

---

# DENOISING GW STRAIN DATA WITH DEEPCLEAN

Muhammed Saleem,  
University of Minnesota

AcceleratingPhysicsWithML@MIT, Jan 28 - Feb 2, 2023.

*Collaborators: Chia-Jui Chou, Alec Gunny, Li-Cheng Yang, Michael Coughlin, Dylan Rankin, Erik Katsavounidis, Ryan Magee, Phil Haris, Will Benoit, Deep Chatterjee, Ethan Marx*

# DENOISING GOALS

- The output reconstructed from an interferometer contains

$$h(t) = s(t) + n(t)$$

Possible GW signal      Detector noise

- Objective: To recover  $s(t)$  with best possible signal-to-noise ratio by minimising the noise  $n(t)$
- Scientific objectives:
  - Signals that are below the noise (un-detectable) becomes detectable
  - Improved SNR improves parameter estimations

# REMovable AND NON-REMovable NOISES

$$n(t) = n_{nw}(t) + n_w(t)$$

- Non-removable (fundamental noise)
- Budgeted by system design
- Eg: photon shot noise, thermal noise
- Can be reduced only with upgraded design and technology
- Source of noise witnessed by dedicated system monitors (witness sensors)
- Environmental contamination or technical noise eg: noise arising from the control of suspended optics

---

# WITNESSED NOISE

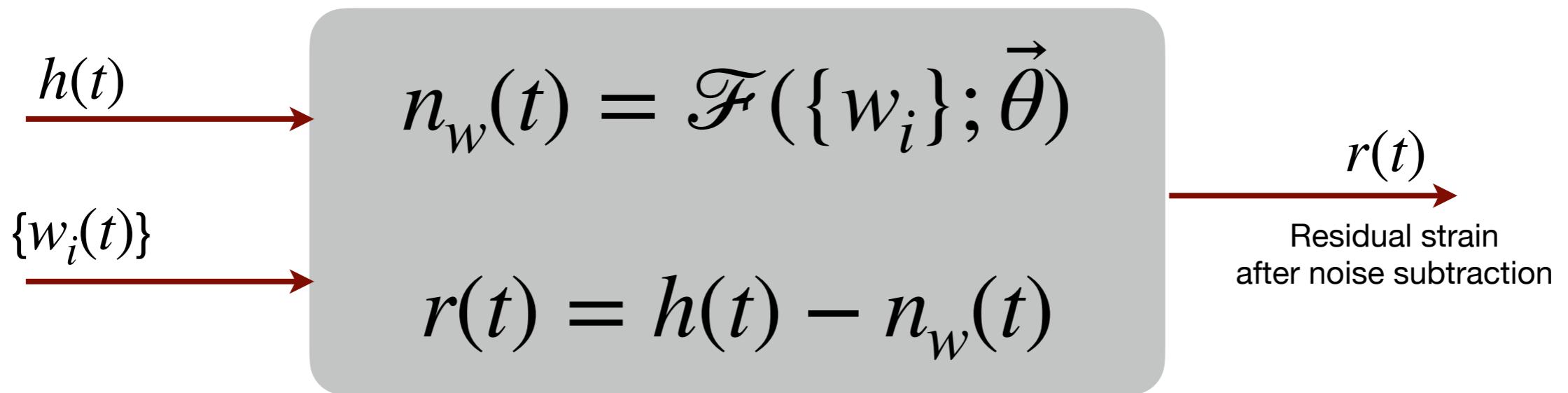
---

- witness sensors or channels - the timeseries denoted by  $\{w_i(t)\}$
- The noise  $n_w(t)$  is the collective contribution from  $\{w(t)\}$
- Schematically,

$$n_w(t) = \mathcal{T}(\{w_i(t)\})$$

where  $\mathcal{T}$  is some activation function representing non-linear or non-stationary coupling of the output of witness channels to the strain channel.

# DEEPCLEAN: A NEURAL NETWORK TO PREDICT $n_w(t)$



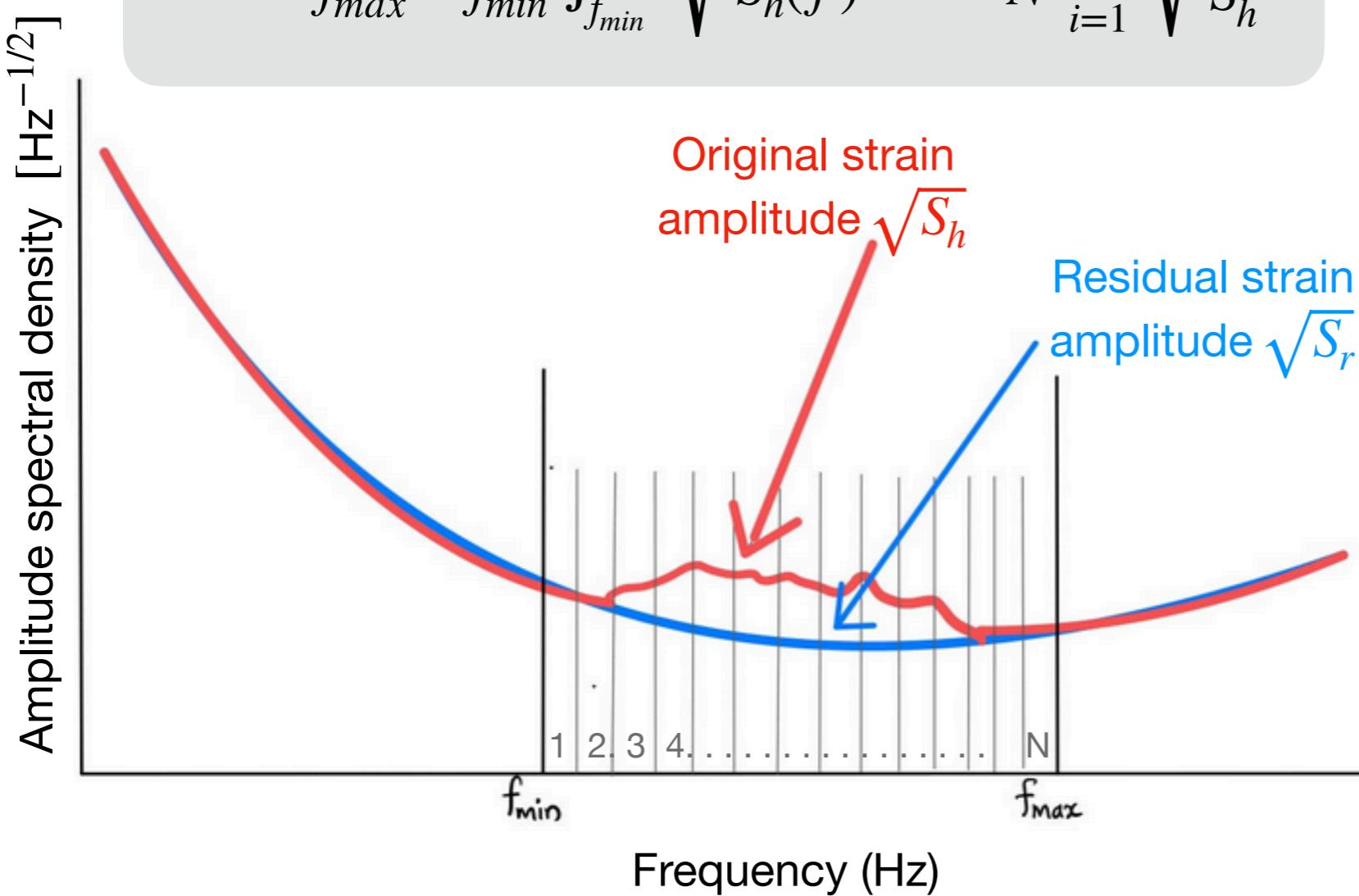
- $\vec{\theta} \rightarrow$  trained weights of the neural network

$$\vec{\theta} = \min_{\vec{\theta}'} \left[ J \left( h(t), \mathcal{F}(\{w_i(t)\}; \vec{\theta}') \right) \right]$$

