA gives the time elapsed since a celestial body's last transit ( $HA > 0$ ), or the time unit the next transit ( $HA < 0$ )
- Hence:

$$HA = LST - RA$$

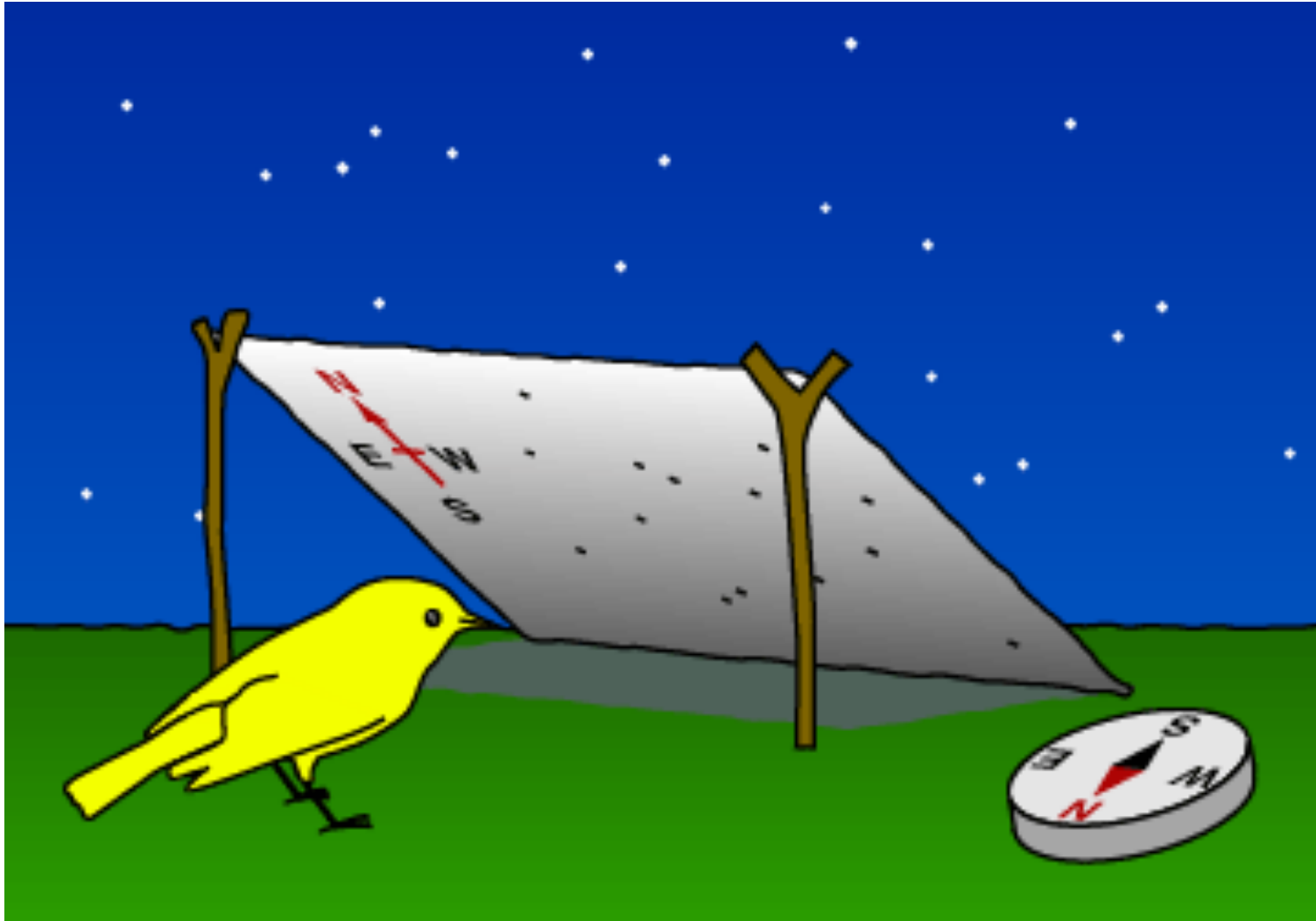
## Celestial coordinates



## Observer's coordinates



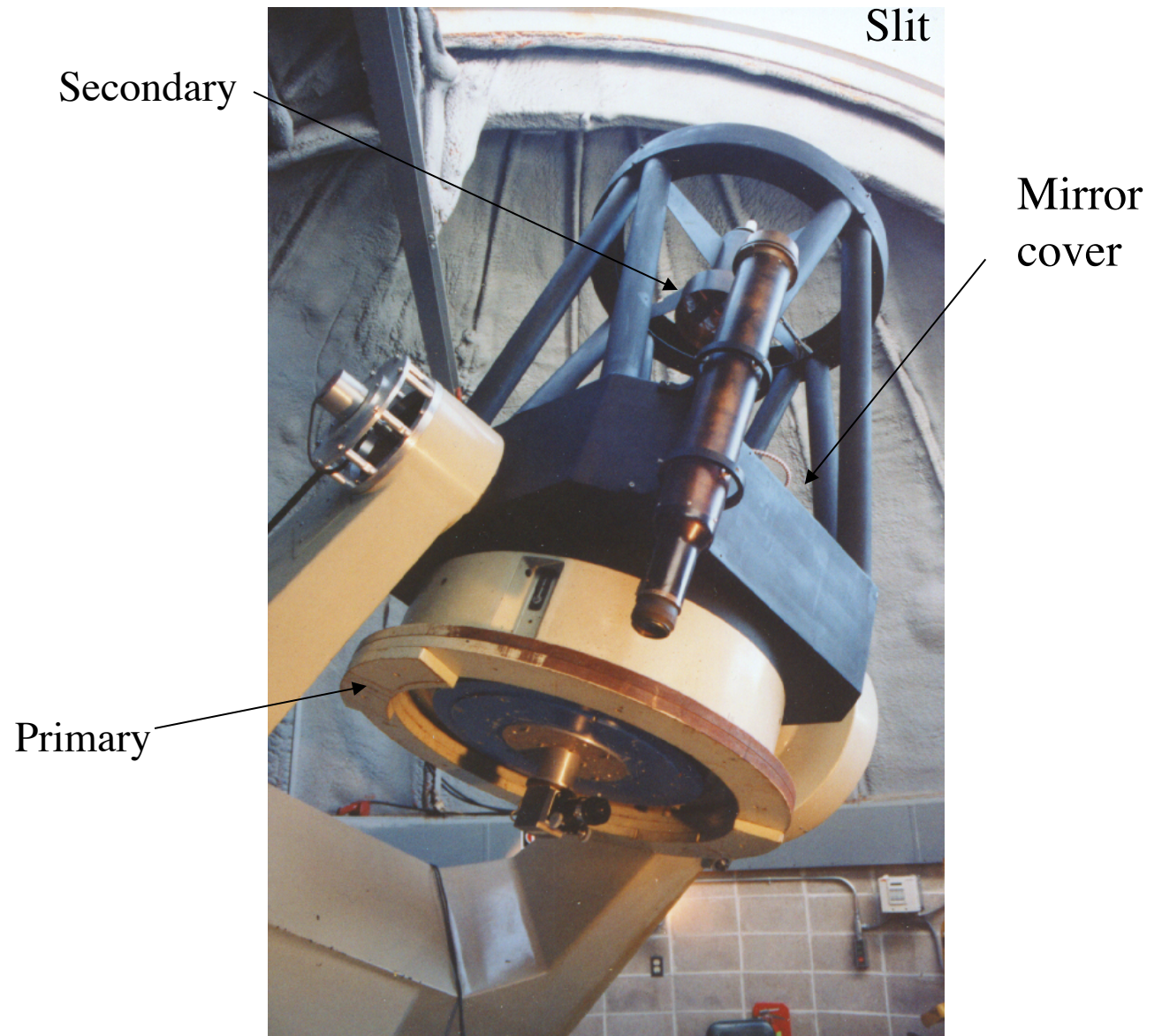
# Why is E on the Left?





