MATH 623: DIFFERENTIAL GEOMETRY II

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1. Introduction

This is the second semester of a two-semester graduate course providing an introduction to differential geometry. The second semester is primarily a study of Riemannian manifolds with a focus on curvature. At the end of the course, we may go in different directions depending on the interests of the class. Possible directions include comparison theorems, principal bundles and the Atiyah–Singer index theorem, Lorentzian manifolds, the Hodge theorem, or the Chern–Gauss–Bonnet theorem.

The topics roughly covered (updated later based on course interest) include:

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Make sure to add topics!

2. Preliminaries and review

Recall from last semester (or your previous experience) the notions of smooth manifold, the tangent and cotangent bundles of a smooth manifold, and tensor fields. Put in definitions to refresh! Maybe recall how to work with the objects!

Sections. Diffeomorphisms.

In coordinates (x^1, \ldots, x^n) on a patch U of M, recall that $\frac{\partial}{\partial x^1}, \ldots, \frac{\partial}{\partial x^n}$ form a basis for T_nM for each $p \in U$.

The following characterization of tensor fields is so useful, let's just get it out of the way.

Lemma 1. Suppose $\overline{T}: \mathcal{X} \times \cdots \times \mathcal{X} \times \Omega^1 \times \cdots \times \Omega^1 \to C^{\infty}(M)$ is an \mathbb{R} -multillinear function of k vector fields and ℓ 1-forms (covector fields). Then \overline{T} arises from a (ℓ, k) -tensor field T if and only if T is multilinear over $C^{\infty}(M)$ in each of its arguments, e.g.,

$$\overline{T}(fX_1,\ldots,X_k,\omega_1,\ldots,\omega_\ell)(p) = f(p)\overline{T}(X_1,\ldots,X_k,\omega_1,\ldots,\omega_\ell)(p).$$

Proof. Given a (ℓ, k) -tensor field T, we form \overline{T} from it by evaluating at each point. We then have

$$\overline{T}(fX_1, \dots, X_k, \omega_1, \dots, \omega_\ell)(p) = T_p(f(p)X_{1,p}, \dots, X_{k,p}, \omega_{1,p}, \dots, \omega_{\ell,p})$$

$$= f(p)T(X_{1,p}, \dots, X_{k,p}, \omega_{1,p}, \dots, \omega_{\ell,p})$$

$$= f(p)\overline{T}(X_1, \dots, X_k, \omega_1, \dots, \omega_\ell)(p).$$

For the other direction, let's just do the case of $\overline{T}: \mathcal{X} \to C^{\infty}(M)$. (This contains the main idea; the general case is an exercise in careful bookkeeping.) Suppose for all $f \in C^{\infty}(M)$ and $X \in \mathcal{X}(M)$ we have

$$\overline{T}(fX)(p) = f(p)\overline{T}(X)(p).$$

Our aim is to show that \overline{T} arises from a 1-form ω . Let's start by defining the purported 1-form. Given $p \in M$ and $v \in T_pM$, choose an $X \in \mathcal{X}(M)$ so that $X_p = v$. We define

$$\omega_p(v) = \overline{T}(X)(p).$$

We must show that ω is well-defined, i.e., that it does not depend on the choice of vector field X. Suppose X_1 and X_2 are two vector fields with $X_{1,p} = X_{2,p} = v$. In particular, the vector field $Y = X_1 - X_2$ satisfies $Y_p = 0$ and so there are smooth vector fields Z_1, \ldots, Z_r and smooth functions f_1, \ldots, f_r with $f_1(p) = \cdots = f_r(p) = 0$ so that $Y = f_1 Z_1 + \cdots + f_r Z_r$. We then have

$$\overline{T}(Y)(p) = \sum_{j=1}^{r} f_j(p)\overline{T}(Z_j)(p) = 0,$$

so that $\overline{T}(X_1)(p) = \overline{T}(X_2)(p)$ and thus $\omega_p : T_pM \to \mathbb{R}$ is well-defined. Its linearity is clear from the linearity of \overline{T} and its smoothness follows from the mapping properties of \overline{T} so it is indeed a 1-form.

2.1. **Notation.** Unless explicitly noted, all manifolds in this course will be smooth (i.e., C^{∞}) manifolds.

We use $\mathcal{X}(M)$ to denote the space of C^{∞} vector fields on M, i.e., smooth sections of TM. For sections of other bundles $E \to M$ we often use $\Gamma(E)$ to denote the space of smooth sections. For vector fields along a curve α we use $\mathcal{X}(\alpha)$ and sections of E above a curve α are denoted $\Gamma(E,\alpha)$.

3. RIEMANNIAN METRICS

Suppose M is a smooth manifold of dimension n (typically $n \geq 2$).

Suppose $g: \mathcal{X}(M) \times \mathcal{X}(M) \to C^{\infty}(M)$ is a symmetric (0,2)-tensor. (In other words, g is a tensor field so that $g_p: T_pM \times T_pM \to \mathbb{R}$ is symmetric at each point $p \in M$.)

Definition 2. We say that a symmetric (0,2)-tensor is a Riemannian metric if $g_p(v,v) > 0$ for all $p \in M$, $v \in T_pM$ with $v \neq 0$. The tensor g is pseudo-Riemannian if for all $p \in M$, if $v \in T_pM$ has $g_p(v,w) = 0$ for all $w \in T_pM$, then v = 0. (In other words, g is pseudo-Riemannian if it is non-degenerate and Riemannian if it is additionally an inner product as you know it from linear algebra.)

A smooth manifold M equipped with a Riemannian metric g is called a Riemannian manifold.

In terms of the coordinate basis induced by a coordinate patch (x^1, \ldots, x^n) in M, we set

$$g_{ij} = g\left(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}\right),$$

so that $g_{ij} \in C^{\infty}(U)$, $g_{ji} = g_{ij}$, and if $v = v^i \frac{\partial}{\partial x^i}$ and $w = w^j \frac{\partial}{\partial x^j}$, then

$$\langle v, w \rangle_p := g_p(v, w) = \sum_{i,j} g_{ij} v^i w^j|_p.$$

A Riemannian metric g then gives an inner product on each tangent space. It also induces metrics on associated bundles (in what follows g is always a metric on the tangent spaces; notation for the induced metrics varies wildly by source and eventually we'll just use g to denote all of them unless there can be confusion as to where various objects live).

One way to view the mechanism for this induction is that g gives a way to "raise and lower indices". More precisely, a Riemannian metric g provides an isomorphism between the tangent and cotangent spaces at each point. Given $\omega \in T_pM$, we associate to it a vector $w_\omega \in T_pM$ by demanding that

$$g_p(w_\omega, v) = \omega(v)$$

for all $v \in T_pM$. Because g is non-degenerate, this uniquely defines the vector w_{ω} . In local coordinates, the displayed equation reads

$$\sum_{i,j} g_{ij}(w_{\omega})^i v^j = \sum_j \omega_j v^j,$$

so that

$$(w_{\omega})^i = \sum_j g^{ij} \omega_j,$$

where g^{ij} are the components of the matrix inverse of (g_{ij}) . Similarly, if $v \in T_pM$, one can identify it with the one-form $\omega_v \in T_p^*M$ so that

$$\omega_v(u) = g_p(v, u)$$

for all $u \in T_p M$. In local coordinates, $(\omega_v)_i = \sum_j g_{ij} v^j$.

(1) Cotangent bundle. Given $\omega, \eta \in T_p^*M$, we define the metric G (sometimes denoted g^{-1} , sometimes still just g) by

$$G(\omega, \eta) = g(w_{\omega}, w_{\eta}),$$

where w_{ω} is the vector associated to ω and w_{η} is the one associated to η . In coordinates, we have that the (i, j)-component of the metric G is the same as the (i, j)-component of the matrix $g^{-1} = (g_{k\ell})^{-1}$, i.e., g^{ij} . To check this, observe that

$$g(w_{\omega}, w_{\eta}) = \sum_{i,j} g_{ij}(w_{\omega})^{i}(w_{\eta})^{j}$$

$$= \sum_{i,j,k,\ell} g_{ij}g^{ik}\omega_{k}g^{j\ell}\eta_{\ell}$$

$$= \sum_{jk\ell} \delta_{j}^{k}g^{j\ell}\omega_{k}\eta_{\ell} = \sum_{i,j} g^{ij}\omega_{i}\eta_{j}.$$

(2) Tensor bundles. For $T_pM \otimes T_pM$ (and higher powers), say that

$$g^{\otimes}(v_1\otimes v_2,w_1\otimes w_2)=g(v_1,w_1)g(v_2w_2),$$

and extend linearly. For factors of T_p^*M , also use the raising/lower operator.

- (3) Exterior powers. Use that $\Lambda^k(TM)$ is a sub-bundle of $(T^*M)^{\otimes k}$ and use above.
- (4) Endomorphism bundle. Identify $\operatorname{End}(TM)$ with $T^*M \otimes TM$.

As an example of a Riemannian metric, suppose $F: M \to \mathbb{R}^N$ is an immersion (so that dF_p is injective for all p and thus has rank dim M). The immersion F (and the ambient inner product on \mathbb{R}^N) induces a Riemannian metric on M by

$$g_p(v, w) = \langle dF_p(v), dF_p(w) \rangle_{\mathbb{R}^N}$$

for all $v, w \in T_pM$. Exercise: Check that this is a Riemannian metric on M.

Definition 3. Two Riemannian manifolds (M, g_M) and (N, g_N) are isometric if there is a diffeomorphism $F: M \to N$ so that

$$\langle dF_p(v), dF_p(w) \rangle_{g_N} = \langle v, w \rangle_{g_M}$$

for all $p \in M$ and $v, w \in T_pM$.

In other words, (M, g_M) is isometric to (N, g_N) if there is a diffeomorphism $F: M \to N$ for which $F^*g_N = g_M$.

More examples:

- (1) \mathbb{R}^n equipped with the dot product. Here $T_p\mathbb{R}^n \cong \mathbb{R}^n$ canonically, and $g(v, w) = v \cdot w$. Once we have the machinery to make this precise, we'll see that this is our model of a *flat* space.
- (2) $\mathbb{S}^n \subset \mathbb{R}^{n+1}$ with the metric induced by the inclusion map. Concretely, we can use coordinates $(\theta^1, \dots, \theta^n) \in [0, \pi)^{n-1} \times [0, 2\pi)$ on a large patch of \mathbb{S}^n together with the map F given by

$$F(\theta^{1}, \dots, \theta^{n}) = \begin{pmatrix} \cos \theta^{1} \\ \sin \theta^{1} \cos \theta^{2} \\ \sin \theta^{1} \sin \theta^{2} \cos \theta^{3} \\ \vdots \\ \sin \theta^{1} \sin \theta^{2} \dots \sin \theta^{n-1} \cos \theta^{n} \\ \sin \theta^{1} \sin \theta^{2} \dots \sin \theta^{n-1} \sin \theta^{n} \end{pmatrix}$$

A straightforward computation shows that

$$dF_{(\theta^1,\dots,\theta^n)} = \begin{pmatrix} -\sin\theta^1 & 0 & \dots & 0 \\ \cos\theta^1\cos\theta^2 & -\sin\theta^1\sin\theta^2 & \dots & 0 \\ \cos\theta^1\sin\theta^2\cos\theta^3 & \sin\theta^1\cos\theta^2\cos\theta^3 & \dots & 0 \\ \vdots & \vdots & \ddots & \vdots \\ \cos\theta^1\sin\theta^2 \dots \sin\theta^n & \sin\theta^1\cos\theta^2 \dots \sin\theta^n & \dots & \sin\theta^1\sin\theta^2 \dots \cos\theta^n \end{pmatrix}$$

In particular, we have

$$g\left(\frac{\partial}{\partial \theta^{i}}, \frac{\partial}{\partial \theta^{j}}\right) = dF\left(\frac{\partial}{\partial \theta^{i}}\right) \cdot dF\left(\frac{\partial}{\partial \theta^{j}}\right) = \begin{cases} 0 & i \neq j \\ 1 & i = j = 1 \\ \sin^{2} \theta^{1} \sin^{2} \theta^{2} \dots \sin^{2} \theta^{j-1} & i = j \neq 1 \end{cases}$$

(This in fact homogeneous, isotropic, etc., but we haven't defined these terms.) The sphere is the nicest example of a *positively curved* space (to be made more precise later).

- (3) Hyperbolic space \mathbb{H}^n . We'll describe three models. (But there are more! Hyperbolic space just keeps on giving.) Fix some R > 0 (this is a parameter going into the metric, just as we could have changed the radius of our sphere in the previous example).
 - (a) The upper half-space $U_R = \{(x, y^1, \dots, y^{n-1}) \in \mathbb{R}^n \mid x > 0\}$ equipped with the metric

$$g = R^2 \frac{dx^2 + (dy^1)^2 + \dots + (dy^{n-1})^2}{x^2},$$

i.e., we take the inner product of two vectors $v, w \in T_{(x,y)}U_R$ by

$$\langle v, w \rangle_{(x,y)} = R^2 \frac{v \cdot w}{x^2}.$$

(b) The Poincaré ball model: Let $B_R^n \subset \mathbb{R}^n$ denote the ball of radius R and equip it with the metric

$$g = 4R^{2} \frac{(du^{1})^{2} + \dots + (du^{n})^{2}}{(R^{2} - |u|^{2})^{2}}.$$

(c) The hyperboloid model: Consider \mathbb{R}^{n+1} equipped with the pseudo-Riemannian metric $\eta_{(t,x^1,\dots,x^n)} = -(dt)^2 + \sum (dx^j)^2$ and let \mathbb{H}^n_R denote one sheet of the two sheeted hyperboloid:

$$\mathbb{H}_{R}^{n} = \{t > 0\} \cap \{-t^{2} + |x|^{2} = -R^{2}\},\,$$

and let $g = i^*\eta$, where $i : \mathbb{H}^n_R \to \mathbb{R}^{n+1}$ is the inclusion.