where  $J$  is some appropriate loss function

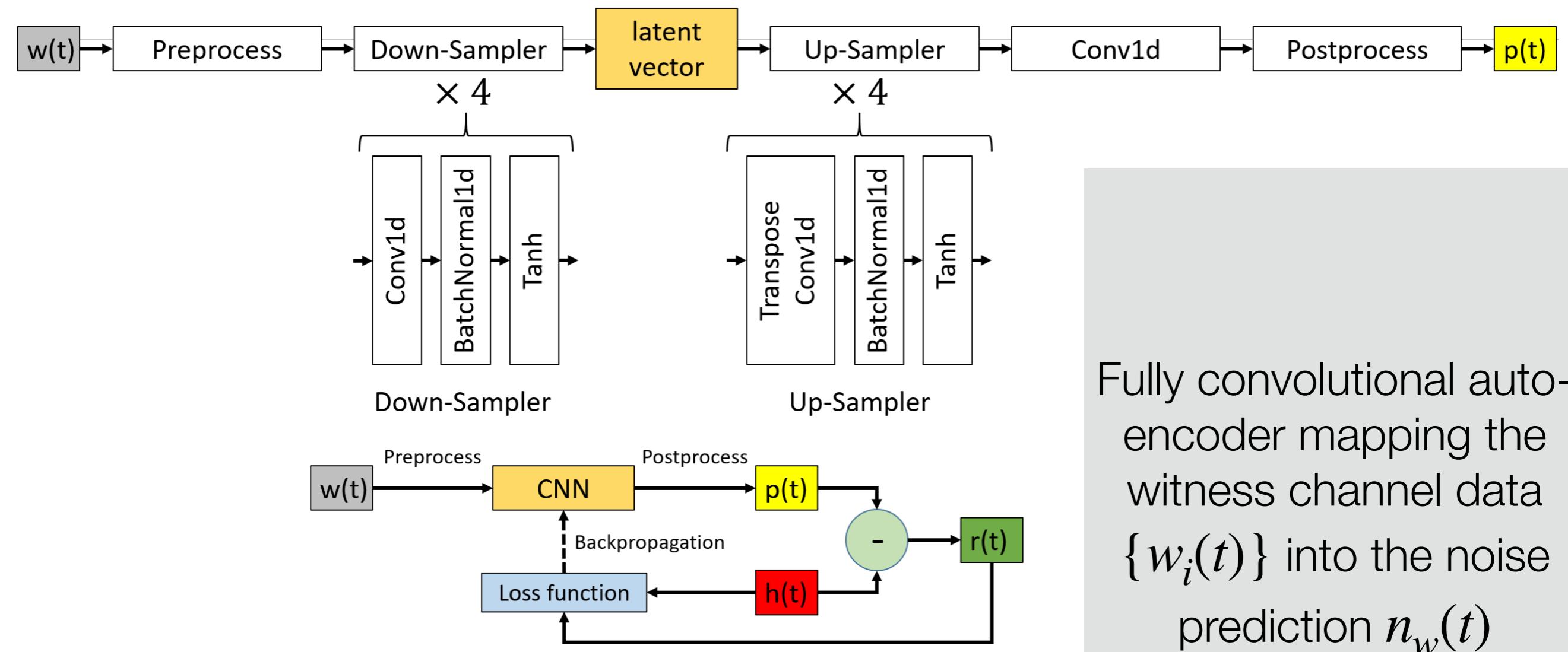
# LOSS FUNCTION IN TERMS OF AMPLITUDE SPECTRAL DENSITY (ASD)

$$J_{asd} = \frac{1}{f_{max} - f_{min}} \int_{f_{min}}^{f_{max}} \sqrt{\frac{S_r(f)}{S_h(f)}} df = \frac{1}{N} \sum_{i=1}^N \sqrt{\frac{S_r^{(i)}}{S_h^{(i)}}}$$



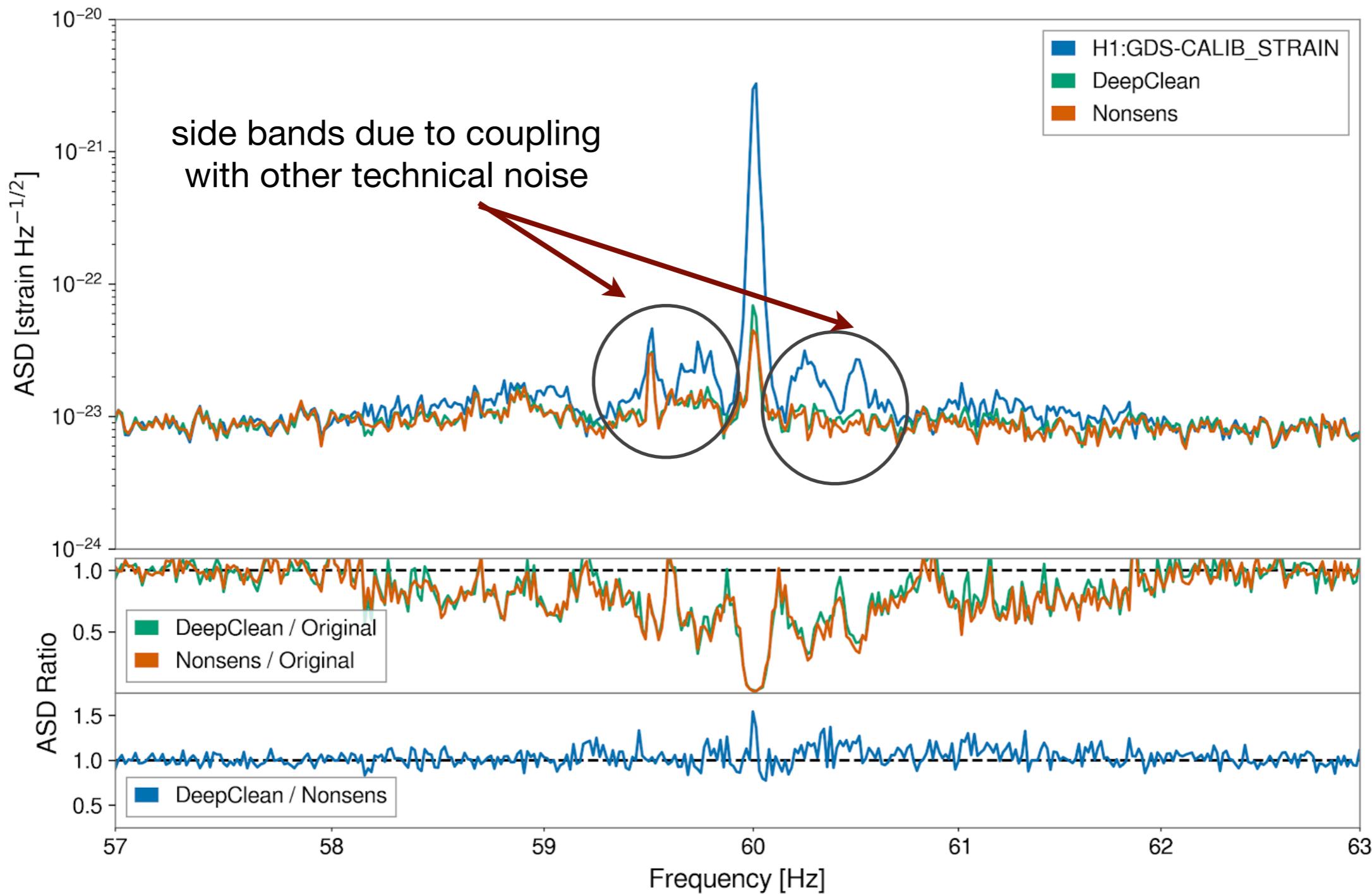
- Minimises the ASD ratio averaged over all the bins in the desired frequency range.
- Ratio of residual ASD to the original ASD
- Involves FFT → the minimum length of the time-domain data dictated by the resolution of the features we are looking for.

