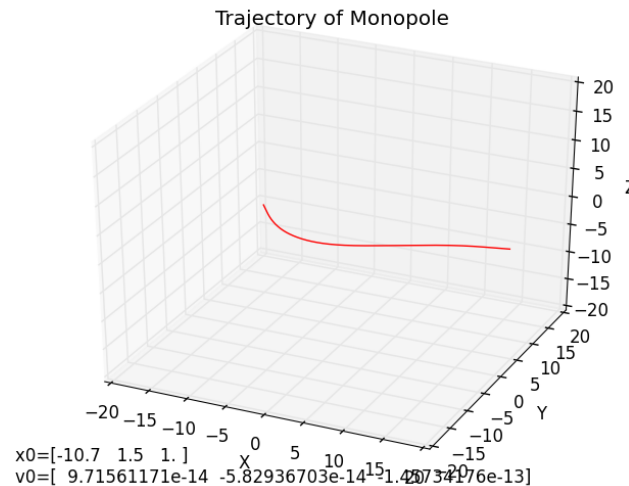


Magnetic Monopoles in the Milky Way Galaxy

Felix Feist, New York University
with Professor Glennys Farrar

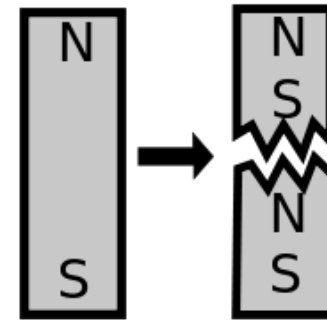


What are magnetic monopoles?

- A theoretical particle whose magnetic field is like the field of a single-pole magnet.
- A monopole would produce a divergent field like that of an electric charge.

Why do we believe in them?

- Dirac quantization condition
- Symmetry of Maxwell's Equations
- Predicted by GUTs



$$2\frac{q_e q_m}{\hbar c} \in \mathbb{Z}$$

Dirac Quantization Condition (cgs)

$$\begin{aligned}\nabla \cdot \vec{\mathbf{E}} &= 4\pi\rho_e \\ \nabla \cdot \vec{\mathbf{B}} &= 0 \\ \nabla \times \vec{\mathbf{E}} &= -\frac{1}{c}\frac{\partial \vec{\mathbf{B}}}{\partial t} \\ \nabla \times \vec{\mathbf{B}} &= \frac{1}{c}\frac{\partial \vec{\mathbf{E}}}{\partial t} + \frac{4\pi}{c}\vec{\mathbf{J}}_e\end{aligned}$$

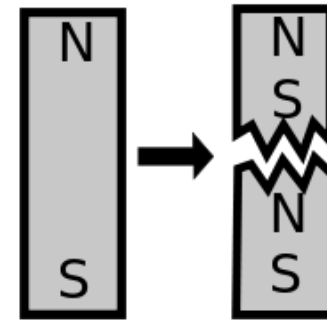
Maxwell's Equations (cgs)

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Maxwell's Equations (cgs)

Searches for the elusive monopole

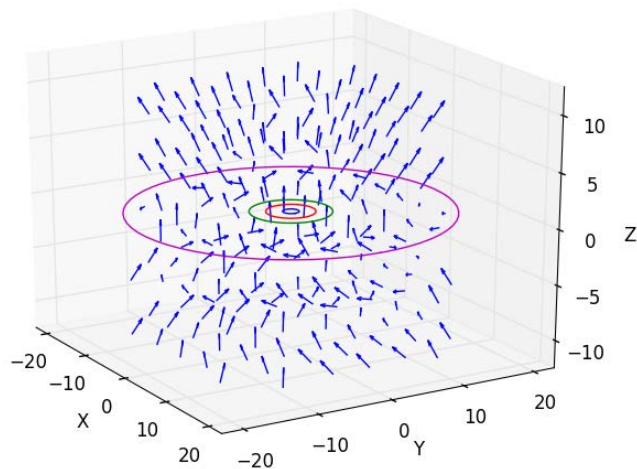
Many searches have been carried out...^[2]

- In cosmic rays
- In the Earth's crust
- Drifting through the atmosphere
- Produced in particle accelerators
- No statistically significant success.
- There is hope: The mass is not known, but GUTs predict $\sim 10^{15} \text{ GeV}/c^2$ ^[3]. If correct, this single particle is 20x heavier than a human blood cell^[4] and 10^{15} times the mass of a proton.