# The 30-inch Telescope

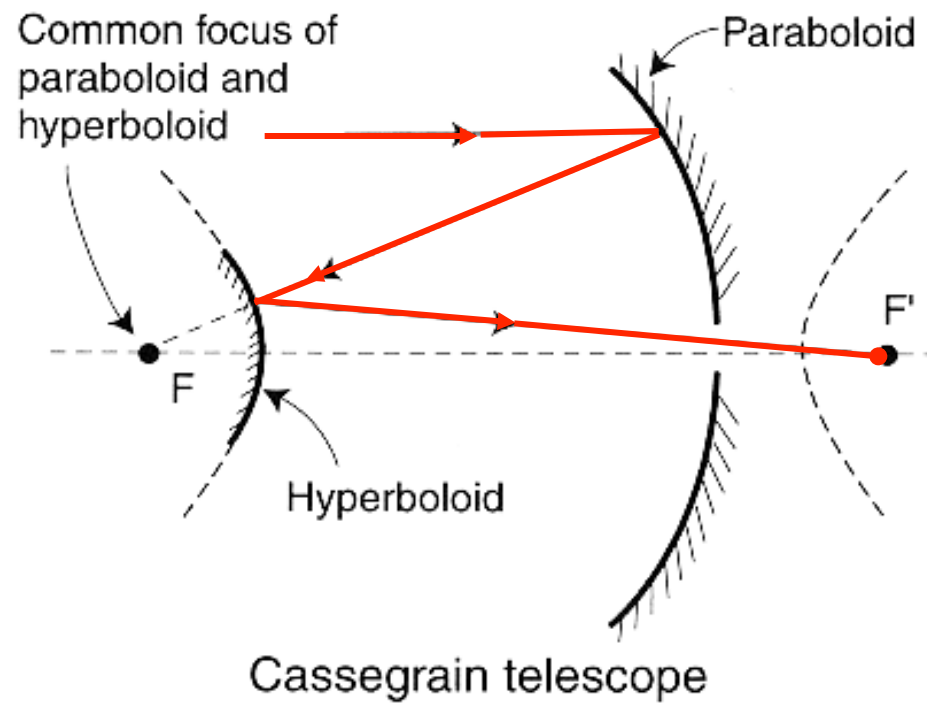
- Cassegrain telescope
  - Two mirror telescope
  - Primary & secondary mirrors