# THE ARCHITECTURE AND THE WORKFLOW



Fully convolutional auto-encoder mapping the witness channel data  $\{w_i(t)\}$  into the noise prediction  $n_w(t)$

# EXAMPLE: SUBTRACTION OF 60 Hz POWER-LINE AND THE SIDEBANDS

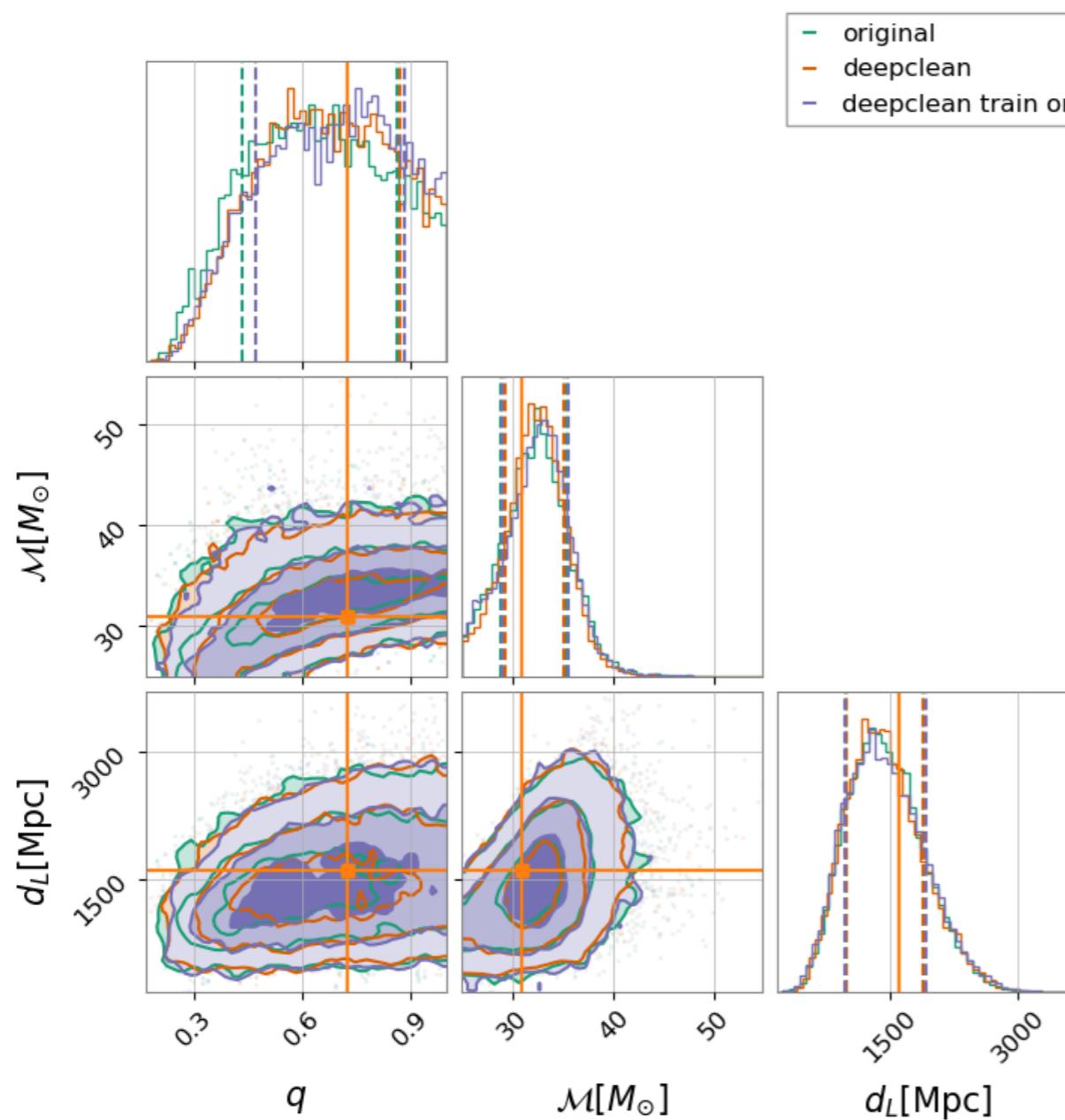


# DEEPCLEAN VALIDATION TESTS WITH MOCK DATA

---

- Twenty days of LIGO data from O3 (between 9/1/2019 - 9/20/2019)
- Injected compact binary coalescence signals (BBH/BNS/NSBH)
  - Drawn from O3-inferred astrophysical population models
- **25000** injections in total, we analyse 266 BBH injections that have coalescence frequency between 55-70 Hz ()
- **Following** analysis performed on the original and cleaned data
  - Parameter estimation of the injections
  - Sensitive volume ( $\langle VT \rangle$ ) estimation
  - Match filter SNR of the injections

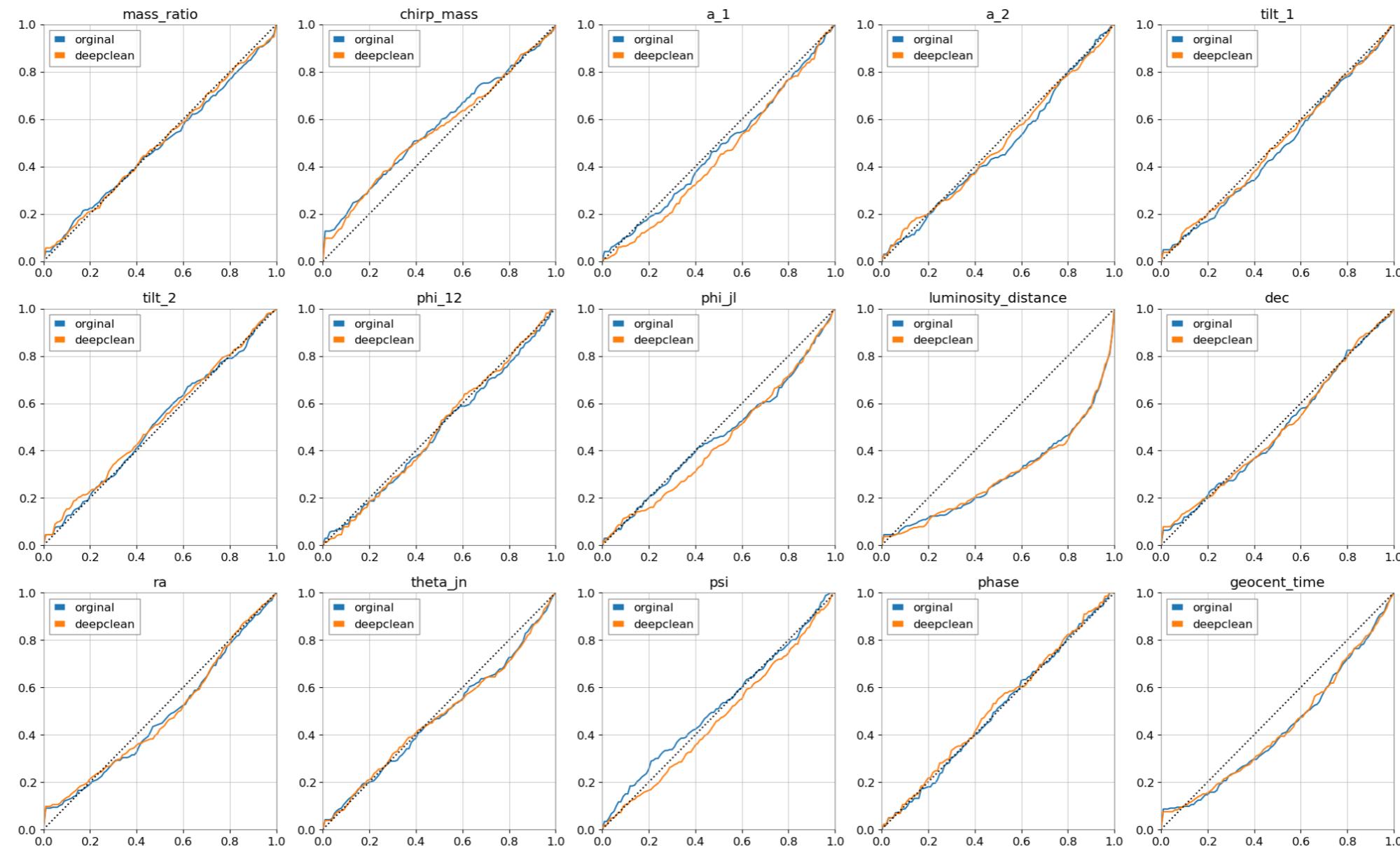
# PE VALIDATION: EXAMPLE FROM A SINGLE INJECTION



- With 60 Hz subtraction, PE improvements are not prominent. This is expected as the signal spends a small fraction of time in the frequency band of 60 Hz noise
- Can we train on data that has GW signals in it?
- Yes, they don't affect the noise prediction.