Theorem 4. All three of the above models of hyperbolic space are isometric.

Proof. You should fill in most of this proof yourself! I'll give you the maps to show it for the hyperboloid and ball models.

Let $S = (-R, 0, ..., 0) \in \mathbb{R}^{n+1}$ and let $P \in \mathbb{H}_R^n$, say $P = (t, x^1, ..., x^n)$. Define the map $\pi : \mathbb{H}_R^n \to B_R^n$ by letting $\pi(P) = u \in \mathbb{R}^n$, where (0, u) is the point where the line from S to P intersects $\{t = 0\}$.

Note that this line is given by

$$(-R, 0, \ldots, 0) + s(t + R, x^1, \ldots, x^n),$$

which hits t = 0 when st + sR - R = 0, i.e., when $s = \frac{R}{t+R}$, so that

$$\pi\left(t, x^{1}, \dots, x^{n}\right) = \frac{R}{t+R}(x^{1}, \dots, x^{n}).$$

As $P \in \mathbb{H}_R^n$, we have that $|x|^2 = t^2 - R^2$ and thus

$$|\pi(P)|^2 = \frac{R^2}{(t+R)^2} |x|^2 = \frac{R^2(t^2-R^2)}{(t+R)^2} = R^2 \frac{t-R}{t+R} < R^2$$

and thus $\pi(P) \in B_R^n$.

The inverse map $\pi^{-1}: B_R^n \to \mathbb{H}_R^n$ is given by

$$\pi^{-1}(z^1,\ldots,z^n) = \left(R\frac{R^2 + |z|^2}{R^2 - |z|^2}, 2\frac{R^2 z}{R^2 - |z|^2}\right).$$

You should check that this is the correct form of the inverse and that both π preserves the inner product.

4. Covariant differentiation and connections

Recall that a (smooth) k-dimensional vector bundle over a smooth manifold M consists of the data $\pi: E \to M$ so that

- (1) π is surjective,
- (2) $\pi^{-1}(p)$ is a k-dimensional vector space for each $p \in M$, and
- (3) for each $p \in M$, there is a chart (x, U) around p in M and a diffeomorphism $\varphi : \pi^{-1}(U) \to x(U) \times \mathbb{R}^k$ that restricts to a vector space isomorphism on each fiber.

One of the first challenges in differential geometry is to determine how to differentiate sections of a vector bundle. For a trivial vector bundle in Euclidean space, you have the "constant sections" and so you just differentiate their coefficients. In general, however, there is no constant section because there is no canonical way of identifying the different fibers of the bundle. One way to get around this is to define the notion of "parallel transport" along a curve. In this view, for each smooth path $\gamma:[a,b]\to M$, we equip the bundle E with linear maps $P(\gamma)_s^t: E_{\gamma(s)}\to E_{\gamma(t)}$ depending smoothly on $s,t\in[a,b]$ (and also on γ in an appropriate sense). We further demand that

$$P(\gamma)_r^t \circ P(\gamma)_s^r = P(\gamma)_s^t$$
.

The maps P provide a way of performing parallel translation along the curve γ . (If E were equipped with a way of measuring distance or angles, we'd also demand that the parallel translation preserve this.) Given these maps, we could differentiate a section Y of E at $p \in M$ in the direction $v \in T_pM$ by taking a curve $\gamma : (-\epsilon, \epsilon) \to M$ with $\gamma(0) = p$ and $\gamma'(0) = v$ and then finding

$$\nabla_v Y = \lim_{s \to 0} \frac{P(\gamma)_s^0(Y_{\gamma(s)}) - Y_p}{s}.$$

One can check that this is a derivation, but showing that in fact it depends only on v and not on the extension γ doesn't quite follow without a more careful accounting of hypotheses.

Instead of defining parallel translation directly, we instead recover it from one of the other related quantities. As is common in many differential geometry texts (especially those focusing on vector bundles like the tangent and cotangent bundles), we'll use the notion of a *Koszul connection*, which we'll just call a *connection*.

Definition 5. A connection ∇ on the vector bundle $\pi: E \to M$ is an \mathbb{R} -linear map $\nabla: \Gamma(E) \to \Gamma(T^*M \otimes E)$ so that the product rule holds, i.e.,

$$\nabla(fs) = df \otimes s + f \nabla s,$$

for all smooth functions $f \in C^{\infty}(M)$ and smooth sections $s \in \Gamma(E)$.

Unwinding this definition, it's the same as providing, for each section $s \in \Gamma(E)$ and $p \in M$, an \mathbb{R} -linear map $(\nabla s)_p : T_pM \to E_p$ so that

- (1) $(\nabla s)_p$ depends smoothly on p,
- (2) for all $a, b \in \mathbb{R}$ and $s_1, s_2 \in \Gamma(E)$, $\nabla(as_1 + bs_2)_p = a(\nabla s_1)_p + b(\nabla s_2)_p$, and
- (3) ∇ satisfies a product rule, so for all smooth functions f on M and $v \in T_pM$,

$$\nabla (fs)_p(v) = df_p(v)s_p + f(p)(\nabla s)_p(v).$$

From now on we'll drop the p subscript and let it be implicit (as $v \in T_pM$). We also typically write $(\nabla s)(v)$ as $\nabla_v s$. When v is the value of a vector field $X \in \mathcal{X}(M)$, we also write $\nabla_X s$, which is the section with value $\nabla_{X_p} s$ at p.

The following lemma tells us that connections are local and so we will not need to worry about whether sections are defined globally or only locally.

Lemma 6. If ∇ is a connection on E and $s_1, s_2 \in \Gamma(E)$ are such that $s_1 \equiv s_2$ in a neighborhood of $p \in M$, then for all $v \in T_pM$, $\nabla_v s_1 = \nabla_v s_2$.

Proof. By linearity it suffices to show that if $s \equiv 0$ in a neighborhood of p then $\nabla_v s = 0$. This statement follows from the product rule. Indeed, for any $f \in C^{\infty}(M)$ so that supp $f \subset \{s = 0\}$, we have $fs \equiv 0$ on M, and

$$0 = \nabla_v(fs) = df_p(v)s_p + f(p)\nabla_v s,$$

so that $f(p)\nabla_v s=0$. This is true for any such f, so $\nabla_v s=0$.

A word of warning: you might think that because a connection eats vector fields and gives you vector fields that it is a tensor, but the product rule (and Lemma 1) tells you that it's not. Connections do, however, lie in an affine space whose underlying linear space is the space of tensors.

Lemma 7. If ∇ and $\widetilde{\nabla}$ are two connections on the tangent bundle TM then the difference $\nabla - \widetilde{\nabla}$ is a (1,2)-tensor.

Proof. Let $T = \nabla - \widetilde{\nabla}$ be the \mathbb{R} -multilinear object $\mathcal{X} \times \mathcal{X} \times \Omega^1 \to C^{\infty}$ given by

$$T(X, Y, \omega) = \omega \left(\nabla_X Y - \widetilde{\nabla}_X Y \right).$$

By Lemma 1, we must only check that T is multilinear over C^{∞} . As it is already multilinear over C^{∞} in X and ω , we need only check in Y, but this follows from the product rule. \square

We can think of $\nabla_v s$ as denoting a directional derivative of s in the direction of v. Just as we did in calculus (and in the previous semester of this course), we'd like to also differentiate along curves. Let's fix a curve $\alpha: (a,b) \to M$.

Definition 8. A section along the curve α is a map $t \mapsto s(t) \in E_{\alpha(t)}$ depending smoothly on t. In an abuse of notation we'll denote the set of smooth sections along α by $\Gamma(E, \alpha)$.

Proposition 9. There is a unique map $\Gamma(E,\alpha) \to \Gamma(E,\alpha)$, denoted $s \mapsto \frac{D}{dt}s$ and called the covariant derivative of s along α , so that

(1)
$$\frac{D}{dt}(s_1(t) + s_2(t)) = \frac{D}{dt}s_1(t) + \frac{D}{dt}s_2(t),$$

(2)
$$\frac{D}{dt}(f(t)s(t)) = f(t)\frac{D}{dt}s(t) + f'(t)s(t)$$
, and

(3) If
$$\widetilde{s} \in \Gamma(E)$$
 satisfies $s(t) = \widetilde{s}_{\alpha(t)} \in E_{\alpha(t)}$, then $\frac{D}{dt}s(t) = \nabla_{\alpha'(t)}\widetilde{s}$.

Proof. By localization we can assume that $\alpha(I)$ is contained in a single coordinate chart on which the bundle E is trivial. We then take a local basis e_1, \ldots, e_k for all E_p for p contained in this chart and write $s(t) = \sum_{j=1}^k s^j(t)e_j$.

For uniqueness, we observe that if $\frac{D}{dt}$ satisfies all three conditions, we must have

(1)
$$\frac{D}{dt}s(t) = \sum_{j=1}^{k} \frac{D}{dt}(s^{j}(t)e_{j})$$

$$= \sum_{j=1}^{k} \left((s^{j})'(t)e_{j} + s^{j}(t)\nabla_{\alpha'(t)}e_{j} \right),$$

as e_j are defined in a neighborhood. The right side does not depend on $\frac{D}{dt}$ so we have uniqueness.

For existence, we now have a formula: we write s in terms of a local basis for the sections and use equation (1) to define the covariant derivative along α .

Another notion of connection is called an *Ehresmann connection* and involves a splitting of the tangent bundle of E into "horizontal" and "vertical" sub-bundles. In particular, there is always a canonical sub-bundle V of TE (called the "vertical bundle") given by the kernel of the pushfoward map (i.e., the differential of the projection) $\pi_*: TE \to TM$. An Ehresmann connection is the data of a "horizontal" sub-bundle complementary to the vertical one, i.e., a sub-bundle $H \subset TE$ so that $TE = H \oplus V$. The notion of connection above induces such a splitting. (To ensure that it is equivalent to the definition above involves another condition that we omit here.)

Proposition 10. A connection ∇ on E induces a splitting $TE = H \oplus V$.

Proof. We must show that ∇ defines a horizontal sub-bundle $H \subset TE$ and that TE splits as the direct sum $H \oplus V$. We start by noting that, for each $e \in E$, $V_e = T_e E_{\pi(e)} \cong E_{\pi(e)}$ because the tangent space of a vector space is canonically isomorphic to the vector space.

We now aim to define the horizontal subspace. We first define a map $K: T_eE \to E_{\pi(e)}$ and then define H_e to be the kernel of K. Given $e \in E$ and $v \in T_eE$, choose $\gamma: (-\epsilon, \epsilon) \to E$ so that $\gamma(0) = e$ and $\gamma'(0) = v$. We now regard γ as a section of E over $\pi \circ \gamma$, i.e., $\gamma \in \Gamma(E, \pi \circ \gamma)$ and set

$$Kv = \frac{D}{dt}\gamma(t)|_{t=0}.$$

We claim that Kv is independent of the choice of γ . By linearity it suffices to show that if $\gamma: (-\epsilon, \epsilon) \to E$ has $\gamma(0) = e$ and $\gamma'(0) = 0$, then $\frac{D}{dt}\gamma(t)|_{t=0} = 0$. We then note that $(\pi \circ \gamma)'(0) = 0$ and appeal to equation (1) after writing γ in terms of a local frame for E to see that indeed $\frac{D}{dt}\gamma(t)|_{t=0} = 0$.

Now, equipped with the map $K: T_eE \to E_{\pi(e)}$, we define $H_e = \ker K$. We now claim that $T_eE \cong H_e + V_e$. Indeed, note that if $v \in V_e$ is a vertical vector, we use the identification $V_e \cong E_{\pi(e)}$ to construct the curve $\gamma: (-\epsilon, \epsilon) \to E_{\pi(e)}$ given by $\gamma(t) = e + tv$. This curve satisfies $\gamma(0) = e$ and $\gamma'(0) = v$ and $\frac{D}{dt}\gamma(t)|_{t=0} = v$, so Kv = v for vertical vectors. The operator K can therefore be regarded as a projection onto V_e and so $T_eE \cong H_e \oplus V_e$.

The smoothness of the sub-bundles follows from the smoothness of the maps $(\pi)_*$ and K; that K depends smoothly on e is a consequence of the identity (1).

We now return to parallel transport. Given a curve $\alpha:[0,1]\to M$ so that $\alpha(0)=p$ and $\alpha(1)=q$, we can construct the parallel translation of a vector $v\in E_p$ along α in two related ways. One way is by solving a differential equation: we say that a section $s\in\Gamma(E,\alpha)$ is parallel if and only if $\frac{D}{dt}s(t)=0$ for all $t\in[0,1]$.

Lemma 11. For every $v \in E_p$, there is a unique $s \in \Gamma(E, \alpha)$ so that s(0) = v and s is parallel.

Proof. Working locally in charts where the bundle is trivial, this again follows from the identity (1), this time interpreted as a linear system of differential equations for the coefficients of the frame. It is not hard to check that existence and uniqueness for ODEs then guarantees a solution.

We then define the parallel translate of v by $P(\alpha)_0^1 v = s(1)$. Note that this value typically depends on the choice of path!

Another way to define the parallel translate of v along α to lift α to a curve $\widetilde{\alpha}:[0,1]\to E$ so that $\pi\circ\widetilde{\alpha}=\alpha$, $\alpha(0)=(p,v)$ and so that $\widetilde{\alpha}'(t)\in H_{\widetilde{\alpha}(t)}$ for all t. The existence of such a lift follows from the observation that $H_e\cong T_{\pi(e)}M$ for all $e\in E$ and the decomposition of $TE=V\oplus H$. (You need $\pi_*\widetilde{\alpha}'(t)=\alpha'(t)$ and you need $\widetilde{\alpha}'(t)$ to be horizontal.)

We'll return to the question of why parallel transport is so called once we start talking about specific connections.

