# A New Novel Trap For Neutral Atoms\*

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A modification to the standard Time-averaged Orbiting Potential (TOP) trap that allows for very isotropic potentials.

Usage: Secondary publications and information retrieval purposes.

PACS numbers: May be entered using the \pacs{#1} command.

Structure: You may use the description environment to structure your abstract; use the optional argument of the \item command to give the category of each item.

#### I. INTRODUCTION

Over a century ago, Boltzmann derived his famous Htheorem to describe entropy using mechanical arguments. His theories met with a vicious opposition because the time reversal symmetry inherent to mechanical systems conflicted with the irreversible nature of entropy. This was said to have aggravated his bouts of depression and ultimately led to his suicide in 1906. In a series of heated articles, Boltzmann did acknowledge that certain exotic cases exist that could lead to bizarre collective behavior but dismissed them as irrelevant to the description of naturally occurring systems. One case in particular was that a gas confined to a perfectly spherical harmonic potential wouldn't necessarily reach equilibrium. Namely, the monopole motion of the gas is undamped. Experimental study of this particular phenomenon has so far been prevented by the difficulty in generating a sufficiently isotropic harmonic potential.

The theoretical issues which fueled the debate stem from the Boltzmann equation. For a phase space distribution,  $f \equiv f(\mathbf{r}, \mathbf{p}, t)$ , the Boltzmann equation in its simplest form is

$$\frac{df}{dt} = I_{coll}[f], \qquad (1)$$

where  $I_{coll}[f]$  is a collision function, referred to as the collision integral. Essentially, the Boltzmann states that the phase space distribution of a gas approaches equilibrium at a rate proportional to the frequency of collisions. At equilibrium, the distribution becomes stationary and the collision integral vanishes. Expanding the derivative and including the explicit definition of  $I_{coll}[f]$  leads to the more standard form of the Boltzmann equation

$$\frac{\partial f}{\partial t} + \mathbf{v}_1 \cdot \nabla_{\mathbf{r}} f + \frac{\mathbf{F}}{m} \cdot \nabla_{\mathbf{v}_1} f =$$

$$\frac{\sigma_0}{4\pi} \int d^2 \Omega d^3 \mathbf{v}_2 |\mathbf{v}_2 - \mathbf{v}_1| \left( f(\mathbf{v}_1') f(\mathbf{v}_2') - f(\mathbf{v}_1) f(\mathbf{v}_2) \right)$$

where  $\sigma_0$  is the collision cross section. Energy and momentum conservation are implicitly assumed, but could easily be included in the collision integral using a delta function. The other major assumption is that the atoms can be treated as hard spheres, but this approximation is valid for s-wave collisions when mean field effects are negligible and the cloud is non-degenerate.

Distributions of the form

$$\log f = \alpha + \beta \cdot \mathbf{v} + \gamma v^2 \tag{3}$$

cause the collision integral to vanish due to conservation of energy and momentum and are referred to as collisionally invariant. These distributions, generally called local equilibrium distributions, do not necessarily satisfy Eq. 2 and constraints must be placed on  $\alpha$ ,  $\beta$ , and  $\gamma$ . For a sufficiently general outside potential, these constants can be constrained to produce an equilibrium distribution equivalent to Maxwell-Boltzmann. For certain special potentials, collisionally invariant distributions exist that satisfy Eq. 2. For an isotropic harmonic potential, the solution is essentially an equilibrium distribution undergoing temperature oscillations where the spatial distribution varies identically such that the distribution satisfies the equilibrium constraint at each point in time. Persistence of monopole oscillations can be described in various ways, but can be elucidated by the two particle case.

In the presence of spherical symmetry, the problem can be treated as analogously one dimensional and the radial motion of a single particle of mass m, energy E and angular momentum L is governed by the effective potential

$$V_e = \frac{L^2}{2mr^2} + \frac{1}{2}m\omega^2 r^2.$$
 (4)

The radial force,

$$m\frac{d^2r}{dt^2} = -\frac{d}{dr}V_e,\tag{5}$$

can be combined with the kinetic energy,

$$\frac{1}{2}m\left(\frac{dr}{dt}\right)^2 = E - V_e,\tag{6}$$

by integrating Eq. 5. This yields

$$\frac{d^2}{dt^2}r^2 = -\Omega^2 \left(r^2 - r_0^2\right),\tag{7}$$

<sup>\*</sup> A footnote to the article title

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where  $\Omega \equiv 2\omega$  and  $r_0^2 = E/(m\omega^2)$ , so that the square radius undergoes sinusoidal oscillations around its mean value  $r_0^2$  at a frequency of  $2\omega$ . If there are 2 particles, 1 and 2, each with individual values of E, L and  $r^2$ , each particle will oscillate at  $2\omega$  and taking the linear combination of their respective differential equations yields