# PE VALIDATION: P-P PLOTS OF ALL THE PARAMETERS

P-P plot shows that deepClean does introduce bias into the PE



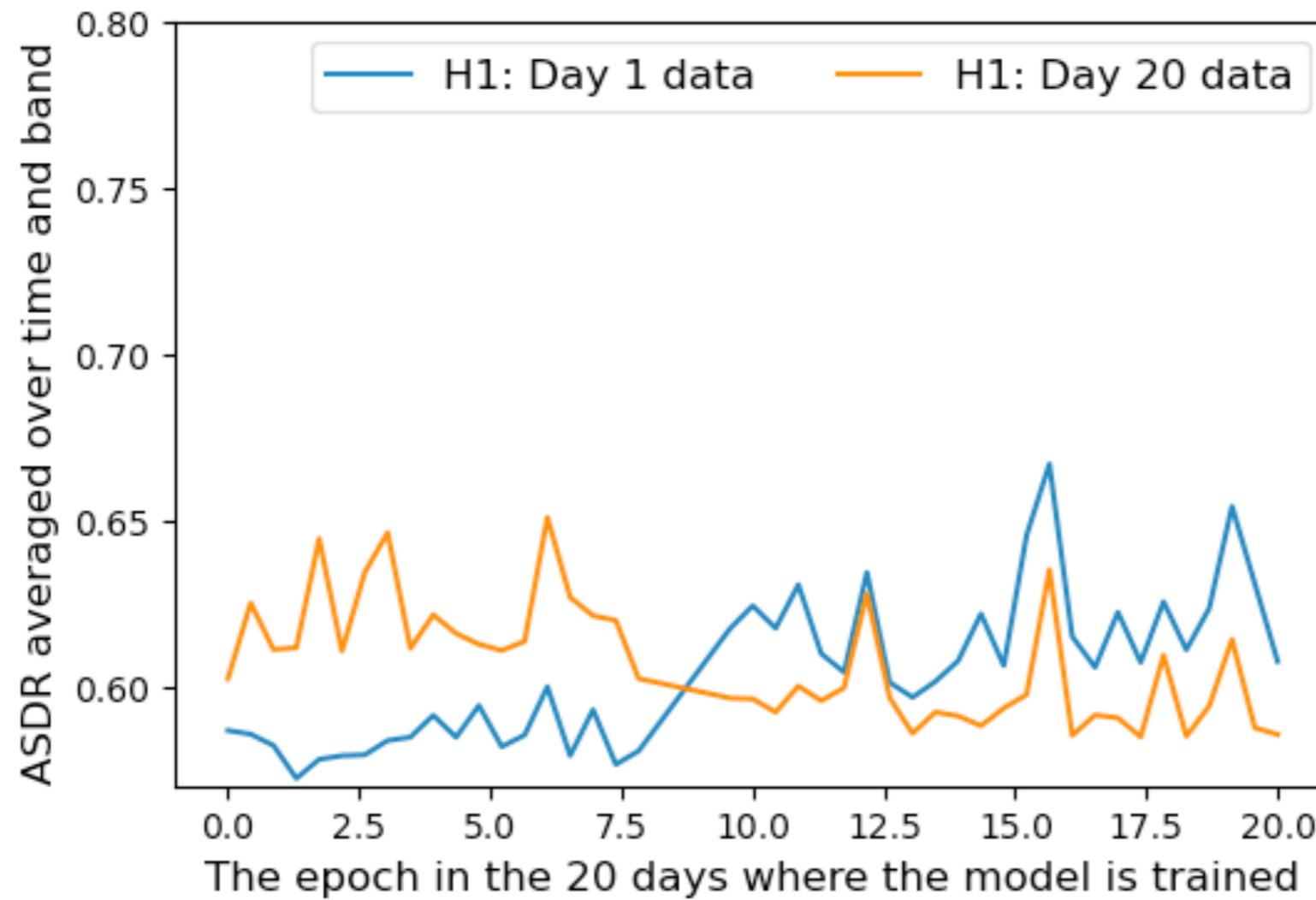
---

# DETECTION IMPROVEMENTS

---

- Results awaited
- GstLAL: Computes the sensitive volume improvements (how many new detections to expect due to denoising)
- Improvements in the matched filter SNR

# HOW OFTEN TO TRAIN DEEPCLEAN



- Once over a few days is found to be enough in O3 data.
- This might change in O4 or for a different noise coupling

---

# ONGOING AND FUTURE WORKS

---

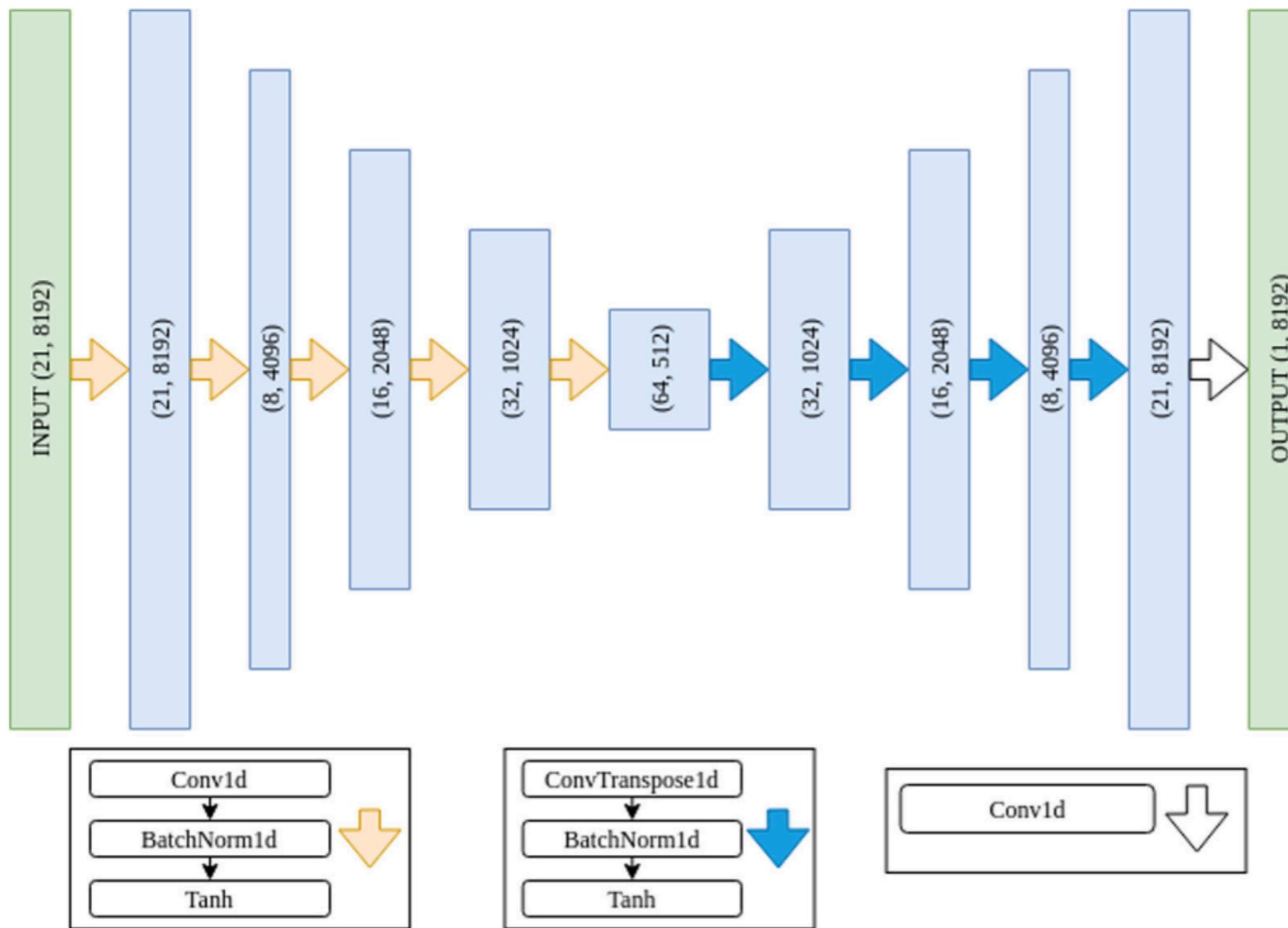
- KAGRA implementation (Chia-Jui's talk up next)
- Online DeepClean on LIGO data
  - Feasibility study done (DeepClean can be performed on frames of 1 s duration and the cleaned frames can be produced in ~1 s latency)
  - Deployment with HERMES taking place (lead by Alec Gunny)
- Low frequency (10-30 Hz) broadband noise subtraction
- LIGO internal review to start soon (to be ready for O4)

---

---

# ADDITIONAL SLIDES

# ARCHITECTURE WITH AN EXAMPLE



Fully convolutional auto-encoder mapping the witness channel data  $\{w_i(t)\}$  into the noise prediction  $n_w(t)$

Input: 8 s time-series from 21 witness channels sampled at 1024 Hz

Output: Noise prediction for 8s sampled at 1024 Hz

---

---

# ONLINE DEEPCLEAN

---

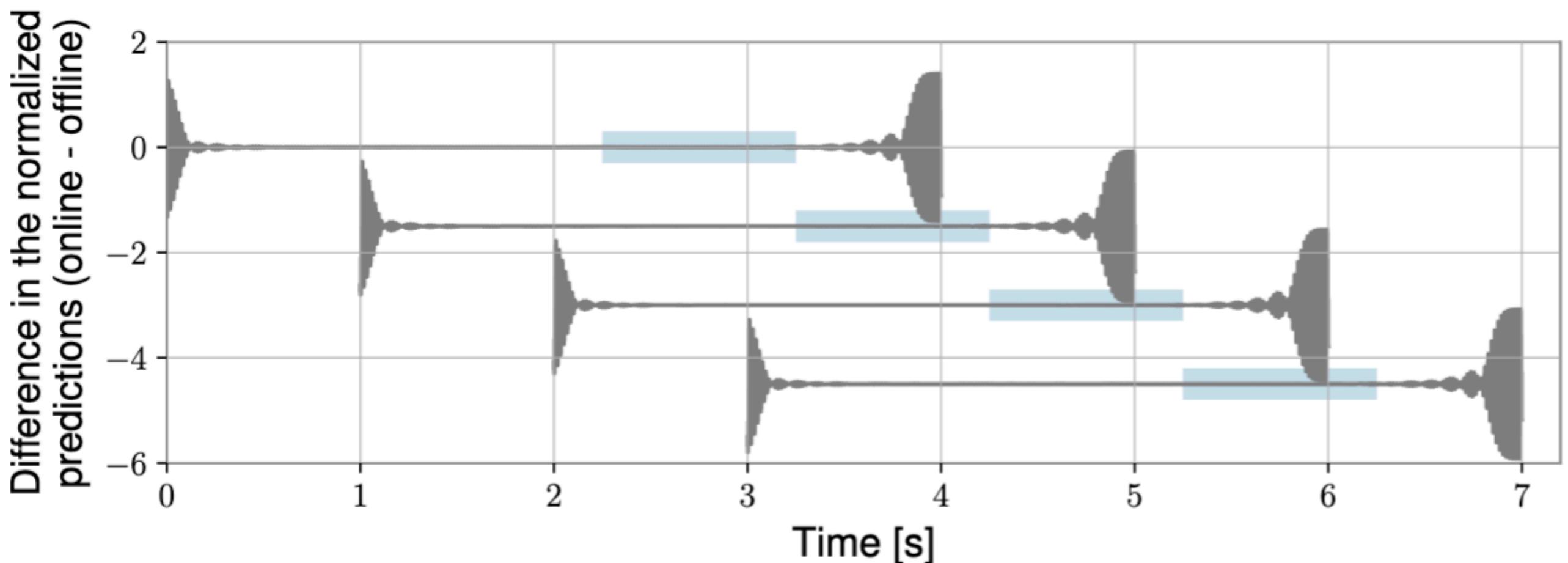
# ONLINE DEEPCLEAN

---

- Aim is to subtract every 1 s frame as soon as they are available
- 1 s cleaned data are noisy at the edges
  - Reason: a combination of the architecture, training and post-processing (the major part).
- Consider 4 s data for cleaning and take 1 s from somewhere in the middle excluding the edges.
  - Causes additional latency