- 4.1. **Induced connections.** A connection ∇ on a vector bundle E over M induces connections over other bundles formed from E. A few examples:
 - (1) Dual bundles. If ∇ is a connection on E, we get a connection ∇^* (we'll later just call this ∇) on the dual bundle E^* by duality. Indeed, if $\xi \in \Gamma(E^*)$, we define $\nabla_v \xi$ by demanding that, for all $s \in \Gamma(E)$,

$$d(\langle \xi, s \rangle)(v) = \langle \nabla_v^* \xi, s \rangle + \langle \xi, \nabla_v s \rangle.$$

(2) If ∇^E and ∇^F are connections on vector bundles E and F over M, then we get a connection $\nabla^E \otimes \nabla^F$ on the vector bundle $E \otimes F$ over M by demanding it satisfy a product rule:

$$\left(\nabla^E \otimes \nabla^F\right)_v(s \otimes t) = \nabla^E_v s \otimes t + s \otimes \nabla^F_v t.$$

(3) A similar construction gives a connection on the exterior powers $\Lambda^k E$ by a product rule:

$$\nabla_v \left(s_1 \wedge \cdots \wedge s_k \right) = \nabla_v s_1 \wedge \cdots \wedge s_k + \cdots + s_1 \wedge \cdots \wedge \nabla_v s_k.$$

(4) Similarly we get a connection $\nabla^E \oplus \nabla^F$ on the direct sum bundle $E \oplus F$ by linearity:

$$(\nabla^E \oplus \nabla^F)_v(s \oplus t) = (\nabla^E_v s) \oplus (\nabla^F_v t)$$
.

(5) By identifying the endomorphisms of E with $E^* \otimes E$ we also get a connection on the endomorphism bundle $\operatorname{End}(E)$.

As a result, if we have a connection on the vector bundle TM then we in fact have a connection on all of the tensor bundles.

4.2. The Levi-Civita connection. We now specialize to the case where E = TM and its associated vector bundles.

Definition 12. Given a connection ∇ on TM, its torsion is given by

$$T(\nabla)(X,Y) = \nabla_X Y - \nabla_Y X - [X,Y] \in C^{\infty}(M)$$

where $X, Y \in \mathcal{X}(M)$ and [X, Y] is the Lie bracket.

Lemma 13. The torsion of a connection on TM is an antisymmetric (0,2)-tensor.

Proof. It's clearly antisymmetric; that it is a tensor follows from the characterization of tensor fields given by Lemma 1. Indeed, we have to check that it is multilinear over $C^{\infty}(M)$. Take $f \in C^{\infty}(M)$, $X, Y \in \mathcal{X}(M)$ and consider

$$T(\nabla)(fX,Y) = \nabla_{fX}Y - \nabla_{Y}(fX) - [fX,Y]$$

= $f\nabla_{X}Y - (Y(f)X + f\nabla_{Y}X) - (f[X,Y] - Y(f)X)$
= $fT(\nabla)(X,Y)$.

Theorem 14. Suppose (M,g) is a Riemannian manifold. There is a unique connection ∇ so that

- (1) ∇ is torsion-free (i.e., $\nabla_X Y \nabla_Y X [X, Y] = 0$), and
- (2) ∇ is compatible with the metric g, i.e., for all $X, Y, Z \in \mathcal{X}(M)$,

$$X (g(Y,Z)) = g (\nabla_X Y, Z) + g (Y, \nabla_X Z).$$

Definition 15. The unique torsion-free connection compatible with g is called the *Levi-Civita connection*.

Proof. To prove uniqueness, suppose that ∇ is torsion-free and compatible with g. Then compatibility yields

$$X \langle Y, Z \rangle + Y \langle Z, X \rangle - Z \langle X, Y \rangle - \langle X, [Y, Z] \rangle + \langle Y, [Z, X] \rangle + \langle Z, [X, Y] \rangle$$

$$= \langle \nabla_X Y, Z \rangle + \langle Y, \nabla_X Z \rangle + \langle \nabla_Y Z, X \rangle + \langle Z, \nabla_Y X \rangle - \langle \nabla_Z X, Y \rangle - \langle X, \nabla_Z Y \rangle$$

$$- \langle X, [Y, Z] \rangle + \langle Y, [Z, X] \rangle + \langle Z, [X, Y] \rangle,$$

while ∇ being torsion-free lets us write, e.g., $\nabla_X Z - \nabla_Z X = [X, Z]$, so that the above expression is equal to (here using that g is symmetric)

$$\begin{split} \langle \nabla_X Y, Z \rangle + \langle Y, [X, Z] \rangle + \langle X, [Y, Z] \rangle \\ + \langle Z, \nabla_X Y + [Y, X] \rangle - \langle X, [Y, Z] \rangle - \langle Y, [Z, X] \rangle + \langle Z, [X, Y] \rangle \\ = 2 \langle \nabla_X Y, Z \rangle \,. \end{split}$$

In other words, we end up with the equality

$$\langle \nabla_X Y, Z \rangle = \frac{1}{2} \left[X \langle Y, Z \rangle + Y \langle Z, X \rangle - Z \langle X, Y \rangle - \langle X, [Y, Z] \rangle + \langle Y, [Z, X] \rangle + \langle Z, [X, Y] \rangle \right].$$

Because g is non-degenerate and this relationship must hold for all Z, any two choices of torsion-free metric-compatible connection must agree.

Similarly, because g is non-degenerate, the formula above also defines the connection ∇ , so we also get existence. (You should check that it is tensorial in X and Z and satisfies a product rule in Y and should also check that it is torsion-free and metric-compatible.)

How does this connection look in local coordinates? Let (x^1, \ldots, x^n) be coordinates in a chart on M and ∂_i denote the corresponding basis for the tangent space. We then have

$$\langle \nabla_{\partial_i} \partial_j, \partial_k \rangle = \frac{1}{2} \left[\partial_i \langle \partial_j, \partial_k \rangle + \partial_j \langle \partial_i, \partial_k \rangle - \partial_k \langle \partial_i, \partial_j \rangle \right]$$
$$= \frac{1}{2} \left[\partial_i g_{jk} + \partial_j g_{ik} - \partial_k g_{ij} \right].$$

If we write $\nabla_{\partial_i}\partial_j = \sum_{\ell} \Gamma_{ij}^{\ell} \partial_{\ell}$, then

$$\sum_{\ell=1}^{n} \Gamma_{ij}^{\ell} g_{\ell k} = \sum_{\ell=1}^{n} \Gamma_{ij}^{\ell} \langle \partial_{\ell}, \partial_{k} \rangle$$
$$= \frac{1}{2} (\partial_{i} g_{jk} + \partial_{j} g_{ik} - \partial_{k} g_{ij}).$$

This is an $n \times n$ system of linear equations, which we solve to find

$$\Gamma_{ij}^{k} = \sum_{\ell=1}^{n} \frac{1}{2} g^{k\ell} \left(\partial_{i} g_{j\ell} + \partial_{j} g_{i\ell} - \partial_{\ell} g_{ij} \right),$$

where $g^{k\ell}$ are the components of the inverse metric (i.e., the inverse of the matrix (g_{ij}) or the components of the induced metric on the cotangent bundle, etc.).

Definition 16. The Γ_{ij}^k are called *Christoffel symbols* or *connection coefficients*.

Word of warning: Γ_{ij}^k are NOT the components of a tensor (they lie in an affine space). As described above, ∇ induces a connection on the cotangent bundle, given for $\omega \in \Omega^1(M)$ by

$$(\nabla_v \omega)(X) = v(\omega(X)) - \omega(\nabla_v X).$$

For more general tensors it has an analogous form, with $\nabla_v A$ of a (k,ℓ) -tensor given by

$$(\nabla_{v}A)(\omega_{1},\ldots,\omega_{k},X_{1},\ldots,X_{\ell}) = v\left(A(\omega_{1},\ldots,\omega_{k},X_{1},\ldots,X_{\ell})\right)$$

$$-\sum_{i=1}^{k}A\left(\omega_{1},\ldots,\omega_{i-1},\nabla_{v}\omega_{i},\omega_{i+1},\ldots,\omega_{k},X_{1},\ldots,X_{\ell}\right)$$

$$-\sum_{j=1}^{\ell}A\left(\omega_{1},\ldots,\omega_{k},X_{1},\ldots,X_{j-1},\nabla_{v}X_{j},X_{j+1},\ldots,X_{\ell}\right).$$

Proposition 17. Parallel transport using the Levi-Civita connection is an isometry.

Proof. Suppose $\alpha:[0,1]\to M$ is smooth and $V,W\in\mathcal{X}(\alpha)$. Then

$$\frac{d}{dt} \langle V(t), W(t) \rangle = \left\langle \frac{D}{dt} V(t), W(t) \right\rangle + \left\langle V(t), \frac{D}{dt} W(t) \right\rangle,$$

so that if V and W are parallel then their inner product is preserved.

5. Geodesics and Hamiltonian flows

Most differential geometry textbooks use the connection directly to define and reason about geodesics. We'll instead take a Hamiltonian approach. To do that, we need some preliminaries about the symplectic structure on the cotangent bundle.

5.1. Symplectic manifolds.

Definition 18. A manifold (M, ω) is called a symplectic manifold if M is a smooth manifold and ω is a non-degenerate closed 2-form on M.

In other words, at each point ω is an alternating (0,2)-tensor so that $d\omega = 0$ and, if $v \in T_pM$ satisfies

$$\omega(v, u) = 0$$
 for all $u \in T_n M$,

then v = 0.

Lemma 19. A symplectic manifold must be even-dimensional.

Proof. Working in local coordinates, this reduces to the statement that if n is odd, then any skew-symmetric $n \times n$ real matrix must have a kernel. To see this, a skew-symmetric matrix A has

$$\det A = \det(A^{\mathsf{T}}) = \det(-A) = (-1)^n \det A,$$

so that $\det A = 0$ if n is odd and thus 0 is an eigenvalue of A, i.e., A must have a kernel. \Box

There is a lot to say about symplectic manifolds, most of which we omit here. One of the most famous is Darboux's theorem, which essentially says that there are no local invariants of symplectic manifolds, i.e., one can always find local coordinates (q, p) so that $\omega = \sum_i dq^i \wedge dp_i$. There's a lovely proof of this using Moser's trick which I'm happy to talk about if you like:

Theorem 20 (Darboux). Suppose (M, ω) is a symplectic 2k-dimensional manifold. Around any point $p \in M$ there is a coordinate chart U and coordinates $(x^1, \ldots, x^k, y^1, \ldots, y^k)$ so that

$$\omega|_{U} = \sum_{i=1}^{k} dx^{i} \wedge dy^{i}.$$

Proof. You can always find a coordinate system $(\widetilde{x}^1, \ldots, \widetilde{x}^k, \widetilde{y}^1, \ldots, \widetilde{y}^k)$ achieving this exactly at the point p. We claim then that there is a local diffeomorphism ϕ of a smaller neighborhood so that

$$\phi^* \left(\sum d\widetilde{x}^i \wedge d\widetilde{y}^i \right) = \omega|_U,$$

and then the desired coordinate system is $x = \tilde{x} \circ \phi$ and $y = \tilde{y} \circ \phi$.

It remains to prove the claim. We first claim that if ω_t is a family of symplectic forms so that $\frac{d}{dt}\omega_t = d\sigma_t$ for some 1-forms σ_t , then there is a family of diffeomorphisms ψ_t so that $\psi_t^*\omega_t = \omega_0$. Indeed, this follows from Lemma 21 below by using Cartan's magic formula

$$\mathcal{L}_X \omega = d \circ i_X + i_X \circ d,$$

where i_X denotes the interior product (i.e., plugging X into the first slot) and observing that the differential equation in Lemma 21 becomes

$$0 = \frac{d}{dt}\omega_t + d(i_X\omega_t) + i_X(d\omega_t) = d(\sigma_t + i_X\omega_t).$$

As ω_t is non-degenerate, one can find X so that $\sigma_t = i_X \omega_t$ and so Lemma 21 applies.

Finally, by the Poincaré lemma (which says that on \mathbb{R}^n , closed forms are exact), the difference $\omega_t - \omega_0 = d\sigma_t$ and so we can apply the claim.

Lemma 21 (Moser's trick). Suppose ω_t , $t \in [0,1]$ is a family of differential forms on M. If there is a solution $X_t \in \mathcal{X}(M)$, $t \in [0,1]$ to the differential equation

$$\frac{d}{dt}\omega_t + \mathcal{L}_{X_t}\omega_t = 0,$$

where \mathcal{L}_X denotes the Lie derivative with respect to X, then there exists a family of diffeomorphisms ψ_t on M so that $\psi_t^* \omega_t = \omega_0$ and $\psi_0 = \operatorname{Id}$.

Proof. Given X_t , let ψ_t be the flow it generates, so

$$\frac{d}{dt}(\psi_t^*\omega_t) = \psi_t^* \left(\frac{d}{dt}\omega_t + \mathcal{L}_{X_t}\omega_t\right) = 0,$$

so
$$\psi_t^* \omega_t = \psi_0^* \omega_0 = \omega_0$$
.

Now, back on track. One of the main things we want to use symplectic structures for is to get *Hamilton vector fields*. Just as we had with Riemannian metrics, we can use the nondegenerate 2-form ω to identify the tangent and cotangent spaces at each point. Indeed, given any 1-form σ , we can find some vector field X associated to it by demanding that

$$\omega(v,X) = \sigma_p(v)$$

for all vectors $v \in T_pM$. In particular, if we have a real-valued function H (called a Hamiltonian) on M, we can find the *Hamilton vector field* of H, which we'll denote X_H , by demanding that X_H be the vector field associated to dH by ω .

As smooth vector fields yield flows, we therefore also obtain a flow ϕ_t from the Hamilton vector field X_H .

Lemma 22. The Hamiltonian H is conserved by the flow ϕ_t .

