## ONE HOUR THIRTY MINUTES

A list of constants is enclosed.

## UNIVERSITY OF MANCHESTER

Statistical Mechanics

7th June 2023, 2.00 p.m. - 3.30 p.m.

Answer  $\underline{\mathbf{ALL}}$  parts of question 1 and  $\underline{\mathbf{TWO}}$  other questions

Electronic calculators may be used, provided that they cannot store text.

The numbers are given as a guide to the relative weights of the different parts of each question.

## Information which may be used in this paper:

The density of states for a spinless particle in two dimensions is  $g(k) = \frac{Ak}{2\pi}$ . The symbol  $\beta$  is defined as  $\beta = 1/(k_B T)$ .

You may use  $k_B = 8.6173 \times 10^{-5} \text{ eV K}^{-1} \text{ and } \hbar c = 1973 \text{ eV Å}.$ 

The symbol  $n_Q$  is defined as  $n_Q = \left(\frac{mk_BT}{2\pi\hbar^2}\right)^{3/2}$ .

The translational partition function for a single non-relativistic particle in a box is  $Z_1 = Vg_s n_Q$ , where  $g_s$  is the spin degeneracy.

In the classical limit, the chemical potential for an ideal gas is given by  $\mu = -k_B T \ln \left( \frac{g_s n_Q}{n} \right)$ .

The Planck formula for the energy density of black-body radiation is

$$u(\omega) = \frac{\hbar}{\pi^2 c^3} \frac{\omega^3}{e^{\hbar \omega \beta} - 1}.$$

Stirling's approximation is  $\ln N! \approx N \ln N - N$  for  $N \gg 1$ .

$$\sum_{n=0}^{\infty} (n+1) x^n = (1-x)^{-2} = \text{for } |x| < 1.$$

The following integrals may be useful (note  $n!! \equiv n(n-2)(n-4)\dots 1$  for odd n.)

$$\int_{0}^{\infty} \frac{x^{1/2}}{e^{x} + 1} dx = 0.678094 \qquad \int_{0}^{\infty} \frac{x^{1/2}}{e^{x} - 1} dx = 2.31516$$

$$\int_{0}^{\infty} \frac{x}{e^{x} + 1} dx = \frac{\pi^{2}}{12} \qquad \int_{0}^{\infty} \frac{x}{e^{x} - 1} dx = \frac{\pi^{2}}{6}$$

$$\int_{0}^{\infty} \frac{x^{3/2}}{e^{x} + 1} dx = 1.15280 \qquad \int_{0}^{\infty} \frac{x^{3/2}}{e^{x} - 1} dx = 1.78329$$

$$\int_{0}^{\infty} \frac{x^{2}}{e^{x} + 1} dx = 1.80309 \qquad \int_{0}^{\infty} \frac{x^{2}}{e^{x} - 1} dx = 2.40411$$

$$\int_{0}^{\infty} \frac{x^{5/2}}{e^{x} + 1} dx = 3.08259 \qquad \int_{0}^{\infty} \frac{x^{5/2}}{e^{x} - 1} dx = 3.74453$$

$$\int_{0}^{\infty} \frac{x^{3}}{e^{x} + 1} dx = \frac{7\pi^{4}}{120} \qquad \int_{0}^{\infty} \frac{x^{3}}{e^{x} - 1} dx = \frac{\pi^{4}}{15}$$

$$\int_{0}^{\infty} x^{n} e^{-x/a} dx = n! \, a^{n+1} \qquad \int_{0}^{\infty} x^{n} e^{-x^{2}/a^{2}} dx = \frac{(n-1)!! \sqrt{\pi} a^{n+1}}{2^{(n/2+1)}} \quad \text{for even } n.$$

2 of 6 P.T.O

1. a) Distinguish between the microstates and macrostates of a many-particle system, giving examples.

[5 marks]

b) One microjoule of energy is added to a kilogram of water at temperature  $27^{\circ}$ C. By what factor does the number of available microstates increase? (Express your answer as  $10^{x}$ .)

[5 marks]

c) The energy E of a photon gas at temperature T is proportional to  $VT^4$ , and its entropy is given by  $S=\frac{4}{3}E/T$ . Derive an expression for the pressure P in terms of E and the volume V.

[4 marks]

d) Two otherwise identical black bodies have different temperatures. The temperature of the hotter body is such that  $k_BT=0.4$  eV. At one particular frequency, such that  $\hbar\omega=0.2$  eV, the intensity of the radiation from the hotter body is twice that from the cooler body. What is the temperature of the cooler body?

[5 marks]

e) Electrons are confined with a density of  $3\times10^{28}$  m<sup>-3</sup>. Calculate the Fermi temperature. Estimate the chemical potential at T=1000 K and at  $T=10^6$  K.

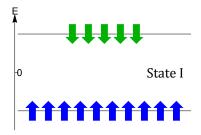
[6 marks]

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2. i) Write down the Boltzmann distribution, clearly explaining the meaning of the symbols you use and stating in what circumstances it holds.

[5 marks]

ii) Consider a system of N distinguishable non-interacting spin- $\frac{1}{2}$  particles, each of magnetic moment  $\mu$ . The following diagram represents the energy levels of each particle in an external field  $B_1$ , and the arrows indicate the occupancy of the two levels at some temperature  $T_1$ . In the diagram, one third of particles are in the higher-energy state. (Only a fraction of the particles are shown.)



Give an expression for  $T_1$  in terms of  $\mu$  and  $B_1$ .

Denoting the state shown by (I), draw two more diagrams indicating energy levels and occupancies for the following pairs of temperature and magnetic field:

(II) 
$$B = 2B_1, T = T_1$$

(III) 
$$B = B_1, T = \frac{1}{2}T_1.$$

For the transitions from (I) to (II) and from (II) to (III), is work done on or by the system? Is heat transferred to or from the system?

[8 marks]

iii) Briefly describe how such a system may be used to reach low temperatures.

[4 marks]

iv) For the system above, find the partition function and hence the magnetisation and entropy of the system as a function of temperature. Sketch the magnetisation as a function of temperature for two values of the magnetic field,  $B_1$  and  $2B_1$ . Hence comment on whether the process described above can be used to reach zero temperature.

[8 marks]

**3.** i) By considering the wave function in two dimensions of a spinless particle free to move within a square of side L, derive the density of states g(k).

[7 marks]

ii) Consider a two-dimensional relativistic spin-zero gas of indistinguishable particles, whose energy and momentum are related by  $E = \sqrt{p^2c^2 + m^2c^4}$ . The particles are free to move within an area A, and the temperature and density are such that the gas is in the classical regime.

Show that the single-particle partition function of the system is given by

$$Z_1 = \frac{A}{2\pi} \left( \frac{\mathrm{e}^{-mc^2\beta} (\beta mc^2 + 1)}{\beta^2 \hbar^2 c^2} \right).$$

What is the partition function of the N-particle system?

[8 marks]

[You may use the following result: 
$$\int_{x_0}^{\infty} x e^{-x} dx = (x_0 + 1) e^{-x_0}.$$
 ]

iii) Show that the average energy of the system is given by  $N\langle E_1\rangle$ , where

$$\langle E_1 \rangle = mc^2 + 2k_B T - \frac{mc^2 k_B T}{mc^2 + k_B T}.$$

Find the limits both for  $mc^2 \gg k_B T$  and for  $mc^2 \ll k_B T$ , and comment on your results in the context of equipartition.

[10 marks]

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**4.** i) Explain the conditions for a system to be described by the grand-canonical distribution. State the associated thermodynamic potential.

Define the grand partition function, defining any symbols you use, and state how it is related to the relevant thermodynamic potential.

[7 marks]

ii) A system has a set of single-particle states of energy  $\{\varepsilon_i\}$ , and is in contact with a reservoir of spin- $\frac{1}{2}$  particles at chemical potential  $\mu$  and temperature T. Find an expression for the average number of particles in a state of energy  $\varepsilon_i$ .

[3 marks]

iii) At very high temperatures, electrons and positrons can be regarded as a dense gas of ultra-relativistic particles. The number of particles is not conserved. Explain why that means that the chemical potential is equal to zero.

Show that the number of electrons is equal to

$$N_{\rm e} = CV \left(\frac{k_{\rm B}T}{\hbar c}\right)^3,$$

and determine the constant C.

[8 marks]

iv) Find an expression for the total energy of the electrons, and comment on its temperature dependence. Qualitatively, how would the result change for spinless particles?

[7 marks]

## END OF EXAMINATION PAPER