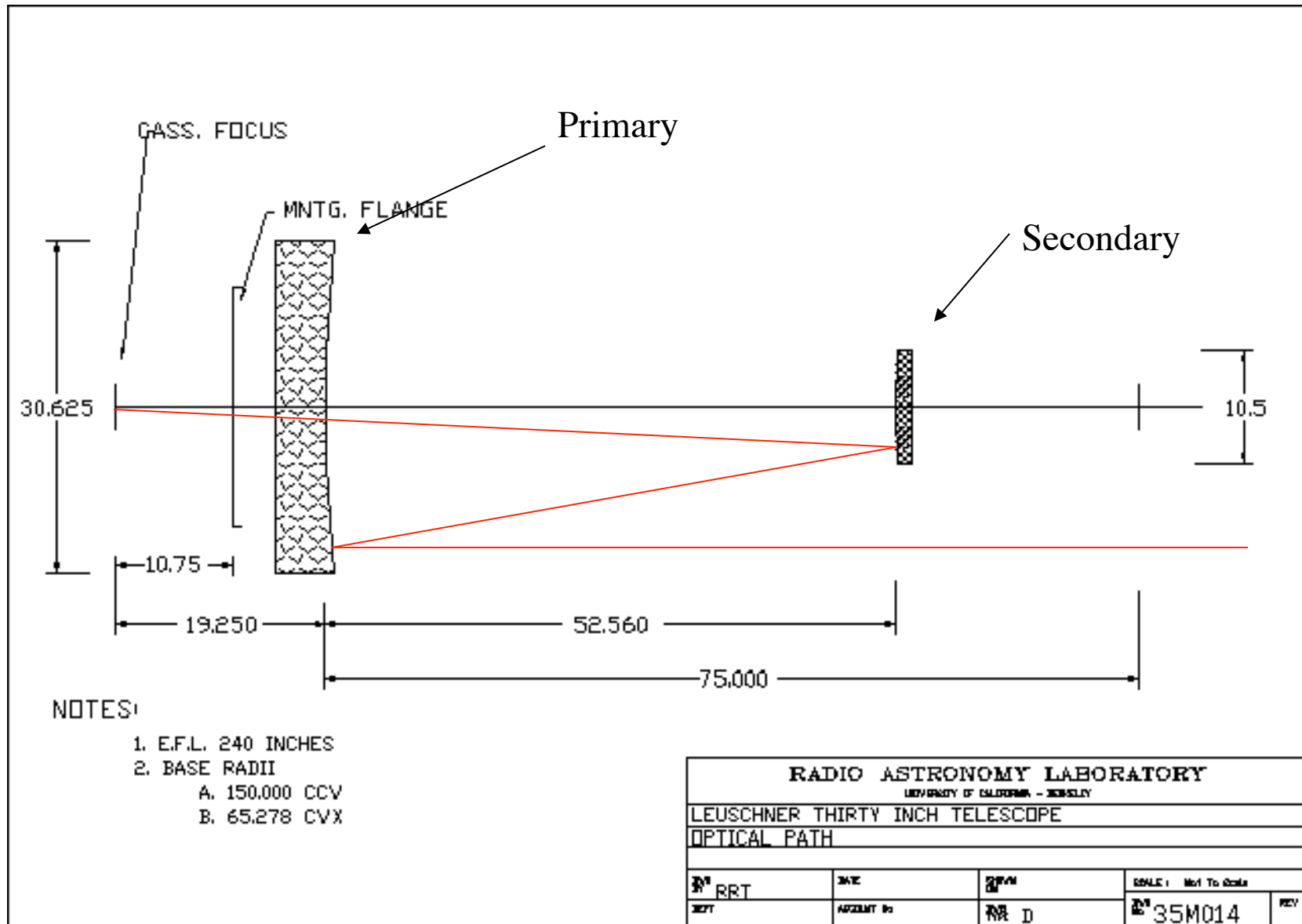


# Two Mirror Telescope

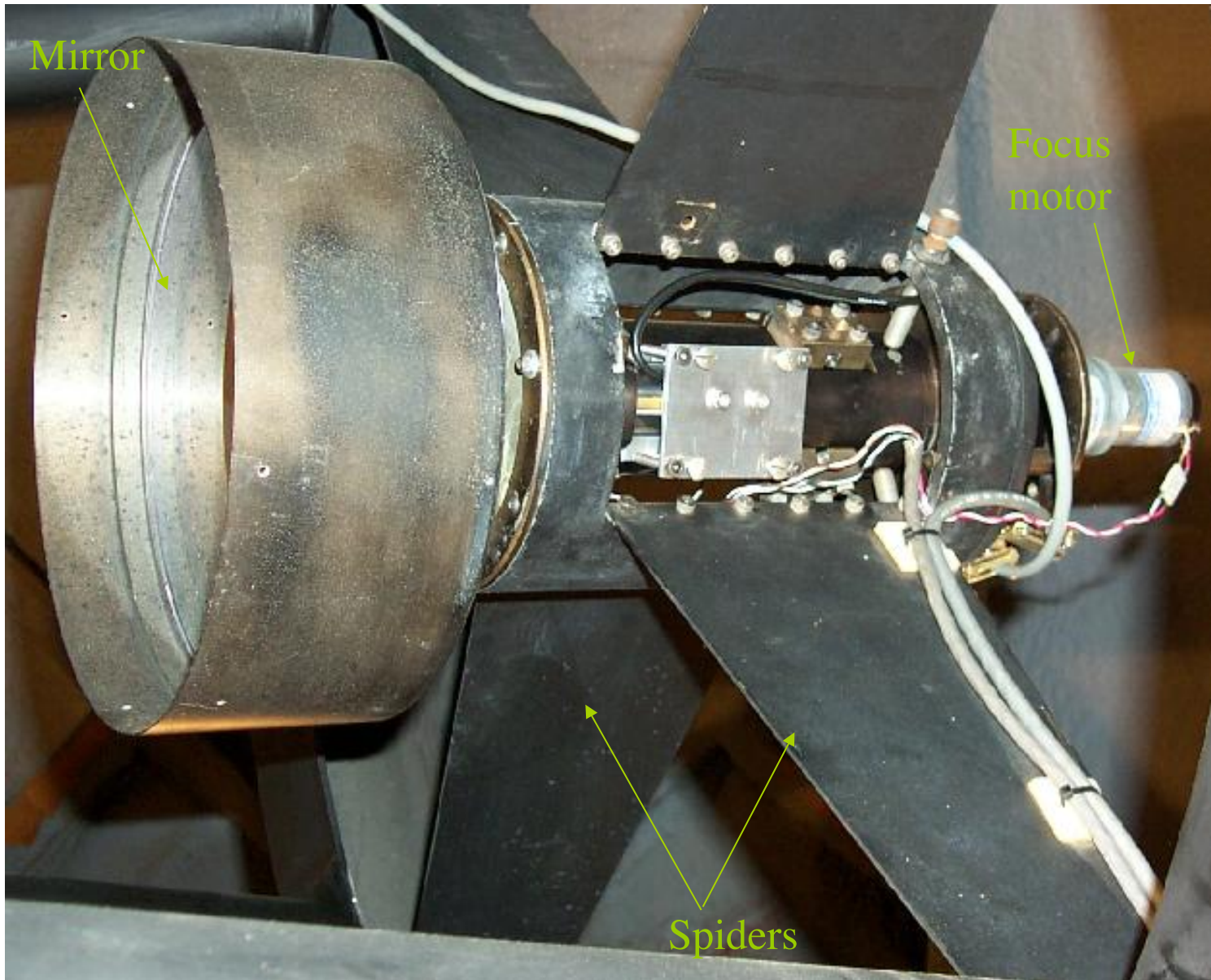
- Concave primary
- Convex secondary



# Telescope Design



# Secondary Mirror



# Telescope Secondary