Proof. This is an exercise in unraveling the definitions and using the chain rule. Indeed, suppose $p \in M$ and let $\gamma(t) = \phi_t(p)$, so that

$$\gamma'(t) = (X_H)_{\gamma(t)},$$

$$\gamma(0) = p.$$

We then differentiate $H(\gamma(t))$:

$$\frac{d}{dt}H(\gamma(t)) = dH(\gamma'(t))$$
$$= dH(X_H) = \omega(X_H, X_H) = 0,$$

so that $H(\gamma(t)) = H(\gamma(0))$.

5.2. A very brief foray into Hamiltonian mechanics. One of the most important examples of a symplectic manifold is the cotangent bundle T^*M of a smooth manifold M. (Note that the dimension of T^*M is always twice the dimension of M and therefore even.) If $\pi: T^*M \to M$ denotes the projection, we can define the *canonical* 1-form α on T^*M by

$$\alpha_{(x,\xi)}(v) = \xi(\pi_* v)$$

where $(x,\xi) \in T^*M$ and $v \in T_{(x,\xi)}(T^*M)$. In other words, the form α acts on a vector v at a point (x,ξ) in the cotangent bundle by evaluating the covector ξ (a covector on M) on the pushvorward of v. In terms of local coordinates (x,ξ) , you can check that

$$\alpha = \sum_{j} \xi_{j} dx^{j}.$$

We then define a symplectic form ω by $\omega = d\alpha$. In a local coordinate system ω has the form

$$\omega = \sum_{j} d\xi_{j} \wedge dx^{j}.$$

It is plainly a 2-form, $d\omega = d(d\alpha) = 0$, and there are several ways to check that it is non-degenerate. One way is to observe that ω^n , the *n*-th wedge power of ω , is a non-vanishing volume form. We can also check it directly by taking $v \in T_{(x,\xi)}(T^*M)$ with $\omega(v,\bullet) = 0$. Writing v in terms of the basis given by the coordinate system, we write

$$v = \sum_{j} v^{j} \frac{\partial}{\partial x^{j}} + \sum_{j} w^{j} \frac{\partial}{\partial \xi_{j}},$$

so that

$$\omega(v, \bullet) = -\sum_{j} v^{j} d\xi_{j} + \sum_{j} w^{j} dx^{j},$$

so that we must have $v^j = w^j = 0$, i.e., v = 0.

¹Recall that a coordinate system x on M induces a coordinate chart (x, ξ) on T^*M by writing covectors in terms of the basis dx^J of each cotangent space; ξ_j are the coefficients here.

From a physical perspective, the symplectic manifold (T^*M, ω) is thought of as a "phase space" for a physical system taking place on the "configuration space" M while a Hamiltonian H is the total energy of the system. If $M = \mathbb{R}$ and

$$H(x,\xi) = \frac{1}{2m} |\xi|^2 + V(x),$$

(i.e., H is kinetic energy plus potential energy), then

$$dH = \frac{1}{m}\xi \, d\xi + V'(x)dx,$$

so that

$$X_{H} = \frac{1}{m} \xi \frac{\partial}{\partial x} - V'(x) \frac{\partial}{\partial \xi},$$

and thus the integral curves of the flow generated by X_H satisfy

$$\frac{dx}{dt} = \frac{1}{m}\xi, \quad \frac{d\xi}{dt} = -V'(x).$$

In particular, x satisfies the second-order differential equation

$$x''(t) = -\frac{1}{m}V'(x(t)),$$

which you might recognize as Newton's second law for the conservative force given by the potential V(x).

5.3. **Geodesics.** Suppose now that (M,g) is a Riemannian manifold and (in a mild abuse of notation) let g^{-1} denote the induced inner product on each cotangent space. Recall from the last section that T^*M is always a symplectic manifold and consider the Hamiltonian function

$$H(x,\xi) = \frac{1}{2} |\xi|_{g^{-1}}^2 = \frac{1}{2} \sum_{i,j} g^{ij}(x) \xi_i \xi_j.$$

By the discussion above, we associate to H a Hamilton vector field X_H and a (very important!) flow ϕ_t .

Definition 23. We say that ϕ_t is the geodesic flow on the cotangent bundle and the integral curves of X_H are called lifted geodesics. If $\gamma(t)$ is a lifted geodesic, its projection $\pi \circ \gamma$ to M is called a geodesic.

In local coordinates (x, ξ) on the cotangent bundle, the Hamilton vector field of H is given by

$$X_{H} = \sum_{i} g^{ii}(x)\xi_{i}\partial_{x_{i}} + \frac{1}{2}\sum_{i\neq j} g^{ij}(x)\xi_{i}\partial_{x_{j}} - \frac{1}{2}\frac{\partial g^{ij}(x)}{\partial x_{k}}\xi_{i}\xi_{j}\partial_{\xi_{k}}$$
$$= w_{\xi} - \frac{1}{2}\frac{\partial g^{ij}(x)}{\partial x_{k}}\xi_{i}\xi_{j}\partial_{\xi_{k}},$$

where, in another abuse of notation, w_{ξ} is the vector field on M (regarded as a vector field on T^*M) associated to ξ by the metric g. In particular, the integral curves $(x(t), \xi(t))$ of X_H satisfy

$$\frac{dx}{dt} = w_{\xi}, \quad \frac{d\xi}{\tilde{d}} = -\frac{1}{2} \sum \frac{\partial g^{ij}(x(t))}{\partial \xi} \xi_i \xi_j.$$

Before we get to examples, let's $connect^2$ this computation with our discussion of connections from earlier.

Lemma 24. A curve $\widetilde{\gamma}:(a,b)\to T^*M$ is a lifted geodesic if and only if $\gamma=\pi\circ\widetilde{\gamma}$ satisfies the geodesic equation, namely,

$$\frac{D}{dt}\gamma'(t) = 0$$

along γ .

In other words, the tangent vector of a geodesic is parallel along the geodesic. This second order equation is the typical way to introduce geodesics; the Hamiltonian formulation is akin to turning a second order equation into a system of first order equations.

Proof. The easiest/fastest way to verify the lemma is to write down the differential equations in local coordinates. Indeed, fix a coordinate system (x^1, \ldots, x^n) .

The coordinates of the lifted geodesic satisfy the following system of equations:

$$\frac{dx^k}{dt} = \frac{\partial}{\partial \xi_k} \left(\frac{1}{2} \sum_{i,j} g^{ij} \xi_i \xi_j \right),$$

$$\frac{d\xi_k}{dt} = -\frac{\partial}{\partial x_k} \left(\frac{1}{2} g^{ij} \xi_i \xi_j \right).$$

Observe that, for fixed k, we may write the sum $g^{ij}\xi_i\xi_j$ as

$$g^{ij}\xi_{i}\xi_{j} = \sum_{i,j\neq k} g^{ij}\xi_{i}\xi_{j} + \sum_{j\neq k} (g^{jk} + g^{kj})\xi_{j}\xi_{k} + g^{kk}\xi_{k}\xi_{k},$$

so that

$$\frac{\partial}{\partial x^k} \left(\frac{1}{2} g^{ij} \xi_i \xi_j \right) = \sum_{j=1}^n g^{jk} \xi_j.$$

In particular, this is the k-th component of the vector associated to ξ . The first half of the differential equation then reads that $\gamma'(t) = v_{\xi}$. We therefore introduce the "vector variables" $v^j = g^{jk}\xi_k$, so that the first half becomes $\frac{dx^j}{dt} = v^j(t)$.

We now turn our attention to the second half. The right side of the second part of the differential equation has the form

$$-\frac{1}{2}\frac{\partial g^{ij}}{\partial x_k}\xi_i\xi_j,$$

which we now rewrite using the fact that g^{ij} are the components of the inverse matrix of g. Recall that for a family of invertible matrices A(s), we have $A(t)A^{-1}(s) = I$, so that $\frac{dA}{ds}A^{-1} + A\frac{dA^{-1}}{ds} = 0$, i.e.,

$$\frac{dA^{-1}}{ds} = -A^{-1}\frac{dA}{ds}A^{-1}.$$

In particular, each component of the matrix must agree. Applying this observation to g^{ij} , we find that

$$-\frac{1}{2}\frac{\partial g^{ij}}{\partial x^k} = \frac{1}{2}g^{i\ell}\frac{\partial g_{\ell m}}{\partial x^k}g^{mj},$$

 $^{^{2}}HA!$

so that the second half of the equation reads

$$\frac{d\xi_k}{dt} = \frac{1}{2}g^{i\ell}(x(t))\frac{\partial g_{\ell m}(t)}{\partial x^k}g^{mj}(t)\xi_i\xi_j.$$

We now aim to turn this into an equation in terms of x and v. We multiply both sides by g^{rk} , sum, and add $\xi_k \frac{g^{rk}}{dt}$ to both sides. Observe that

$$\frac{dg^{rk}}{dt} = \sum_{s} \frac{\partial g^{rk}}{\partial x^{s}} \frac{dx^{s}}{dt} = -g^{ra} \frac{\partial g_{ab}}{\partial x^{s}} g^{bk} \frac{dx^{s}}{dt},$$

so that the equation reads (after using the first half for dx^s/dt)

$$\frac{d}{dt}\left(g^{rk}\xi_k\right) = \frac{1}{2}g^{rk}\frac{\partial g_{\ell m}}{\partial x^k}g^{i\ell}g^{mj}\xi_i\xi_j - g^{ra}\frac{\partial g_{ab}}{\partial x^s}g^{bk}g^{sj}\xi_j\xi_k.$$

In terms of v^j , the equations then read (after re-indexing)

$$\begin{split} \frac{dx^k}{dt} &= v^k, \\ \frac{dv^k}{dt} &= v^i v^j g^{k\ell} \left(\frac{1}{2} \frac{\partial g_{ij}}{\partial x^\ell} - \frac{\partial g_{i\ell}}{\partial x^j} \right) \\ &= -\frac{1}{2} v^i v^j g^{k\ell} \left(\frac{\partial g_{i\ell}}{\partial x^j} + \frac{\partial g_{j\ell}}{\partial x^i} - \frac{\partial g_{ij}}{\partial x^\ell} \right) = -v^i v^j \Gamma^k_{ij}. \end{split}$$

Notice that this equation is the first-order rewriting of the following system of second-order equations:

$$\frac{d^2x^k}{dt^2} + \Gamma_{ij}^k(x(t)) \left(\frac{dx^i}{dt}\right) \left(\frac{dx^j}{dt}\right) = 0.$$

We now turn to the other system of equations, namely

$$\frac{D}{dt}\gamma'(t) = 0.$$

As $\gamma'(t) = \frac{dx}{dt}$, we can regard this as a second-order equation of the form (where we have written out the basis elements explicitly)

$$\frac{D}{dt} \left(\frac{dx^k}{dt} \frac{\partial}{\partial x^k} \right) = 0.$$

By our construction of the covariant derivative, we have

$$\frac{D}{dt} \left(\frac{dx^k}{dt} \frac{\partial}{\partial x^k} \right) = \frac{d^2 x^k}{dt^2} \frac{\partial}{\partial x^k} + \frac{dx^k}{dt} \nabla_{\frac{dx^j}{dt} \frac{\partial}{\partial x^j}} \frac{\partial}{\partial x^k}
= \frac{d^2 x^k}{dt^2} \frac{\partial}{\partial x^k} + \frac{dx^k}{dt} \frac{dx^j}{dt} \Gamma_{jk}^{\ell} \frac{\partial}{\partial x^{\ell}}.$$

After re-indexing, we see that the equation reads

$$\frac{d^2x^k}{dt^2}\frac{\partial}{\partial x^k} + \frac{dx^i}{dt}\frac{dx^j}{dt}\Gamma^k_{ij}\frac{\partial}{\partial x^k},$$

i.e., the same system as above.

In fact, we can recover the entire connection from the symplectic structure and the Hamiltonian. For convenience we discuss only how to recover the splitting of $T_{(x,\xi)}(T^*M)$ into horizontal and vertical subspaces. (The following discussion is adapted from Robert Bryant's answer here: https://mathoverflow.net/questions/127319/intuition-for-levi-civita-connection-Indeed, we regard the Hamiltonian function $g^{ij}\xi_i\xi_j$ in two different respects. The first is that it provides our Hamiltonian vector field X_H and the second is that it induces a quadratic form γ_H on each $T_{(x,\xi)}T^*M$. In our Riemannian case, the quadratic form it induces is simply π^*g , where g is regarded as a quadratic form on each T_xM . We then consider the symmetric quadratic form

$$\dot{\gamma}_H = \mathcal{L}_{X_H} \gamma_H.$$

A quick computation in coordinates shows that in terms of the basis $\partial/\partial x^j$, $\partial/\partial \xi_k$, $\dot{\gamma}_H$ corresponds to a matrix

$$\begin{pmatrix} * & I_n \\ I_n & 0 \end{pmatrix},$$

where I_n is the $n \times n$ identity matrix and the entries in the upper left can be computed in terms of Christoffel symbols but are irrelevant here. As a result, $\dot{\gamma}_H$ is non-degenerate and has n positive and n negative eigenvalues. Using this form (and its matrix realization), we can see that at each (x,ξ) there is a unique n-plane that is copmlementary to $V_{(x,\xi)}$, null with respect to the symplectic form ω , and also null with respect to $\dot{\gamma}_H$. These n-planes fit smoothly into a bundle, which a coordinate computation shows is the same horizontal bundle as the one associated to the Levi-Civita connection.

5.4. The exponential map. Back to the main thread, given any point $(x, \xi) \in T^*M$, there is a unique lifted geodesic through that point, namely, $t \mapsto \phi_t(x, \xi)$. From our ODE discussion last semester, we know that it also has a maximal connected interval of existence. By using the Riemannian metric to identify tangent and cotangent vectors, we then have

Proposition 25. Given any point $p \in M$ and any $v \in T_pM$, there is an open interval I containing 0 and a unique geodesic $\gamma: I \to M$ so that $\gamma(0) = p$ and $\gamma'(0) = v$.

