

# Asteroid Astrometry from CCD Images

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## Abstract

In this experiment, we——

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## 1. Introduction

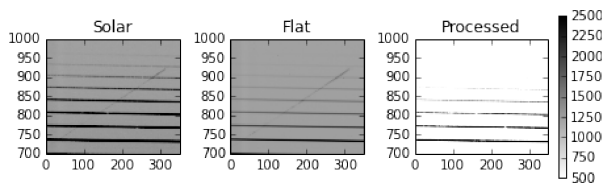


Figure 1: Both continuous sources shows the same artifact in a corner of the CCD image. The figures below are zoomed in — on this feature, which possibly from the reflected light . The halogen exposure is used for flat correction. Along with subtracting the background from the non-incident solar spectrum, the artifact is removed in the processed image shown in the rightmost figure.

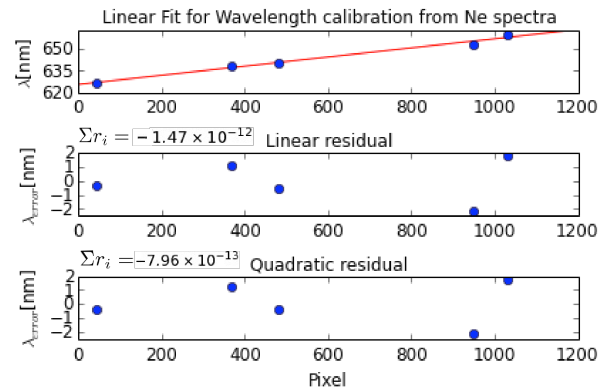


Figure 2: The first order fit in the top figure shows that the dispersion is approximately linear ( $\frac{d\lambda}{d\text{pixel}}=0.0315$  nm/pixel ). Since there is no notable pattern in linear residual, and the magnitude of the residual is small. From the bottommost quadratic figure, we see that the residual decrease by an order of magnitude.

## 2. Wavelength Calibration

We dark subtract the image

## 3. Doppler Shift Determination

## 4. Conclusion

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