The Galactic Magnetic Field

- In 2012, Professor Glennys Farrar came up with an improved model of the Milky Way magnetic field.^[5]
- The motion of monopoles can now be studied more precisely.
- Does the magnetic field shield us from monopoles? Simulations can provide an answer.

$$\left(\frac{b_j}{r}\right) (1 - L(z, h_{disk}, w_{disk})) (\sin(11.5^\circ) \hat{r} - \cos(11.5^\circ) \hat{\phi}_p)$$



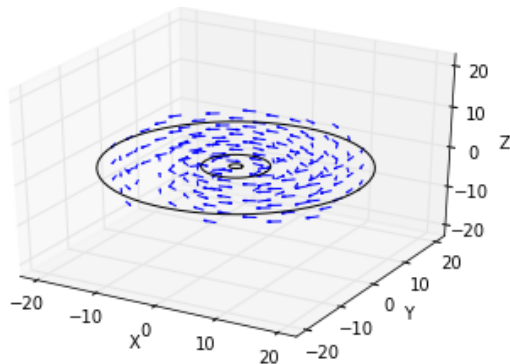
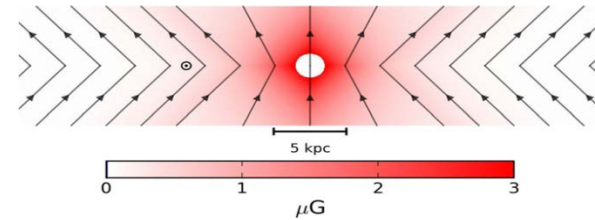
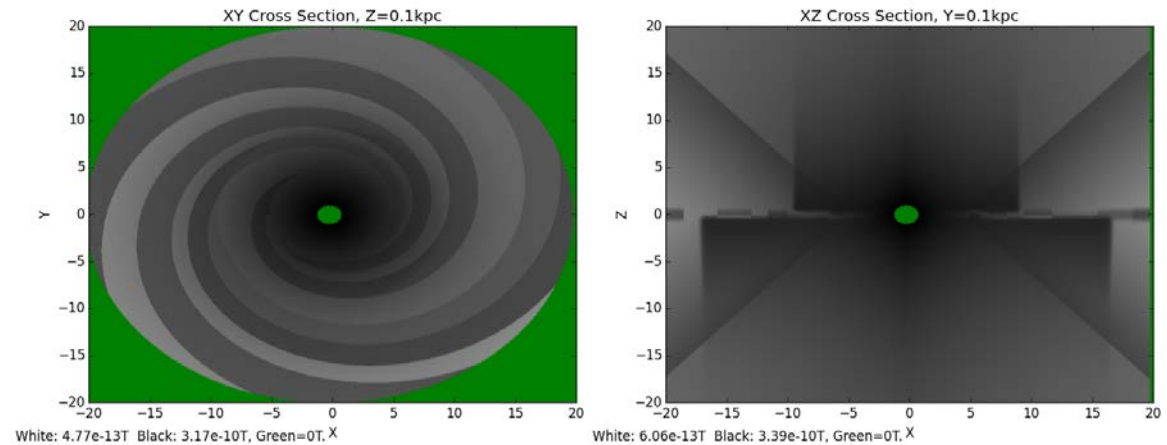
$$\begin{cases} B_X e^{\frac{-r_p}{r_X}} \left(\frac{r_p}{r}\right)^2 \hat{\mathbf{b}} & |r_p| < r_X^c \\ B_X e^{\frac{-r_p}{r_X}} \left(\frac{r_p}{r}\right) \hat{\mathbf{b}} & |r_p| \geq r_X^c \end{cases}$$

$$\hat{\mathbf{b}} = \cos(\Theta_X) \cos(\phi) \hat{\mathbf{x}} + \cos(\Theta_X) \sin(\phi) \hat{\mathbf{y}} + \sin(\Theta_X) \hat{\mathbf{z}}$$