$$\frac{d^2}{dt^2}r_t^2 = -\Omega^2 \left(r_t^2 - r_{0t}^2\right) \tag{8}$$

so their combined square radius,  $r_t^2 = r_1^2 + r_2^2$ , oscillates around its mean value,  $r_{0t}^2 = (E_1 + E_2)/(m\omega^2)$ . The magnitude of the monopole motion depends on the magnitude and relative phase of the individual particle trajectories. These individual quantities will abruptly change in the event of a collision. Assuming the collisions are local,  $r_1$ ,  $r_2$  and thus  $r_t^2$  will not change from the instant before to the instant after the collision. Similarly, momentum and energy conservation imply that  $\frac{d}{dt}r_t^2$  and  $r_{0t}^2$  are unchanged by the collision. These three continuities imply that the parameters and boundary conditions of Eq. 8 are matched directly before and after a collision. This ensures that neither the magnitude or phase of the oscillation will change as the result of a pairwise collsion. Generalizing to N atoms, so that

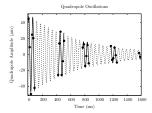
$$r_t^2 = \sum_{i=1}^{N} r_i^2 \tag{9}$$

and

$$r_{0t}^2 = \frac{1}{m\omega^2} \sum_{i=1}^{N} E_i, \tag{10}$$

one can see that the monopole mode is left unperturbed – and in particular undamped – by local, pairwise, momentum- and energy-conserving collisions.

The system becomes integrable and the energy in the monopole mode is a conserved quantity.



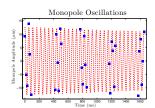


FIG. 1. Sample data for a driven quadrupole mode and monopole mode in a spherical trap. Black lines on the quadrupole data indicate a typical fitting procedure where individual periods taken in a single run are fit with an undamped sine wave to extract the amplitude.

a. A few notes on tags \tag{#1} requires the amsmath package. Place the \tag{#1} command before the \label{#1}, if any. The numbering produced by Damping vs. Collision Rate

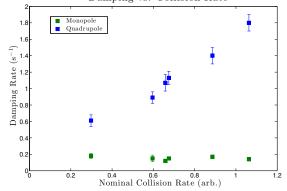


FIG. 2. Sample data for a driven quadrupole mode and monopole mode in a spherical trap. Black lines on the quadrupole data indicate a typical fitting procedure where individual periods taken in a single run are fit with an undamped sine wave to extract the amplitude.

\tag{#1} does not affect the automatic numbering in REVTEX; therefore, the number must be known ahead of time, and it must be manually adjusted if other equations are added. \tag{#1} works with both single-line and multiline equations. \tag{#1} should only be used in exceptional cases—do not use it to number many equations in your paper. Please note that this feature of the amsmath package is not compatible with the hyperref (6.77u) package.

Enclosing display math within \begin{subequations} and \end{subequations} will produce a set of equations that are labeled with letters, as shown in Eqs. (11b) and (11a) below. You may include any number of single-line and multiline equations, although it is probably not a good idea to follow one display math directly after another.

$$\mathcal{M} = ig_Z^2 (4E_1 E_2)^{1/2} (l_i^2)^{-1} (g_{\sigma_2}^e)^2 \chi_{-\sigma_2}(p_2)$$

$$\times [\epsilon_i]_{\sigma_1} \chi_{\sigma_1}(p_1).$$
(11a)

$$\left\{abc123456abcdef\alpha\beta\gamma\delta1234556\alpha\beta\frac{1\sum_{b}^{a}}{A^{2}}\right\},\qquad(11b)$$

Giving a \label{#1} command directly after the \begin{subequations}, allows you to reference all the equations in the subequations environment. For example, the equations in the preceding subequations environment were Eqs. (11).

## 1. Wide equations

The equation that follows is set in a wide format, i.e., it spans the full page. The wide format is reserved for long equations that cannot easily be set in a single column:

$$\mathcal{R}^{(d)} = g_{\sigma_2}^e \left( \frac{[\Gamma^Z(3,21)]_{\sigma_1}}{Q_{12}^2 - M_W^2} + \frac{[\Gamma^Z(13,2)]_{\sigma_1}}{Q_{13}^2 - M_W^2} \right) + x_W Q_e \left( \frac{[\Gamma^{\gamma}(3,21)]_{\sigma_1}}{Q_{12}^2 - M_W^2} + \frac{[\Gamma^{\gamma}(13,2)]_{\sigma_1}}{Q_{13}^2 - M_W^2} \right) . \tag{12}$$

This is typed to show how the output appears in wide format. (Incidentally, since there is no blank line between the equation environment above and the start of this paragraph, this paragraph is not indented.)

#### II. CROSS-REFERENCING

REVT<sub>E</sub>X will automatically number such things as sections, footnotes, equations, figure captions, and table captions. In order to reference them in text, use the \label{#1} and \ref{#1} commands. To reference a particular page, use the \pageref{#1} command.

The \label{#1} should appear within the section heading, within the footnote text, within the equation, or within the table or figure caption. The \ref{#1} command is used in text at the point where the reference is to be displayed. Some examples: Section I on page 1, Table I, and Fig. 3.

# III. FLOATS: FIGURES, TABLES, VIDEOS, ETC.

Figures and tables are usually allowed to "float", which means that their placement is determined by LATEX, while the document is being typeset.

Use the figure environment for a figure, the table environment for a table. In each case, use the \caption command within to give the text of the figure or table caption along with the \label command to provide a key for referring to this figure or table. The typical content of a figure is an image of some kind; that of a table is an alignment.

Insert an image using either the graphics or graphix packages, which define the \includegraphics{#1} com-

TABLE I. A table that fits into a single column of a two-column layout. Note that REVTEX 4 adjusts the intercolumn spacing so that the table fills the entire width of the column. Table captions are numbered automatically. This table illustrates left-, center-, decimal- and right-aligned columns, along with the use of the ruledtabular environment which sets the Scotch (double) rules above and below the alignment, per APS style.

Left <sup>a</sup>	$Centered^b$	Decimal	Right
1	2	3.001	4
10	20	30	40
100	200	300.0	400

<sup>&</sup>lt;sup>a</sup> Note a.

mand. (The two packages differ in respect of the optional arguments used to specify the orientation, scaling, and translation of the image.) To create an alignment, use the tabular environment.

The best place to locate the figure or table environment is immediately following its first reference in text; this sample document illustrates this practice for Fig. 3, which shows a figure that is small enough to fit in a single column.

In exceptional cases, you will need to move the float earlier in the document, as was done with Table II: LATEX's float placement algorithms need to know about a full-page-width float earlier.

Fig. 4 has content that is too wide for a single column, so the figure\* environment has been used.

The content of a table is typically a tabular environment, giving rows of type in aligned columns. Column entries separated by &'s, and each row ends with \\. The required argument for the tabular environment specifies how data are aligned in the columns. For instance, entries may be centered, left-justified, right-justified, aligned on a decimal point. Extra column-spacing may be be specified as well, although REVT<sub>E</sub>X 4 sets this spacing so that the columns fill the width of the table. Horizontal rules are typeset using the \hline command. The doubled (or Scotch) rules that appear at the top and bottom of a table can be achieved enclosing the tabular environment within a ruledtabular environment. Rows whose columns span multiple columns can be typeset using the \multicolumn{#1}{#2}{#3} command (for example, see the first row of Table II).

Tables I, II, III, and IV show various effects. A table that fits in a single column employs the table environment. Table II is a wide table, set with the table\* environment. Long tables may need to break across pages. The most straightforward way to accomplish this is to specify the [H] float placement on the table or table\* environment. However, the LATEX  $2_{\varepsilon}$  package longtable allows headers and footers to be specified for each page of the table. A simple example of the use of longtable can

Test Figure

FIG. 3. A figure caption. The figure captions are automatically numbered.

<sup>&</sup>lt;sup>b</sup> Note b.

### Wide Test Figure

FIG. 4. Use the figure\* environment to get a wide figure that spans the page in twocolumn formatting.

TABLE II. This is a wide table that spans the full page width in a two-column layout. It is formatted using the table\* environment. It also demonstates the use of \multicolumn in rows with entries that span more than one column.

	D	$1\atop 4h$	L	$O_{4h}^5$
Ion	1st alternative	2nd alternative	lst alternative	2nd alternative
K	(2e) + (2f)	(4i)	(2c) + (2d)	(4f)
Mn	$(2g)^{\mathbf{a}}$	(a) + (b) + (c) + (d)	(4e)	(2a) + (2b)
Cl	(a) + (b) + (c) + (d)	$(2g)^{\mathrm{a}}$	$(4e)^{a}$	
He	$(8r)^{a}$	$(4j)^{\mathrm{a}}$	$(4g)^{a}$	
Ag	. ,	$(4k)^{\mathrm{a}}$	/	$(4h)^{a}$

<sup>&</sup>lt;sup>a</sup> The z parameter of these positions is  $z \sim \frac{1}{4}$ .

be found in the file summary.tex that is included with the REVTEX 4 distribution.

There are two methods for setting footnotes within a table (these footnotes will be displayed directly below the table rather than at the bottom of the page or in the bibliography). The easiest and preferred method is just to use the \footnote{#1} command. This will automatically enumerate the footnotes with lowercase roman letters. However, it is sometimes necessary to have multiple entries in the table share the same footnote. In this case, there is no choice but to manually create the footnotes using \footnotemark[#1] and \footnotetext[#1]{#2}. #1 is a numeric value. Each time the same value for #1 is used, the same mark is produced in the table. The \footnotetext[#1]{#2} commands are placed after the tabular environment. Examine the LATEX source and output for Tables I and IV for examples.

Video 1 illustrates several features new with REVTEX4.1, starting with the video environment, which is in the same category with figure and table. The \setfloatlink command causes the title of the video to

TABLE III. Numbers in columns Three—Five are aligned with the "d" column specifier (requires the dcolumn package). Nonnumeric entries (those entries without a ".") in a "d" column are aligned on the decimal point. Use the "D" specifier for more complex layouts.

One	Two	Three	Four	Five
one	two	three	four	five
$_{\mathrm{He}}$	2	2.77234	45672.	0.69
$C^{a}$	$C_{\rm p}$	12537.64	37.66345	86.37

<sup>&</sup>lt;sup>a</sup> Some tables require footnotes.

be a hyperlink to the indicated URL; it may be used with any environment that takes the \caption command. The \href command has the same significance as it does in the context of the hyperref package: the second argument is a piece of text to be typeset in your document; the first is its hyperlink, a URL.

Physical Review style requires that the initial citation of figures or tables be in numerical order in text, so don't cite Fig. 4 until Fig. 3 has been cited.

TABLE IV. A table with numerous columns that still fits into a single column. Here, several entries share the same footnote. Inspect the LATEX input for this table to see exactly how it is done.

	$r_c$ (Å)	$r_0$ (Å)	$\kappa r_0$		$r_c$ (Å)	$r_0$ (Å)	$\kappa r_0$
Cu	0.800	14.10	2.550	Sna	0.680	1.870	3.700
Ag	0.990	15.90	2.710	$\mathrm{Pb^{b}}$	0.450	1.930	3.760
Au	1.150	15.90	2.710	$Ca^{c}$	0.750	2.170	3.560
Mg	0.490	17.60	3.200	$\mathrm{Sr^d}$	0.900	2.370	3.720
Zn	0.300	15.20	2.970	$\mathrm{Li}^{\mathrm{b}}$	0.380	1.730	2.830
$\operatorname{Cd}$	0.530	17.10	3.160	$Na^{e}$	0.760	2.110	3.120
$_{\mathrm{Hg}}$	0.550	17.80	3.220	$K^{e}$	1.120	2.620	3.480
Al	0.230	15.80	3.240	$Rb^{c}$	1.330	2.800	3.590
Ga	0.310	16.70	3.330	$Cs^d$	1.420	3.030	3.740
In	0.460	18.40	3.500	$\mathrm{Ba^e}$	0.960	2.460	3.780
Tl	0.480	18.90	3.550				

<sup>&</sup>lt;sup>a</sup> Here's the first, from Ref. ? .

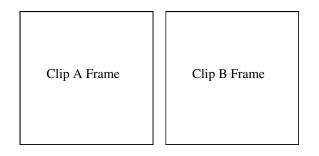
<sup>&</sup>lt;sup>b</sup> Some tables need more than one footnote.

<sup>&</sup>lt;sup>b</sup> Here's the second.

<sup>&</sup>lt;sup>c</sup> Here's the third.

<sup>&</sup>lt;sup>d</sup> Here's the fourth.

<sup>&</sup>lt;sup>e</sup> And etc.



Video 1. Students explain their initial idea about Newton's third law to a teaching assistant. Clip (a): same force. Clip (b): move backwards.

#### ACKNOWLEDGMENTS

We wish to acknowledge the support of the author community in using REVTEX, offering suggestions and encouragement, testing new versions, ....

### Appendix A: Appendixes

To start the appendixes, use the \appendix command. This signals that all following section commands refer to appendixes instead of regular sections. Therefore, the \appendix command should be used only once—to setup the section commands to act as appendixes. Thereafter normal section commands are used. The heading for a section can be left empty. For example,

\appendix
\section{}

will produce an appendix heading that says "APPENDIX A" and

\appendix \section{Background}

will produce an appendix heading that says "APPENDIX A: BACKGROUND" (note that the colon is set automatically).

If there is only one appendix, then the letter "A" should not appear. This is suppressed by using the star version of the appendix command (\appendix\* in the place of \appendix).

#### Appendix B: A little more on appendixes

Observe that this appendix was started by using

\section{A little more on appendixes}

Note the equation number in an appendix:

$$E = mc^2. (B1)$$

# 1. A subsection in an appendix

You can use a subsection or subsubsection in an appendix. Note the numbering: we are now in Appendix B 1.

Note the equation numbers in this appendix, produced with the subequations environment:

$$E = mc,$$
 (B2a)

$$E = mc^2, (B2b)$$

$$E \gtrsim mc^3$$
. (B2c)

They turn out to be Eqs. (B2a), (B2b), and (B2c).