It's traditional to let the exponential map denote the mapping from the tangent space/bundle (rather than cotangent space/bundle) to the manifold, so we'll do that. For a point $(p, v) \in TM$, let $\gamma_{(p,v)}$ denote the unique geodesic with $\gamma(0) = p$ and $\gamma'(0) = v$, and let $I_{(p,v)}$ denote its maximal interval of existence.³

Definition 26. Define the exponential map $\exp(p,v) = \gamma_{(p,v)}(1)$ if $1 \in I_{(p,v)}$. Define the restricted exponential map $\exp_p(v) = \gamma_{(p,v)}(1)$ if $1 \in I_{(p,v)}$.

Let's let $\mathcal{E} \subset TM$ denote the domain of exp, i.e., $(p, v) \in \mathcal{E}$ if and only if $1 \in I_{(p,v)}$.

Proposition 27. (1) \mathcal{E} is open, contains the zero section, and is star-shaped.

- (2) For each $(p,v) \in TM$, the geodesic $\gamma_{(p,v)}$ is given by $\gamma_{(p,v)}(t) = \exp_p(tv)$ whenever either side is defined.
- (3) $\exp is smooth$.

Lemma 28. For any $(p, v) \in TM$, and $c, t \in \mathbb{R}$, $\gamma_{(p,cv)}(t) = \gamma_{(p,v)}(ct)$ whenever either side is defined.

³In keeping with the view of the author, we should instead define the exponential map from T^*M to M by $\exp(x,\xi) = \phi_1(x,\xi)$ whenever it is defined. The restricted exponential map would then be $\exp_x(\xi) = \phi_1(x,\xi)$, but we do not pursue this here for ease of reading other texts.

Proof. It's enough to show that $\gamma_{(p,cv)}(t)$ exists and equality holds whenever $\gamma_{(p,v)}(ct)$ is defined. (Replace v by $c^{-1}v$ and t by $c^{-1}t$ for the other direction.)

Suppose I is the maximal interval of definition for $\gamma = \gamma_{(p,v)}$, and define

$$\widetilde{\gamma}(t) = \gamma(ct).$$

The maximal interval of definition for $\tilde{\gamma}$ is

$$c^{-1}I = \{t : ct \in I\}.$$

We now observe that $\widetilde{\gamma}$ satisfies the geodesic equation (by the chain rule) with $\widetilde{\gamma}(0) = p$ and $\widetilde{\gamma}'(0) = cv$ (so that $\widetilde{\gamma} = \gamma_{(p,cv)}$).

Proof of proposition. By the lemma we have that

$$\exp(p, cv) = \gamma_{(p,cv)}(1) = \gamma_{(p,v)}(c),$$

which proves the second statement. If $v \in \mathcal{E}_p$ (where $\mathcal{E}_p = \{v \in T_pM : (p,v) \in \mathcal{E}\}$), then $\gamma_{(p,v)}$ is defined at least on [0,1]. Then, for $t \in [0,1]$, we have $\exp_p(tv) = \gamma_{(p,tv)}(1) = \gamma_{(p,v)}(t)$ and thus \mathcal{E} is star-shaped.

We still need to show that \mathcal{E} is open and that exp is smooth. For now we use the notation $e: TM \to T^*M$ to denote the bundle isomorphism given by "lowering indices". By Lemma 24, we see that if $\gamma(t)$ is a geodesic, then

$$t \mapsto (\gamma(t), e(\gamma'(t)))$$

is the lifted geodesic that projects to γ . By the definition of the flow ϕ_t , we then have

$$\left(\gamma_{(p,v)}(t), e\left(\gamma'_{(p,v)}(t)\right)\right) = \phi_t\left(p, e(v)\right).$$

By the proof last semester⁴ of existence and uniqueness of ODEs, there is an open neighborhood U of $\{0\} \times T^*M$ in $\mathbb{R} \times T^*M$ on which $\phi_{\bullet}(\bullet, \bullet)$ is defined and the smooth dependence on parameters shows that ϕ is smooth where it is defined.

So, if $(p, v) \in \mathcal{E}$, the geodesic $\gamma_{(p,v)}$ is defined at least on [0, 1], so $(1, (p, e(v))) \in U$ and it has an open neighborhood around it, so there is an open neighborhood around (p, v) for which the flow exists for all $t \in [0, 1]$, so \mathcal{E} is open.

Finally, $\exp(p, v) = \pi \circ \phi_1(p, e(v))$, so exp is smooth as a composition of smooth functions.

Lemma 29. For any $p \in M$, there is a neighborhood $V \subset T_pM$ of $0 \in T_pM$ and a $U \subset M$ so that $\exp_p : V \to U$ is a diffeomorphism.

Proof. This follows from the inverse function theorem as soon as we show that the derivative of the restricted exponential map at 0 is invertible. As T_pM is a vector space, we have $T_0(T_pM) \cong T_pM$ canonically so we can think of

$$(D\exp_p)_0: T_pM \to T_pM.$$

Let $v \in T_pM$ and take $\tau : (-\epsilon, \epsilon) \to T_0(T_pM) \cong T_pM$ to be $\tau(t) = tv$, so that $\tau(0) = 0$ and $\tau'(0) = v$. We now compute

$$(D \exp_p)_0(v) = \frac{d}{dt}|_{t=0} \exp_p \circ \tau(t)$$
$$= \frac{d}{dt}|_{t=0} \gamma_{(p,v)}(t) = v,$$

⁴Soon to be recreated and improved in Appendix A.

so indeed $(D \exp_p)_0 = \operatorname{Id}_{T_p M}$ and is therefore invertible.

6. Curvature

One question we might ask is to what degree we can make a Riemannian metric "look like" the Euclidean metric at a point. Can we make its value agree at a given point? (Yes.) Its first derivatives? (Again, yes.) What about its second derivatives?

6.1. **Geodesic normal coordinates.** The invertibility of the exponential map at the origin gives a way of picking distinguished coordinates at a point. Given $p \in M$, let e_1, \ldots, e_n be an orthonormal basis for T_pM and consider the map $V \to M$, where $V \subset \mathbb{R}^n$, given by

$$(x^1,\ldots,x^n)\mapsto \exp_p\left(x^1e_1+\cdots+x^ne_n\right).$$

This map is a diffeomorphism from a neighborhood of $0 \in \mathbb{R}^n$ to a neighborhood of p and so it gives a local coordinate system called *geodesic normal coordinates*.

What does the metric look like in these coordinates?

We first note that $p \leftrightarrow (0, ..., 0)$ and then claim that, in this coordinate system, $g_{ij}(0) = \delta_{ij}$. Indeed, we computed earlier that $(D \exp_p)_0 = \operatorname{Id}_{T_p M}$, so

$$g_{ij}(x) = \left\langle \frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j} \right\rangle$$
$$= \left\langle (D \exp_p)_{\sum x^k e_k}(e_i), (D \exp_p)_{\sum x^k e_k}(e_j) \right\rangle,$$

so at x = 0, this is $g_{ij}(0) = \langle e_i, e_j \rangle = \delta_{ij}$.

We now turn our attention to the Christoffel symbols. Fix $(x^1, x^2, ..., x^n) \in V$ and take $\alpha(t) = (tx^1, ..., tx^n)$; α is a geodesic in U because $\alpha(t) = \exp_p(t(x^1, ..., x^n))$ (in this coordinate system). Because α is a geodesic, we have

$$\frac{d^2}{dt^2}\alpha^k(t) + \sum_{i,j=1}^n \Gamma_{ij}^k(\alpha(t)) \frac{d\alpha^i}{dt}(t) \frac{d\alpha^j}{dt}(t) = 0$$

for all k and all t sufficiently small. As we can see that $\frac{d^2}{dt^2}\alpha(t)=0$, we see that

$$\sum_{i,j=1}^{n} \Gamma_{ij}^{k}(\alpha(t)) \frac{d\alpha^{i}}{dt}(t) \frac{d\alpha^{j}}{dt}(t) = 0.$$

Setting t = 0, we have that

$$\sum_{ij} \Gamma_{ij}^k(0) x^i x^j = 0$$

for all k and all x. As it is true for all x, we may apply it to $x = e_i$ to see that $\Gamma_{ii}^k(0) = 0$ for all i. We then apply it with $x = e_i + e_j$ and see that $\Gamma_{ij}^k(0) + \Gamma_{ji}^k(0) = 0$ and thus $\Gamma_{ij}^k(0)$ is antisymmetric in i, j. On the other hand, we also knew that $\Gamma_{ij}^k(0)$ was symmetric in i, j (indeed, this follows from the coordinate formula for Γ_{ij}^k), so they must all vanish at 0.

What does this fact about the Christoffel symbols mean for the first derivative of the metric? Well, at x = 0, we have $\nabla_{\partial_k} \partial_i = 0$, so $\partial_k g_{ij}|_{x=0} = 0$, i.e., $g_{ij}(x) = \delta_{ij} + \text{quadratic}$ and higher terms.

What do these quadratic terms represent?

6.2. **Dimension counting.** Let's take a more general approach and try to pick a new system of coordinates x^1, \ldots, x^n so that our point p is x = 0 and so that we arrange for the vanishing of as many terms in the Taylor series of g_{ij} at zero as possible.

Let's assume we already have a coordinate system y^1, \ldots, y^n sending p to y = 0 and let $g_{ij} = \left\langle \frac{\partial}{\partial y^i}, \frac{\partial}{\partial y^j} \right\rangle$. If we change coordinates to x^1, \ldots, x^n , we know that

$$\frac{\partial}{\partial y^j} \leadsto \sum_i \frac{\partial x^i}{\partial y^j} \frac{\partial}{\partial x^i},$$

i.e.,

$$\frac{\partial}{\partial x^i} = \sum \frac{\partial y^j}{\partial x^i} \frac{\partial}{\partial y^j}.$$

Letting $\widetilde{g}_{ij} = \left\langle \frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j} \right\rangle$, we see that

$$\widetilde{g}_{ij} = \left\langle \frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j} \right\rangle = \sum_{k,\ell} \left(\frac{\partial y^k}{\partial x^i} \right) \left(\frac{\partial y^\ell}{\partial x^j} \right) g_{k\ell}.$$

We therefore use the values of the Jacobian, i.e., the first derivatives of the coordinate changes, to put \widetilde{g}_{ij} into a model form. We have the freedom to choose all n^2 values of $\frac{\partial y^k}{\partial x^j}$, so we think of this as n^2 unknowns. Now \widetilde{g}_{ij} is always a symmetric matrix, so there are n(n+1)/2 independent components of \widetilde{g}_{ij} and so if we want to make $\widetilde{g}_{ij} = \delta_{ij}$, we have a system of n(n+1)/2 equations in n^2 unknowns. Unless n=1, this is an underdetermined system (there are n(n-1)/2 more unknowns than equations), so we should expect to be able to solve it. Indeed, the number of excess degrees of freedom here is the dimension of SO(n) and corresponds to the rotational symmetry enjoyed by Euclidean space.

We'd now like to arrange that the first derivatives of \tilde{g} vanish at 0. Differentiating the equation above, we have

$$\frac{\partial}{\partial x^k} \widetilde{g}_{ij} = \frac{\partial}{\partial x^k} \left(\sum_{p,q} \frac{\partial y^p}{\partial x^i} \frac{\partial y^q}{\partial x^j} g_{pq} \right),$$

which yields a total of n equations for each component of g, i.e., $n^2(n+1)/2$ total equations. We have the freedom to pick the second derivatives of our coordinate changes at the point 0, but these are subject to the constraint that

$$\frac{\partial^2 y^p}{\partial x^i \partial x^j} = \frac{\partial^2 y^p}{\partial x^j \partial x^i},$$

so we have n(n+1)/2 choices for each y^p , i.e., a total of $n^2(n+1)/2$ unknowns. This set of equations is formally determined (number of equations is the same as number of unknowns), so we'd expect it to have a unique solution. This is essentially what we found in our discussion of geodesic normal coordinates.

What about the second derivatives? Now the equation for the second derivative of the metric tensor involves third derivatives of our coordinate change. The second derivatives

$$\frac{\partial^2}{\partial x^k \partial x^\ell} \widetilde{g}_{ij}$$

are again subject to the constraint that mixed partials commute, so we have n(n+1)/2 equations for each component of \tilde{g} , i.e., a total of $n^2(n+1)^2/4$ equations. How many new

unknowns do we have? In other words, how many different third partial derivatives $\frac{\partial^3 y^k}{\partial x^k \partial x^\ell \partial x^r}$ does each component y^k have? We consider in cases. When all three of k, ℓ , and r are distinct, there are $\binom{n}{3} = n(n-1)(n-2)/6$ such possibilities. When two are distinct, we have n choices for which derivative we take twice and then n-1 for the third derivative, i.e., n(n-1) choices. Finally, when all three indices agree, we have n choices for their common value. This gives a total of

$$\frac{n(n-1)(n-2)}{6} + n(n-1) + n = \frac{n}{6}\left((n-1)(n-2) + 6n\right) = \frac{n(n+1)(n+2)}{6}.$$

Alternatively, we could have used a counting argument that labels how many indices take each value. More precisely, we think of having three identical elements representing the three derivatives we can take and then divide them up into n buckets (which we realize by n-1 dividers). If two elements land in the j-th bucket, we take two derivatives in x^j . If our derivatives and buckets are distinguishable, then there are (n-1+p)! ways to order them. The dividers are definitely not distinguishable, so we must divide by (n-1)!. Similarly, our mixed partials are symmetric, so we should also think that we can't distinguish the order in which we take derivatives, so we should also divide by 3!. This leaves

$$\binom{n-1+3}{3}$$

degrees of freedom for each y^k . (In general you'd have $\binom{n-1+p}{p}$ degrees of freedom for p-th derivatives.) As we have n different functions y^k , we then have $n^2(n+1)(n+2)/6$ total degrees of freedom.