Fun fact: it is piecewise in ~30 pieces

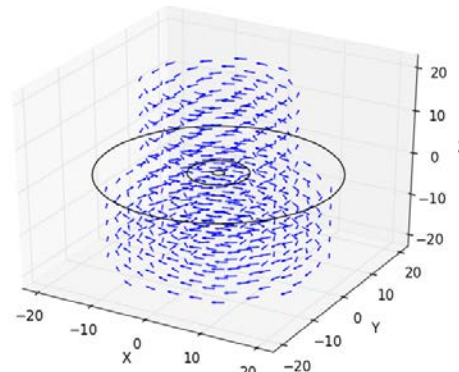
The field repels positive monopoles from the upper hemisphere & pulls them in from the lower hemisphere. It works oppositely for negative monopoles.

Thus an isotropic flux to the Earth is reduced by at most 50%.



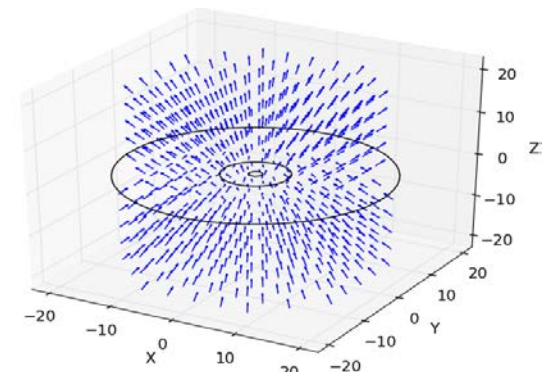
Disk Field

+



Halo Field

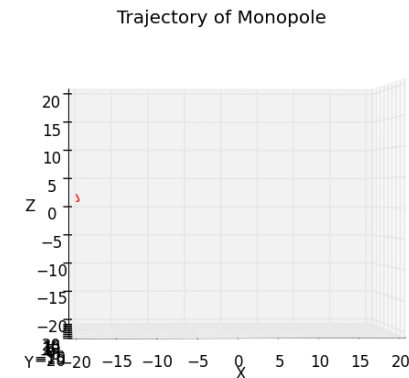
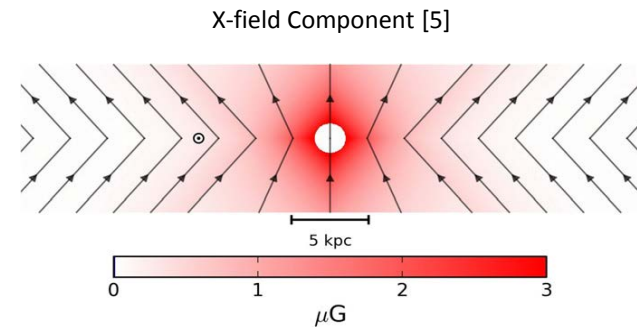
+



X-Field

The Simulation

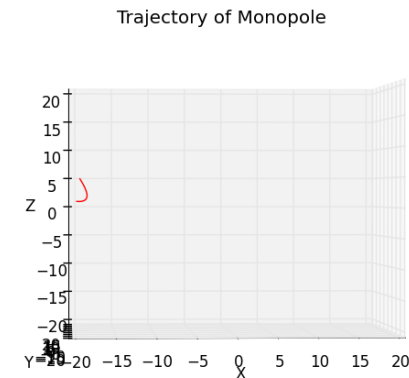
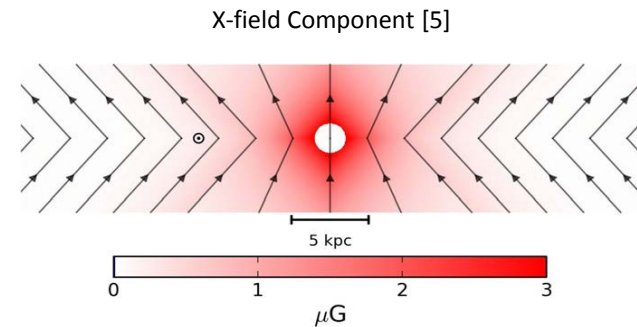
- Wrote 4th order Runge-Kutta solver in Python to find out the energy at which monopoles can tunnel through to the earth
- Dynamics had to be relativistic
- Simple method: increase starting velocity until the monopole reaches $r_{\text{earth}} = 8.5\text{kpc}$
- Gravitational field is negligible for monopoles with $m \ll 1\text{e}17\text{GeV}/c^2$ so it is not considered



```
x0=[-19.8 1.5 1.]  
v0=[ 3.83771232e-15 -5.09430839e-16 -3.39620559e-16]
```