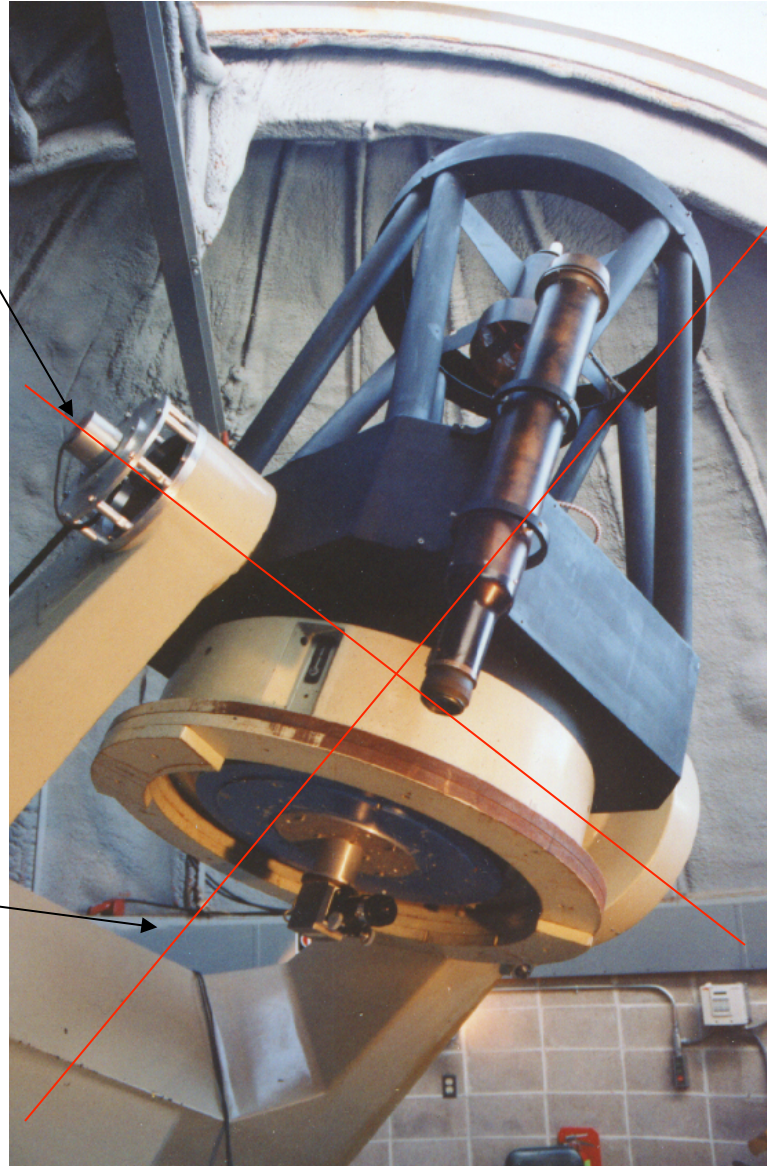


# The 30-inch Telescope

- Equatorial mount
  - Hour angle axis rotates 360 degrees in 23 hours 56 minutes

Declination axis

Hour  
angle  
axis



# The Dome



# Dome & Slit