The thing to note here is that the problem of annihilating the second derivatives at a point is a formally overdetermined problem, as there are

$$\frac{n^2(n+1)^2}{4} - \frac{n^2(n+1)(n+2)}{6} = \frac{n^2(n+1)}{12} \left(3(n+1) - 2(n+2) \right) = \frac{n^2(n^2-1)}{12}$$

more equations than unknowns. We therefore do not expect be able to solve this problem generally. Encoding this obstruction is the *curvature tensor*.

6.3. The curvature tensor. As before we suppose (M, g) is a Riemannian manifold equipped with the Levi-Civita connection.

Definition 30. For $X, Y, Z \in \mathcal{X}(M)$, define

$$R(X,Y)Z = \nabla_X \nabla_Y Z - \nabla_Y \nabla_X Z - \nabla_{[X,Y]} Z,$$

so that $R: \mathcal{X}(M) \times \mathcal{X}(M) \times \mathcal{X}(M) \to \mathcal{X}(M)$.

We call R "the" curvature tensor; we should verify that it is a tensor.

Proposition 31. R is a (1,3)-tensor.

Proof. As before, we need to show that for $X, Y, Z \in \mathcal{X}(M)$ and $f \in C^{\infty}(M)$, that

- (i) R(fX,Y)Z = fR(X,Y)Z,
- (ii) R(X, fY)Z = fR(X, Y)Z, and
- (iii) R(X,Y)(fZ) = fR(X,Y)Z.

In fact, R is clearly antisymmetric in X and Y, so the first two are equivalent. We now compute:

$$\begin{split} R(fX,Y)Z &= \nabla_{fX}\nabla_{Y}Z - \nabla_{Y}\nabla_{fX}Z - \nabla_{[fX,Y]}Z \\ &= f\nabla_{X}\nabla_{Y}Z - \nabla_{Y}\left(f\nabla_{X}Z\right) - \nabla_{f[X,Y]-Y(f)X}Z \\ &= f\nabla_{X}\nabla_{Y}Z - f\nabla_{Y}\nabla_{X}Z - Y(f)\nabla_{X}Z - f\nabla_{[X,Y]}Z + Y(f)\nabla_{X}Z \\ &= fR(X,Y)Z. \end{split}$$

Similarly,

$$\begin{split} R(X,Y)(fZ) &= \nabla_X \nabla_Y (fZ) - \nabla_Y \nabla_X (fZ) - \nabla_{[X,Y]} (fZ) \\ &= \nabla_X \left(f \nabla_Y Z + Y(f) Z \right) - \nabla_Y \left(f \nabla_X Z + X(f) Z \right) - f \nabla_{[X,Y]} Z - [X,Y](f) Z \\ &= f \nabla_X \nabla_Y Z + X(f) \nabla_Y Z + Y(f) \nabla_X Z + X(Y(f)) Z \\ &- f \nabla_Y \nabla_X Z - Y(f) \nabla_X Z - X(f) \nabla_Y Z - Y(X(f)) Z - f \nabla_{[X,Y]} Z - [X,Y](f) Z \\ &= f R(X,Y) Z + (X(Y(f)) - Y(X(f)) - [X,Y](f)) Z = f R(X,Y) Z. \end{split}$$

In local coordinates, we write

$$R(\partial_i, \partial_j)\partial_k = \sum_{\ell=1}^n R_{ijk}^{\ell}\partial_{\ell}.$$

Since R is a tensor, these components tell us all of the information about R, and so

$$R(X,Y)Z = \sum_{i,j,k,\ell} X^i Y^j Z^k R^{\ell}_{ijk} \partial_{\ell}.$$

The tensor R is often called the Riemann curvature tensor and it is independent of coordinate choices (indeed, we defined it intrinsically). It's often convenient to turn R from a (1,3)-tensor into a (0,4)-tensor using the metric and we use the same letter R to denote this. Indeed, we define

$$R(X, Y, Z, W) = \langle R(X, Y)Z, W \rangle$$
,

so that

$$R_{ijk\ell} = \sum_{p} R_{ijk}^{p} g^{p\ell}.$$

The Riemann curvature tensor has several symmetries, and these are easiest to state for the (0,4)-tensor.

Proposition 32. (i) R(Y, X, Z, W) = -R(X, Y, Z, W),

- (ii) R(X, Y, W, Z) = -R(X, Y, Z, W),
- (iii) R(Z, W, X, Y) = R(X, Y, Z, W), and
- (iv) R(X, Y, Z, W) + R(Y, Z, X, W) + R(Z, X, Y, W) = 0.

The last identity (the cyclic one) is called the first Bianchi identity.

Proof. The first identity is obvious from the definition, which is antisymmetric in X and Y. For the second identity, we let f denote the smooth function $f = \langle Z, W \rangle$, so that

$$X(Y(f)) - Y(X(f)) - [X, Y](f) = 0.$$

We now use that the Levi-Civita connection is compatible with the metric to write

$$0 = X (Y \langle Z, W \rangle) - Y (X \langle Z, W \rangle) - [X, Y] \langle Z, W \rangle$$

$$= X (\langle \nabla_Y Z, W \rangle + \langle Z, \nabla_Y W \rangle) - Y (\langle \nabla_X Z, W \rangle + \langle Z, \nabla_X W \rangle) - \langle \nabla_{[X,Y]} Z, W \rangle - \langle Z, \nabla_{[X,Y]} W \rangle$$

$$= \langle \nabla_X \nabla_Y Z, W \rangle + \langle \nabla_Y Z, \nabla_X W \rangle + \langle \nabla_X Z, \nabla_Y W \rangle + \langle Z, \nabla_X \nabla_Y W \rangle - \langle \nabla_Y \nabla_X Z, W \rangle$$

$$- \langle \nabla_X Z, \nabla_Y W \rangle - \langle \nabla_Y Z, \nabla_X W \rangle - \langle Z, \nabla_Y \nabla_X W \rangle - \langle \nabla_{[X,Y]} Z, W \rangle - \langle Z, \nabla_{[X,Y]} W \rangle$$

$$= \langle R(X, Y) Z, W \rangle - \langle Z, R(X, Y) W \rangle = R(X, Y, Z, W) - R(X, Y, W, Z),$$

as desired.

To prove the third identity, we'll use the fourth identity (the Bianchi identity), so let's prove the fourth one now.

$$\begin{split} R(X,Y)Z + R(Y,Z)X + R(Z,X)Y &= \nabla_X \nabla_Y Z - \nabla_Y \nabla_X Z - \nabla_{[X,Y]} Z + \nabla_Y \nabla_Z X - \nabla_Z \nabla_Y X \\ &- \nabla_{[Y,Z]} X + \nabla_Z \nabla_X Y - \nabla_X \nabla_Z Y - \nabla_{[Z,X]} Y \\ &= \nabla_X \left(\nabla_Y Z - \nabla_Z Y \right) - \nabla_{[Y,Z]} X + \nabla_Y \left(\nabla_Z X - \nabla_X Z \right) - \nabla_{[Z,X]} Y \\ &+ \nabla_Z \left(\nabla_X Y - \nabla_Y X \right) - \nabla_{[X,Y]} Z. \end{split}$$

Because the Levi-Civita connection is torsion-free, we know that for any X, Y, we have $\nabla_X Y - \nabla_Y X = [X, Y]$, so that the above sum is given by

$$R(X,Y)Z + R(Y,Z)X + R(Z,X)Y = \nabla_X[Y,Z] - \nabla_{[Y,Z]}X + \nabla_Y[Z,X] - \nabla_{[Z,X]}Y + \nabla_Z[X,Y] - \nabla_{[X,Y]}Z$$
$$= [X,[Y,Z]] + [Y,[Z,X]] + [Z,[X,Y]] = 0,$$

by the Jacobi identity, thus proving the Bianchi identity.

We finally turn to the third identity. We use the first two identities and the Bianchi identity to write R(Z, W, X, Y) in two ways:

$$R(Z, W, X, Y) = -R(W, Z, X, Y) = R(Z, X, W, Y) + R(X, W, Z, Y),$$

$$R(Z, W, X, Y) = -R(Z, W, Y, X) = R(W, Y, Z, X) + R(Y, Z, W, X),$$

so that

$$2R(Z, W, X, Y) = R(Z, X, W, Y) + R(X, W, Z, Y) + R(W, Y, Z, X) + R(Y, Z, W, X).$$

An identical calculation gives

$$2R(X,Y,Z,W) = R(X,Z,Y,W) + R(Z,Y,X,W) + R(Y,W,X,Z) + R(W,X,Y,Z).$$

Using antisymmetry twice on each term (e.g., writing R(W, X, Y, Z) = -R(X, W, Y, Z) = R(X, W, Z, Y)) shows that these two sums agree, establishing the third identity.

6.3.1. Another counting argument. How many degrees of freedom does R have? In other words, how many independent components $R_{ijk\ell}$ can be prescribed without the above symmetries forcing our hands?

To be more precise, we ask for the dimension of the space of (0,4) tensors satisfying the following four identities:

$$R_{ijk\ell} = R_{k\ell ij},$$

$$R_{ijk\ell} = -R_{jik\ell},$$

$$R_{ijk\ell} = -R_{ij\ell k},$$

$$R_{ijk\ell} + R_{jki\ell} + R_{kij\ell} = 0.$$

We'll now count by cases.⁵ Antisymmetry demands vanishing if all indices are repeated, i.e., $R_{aaaa} = 0$. For two distinct indices, we have three possibilities, namely R_{aaab} , R_{aabb} , and R_{abab} , and all others can be put into one of these forms by the identities. Of these, only R_{abab} might be non-vanishing, giving us $\binom{n}{2} = n(n+1)/2$ components with two distinct indices.

For three distinct indices, terms of the form R_{aabc} must vanish, and all other terms can be put into the form R_{abac} . For these, we have n choices for the repeated index a and $\binom{n-1}{2} = n(n-1)/2$ choices for the other two indices, giving a total of $3\binom{n}{3}$ of these components.

Finally, for four distinct indices, we start by noticing that we have

$$R_{abcd} = -R_{bacd} = -R_{abdc} = R_{badc} = R_{cdab} = -R_{dcab} = -R_{cdba} = R_{dcba}$$

so, given a choice of a, b, c, d (with, say a < b < c < d for definiteness), there are only three independent components with these indices and we can take as their representatives R_{abcd} , R_{bcad} , and R_{cabd} . The Bianchi identity tells us that the third of these is determined by the other two, so we have in fact two components for each choice of four distinct indices, i.e., a total of

$$2\binom{n}{4} = \frac{1}{12}n(n-1)(n-2)(n-3)$$

independent components.

Summing these up, we have a total of

$$\frac{1}{12}n(n-1)(n-2)(n-3) + \frac{1}{2}n(n-1)(n-2) + \frac{1}{2}n(n+1)$$

$$= \frac{1}{12}n(n-1)\left(n^2 - 5n + 6 + 6n - 12 + 6\right) = \frac{1}{12}n^2(n^2 - 1)$$

independent components. It is no coincidence that this is the same number we found in our counting argument earlier in this section, as the curvature tensor is the obstruction to looking metrically Euclidean.

In particular, note this size in low dimensions. In one dimension, there are no independent components; the curvature tensor always vanishes there. Indeed, you've seen that any curve can be parametrized by arc length, which provides you a local isometry with \mathbb{R} .

In two dimensions, the curvature tensor has a single independent component, which you've seen in another form as the Gaussian curvature. In this sense the curvature tensor generalizes the curvature we defined for surfaces.

In three dimensions, the curvature has 6 independent components and in four dimensions it has 20.

⁵If you have a cleaner way to do this please let me know.

6.4. **Sectional curvature.** Another generalization of the Gaussian curvature is called the sectional curvature, which essentially is the Gaussian curvature of a submanifold.

Definition 33. Let $p \in M$ and let $\Pi \subset T_pM$ be a two-dimensional subspace with basis v and w. The sectional curvature of Π is

$$K(\Pi) = \frac{R(v, w, w, v)}{|v|^2 |w|^2 - \langle v, w \rangle^2}.$$

We start by claming that $K(\Pi)$ depends only on Π and not on the choice of basis v, w. Indeed, suppose \widetilde{v} and \widetilde{w} is another basis, so

$$\widetilde{v} = av + bw,$$

 $\widetilde{w} = cv + dw.$

with $\begin{pmatrix} a & b \\ c & d \end{pmatrix}$ invertible. We then use multilinearity and antisymmetry to see that

$$R(\widetilde{v}, \widetilde{w}, \widetilde{w}, \widetilde{v}) = (ad - bc)R(v, w, \widetilde{w}, \widetilde{v}) = (ad - bc)^2 R(v, w, w, v),$$

while we also have

$$\left|\widetilde{v}\right|^{2}\left|\widetilde{w}\right|^{2} - \left\langle\widetilde{v},\widetilde{w}\right\rangle^{2} = (ad - bc)^{2} \left(\left|v\right|^{2}\left|w\right|^{2} - \left\langle v,w\right\rangle^{2}\right),\,$$

so that $K(\Pi)$ is well-defined.

One interpretation of $K(\Pi)$ is as the Gaussian curvature at p of the surface $\exp_p(\Pi)$. In two dimensions, there is only one sectional curvature K at each point and the Riemann tensor is given by

$$R(X, Y, Z, W) = K(\langle X, W \rangle \langle Y, Z \rangle - \langle X, Z \rangle \langle Y, W \rangle).$$

It's also true that in higher dimensions the entire Riemann tensor can be recovered from the sectional curvatures, but the reconstruction formula is long and we probably won't use it in this course.

Definition 34. A Riemannian manifold (M, g) has constant sectional curvature k if $K(\Pi) = k$ for all $p \in M$ and all $\Pi \subset T_pM$.

As examples, you should check that $\mathbb{S}^n = \{x \in \mathbb{R}^{n+1} \mid |x| = 1\}$, equipped with the pullback of the Euclidean metric, has constant sectional curvature 1.

6.5. Curvature as an operator on tensors. Given an (r, s)-tensor T, for $X, Y \in \mathcal{X}(M)$, we obtain another (r, s)-tensor that we call R(X, Y)T by

$$R(X,Y)T = \nabla_X \nabla_Y T - \nabla_Y \nabla_X T - \nabla_{[X,Y]} T.$$

Proposition 35. If $\theta \in \Omega^1(M)$, and $X, Y, Z \in \mathcal{X}(M)$, then

$$(R(X,Y)\theta)(Z) = -\theta(R(X,Y)Z).$$

Proof. By definition of $\nabla_Y \theta$, we have

$$Y(\theta(Z)) = \nabla_Y \theta(Z) + \theta(\nabla_Y Z),$$

so that

$$X (Y (\theta(Z))) = X (\nabla_Y \theta(Z) + \theta(\nabla_Y Z))$$

= $(\nabla_X \nabla_Y \theta) (Z) + \nabla_Y \theta(\nabla_X Z) + \nabla_X \theta(\nabla_Y Z) + \theta(\nabla_X \nabla_Y Z).$

Now, applying X(Y(f)) - Y(X(f)) - [X, Y](f) = 0 with the smooth function $f = \theta(Z)$, we have

$$0 = X(Y(\theta(Z))) - Y(X(\theta(Z))) - [X, Y](\theta(Z))$$

= $(\nabla_X \nabla_Y \theta - \nabla_Y \nabla_X \theta - \nabla_{[X,Y]} \theta) (Z) + \theta (\nabla_X \nabla_Y Z - \nabla_Y \nabla_X Z - \nabla_{[X,Y]} Z),$

finishing the proof.

In general, for an (r, s)-tensor T, we have

$$(R(X,Y)T) (\theta_1, \dots, \theta_r, Z_1, \dots, Z_s)$$

$$= -\sum_{k=1}^r T(\theta_1, \dots, R(X,Y)\theta_k, \dots, \theta_r, Z_1, \dots, Z_s)$$

$$+ \sum_{k=1}^s T(\theta_1, \dots, \theta_r, Z_1, \dots, R(X,Y)Z_k, \dots, Z_s).$$

Applying this to the metric tensor g, we already know $\nabla_X g = 0$, so R(X,Y)g = 0 and thus

$$0 = [R(X,Y)g](Z,W) = -g(R(X,Y)Z,W) - g(Z,R(X,Y)W)$$

= $-R(X,Y,Z,W) - R(X,Y,W,Z)$,

which yields one of our earlier properties (nothing new here).

There is, however, a nontrivial property of ∇R , though:

Proposition 36 (Second Bianchi identity). For $X, Y, Z, V, W \in \mathcal{X}(M)$,

$$\left(\nabla_X R\right)\left(Y,Z,V,W\right) + \left(\nabla_Y R\right)\left(Z,X,V,W\right) + \left(\nabla_Z R\right)\left(X,Y,V,W\right) = 0.$$

Proof. One (tedious) way to check this is to use the definnition and express the curvature in terms of commutators and eventually appeal to the Jacobi identity as we did for the first Bianchi identity. It's convenient to take a shortcut.

The equation above is tensorial so it is enough to show it in a convenient choice of coordinates. Fix $p \in M$; we want to check the identity at p. We work in geodesic normal coordinates at p, so $\Gamma_{ij}^k(p) = 0$. It's enough to check it at p for $(X, Y, Z, V, W) = (\partial_i, \partial_j, \partial_k, \partial_\ell, \partial_m)$. Note that because we are working in geodesic normal coordinates, we have $\nabla_{\partial_*}\partial_* = 0$ at p. We therefore conclude that

$$\nabla_X R(Y, Z, V, W) = X(R(Y, Z, V, W))$$

at p because all other terms vanish.

We now compute. One shortcut we use is to keep two of the terms that vanish at p. Indeed, we know that at p,

$$-R(\partial_j, \partial_k, \nabla_{\partial_i}\partial_\ell, \partial_m) - R(\partial_j, \partial_k, \partial_\ell, \nabla_{\partial_i}\partial_m) = 0.$$

Our other shortcut is that we are using coordinate vector fields, and so $[\partial_*, \partial_*] = 0$. We use this in two ways. First, there is no third term in the expression for $R(\partial_j, \partial_k)\partial_\ell$. Second, because the metric is torsion-free, we get to conclude that $\nabla_{\partial_a}\partial_b = \nabla_{\partial_b}\partial_a$. We then see that,

at p,

$$X(R(Y,Z,V,W)) = \partial_{i} \langle R(\partial_{j},\partial_{k})\partial_{\ell},\partial_{m} \rangle - R(\partial_{j},\partial_{k},\nabla_{\partial_{i}}\partial_{\ell},\partial_{m}) - R(\partial_{j},\partial_{k},\partial_{\ell},\nabla_{\partial_{i}}\partial_{m})$$

$$= \langle \nabla_{\partial_{i}} \left(\nabla_{\partial_{j}} \nabla_{\partial_{k}} \partial_{\ell} - \nabla_{\partial_{k}} \nabla_{\partial_{j}} \partial_{\ell} \right), \partial_{m} \rangle + \langle \nabla_{\partial_{j}} \nabla_{\partial_{k}} \partial_{\ell} - \nabla_{\partial_{k}} \nabla_{\partial_{j}} \partial_{\ell}, \nabla_{\partial_{i}} \partial_{m} \rangle$$

$$- \langle \nabla_{\partial_{j}} \nabla_{\partial_{k}} \nabla_{\partial_{i}} \partial_{\ell} - \nabla_{\partial_{k}} \nabla_{\partial_{j}} \nabla_{\partial_{i}} \partial_{\ell}, \partial_{m} \rangle - \langle \nabla_{\partial_{j}} \nabla_{\partial_{k}} \partial_{\ell} - \nabla_{\partial_{k}} \nabla_{\partial_{j}} \partial_{\ell}, \nabla_{\partial_{i}} \partial_{m} \rangle$$

$$= \langle \left[\nabla_{\partial_{i}}, \left[\nabla_{\partial_{i}}, \nabla_{\partial_{k}} \right] \right] \partial_{\ell}, \partial_{m} \rangle.$$

Now we take the cyclic sum over i, j, k and use the Jacobi identity (which is a theorem about letters and so valid for commutators) to finish the proof.

6.6. Ricci and scalar curvature. Given the Riemann tensor R, we can define a (0, 2)-tensor called the Ricci tensor by taking a trace. You should check for yourself that the following definition is in fact the trace of $R(\bullet, X)Y$ regarded as a linear transformation $T_pM \to T_pM$.

Definition 37. The Ricci curvature is the (0,2)-tensor given in any coordinate system by

$$\operatorname{Ric}(X,Y) = \sum_{j=1}^{n} g^{ij} R\left(X, \frac{\partial}{\partial x^{i}}, \frac{\partial}{\partial x^{j}}, Y\right).$$

The scalar curvature is a function given by contracting the Ricci curvature:

$$R_{p} = \sum_{k,\ell} g^{k\ell} \operatorname{Ric} \left(\partial_{k}, \partial_{\ell} \right) = \sum_{i,j,k,\ell} g^{ij} g^{k\ell} R \left(\partial_{k}, \partial_{i}, \partial_{j}, \partial_{\ell} \right).$$

Note that the Ricci tensor is a symmetric (0, 2)-tensor.

The Ricci tensor and the scalar curvature satisfy a nice relationship called the contracted second Bianchi identity.⁶

Theorem 38 (Contracted second Bianchi identity). For $p \in M$ and $X \in T_pM$, we have, for any basis v_i of T_pM ,

$$X(R) = 2\sum_{i,j=1}^{n} g^{ij} \left(\nabla_{v_i} \operatorname{Ric} \right) (X, v_i).$$

where R denotes the scalar curvature and Ric the Ricci tensor.

Proof. The key idea is that the covariant derivative commutes with contraction. We'll prove it only for the two cases we need here, but it is true in general. ⁷ Geodesic normal coordinates make this much easier to see.

Fix $p \in M$ and let x^1, \ldots, x^n be geodesic normal coordinates at p (so that $\Gamma_{ij}^k(p) = 0$). Recall that $\nabla_X g = 0$, so that, in particular,

$$\partial_k g_{\ell m} - \Gamma^r_{k\ell} g_{rm} - \Gamma^r_{km} g_{\ell r} = 0.$$

As we are working in geodesic normal coordinates, we have that at p, $\partial_k g_{\ell m} = 0$. In particular, because $g_{\ell m} g^{\ell m} = n$, we also have $\partial_k g^{\ell m} = 0$ at p. We then have, for a (0,2)-tensor T,

⁶This relationship comes up in relativity and tells you that the left hand side of the Einstein equations (which describe the curvature of the spacetime) must be divergence free and therefore impose a constraint on the right hand side (which describes the contribution from the matter fields).

⁷The more general statement follows by an inductive procedure together with the base cases of a (2,0) and (1,1)-tensor.

that at p (again, we are using geodesic normal coordinates; ordinarily there would be more terms!)

$$\partial_k \left(g^{ij} T_{ij} \right) = g^{ij} \partial_k T_{ij} = g^{ij} (\nabla_{\partial_k} T)_{ij}.$$

Similarly, for a (0,4)-tensor T, we have at p

$$\partial_k \left(g^{ij} T_{ij\ell m} \right) = g^{ij} \partial_k T_{ij\ell m} = g^{ij} (\nabla_{\partial_k} T)_{ij\ell m}.$$

We now turn to the identity. By the second Bianchi identity, we have

$$0 = (\nabla_{\partial_i} R)(\partial_j, \partial_k, \partial_\ell, \partial_m) + (\nabla_{\partial_j} R)(\partial_k, \partial_i, \partial_\ell, \partial_m) + (\nabla_{\partial_k} R)(\partial_i, \partial_j, \partial_\ell, \partial_m),$$

so we contract by multiplying by $g^{i\ell}$ and g^{jm} and then summing. Because contraction commutes with covariant derivatives, we know that, e.g., at p

$$(\nabla_{\partial_i} R)(\partial_i, \partial_k, \partial_\ell, \partial_m) = \partial_i (R(\partial_i, \partial_k, \partial_\ell, \partial_m)),$$

so that we obtain at p

$$0 = g^{i\ell} \partial_i \left(g^{jm} R(\partial_i, \partial_k, \partial_\ell, \partial_m) \right) + g^{jm} \partial_i \left(g^{i\ell} R(\partial_k, \partial_i, \partial_\ell, \partial_m) \right) + \partial_k \left(g^{i\ell} g^{jm} R(\partial_i, \partial_i, \partial_\ell, \partial_m) \right).$$

Using the definition of the Ricci and scalar curvatures, we see that this equation yields

$$\partial_k R = g^{jm} \partial_j \left(\operatorname{Ric}(\partial_k, \partial_m) \right) + g^{i\ell} \partial_i \left(\operatorname{Ric}(\partial_k, \partial_\ell) \right)$$

= $2g^{ij} \partial_j \left(\operatorname{Ric}(\partial_k, \partial_i) \right) = 2g^{ij} \left(\nabla_{\partial_i} \operatorname{Ric} \right) \left(\partial_k, \partial_i \right),$

as desired. \Box

A significant application of the contracted second Bianchi identity is Schur's theorem:

Theorem 39. Suppose $n = \dim M \geq 3$ and M is connected.

(1) Assume there is a smooth function $f \in C^{\infty}(M)$ so that

$$\operatorname{Ric}_p(X,Y) = f(p)g(X,Y)$$

for all $p \in M$ and $X, Y \in T_pM$. Then f is constant.

(2) Assume there is a smooth $f \in C^{\infty}(M)$ so that

$$R(X, Y, Z, W) = f(p) (g(X, W)g(Y, Z) - g(X, Z)g(Y, W))$$

for all $p \in M$ and $X, Y, Z, W \in T_pM$. (In other words, the sectional curvatures at p are all f(p)/(n-1).) Then f is constant.

Proof. (1) We note that Ric = fg, so

$$\nabla_{v_i} \operatorname{Ric} = \nabla_{v_i}(fg) = v_i(f)g.$$

The scalar curvature is the trace of Ric, so R=nf. By the contracted second Bianchi identity, we have

$$nX(f) = 2\sum g^{ij}\nabla_{v_j}\operatorname{Ric}(X, v_i) = 2X(f).$$

As $n \neq 2$, we conclude X(f) = 0 and so f is constant.

(2) Now observe

$$\operatorname{Ric}_{p}(X,Y) = \sum_{i,j} g^{ij} R(X, v_{i}, v_{j}, Y)$$

$$= \sum_{i,j} f(p) g^{ij} (g(X,Y)g_{ij} - g(X, v_{i})g(Y, v_{j}))$$

$$= \sum_{i,j} f(p) \left(ng(X,Y) - \sum_{k,\ell} g^{ij} g_{ik} X^{k} g_{j\ell} Y^{\ell} \right)$$

$$= \sum_{i,j} f(p) (ng(X,Y) - g(X,Y)) = (n-1)f(p)g,$$

putting us in the setting of the first part.

Not so relevant here, but an important tensor in relativity is the Einstein tensor, given by $\operatorname{Ric} - \frac{1}{2} Rg$. It is divergence-free by the contracted second Bianchi identity.

7. RIEMANNIAN DISTANCE

Definition 40. For a piecewise smooth path $\gamma:[0,t_0]\to M$, define its length

$$L(\gamma) = \int_0^{t_0} |\gamma'(t)|_g dt = \int_0^{t_0} \sqrt{g(\gamma'(t), \gamma'(t))} dt$$

and its energy

$$E(\gamma) = \int_0^{t_0} |\gamma'(t)|_g^2 dt = \int_0^{t_0} g(\gamma'(t), \gamma'(t)) dt.$$

A quick exercise in the chain rule proves the following:

Lemma 41. $L(\gamma)$ is independent of orientation-preserving parametrization.

Proof. Suppose $\gamma:[0,t_0]\to M$ is smooth (for piecewise smooth, break into pieces) and $\alpha:[0,s_0]\to[0,t_0]$ is an orientation-preserving diffeomorphism. Let $\tilde{\gamma}=\gamma\circ\alpha$, so then $\tilde{\gamma}'(s)=\gamma'(\alpha(s))\alpha'(s)$, and

$$L(\tilde{\gamma}) = \int_0^{s_0} |\tilde{\gamma}'(s)|_g \, ds = \int_0^{s_0} |\gamma'(\alpha(s))| \, \alpha'(s) \, ds$$
$$= \int_0^{s_0} |\gamma'(\alpha(s))| \, \alpha'(s) \, ds = \int_0^{t_0} |\gamma'(t)| \, dt = L(\gamma).$$

We now make our Riemannian manifold into a metric space.

Definition 42. For (M, g) Riemannian, and $p, q \in M$, define

$$d(p,q) = \inf \{ L(\gamma) \mid \gamma : [0,1] \to M \text{ piecewise smooth, } \gamma(0) = p, \gamma(1) = q \}.$$

Note that we could equivalently minimize the energy over paths from p to q – one inequality is Cauchy–Schwarz and the other is reparametrizing by arc length.

Our aims for this section include the following:

(1) d is a metric,

- (2) If the infimum is attained, the minimizer is a geodesic,
- (3) Conditions so that the infimum is attained, and
- (4) A second derivative test for minimization.

Lemma 43 (Gauss lemma). Let $P \in M$ and $v \in T_pM$, $v \neq 0$. For any $w \in T_v(T_pM) \cong T_pM$, we have

$$\langle (D \exp_p)_v(v), (D \exp_p)_v(w) \rangle = \langle v, w \rangle.$$

Proof. Let $\tilde{F}(t,s) = t(v+sw)$ in T_pM and $F(t,s) = \exp_p(\tilde{F}(t,s))$. Observe that

$$\begin{split} \frac{\partial \tilde{F}}{\partial t}(0,0) &= v, & \frac{\partial \tilde{F}}{\partial s}(0,0) &= 0, \\ \frac{\partial \tilde{F}}{\partial t}(1,0) &= v, & \frac{\partial \tilde{F}}{\partial s}(1,0) &= w, \end{split}$$

so that

$$\begin{split} \frac{\partial F}{\partial t}(0,0) &= v, & \frac{\partial F}{\partial s}(0,0) &= 0, \\ \frac{\partial F}{\partial t}(1,0) &= (D\exp_p)_v(v), & \frac{\partial F}{\partial s}(1,0) &= (D\exp_p)_v(w). \end{split}$$

The curve $t \mapsto F(t, s)$ is a geodesic with initial velocity vector v + sw, so

$$\frac{D}{dt}\left(\frac{\partial F}{\partial t}\right) = 0$$
 and $\left\langle\frac{\partial F}{\partial t}, \frac{\partial F}{\partial t}\right\rangle = \left\langle v + sw, v + sw\right\rangle$.

We now claim that

$$\frac{D}{dt} \left(\frac{\partial F}{\partial s} \right) = \frac{D}{ds} \left(\frac{\partial F}{\partial t} \right).$$

Assuming this claim, we have

$$\begin{split} \frac{\partial}{\partial t} \left\langle \frac{\partial F}{\partial t}, \frac{\partial F}{\partial s} \right\rangle &= \left\langle \frac{D}{dt} \frac{\partial F}{\partial t}, \frac{\partial F}{\partial s} \right\rangle + \left\langle \frac{\partial F}{\partial t}, \frac{D}{dt} \frac{\partial F}{\partial s} \right\rangle \\ &= 0 + \left\langle \frac{\partial F}{\partial t}, \frac{D}{ds} \frac{\partial F}{\partial t} \right\rangle \\ &= \frac{1}{2} \frac{\partial}{\partial s} \left\langle \frac{\partial F}{\partial t}, \frac{\partial F}{\partial t} \right\rangle = \left\langle v, w \right\rangle. \end{split}$$

So, since the value of the inner product at (0,0) is 0, its value at (1,0) must be $\langle v,w\rangle$. We now prove the claim. Suppose $F:[0,t_0]\times[0,s_0]\to M$ is smooth. Note that

$$\frac{D}{ds}\left(\frac{\partial F}{\partial t}\right) = \nabla_{\frac{\partial F}{\partial s}}\frac{\partial F}{\partial t} = \nabla_{\frac{\partial F}{\partial t}}\frac{\partial F}{\partial t} + \left[\frac{\partial F}{\partial s},\frac{\partial F}{\partial t}\right] = \frac{D}{dt}\left(\frac{\partial F}{\partial s}\right) + \left[\frac{\partial F}{\partial s},\frac{\partial F}{\partial t}\right],$$

so it remains to see that the commutator term vanishes, which can be checked in coordinates. (Alternatively, we could go immediately to coordinates to see that

$$\left(\frac{D}{ds}\left(\frac{\partial F}{\partial t}\right)\right)^k = \frac{\partial^2 F^k}{\partial s \partial t} + \sum_{i,j} \Gamma^k_{ij}(F(t,s)) \frac{\partial F^i}{\partial s} \partial F^j \partial t,$$

which is symmetric in the roles of s and t.

Now we construct a "local position vector". Fix $p \in M$, and for q close to p we write $q = \exp_p(v)$ for some $v \in T_pM$ (in the neighborhood of 0 where \exp_p is a diffeomorphism). Let

$$P = (D \exp_p)_v(v) \in T_{\exp_p(v)}M = T_qM.$$

Note that P is a smooth vector field defined in a neighborhood of p, and define the function Q in this neighborhood by

$$Q(q) = |v|_q^2,$$

where $v = \exp_p^{-1}(q)$. We claim that the gradient of Q is 2P.

Proof of claim. Define $\tilde{Q}: T_pM \to \mathbb{R}$ by $\tilde{Q}(v) = |v|_g^2$, so that in a neighborhood of p, $\tilde{Q} = Q \circ \exp_p$.

Take q close to p and $w \in T_qM$, and let $v = \exp_p^{-1}(q)$, and $\tilde{w} \in T_v(T_pM) \cong T_pM$ be defined by $(D \exp_p)_v(\tilde{w}) = w$.

Since $\tilde{Q} = \exp_n^* Q$, we have

$$\begin{split} \langle (\operatorname{grad} Q)_q, w \rangle &= w(Q) = (D \exp_p)_v(\tilde{w})(Q) \\ &= \tilde{w} \left(\exp_p^* Q \right) = \tilde{w}(\tilde{Q}) \\ &= \left\langle (\operatorname{grad} \tilde{Q})_v, \tilde{w} \right\rangle = 2 \left\langle v, \tilde{w} \right\rangle. \end{split}$$

Now by Gauss we also have $\langle P, w \rangle = \langle v, \tilde{w} \rangle$, so we must have $2P = \operatorname{grad} Q$ because w (and hence \tilde{w}) was arbitrary.

Proposition 44. Let $p \in M$ and $B_r(0) = \{v \in T_pM \mid |v|_g^2 < r^2\}$. Assume r > 0 is sufficiently small that $\exp_p|_{B_r(0)}$ is a diffeomorphism and let $U = \exp_p(B_r(0))$. For $q \in U$, the radial geodesic γ from p to q is the unique shortest curve (up to reparametrization) in U from p to q.

Proof. We have to show that for $\alpha:[0,t_0]\to U$ with $\alpha(0)=p$ and $\alpha(t_0)=q$, we have

- (1) $L(\alpha) \ge L(\gamma)$, and
- (2) If $L(\alpha) = L(\gamma)$, then α is a reparametrization of γ .

In the notation from above, define a function $R = \sqrt{Q}$ on U, so that V = P/R is the outward radial unit vector on $U \setminus \{p\}$, and grad $r = \frac{1}{2R} \operatorname{grad} Q = P/R = V$.

Without loss of generality we can assume that $\alpha(t) \neq p$ for t > 0 (or else we could start later with a path that was no longer). Write α' in terms of its radial and orthogonal components, i.e.,

$$\alpha' = \langle \alpha', V \rangle V + N,$$

where N is orthogonal to V. (At t = 0, it doesn't really matter what you do, so maybe take $V = \alpha'(0)$ and N = 0.) We compute, for t > 0,

$$|\alpha'(t)| = \sqrt{\langle \alpha', \alpha' \rangle} = \sqrt{\langle \alpha', V \rangle^2 + |N|^2} \ge |\langle \alpha', V \rangle|.$$

V is the gradient of R, so $\langle \alpha', V \rangle = \frac{d}{dt}(R \circ \alpha)$, and thus

$$L(\alpha) = \int_0^{t_0} |\alpha'(t)| \ dt \ge \int_0^{t_0} \frac{d}{dt} (R \circ \alpha) \ dt = R(\alpha(t_0)) = R(q).$$

On the other hand, we compute the length of γ :

$$L(\gamma) = \int_0^1 |\gamma'(t)| \ dt = \int_0^1 |v| \ dt = |v| = R(q),$$

so $L(\alpha) \geq L(\gamma)$. To get equality in the above inequality, we must have |N| = 0 and $|\langle \alpha', V \rangle| = \langle \alpha', V \rangle$, so that $\alpha'(t) = \left(\frac{d}{dt}(R \circ \alpha)\right) V$, and thus α travels monotonically along γ and is a reparametrization.

Proposition 45. Let $U = \exp_p(B_r(0))$ be as above. For any $q \in U$, the radial geodesic γ from p to q is the unique (up to reparametrization) shortest curve in M from p to q, i.e., $d(p,q) = L(\gamma)$.

Proof. We know it's the shortest curve within U from p to q, and $L(\gamma) < r$, so it remains to show that if α is a curve in M starting at p and leaving U, then $L(\alpha) \ge r$. Since α leaves U, it meets every geodesic sphere $S(a) = \exp_p(\partial B_a(0))$ with a < r. If α_1 is the shortest inital sement from p to S(a), then, because α lies in U initially, we must have $L(\alpha) \ge L(\alpha_1) = a$ for all a < r, i.e., $L(\alpha) \ge r$.

Proposition 46. If M is connected then d is a metric, i.e.,

- (i) d(p,q) = 0 if and only if p = q,
- (ii) d(p,q) = d(q,p), and
- (iii) $d(p,q) \le d(p,z) + d(z,q)$ for any $z \in M$.

Proof. Note that if M is connected then it is path connected (because it is a manifold), so $d(p,q) < \infty$. The second statement is easy by reversing paths, while the third follows by concatenating them. One implication in the first statement is trivial, so it remains to show that if d(p,q) = 0, then p = q. Let $p, q \in M$, $p \neq q$. Let $U = \exp_p(B_r(0))$ be as before. If $q \in U$, then we write $q = \exp_p(v)$ and then $d(p,q) = |v| \neq 0$, while if $q \notin U$, we saw that $d(p,q) \geq r$, so we're done.

Proposition 47. If (M, g) is Riemannian, then the distance function d induces a topology on M that agrees with the original one. In other words, U is open in M if and only if for all $p \in U$, there is an $\epsilon > 0$ so that $d(p, q) \ge \epsilon$ for all $q \notin U$.

Proof. Need to check:

- (1) Given any open $U \subset M$ and $p \in U$, there is some r > 0 so $d(p,q) \ge r$ for all $q \notin U$, and
- (2) Given any r > 0 and $p \in M$, there is an open set U so that d(p,q) < r for all $q \in U$.

The second statement follows from the fact that \exp_p is a local diffeomorphism around 0. The first statement follows from the proof of Proposition 45.

Proposition 48. Let $p, q \in M$ and suppose that $\alpha : [0,1] \to M$ is a path from p to q with $L(\alpha) = d(p,q)$. Then α is a geodesic (up to reparametrization).

You might try to reparametrize α by arc length, but the worry is that it might no longer be smooth. (Picture sharp corners you can slow down enough for.) Instead we'll cut it into small pieces and use Proposition 44.

Proof. Split [0,1] into finitely many intervals $[t_i,t_{i+1}]$ so that $\alpha([t_i,t_{i+2}]) \subset U_i$, where U_i is a normal coordinate neighborhood of $\alpha(t_{i+1})$. Because α minimizes distance between its

endpoints it must also minimize distance between the points in between by the triangle inequality, so that

$$L\left(\alpha|_{[t_i,t_{i+2}]}\right) = d(\alpha(t_i),\alpha(t_{i+2})),$$

and thus $\alpha|_{[t_i,t_{i+2}]}$ is a reparametrization of a geodesic. Write

$$\alpha|_{[t_i,t_{i+2}]} = \gamma_i \circ \varphi_i,$$

where γ_i is a unit speed geodesic and φ_i is monotone. Say $\gamma_i : [0, T_i] \to U_i$ and $\varphi_i : [t_i, t_{i+2}] \to [0, T_i]$ and let $\tau_i = \varphi_i(t_{i+1}) \in [0, T_i]$, $\gamma_i(\tau_i) = \alpha(t_{i+1})$.

Consider, for $s \in [0, 1]$, the geodesics $s \mapsto \gamma_i((T_i - \tau_i)s + \tau_i)$ and $s \mapsto \gamma_{i+1}(\tau_{i+1}s)$. Note that these have the same endpoints and are both contained in the normal coordinate neighborhood U_{i+1} , and are both radial geodesics, so the must agree. We can then define γ by

$$\gamma(s) = \gamma_i \left(s - \sum_{j=1}^{i-1} \tau_j \right)$$

for all s so that $\sum_{j=1}^{i-1} \tau_j \leq s \leq \sum_{j=1}^{i} \tau_i$, so that γ is a unit speed geodesic connecting p and q and a reparametrization of α .

APPENDIX A. SMOOTH DEPENDENCE ON PARAMETERS

Last semester we had a more-difficult-than-necessary discussion of how solutions of ODEs depend smoothly on parameters.