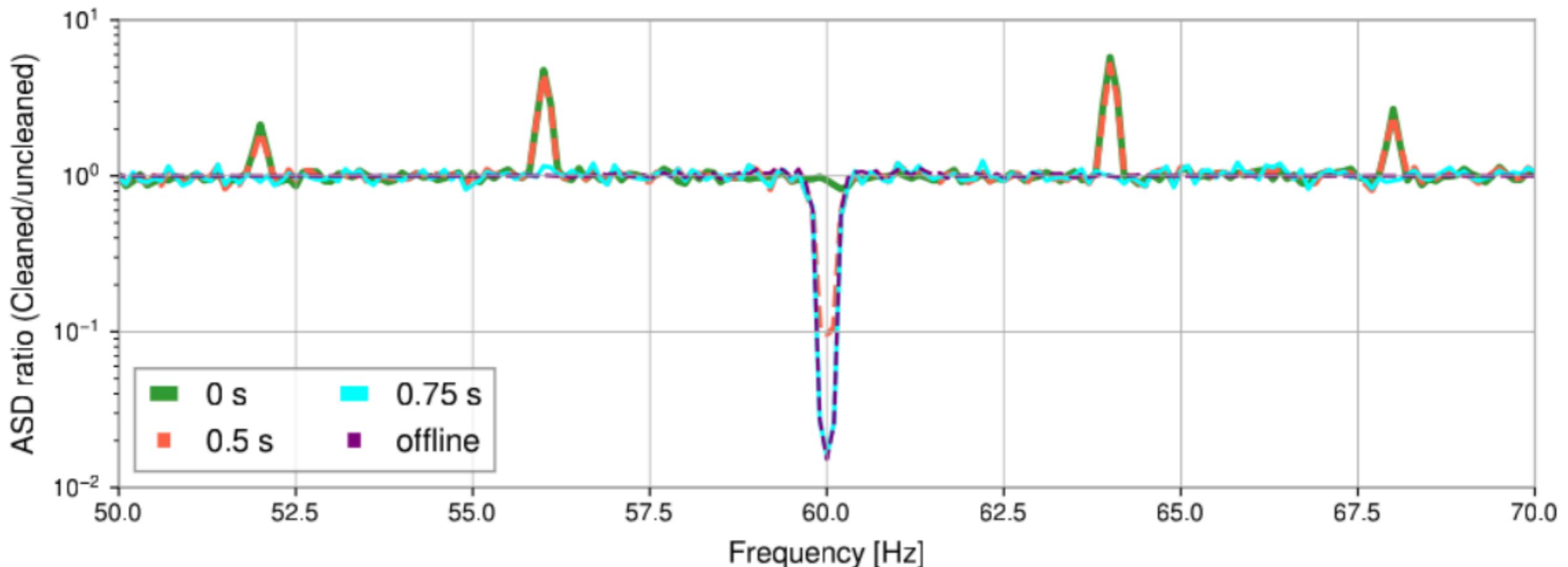
*Will be discussed in other talks*

# SLIDING WINDOWS: SUPPRESSING THE EDGE EFFECTS



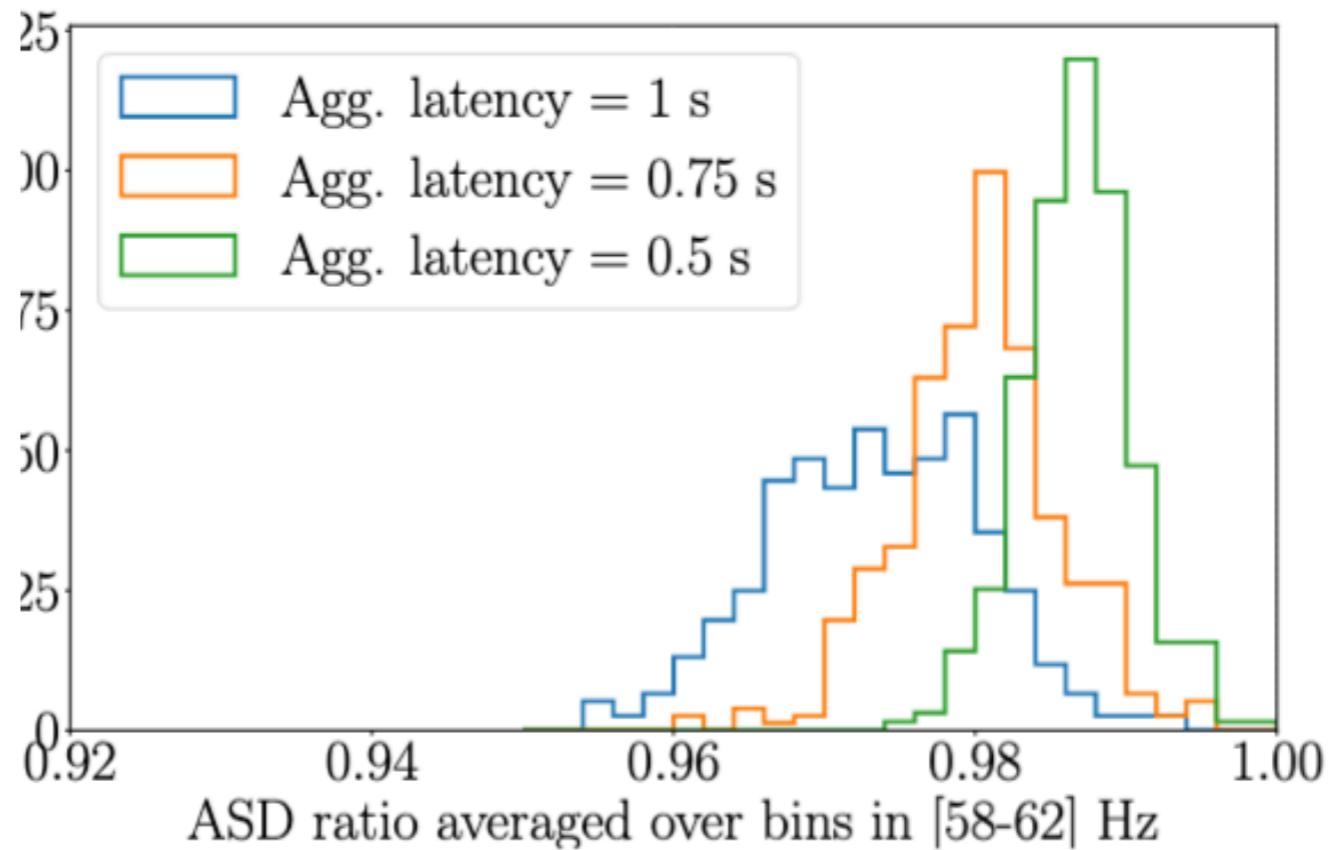
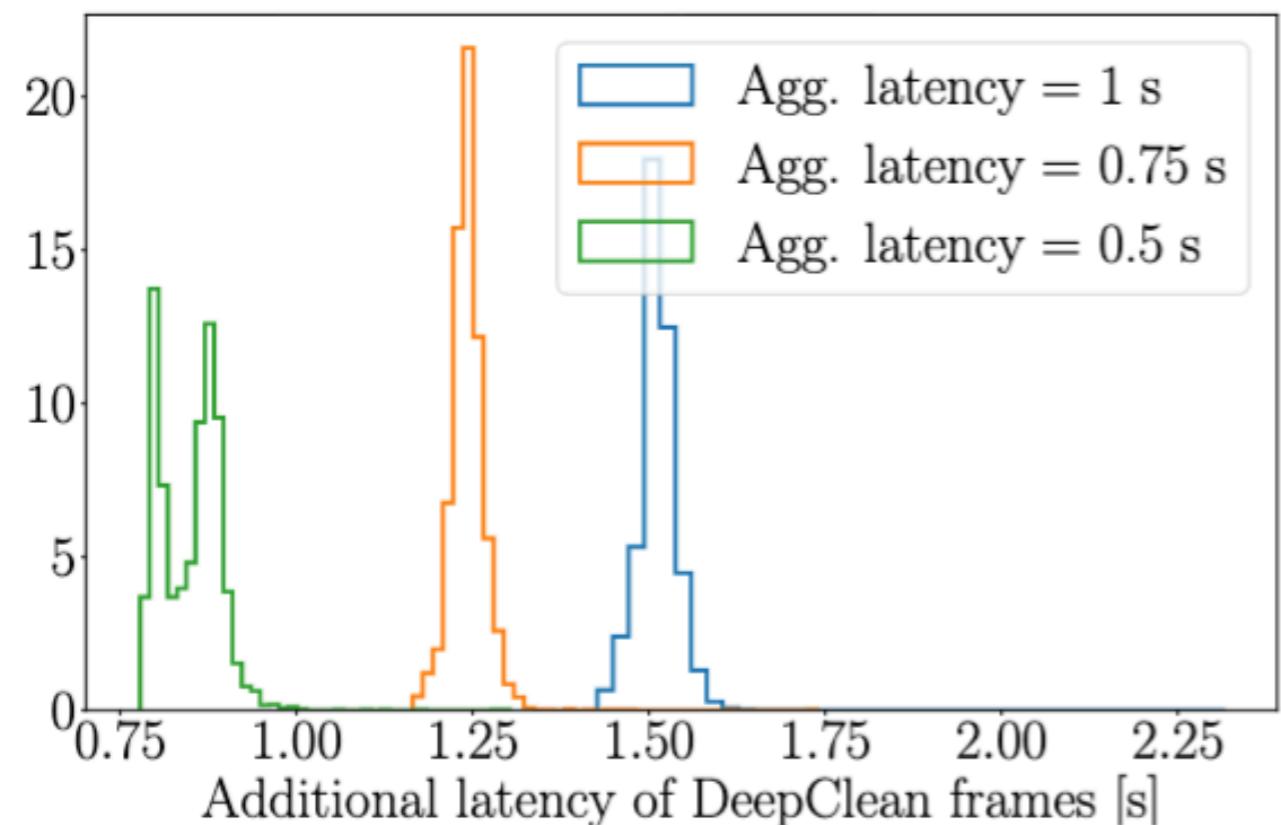
- Take unaffected 1s from the 4s cleaned segment and run cleaning on overlapping windows

# ONLINE DEEPCLEAN: LATENCY VS SUBTRACTION QUALITY

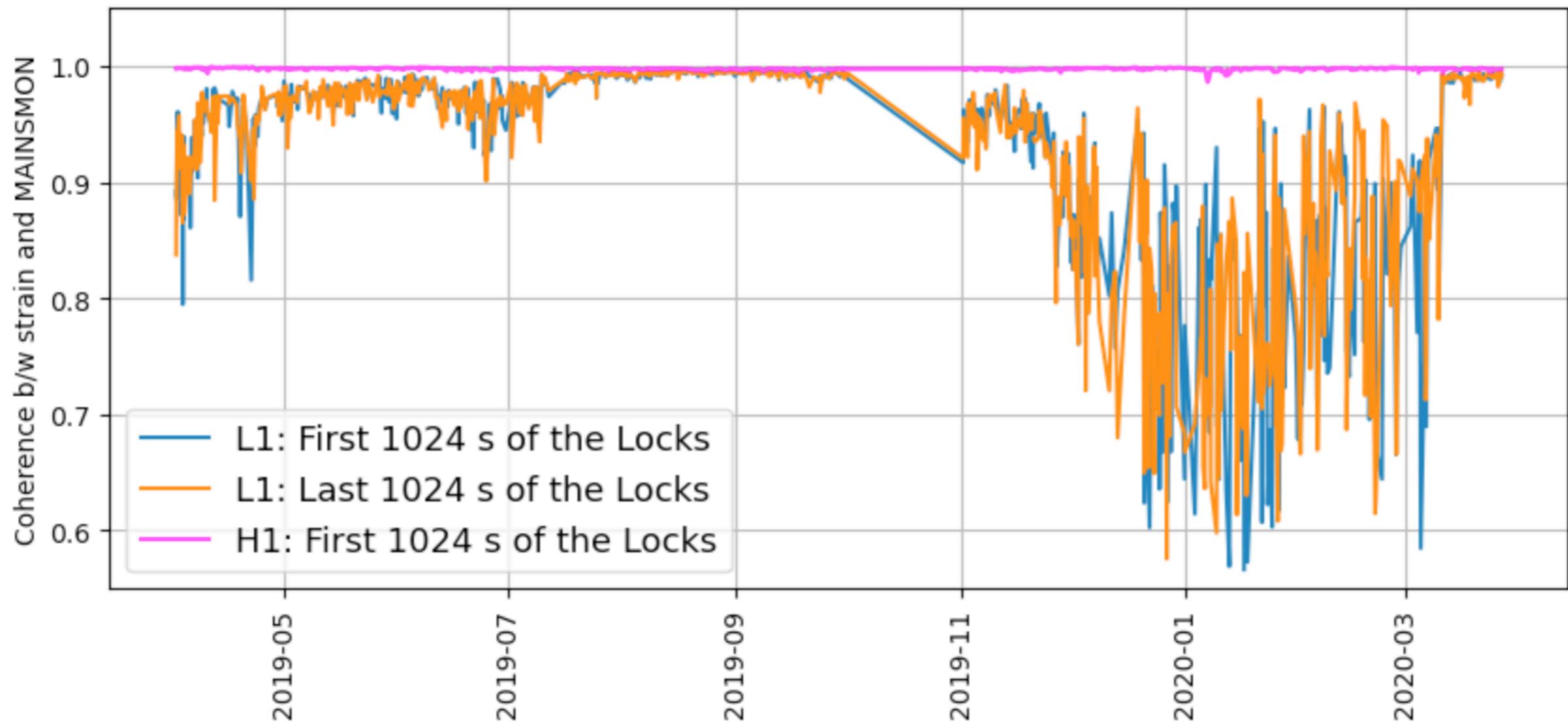


- Longer the excluded edge, better the quality of the subtraction

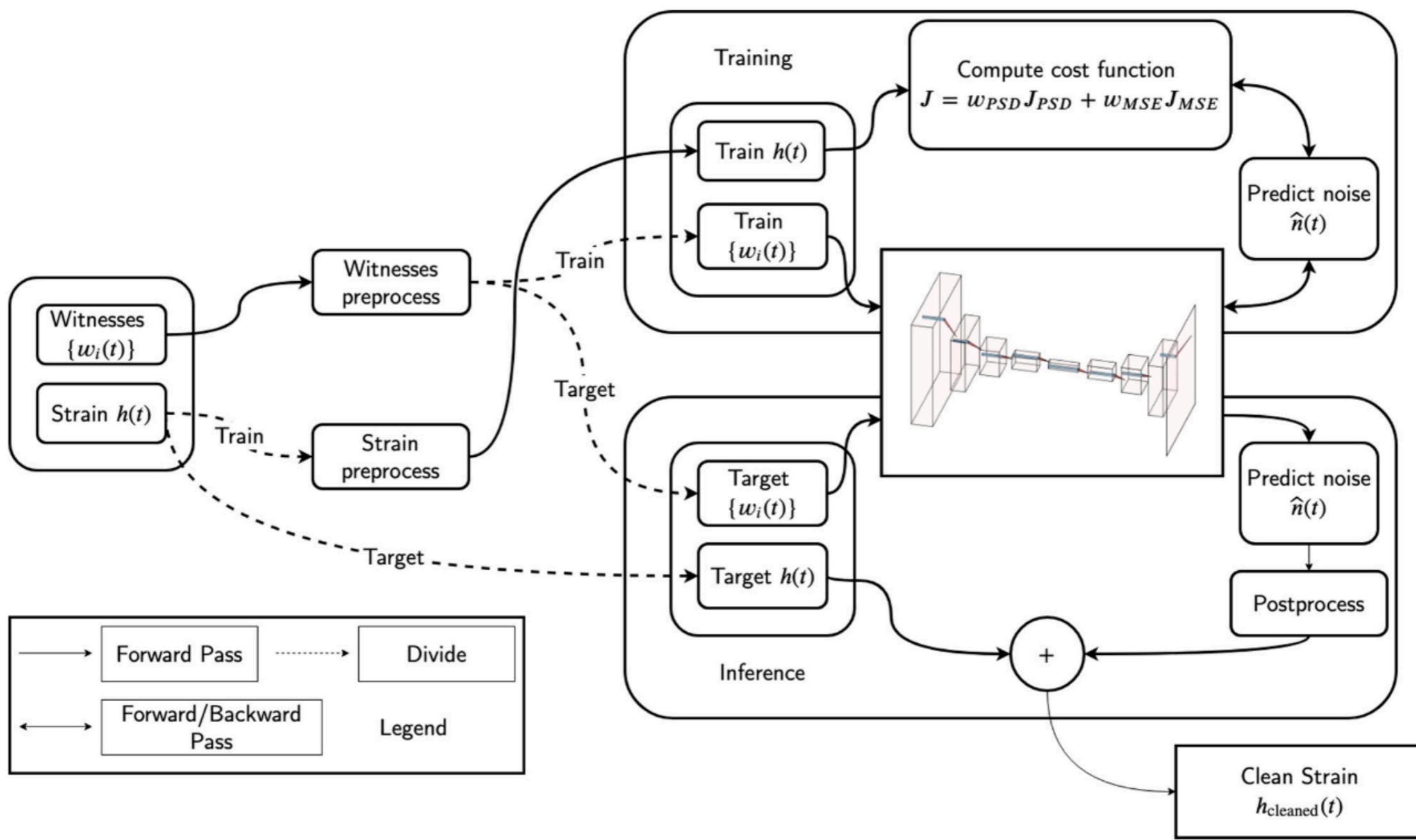
# ONLINE DEEPCLEAN: LATENCY VS IMPROVEMENTIN ASD RATIO



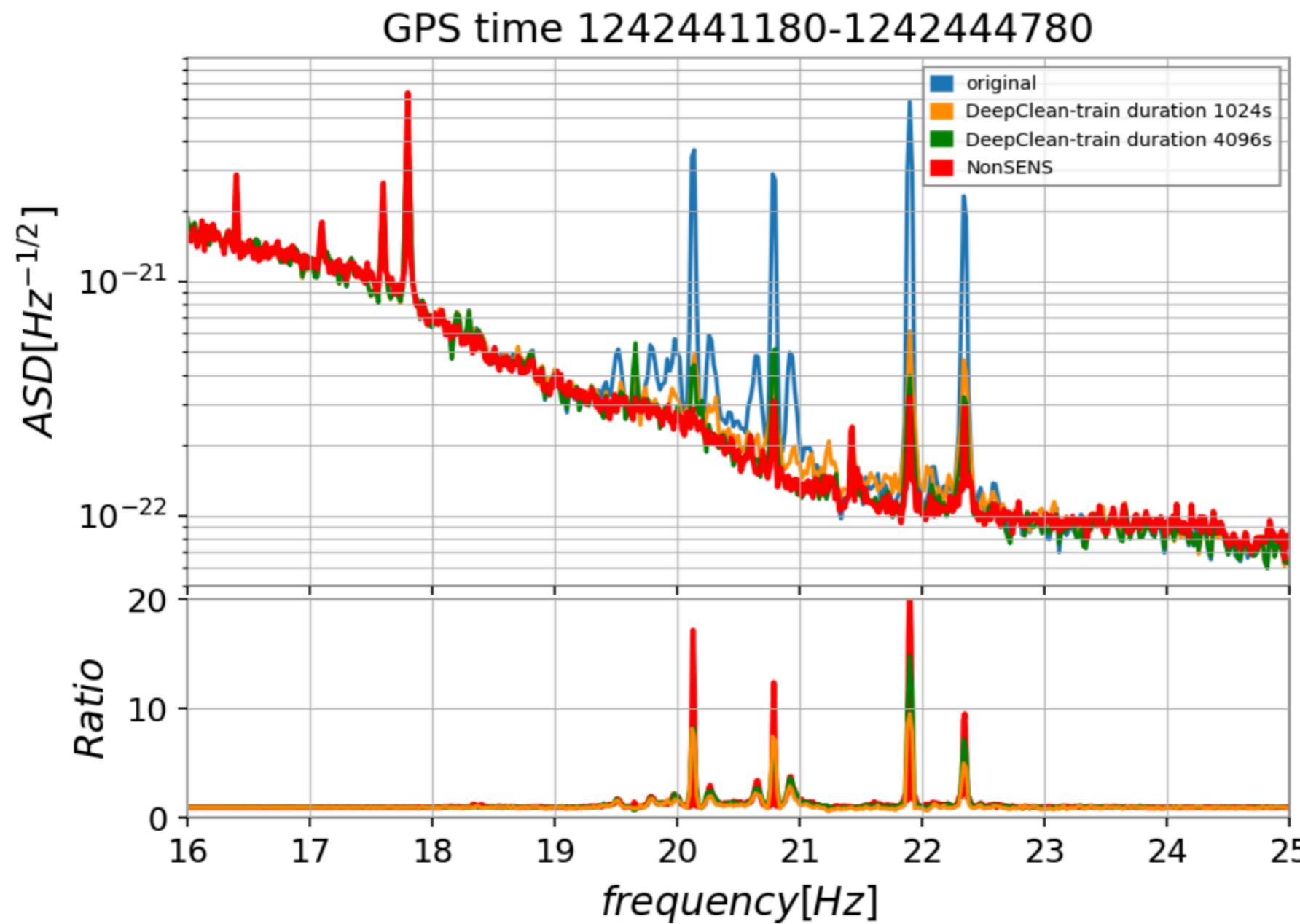
# COHERENCE BETWEEN THE STRAIN AND POWER-LINE WITNESS