The Simulation

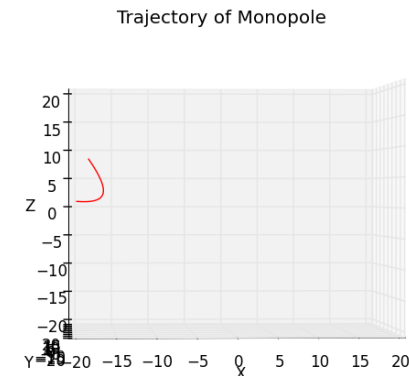
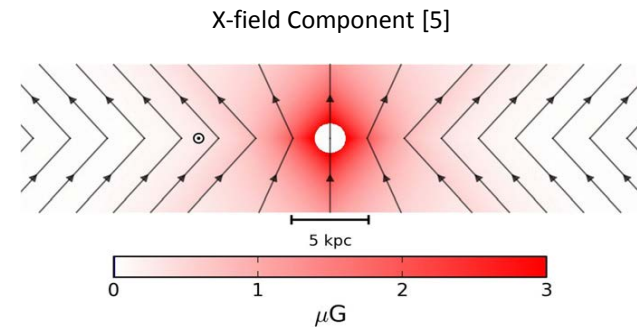
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```
x0=[-19.8 1.5 1.]  
v0=[ 9.59428080e-15 -1.27357710e-15 -8.49051398e-16]
```


The Simulation

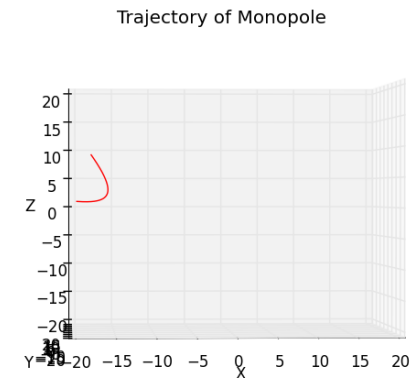
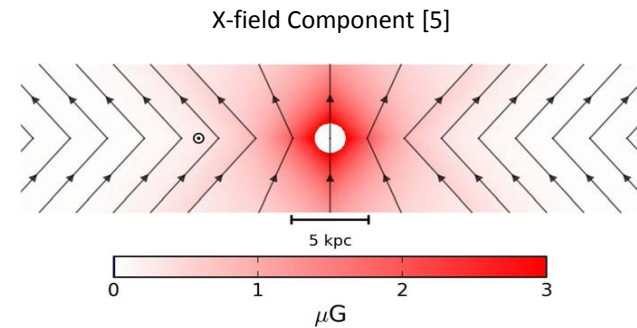
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```
x0=[-19.8 1.5 1.]  
v0=[ 1.91885616e-14 -2.54715419e-15 -1.69810280e-15]
```

