# Exploring Monte-Carlo-integration techniques in Bayesian model selection

Jakob Krause\* and Dominic Schüchter<sup>†</sup> (Dated: March 2, 2021)

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#### I. INTRODUCTION

In physics, one is often faced with the problem of *Model Selection* for a given data set. That means finding a mathematical description of the data, which on the one hand sufficiently characterizes the data structure and on the other hand satisfies the expected dependencies. This is by no means a trivial task; one has to understand the underlying physical model beforehand to not make the mistake of choosing a too complicated model although it may seemingly fit the data. At the same time too trivial assumptions can also lead in the wrong direction. Compactly this problem can be formulated in the following way (adapted from [5, Chap. 4]):

Alice has a theory; Bob also has a theory, but with an adjustable parameter  $\lambda$ . Whose theory should we prefer on the basis of data D?

BAYESian inference provides quantitative measures for this model-selection problem, e.g. the BAYES-factor and the BAYES-complexity, these have among others been successfully used in astronomy as can be found in [4, 7, 8] respectively. In this paper we will investigate two simulated example problems and apply various measures of bayesian model selection as to find out the true underlying model which was used to generate the simulated data. We will focus on the numerical evaluation of such problems, especially Monte-Carlo techniques.

## II. THEORY

In the following we will give a short introduction into BAYES THEORY and the underlying concepts of model selection.

#### A. Bayes' Theorem

The fundamental equation of BAYESIAN statistics – for a dataset y and parameters  $\theta$  – is given by BAYES The-

orem.

$$prob(\boldsymbol{\theta}|y) = p(\boldsymbol{\theta}|y) = \frac{p(y|\boldsymbol{\theta}) \cdot p(\boldsymbol{\theta})}{p(y)}$$
(1)

Where  $p(\theta|y)$  is the posterior probability for parameters  $\theta$  given the data y,  $p(y|\theta)$  is the likelihood that the data fits a model with parameters  $\theta$ ,  $p(\theta)$  is the prior probability of  $\theta$  and  $p(y) = \int_{-\infty}^{+\infty} \mathrm{d}\theta p(y|\theta) p(\theta)$  is the marginal likelihood which acts as a normalization. In the case of parameter selection, the normalization can often be neglected since it is only a constant. In the case of the model comparison it is a crucial quantity, as we will disuss in section (II C) [5, Chap. 2].

### B. Parameter estimation

Assume we want to find the best parameters for a given dataset y and model M. Eq. (1) then can be written as

$$p(\boldsymbol{\theta}|y, M) = \frac{p(y|\boldsymbol{\theta}, M) \cdot p(\boldsymbol{\theta}|M)}{p(y|M)}.$$
 (2)

Evaluating this equation will give probability density functions (PDFs)

$$p(\theta_i|y, M) = \int p(\boldsymbol{\theta}|y, M) \prod_{i \neq i} d\theta_j$$
 (3)

for each parameter  $\theta_i$ . Because of equation (3)  $p(\theta|D, M)$  is also called *marginal posterior*. From this one can find the best fit value either by finding the value of  $\theta$  for which the marginal posterior is maximized or by calculating the mean with respect to the marginal posterior.

$$p(\theta|y, M) = \max \Leftrightarrow \theta = \hat{\theta}$$
$$\langle \theta \rangle = \int_{-\infty}^{\infty} d\theta p(\theta|y, M) \cdot \theta$$
(4)

Throughout this paper we will use  $\langle \theta \rangle$  as our best fit estimate. [5]

#### C. Model comparison

To compare two (or more) competing models  $M_i$  that describe a dataset D let us write BAYES theorem once

<sup>\*</sup> http://www.github.com/krausejm; krause@hiskp.uni-bonn.de

<sup>†</sup> http://www.github.com/dschuechter; dschuechter@uni-bonn.de

again

$$p(M_i|y) = \frac{p(M_i) \cdot p(y|M_i)}{p(y)}.$$
 (5)

We recognize  $p(y|M_i)$  as the marginal likelihood of the model  $M_i$ . Let now the two competing models be those of Alice and Bob introduced in section (I). Thus,

$$p(y|M_i) = \int_{-\infty}^{\infty} d\lambda p(y|\lambda, M) \cdot p(\lambda|M)$$

follows [5, Chap. 3].

# 1. Bayes Factor

One way of comparing two models is to compare their posterior odds, that is the ratio quote lecture A. Wirzba?

$$O_{ij} := \underbrace{\frac{p(M_i|y)}{p(M_j|y)}}_{\text{posterior odds}} = \underbrace{\frac{p(y|M_i)}{p(y|M_j)}}_{\text{BAYES Factor prior odds}} \cdot \underbrace{\frac{p(M_i)}{p(M_j)}}_{\text{prior odds}} = B_{ij} \cdot \frac{p(M_i)}{p(M_j)}.$$
(6)

Usually we can assume the prior odds to be  $\sim 1$  because we do not favor one model over the other just on prior beliefs. With this equation (6) simply becomes

$$O_{ij} = B_{ij}$$
.

Calculating  $B_{ij}$  will be referred to as calculating the BAYES factor in favor of model  $M_i$ . We can then evaluate the result of  $B_{ij}$  as follows

$$B_{ij} = B_{ji}^{-1} = \begin{cases} \geq 1 \mod M_i \text{ favored} \\ < 1 \mod M_j \text{ favored} \end{cases}$$

Nevertheless a BAYES Factor  $\geq 1$  does not mean that the favored model is far superior. There is an empirical scale which points to the strength of evidence for a given  $B_{ij}$ . This is displayed in table (I).

$ \ln B_{ij} $	Odds	Probability	Strength of evidence
< 1.0	$\lesssim 3:1$	< 0.750	Inconclusive
1.0	$\sim 3:1$	0.750	Weak evidence
2.5	$\sim 12:1$	0.923	Moderate evidence
5.0	$\sim 150:1$	0.993	Strong evidence

**TABLE I:** Empirical scale for evaluating the strength of evidence when comparing two models  $M_i$  vs.  $M_j$ , adapted from [8]

Bayes factor vor allem attraktiv für genestete modelle

# 2. Bayes Complexity

When comparing models with a different amount of parameters, we obviously denote more parameters as more

complex. A fundamental question of model-selection is then if the given data supports more or less parameters. A naïve measure of complexity would be the number of free parameters a model has, we denote this as  $\mathcal{C}_0$ . A more sophisticated measure of complexity is given by the so called BAYESIAN complexity, first introduced by Spiegelhalter et al [6], which can be written as [4]

$$C_b = -2 \int d\boldsymbol{\theta} p(\boldsymbol{\theta}|y, M) \log(\mathcal{L}(\boldsymbol{\theta})) + 2 \log(\mathcal{L}(\tilde{\boldsymbol{\theta}})), \quad (7)$$

with the likelihood  $\mathcal{L}(\boldsymbol{\theta}) = p(D|\boldsymbol{\theta}, M)$  and  $\tilde{\boldsymbol{\theta}} = \langle \boldsymbol{\theta} \rangle$ .  $C_b$  describes how many model parameters the data is able to constrain [4] and is thus a useful tool for examining models with an increasing number of parameters. If we define a  $\chi^2$  as  $\mathcal{L}(\boldsymbol{\theta}) \propto \exp(-\chi^2/2)$ , we can write

$$C_b = \overline{\chi^2(\boldsymbol{\theta})} - \chi^2(\tilde{\boldsymbol{\theta}}), \tag{8}$$

where  $\overline{\chi}$  denotes the mean taken over the posterior PDF. The definition of  $C_b$  is chosen such that  $C_b \to C_0$  for highly informative data [4].

#### III. METHODS

We will now describe the functions of the used algorithms for the model selection. Since the implementation of the so called nested sampling is sufficiently dealt with by Gruppe - Bayesian parameter fitting our implementation uses the PyMC3-Python Library [2]. PyMC3 offers simple solutions to create models and a wide variety of sampling algorithms. Furthermore the ArviZ-Library provides methods for posterior analysis and visualization [1].

### A. Monte-Carlo Sampling

Often equations (3), (7) and (8) have no closed analytical solutions, so we are left with a numerical ansatz (or analytical approximations) [9]. One possibility is given by Markov-Chain-Monte-Carlo Sampling (MCMC) [?]. The main idea is to generate a chain of parameter samples  $\boldsymbol{\theta}^{(t)}$  that follow the posterior PDF (2). For such a chain the mean with respect to the posterior is given by

$$\langle \boldsymbol{\theta} \rangle \approx \int p(\boldsymbol{\theta}|y) \boldsymbol{\theta} d\boldsymbol{\theta} = \frac{1}{N} \sum_{t=0}^{N-1} \boldsymbol{\theta}^{(t)},$$
 (9)

since the samples follow the PDF  $p(\theta|y)$ . Analogously the expectation value of any function with respect to the posterior is

$$\langle f(\boldsymbol{\theta}) \rangle \approx \frac{1}{N} \sum_{t=0}^{N-1} f(\boldsymbol{\theta}^{(t)}).$$
 (10)

Having established a MARKOV chain, one can compute the marginal posterior (3) by binning  $\theta_i$  in the given parameter range and ignoring all other parameters [8].

#### B. Metropolis-Hastings Algorithm

### C. Sequential Monte Carlo (SMC)

For our purposes we chose the *Sequential Monte Carlo* sampling algorithm provided by PyMC3 [3], because it grants easy access to the marginal likelihood.

# D. Savage Dickey Density Ration (SDDR)

## IV. EXAMPLES

# A. Betabinomial example (coin flip)

Let us now consider as a starting example, the flipping of a two-sided coin, i.e. an experiment where we can measure either heads (H) or tails (T) with 50% probability, respectively. This, while simple, allows us an intuitive approach to Bayesian inference and model selection as well as to the MCMC techniques discussed before. Furthermore is this example easily altered to many real-life problems, such as birth rates, ..., or anything with the

option of either success or failure.

### 1. Analytical approach?

Assume we throw a coin 20 times. We observe 6 H and 14 T. "Is this a fair coin?" might be a question to ask yourself since the bias in outcome is quite large. Naively expecting a fair coin we could assign a *prior* to the probability of heads  $\theta$  as centred around 0.5, so for example a gaussian with  $\mu = 0.5$ ,  $\sigma = 0.1$ .

# 2. Numerical approach

# B. Fitting a polynomial of unknown degree

- 1. Analytical approach?
- 2. Numerical approach

#### V. DISCUSSION

#### VI. SUMMARY

<sup>[1]</sup> Arviz: Exploratory analysis of bayesian models — arviz dev documentation. URL https://arviz-devs.github.io/arviz/. last visit: March 2, 2021.

<sup>[2]</sup> Pymc3 documentation — pymc3 3.10.0 documentation, . URL https://docs.pymc.io/. last visit: March 2, 2021.

<sup>[3]</sup> Pymc3 documentation — sequential monte carlo, . URL https://docs.pymc.io/notebooks/SMC2\_gaussians. html. last visit: March 2, 2021.

<sup>[4]</sup> M. Kunz, R. Trotta, and D. R. Parkinson. Measuring the effective complexity of cosmological models. *Phys. Rev. D*, 74:023503, Jul 2006. doi: 10.1103/PhysRevD.74.023503. URL https://link.aps.org/doi/10.1103/PhysRevD.74.023503.

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