# DEEPCLEAN: FULL WORKFLOW



# SUBTRACTION OF LOW-FREQUENCY BROADBAND NOISE (10-30 Hz)



---

# A TYPICAL OFFLINE DEEPCLEAN ANALYSIS

---

- Training:
  - About 1000 s of data used to train
  - Sent to the CNN in batches of segments and the ASD computed by averaging over all segments.
  - Length of the segment determined by the desired resolution (for 1/4 Hz resolution, the segments should be at least 4 s longer)
- Inference:
  - No restrictions in particular on the length of the data cleaned
    - (Too shorter segments (< 10 s) not recommended, due to some edge effects)
  -

# OFFLINE DEEPCLEAN: SPECIFICATIONS

---

- Network trained with 2000 s (or 4000 s) data from around (up to a few days before or after) the target segment.
- Target segment is chosen to be 4096 s
- Batch: 32 overlapping kernels, with each kernel being  $22 \times 32768$  matrices
  - Assuming there are 22 channels (for 60 Hz subtraction) and one kernel is 8 s and data is sampled at 4096 Hz
  - Batch size of 16 and kernels of 4s or 2s have been found to be equally effective in cleaning, however for training, the kernels need to be at least 8 s long as the loss function is PSD-based (a mean PSD with 4 FFTs and 1/2 Hz resolution needs 8s data)

# WITNESS CHANNELS USED IN 60 Hz SUBTRACTION

HI:GDS-CALIB\_STRAIN

HI:PEM-CS\_MAINSMON\_EBAY\_I\_DQ

HI:ASC-INPI\_P\_INMON

HI:ASC-INPI\_Y\_INMON

HI:ASC-MICH\_P\_INMON

HI:ASC-MICH\_Y\_INMON

HI:ASC-PRCI\_P\_INMON

HI:ASC-PRCI\_Y\_INMON

HI:ASC-PRC2\_P\_INMON

HI:ASC-PRC2\_Y\_INMON

HI:ASC-SRCI\_P\_INMON

HI:ASC-SRCI\_Y\_INMON

HI:ASC-SRC2\_P\_INMON

HI:ASC-SRC2\_Y\_INMON

HI:ASC-DHARD\_P\_INMON

HI:ASC-DHARD\_Y\_INMON

HI:ASC-CHARD\_P\_INMON

HI:ASC-CHARD\_Y\_INMON

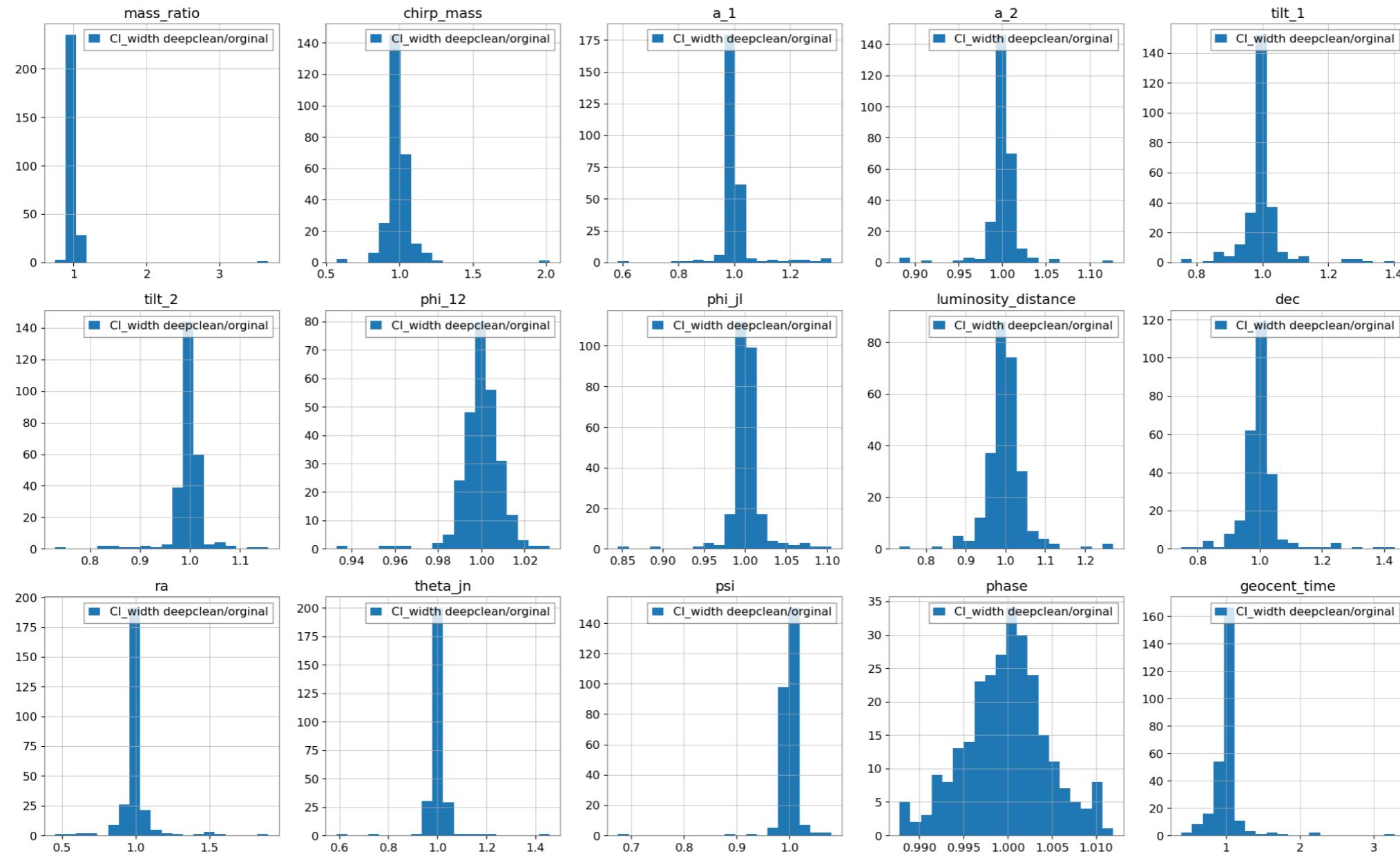
HI:ASC-DSOFT\_P\_INMON

HI:ASC-DSOFT\_Y\_INMON

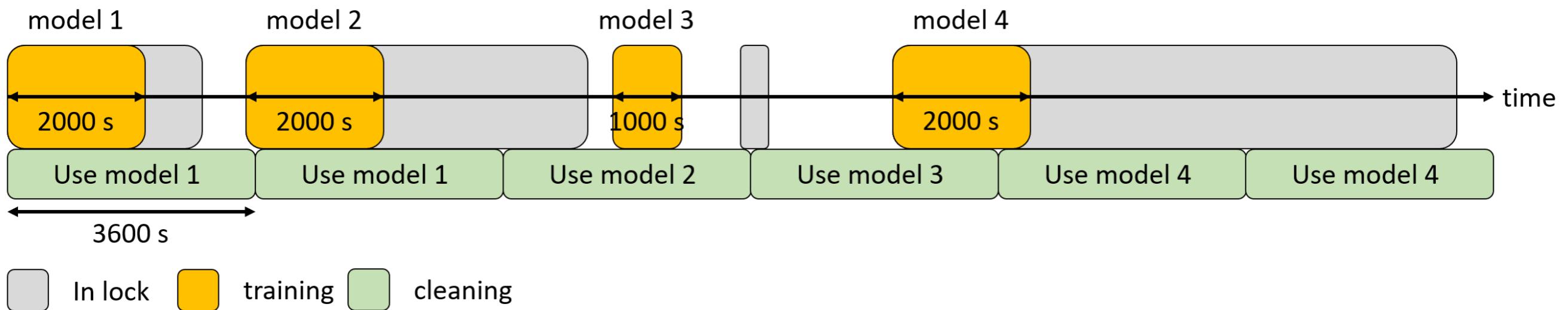
HI:ASC-CSOFT\_P\_INMON

HI:ASC-CSOFT\_Y\_INMON

# PE IMPROVEMENTS: RATIO OF ERROR BARS



# TRAINING STRATEGY FOLLOWED FOR MDC ANALYSIS



# TRAINING DEEPCLEAN

---

- Preprocess: normalize (zero mean and unit variance) and bandpass to the desired frequency band.
- Data sent to the network in mini-batches with one batch typically having 32 samples
- Loss function computed for the network prediction by averaging over the mini-batch.
- Loss is minimized using ADAM (the Gradient Descent algorithm) using both Forward pass and backward pass