The Simulation

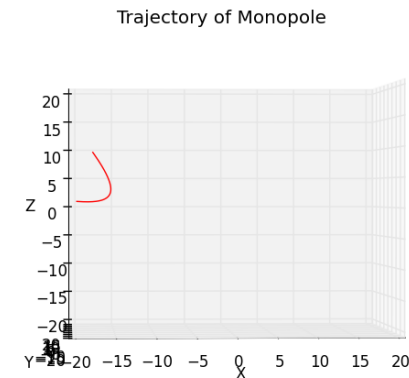
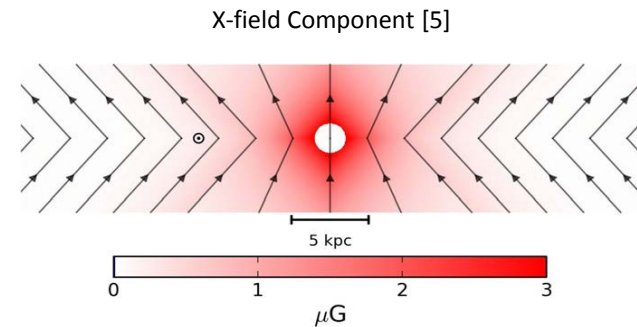
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```
x0=[-19.8 1.5 1.]  
v0=[ 2.20668458e-14 -2.92922732e-15 -1.95281822e-15]
```

The Simulation

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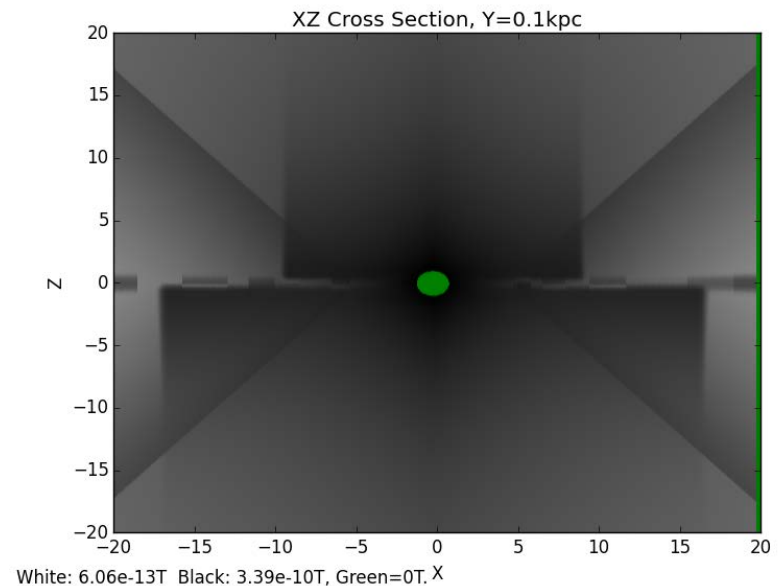


```
x0=[-19.8 1.5 1.]  
v0=[ 2.39857020e-14 -3.18394274e-15 -2.12262850e-15]
```

Results

- Monopoles can tunnel through to the earth at initial kinetic energies of $\sim 5e10$ GeV or more, but only from ideal angles of entry.
- Starting at kinetic energies of $\sim 1.3e12$ GeV, monopoles can reliably reach the earth from any angle.
- Consistent with order-of-magnitude calculations.
- Independent of monopole mass!


=> Total flux to the earth is halved for monopoles with kinetic energy less than $5e10$ GeV. It then increases with monopole energy until $1e12$ GeV, at which point it is unaffected by the magnetic field.



Discussion

- Wick (2002) estimates that relic monopoles are distributed with mean Kinetic Energy $\sim 7e12$ GeV
- The breaking point at which the Flux is unaffected is $1.3e12$ GeV
=> Relic monopoles are unaffected by the galactic field.
- Lower energy monopoles may be produced by other sources. If less than $5e10$ GeV, we are being shielded from them. This means that if cosmic ray monopoles exist, detection experiments may receive a reduced flux and a bias of monopole charge which depends on the latitude/longitude of the detector.
- Error due to $\sim 10\%$ uncertainty in field. No timestep or machine precision error.

References

- [1]  image from wikipedia: Magnetic Monopoles
- [2] List of monopole searches from Sloan, Non-collider searches for stable massive particles arXiv:1410.1374v1 [hep-ph] 6 Oct 2014
- [3] <http://moedal.web.cern.ch/content/search-magnetic-monopole>
- [4] Red blood cell mass from Wolfram Alpha search for “ $1e15\text{GeV}/c^2$ ”
- [5] Galactic field model from “A New Model of the Galactic Magnetic Field” (Jansson, Farrar 2012): The Astrophysical Journal, 757:14 (13pp), 2012 September 20

I produced all other diagrams in Python using matplotlib:

