

Electromagnetism 2

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Chapter 1

Sources of Radiation

Radiation is the phenomenon of energy being transported to an observer. Radiation must be disconnected from its source. Plane electromagnetic waves satisfy conditions of being radiation, and are thus known as free fields. The source of electromagnetic radiation must be charges as those generate the electric and magnetic fields. However, electromagnetic radiation is only generated through accelerating charges. For a free field, there must be locations where radiation propagates while $\rho = 0$.

1.1 Charges

1.1.1 Stationary Charges

For a stationary charge, we can write,

$$\mathbf{E} = \frac{1}{4\pi\epsilon_0} \frac{q(\mathbf{r} - \mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|^3} \propto \frac{1}{|\mathbf{r} - \mathbf{r}'|^2} \quad \mathbf{B} = 0 \quad (1.1)$$

We can clearly see that a stationary charge cannot radiate as it has no component in \mathbf{b} . Furthermore, the energy flow varies as $1/r^2$, which require $\mathbf{E} \propto \frac{1}{r}$ and $\mathbf{B} \propto \frac{1}{r}$ which neither field satisfies.

1.1.2 Charges with Uniform Velocity

We can perform a Lorentz boost in the case of a moving charge,

$$\mathbf{E}(\mathbf{r}) = \frac{q}{4\pi\epsilon_0} \frac{1 - \beta^2}{(1 - \beta^2 \sin^2 \theta)^{3/2}} \frac{\hat{\mathbf{r}}}{|\mathbf{r}|^2}. \quad (1.2)$$

We find that the electric flux still varies as $\frac{1}{r^2}$, thus cannot . Furthermore, we can move into the charge's rest frame where it still varies as $\frac{1}{r^2}$. This is the case as charge is both a conserved and Lorentz invariant quantity.

We could also make the an argument using the Biot-Savarte law,

$$\mathbf{B}(\mathbf{r}) = \frac{\mu_0}{4\pi} \frac{q\mathbf{v} \times \hat{\mathbf{r}}}{|\mathbf{r}|^2} \quad (1.3)$$

which still varies as $\frac{1}{r^2}$. Thus, we require acceleration of the charge for radiation to occur.

The reason we require that the electric field and magnetic field be proportional to $1/r$ is because of energy conservation. Let us recall that

$$\mathbf{S} = \frac{1}{\mu_0} \mathbf{E} \times \mathbf{B} \quad (1.4)$$

thus $S \propto EB$. We require,

$$\int S \cdot d\mathbf{A} = \text{Constant} \quad (1.5)$$

for energy conservation. Since $dA \propto r^2$, we require $S \propto 1/r^2$, and thus $E, B \propto 1/r$.

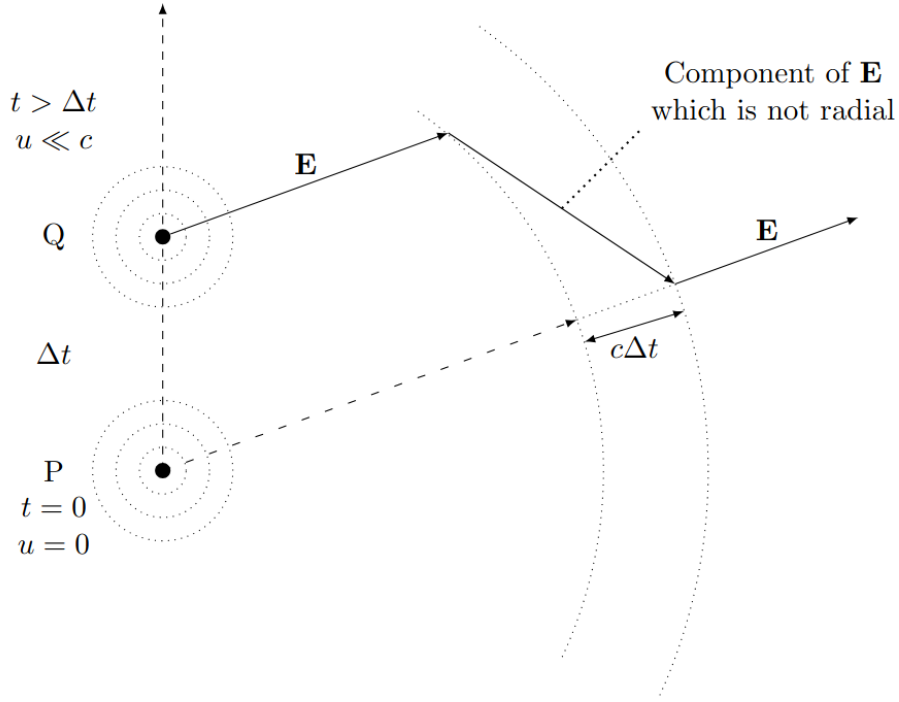


Figure 1.1

1.2 Retarded Potentials

If we recall eqs. (A.6) and (A.7) of the sourced potentials, we can analyse more closely the static and dynamic cases to gain more insight into radiation phenomena.

1.2.1 Statics

In the static case, we clearly see that we recover Poisson's equation, which we can solve trivially.

1.2.2 Time-Dependent case

In the time dependent case, we require that the electric field propagates at the speed of light c . Thus, we wish to obtain a scalar potential $\phi(\mathbf{r}, t)$ due to a volume of charge $\rho dV'$ at \mathbf{r}' . We must use the charge density which existed in that volume at an earlier time τ ,

$$\tau = t - \frac{|\mathbf{r} - \mathbf{r}'|}{c} \quad (1.6)$$

where $\frac{|\mathbf{r} - \mathbf{r}'|}{c}$ is the time taken for information to travel from \mathbf{r}' to \mathbf{r} . We can call τ the *retarded time*. We can thus write down the retarded potentials,

$$\phi(\mathbf{r}, t) = \frac{1}{4\pi\epsilon_0} \int_V \frac{\rho(\mathbf{r}', \tau)}{|\mathbf{r} - \mathbf{r}'|} dV' \quad (1.7)$$

$$\mathbf{A}(\mathbf{r}, t) = \frac{\mu_0}{4\pi} \int_V \frac{\mathbf{J}(\mathbf{r}', \tau)}{|\mathbf{r} - \mathbf{r}'|} dV' \quad (1.8)$$

1.3 Radiation from an accelerated charge

1.3.1 Displaced point charge

Consider figure 1.1, initially at rest at point P in free space. It is given a sudden impulse and undergoes a sudden acceleration a over a short time Δt , such that when the charge reaches Q it is moving at a constant velocity $u = a\Delta t \ll c$. If we consider the observer at a far away distance r , they will still see the charge at its original, stationary position. In order to see the charge move, the observer must wait a

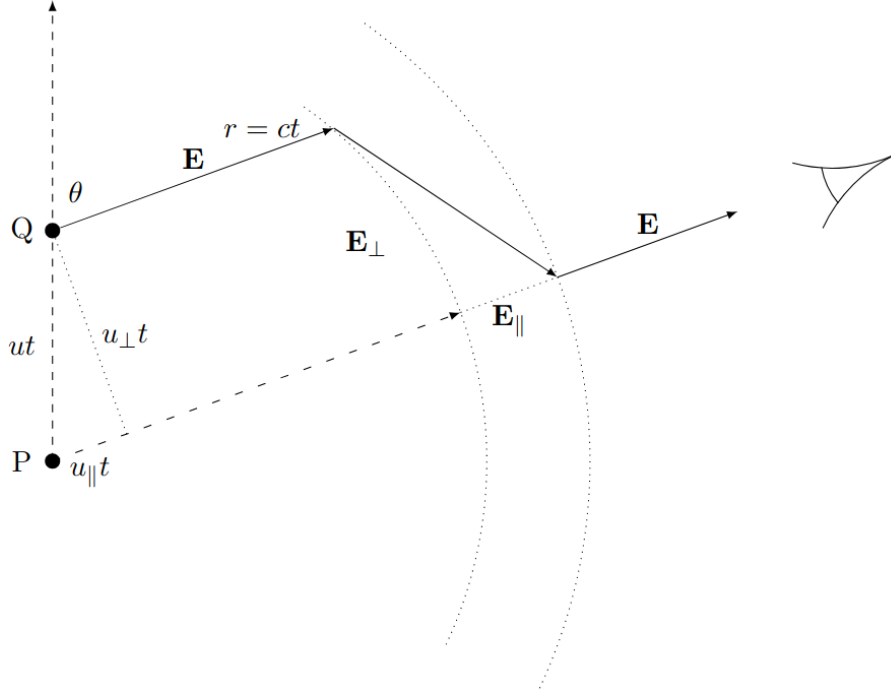


Figure 1.2

time $t = r/c$. In this time, there must be a boundary where the field configurations move from the old stationary state to the moving state. This boundary will have a width $c\Delta t$ and will must move away from the charge's location at a speed c .

Let us recall that $\nabla \cdot \mathbf{E} = 0$, so there can be no discontinuities in the field lines generated by the charge. The boundary thus presents itself as a *kink* in the electric field lines, as shown in figure 1.1. This is a component of \mathbf{E} which is not radial.

Let us now consider the observer looking at the charge at some angle θ while it undergoes its final motion, reaching a position ut . Let us note that the boundary of observing the charge's motion is

$$r = ct. \quad (1.9)$$

According to the observer, the charge appears at $u_{\parallel}t$ parallel to their line of sight and $u_{\perp}t$ perpendicular to their line of sight. We can then relate these to the parallel and perpendicular components of the kink E_{\parallel} and E_{\perp} ,

$$\frac{E_{\perp}}{E_{\parallel}} = \frac{u_{\perp}t}{c\Delta t}. \quad (1.10)$$

If we note that $u_{\perp} = a_{\perp}\Delta t$, we can write,

$$E_{\perp} = E_{\parallel} \frac{a_{\perp}r}{c^2}. \quad (1.11)$$

Furthermore, since the charge's field is radial, we can say,

$$E_{\parallel} = \frac{1}{4\pi\epsilon_0} \frac{q}{r^2} \quad (1.12)$$

so we can write,

$$E_{\perp} = \frac{1}{4\pi\epsilon_0} \frac{qa_{\perp}}{c^2 r}. \quad (1.13)$$

Let us summarise what we have found using this thought experiment,

1. The field $E_{\perp} \propto 1/r$ in the kink region.

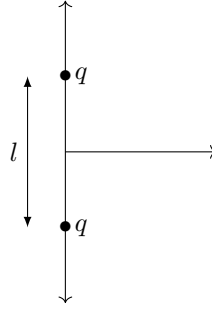


Figure 1.3

2. Recalling equation (A.11), we see $B_{\perp} = E_{\perp}/c$ in the kink region.

3. The magnitude of the Poynting vector $\mathbf{S} = \frac{1}{\mu_0} \mathbf{E} \times \mathbf{B}$ is, $S \propto EB \propto 1/r^2$

which satisfy our conditions for radiation.

1.3.2 Radiation pattern from a displaced point charge

We wish to now consider the electric field and the power flux (Poynting vector) as time dependent functions, noting that we will require using the retarded time. By noting that $a_{\perp} = a \sin \theta$, we can simply write,

$$|\mathbf{E}(\mathbf{r}, t)| = \frac{1}{4\pi\epsilon_0} \frac{q|\mathbf{a}(t - \frac{r}{c})| \sin \theta}{c^2 r}. \quad (1.14)$$

Then, we can write the Poynting vector as,

$$|\mathbf{S}(\mathbf{r}, t)| = \frac{1}{16\pi^2\epsilon_0} \frac{q^2 |\mathbf{a}^2(t - \frac{r}{c})| \sin^2 \theta}{c^3 r^2}. \quad (1.15)$$

1.3.3 Total Power Radiated

The Poynting vector is the directional energy flux per unit area, per unit time. Thus, to find the power radiated, we simply compute the integral,

$$P(t) = \oint \mathbf{S} \cdot d\mathbf{A} = \int_0^{2\pi} \int_0^{\pi} S r^2 \sin \theta d\theta d\phi. \quad (1.16)$$

Substituting expression (1.15), we obtain Larmor's formula,

$$P(t) = \frac{q^2 [a(t - \frac{r}{c})]^2}{6\pi\epsilon_0 c^3}. \quad (1.17)$$

1.4 Radiation from an Oscillating Electric Dipole

Consider an oscillating dipole, with charges separated by a length l and oscillating along the z axis. From Larmor's formula, $P(t) \propto (qa(\tau))^2$. We can approximate the oscillating charges as a charge distribution with charge per unit length μ and length l . We can write,

$$\begin{aligned} qa(\tau) &= \mu l a(\tau) \\ &= \mu \frac{dv}{d\tau} l \\ &= l \frac{d(\mu v(\tau))}{d\tau} \end{aligned} \quad (1.18)$$

where we find $\mu v(\tau)$ to have units of current, i.e., $[\mu v(\tau)] = [A]$. We can represent this as a current,

$$\mu v(\tau) = I(\tau) = I_0 \cos \omega t \quad (1.19)$$

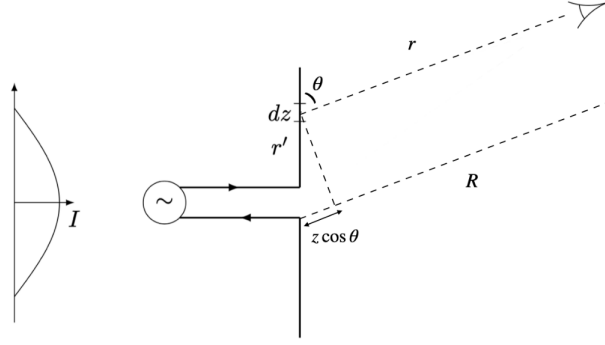


Figure 1.4

where I_0 and ω are the corresponding amplitude and angular frequency of oscillation. Thus, we can write $qa(\tau)$ as,

$$qa(\tau) = -\omega L I_0 \sin(\omega t - kr). \quad (1.20)$$

By substitution, we can find the corresponding electric and magnetic fields,

$$E_{\perp} = \frac{I_0 l \sin \theta}{4\pi \epsilon_0 r c^2} \omega \sin(\omega t - kr) \quad B_{\perp} = \frac{I_0 l \sin \theta}{4\pi \epsilon_0 r c^3} \omega \sin(\omega t - kr) \quad (1.21)$$

and the corresponding power emitted is,

$$P = \left[\frac{l^2 \omega^2}{6\pi \epsilon_0 c^3} \right] I_0^2 \sin^2(\omega t - kr). \quad (1.22)$$

1.4.1 Radiation Resistance

If we recall the formula $P = I^2 R$, we find that we can extract a resistance from equation (1.22). This is known as the *radiation resistance*, given by,

$$R_{\text{rad}} = \frac{l^2 \omega^2}{6\pi \epsilon_0 c^3} \quad (1.23)$$

1.5 Half-Wave Antenna

Let us consider a more realistic antenna, as in figure 1.4. This is a centre fed, thin, straight antenna of length l . The solution for the current distribution of this antenna can be obtained from Maxwell's equations, whose solution is,

$$I(z) = I_0 \sin \left(m\pi \left(\frac{|z|}{l} + \frac{1}{2} \right) \right) \quad (1.24)$$

where m is an integer. We can integrate the Hertzian dipole over the length of the antenna in order to obtain an expression for the electric field, we have

$$dE = \frac{I(z) dz \sin \theta}{4\pi \epsilon_0 r c^2} \omega \sin(\omega t - kr). \quad (1.25)$$

Before we integrate this, we must consider the phase $\omega t - kr$. We find that there will be a time-lag of $\Delta t \frac{z \cos \theta}{c}$, with a resultant lag in phase of $kz \cos \theta$. Furthermore, we can approximate $r \approx R - z \cos \theta$, thus our integral becomes,

$$E = \frac{\omega \sin \theta}{4\pi \epsilon_0 R c^2} \int_{-l/2}^{l/2} I(z) \sin(\omega t - kR + kz \cos \theta) dz \quad (1.26)$$

and computing this,

$$E = \frac{I_0 l}{4\pi \epsilon_0 R c^2} \left[\frac{2 \cos \left(\frac{\pi}{2} \cos \theta \right)}{\pi \sin \theta} \right] \omega \sin(\omega t - kR). \quad (1.27)$$

We wish to then find the time-averaged poynting vector, computed by $\langle S \rangle = \frac{1}{\mu_0} \langle \mathbf{E} \times \mathbf{B} \rangle$, which we find to be,

$$\langle S \rangle = \frac{I_0^2}{8\pi^2 \mu_0 \varepsilon_0^2 R^2 c^3} \frac{\cos^2 \left(\frac{\pi}{2} \cos \theta \right)}{\sin^2 \theta}. \quad (1.28)$$

We can then compute the time average of the power radiated by $\langle P \rangle = \oint \langle S \rangle dA$, which requires us to compute,

$$\langle P \rangle = \frac{I_0^2}{4\pi \mu_0 \varepsilon_0^2 R^2 c^3} \int_0^\pi \frac{\cos^2 \left(\frac{\pi}{2} \cos \theta \right)}{\sin \theta} d\theta \quad (1.29)$$

We then find,

$$\langle P \rangle \simeq 1.22 \frac{I_0^2}{r\pi \varepsilon_0 c} \simeq 0.194 Z_0 I_{\text{rms}}^2 \quad (1.30)$$

where we define $Z_0 = \frac{1}{\varepsilon_0 c} = \sqrt{\frac{\mu_0}{\varepsilon}}$ and $I_{\text{rms}} = \frac{I_0}{2}$.

Chapter 2

Radiation in Matter

2.1 Maxwell's Equations in a Medium

Let us recall electromagnetic fields in materials.

2.1.1 Electrics

The \mathbf{P} field is the dipole moment per unit volume, and describes bound charges, such that,

$$\nabla \cdot \mathbf{P} = \rho_b \qquad \mathbf{P} \cdot \hat{\mathbf{n}} = \sigma_b \qquad (2.1)$$

where $_b$ indicates a bound charge, and σ_b is the bound surface charge. The total charge is composed of bound charge and free charge, such that $\rho = \rho_b + \rho_f$. We can then define the \mathbf{D} field, such that,

$$\nabla \cdot \mathbf{D} = \rho_f \qquad (2.2)$$

for linear dielectrics, and,

$$\mathbf{D} = \epsilon_0 \mathbf{E} + \mathbf{P} = \epsilon_0 \epsilon_r \mathbf{E} \qquad (2.3)$$

where $\epsilon_r = 1 + \chi_E$.

2.1.2 Magnetics

In magnetics, materials can have different types of magnetism:

- *Paramagnetism*: Magnetisation parallel to \mathbf{B} , dipoles align with \mathbf{H} .
- *Diamagnetism*: Magnetisation is antiparallel with \mathbf{B} , dipoles anti-align with \mathbf{H} .
- *Ferromagnetism*: Magnetisation is retained even after \mathbf{B} is removed, regions of dipoles align together with \mathbf{H} .

We can define the magnetisation \mathbf{M} as the magnetic dipole moment per unit volume, from which we can obtain the bound and surface currents,

$$\mathbf{j}_b = \nabla \times \mathbf{M} \qquad \mathbf{K}_b = \mathbf{M} \times \hat{\mathbf{n}}. \qquad (2.4)$$

We can then define a \mathbf{H} field for free currents such that,

$$\nabla \times \mathbf{H} = \mathbf{J}_f \qquad (2.5)$$

where,

$$\mathbf{H} = \frac{1}{\mu_0} \mathbf{B} - \mathbf{M} = \frac{1}{\chi_M} \mathbf{M} = \frac{1}{\mu_0 \mu_r} \mathbf{B}. \qquad (2.6)$$

2.1.3 Time dependent case

$$\nabla \cdot \mathbf{D} = \rho_f \qquad \nabla \cdot \mathbf{B} = 0 \qquad (2.7)$$

$$\nabla \times \mathbf{H} = \frac{\partial \mathbf{D}}{\partial t} + \mathbf{j}_f \qquad \nabla \times \mathbf{E} = -\frac{\partial \mathbf{B}}{\partial t} \qquad (2.8)$$

2.2 Plane Waves in Lossy Media

Consider electromagnetic waves travelling in the z direction such that,

$$\mathbf{E} = (E_x, 0, 0) \quad E_x = E_0 e^{i(\omega t - kz)} \quad (2.9)$$

$$\mathbf{B} = (0, B_y, 0) \quad B_y = B_0 e^{i(\omega t - kz)} \quad (2.10)$$

Considering Faraday's law, we get,

$$-ikE_0 e^{i(\omega t - kz)} = -i\omega B_0 e^{i(\omega t - kz)} \quad (2.11)$$

and thus,

$$\frac{B_0}{E_0} = \frac{k}{\omega}. \quad (2.12)$$

Recalling Ohm's law $\mathbf{J} = \sigma \mathbf{E}$ where σ is the conductivity, and Ampere's law, we obtain,

$$k^2 = \frac{\omega^2}{c^2} \epsilon_r \mu_r - i\omega \mu_r \mu_0 \sigma \quad (2.13)$$

which is the dispersion relation for a linear, isotropic, dielectric medium. We see that this expression is complex, and reduces to the vacuum condition under vacuum.

2.2.1 Conducting Media

Let us consider a conducting media with dielectric properties and free current \mathbf{J}_f . Using Ampere's law,

$$\begin{aligned} \nabla \times \mathbf{H} &= \frac{\partial \mathbf{D}}{\partial t} + \mathbf{J}_f \\ &= i\omega \epsilon_0 \epsilon_r \mathbf{E} + \sigma \mathbf{E} \\ &= (i\omega \epsilon_0 \epsilon_r + \sigma) \mathbf{E} \\ &= i\omega \left(\epsilon_0 \epsilon_r + \frac{\sigma}{i\omega} \right) \mathbf{E}. \end{aligned} \quad (2.14)$$

We can then define the *complex permittivity*,

$$\epsilon_c = \epsilon_0 \epsilon_r - \frac{i\sigma}{\omega} \quad (2.15)$$

which we apply to conducting media. In the case of complex media, Ampere's law becomes $\nabla \times \mathbf{H} = i\omega \epsilon_c \mathbf{E}$, but otherwise, we simply replace $\epsilon_0 \epsilon_r$ with ϵ_c to retrieve Maxwell's equations for conducting media.

Appendix A

Revision Equations

A.1 Potentials

Static Electric Potential

$$\mathbf{E} = -\nabla\phi \quad (\text{A.1})$$

$$\nabla^2\phi = -\frac{\rho}{\varepsilon_0} \quad (\text{A.2})$$

Static Magnetic Potential

$$\mathbf{B} = \nabla \times \mathbf{A} \quad (\text{A.3})$$

$$-\nabla \times \mathbf{B} = \nabla^2 \mathbf{A} = -\mu_0 \mathbf{J} \quad (\text{A.4})$$

if we choose,

$$\mathbf{A} \rightarrow \mathbf{A} + \nabla\psi \quad \nabla \cdot \mathbf{A} = 0. \quad (\text{A.5})$$

Dynamic Potentials

$$\nabla^2 \mathbf{A} - \mu_0 \varepsilon_0 \frac{\partial \mathbf{A}}{\partial t} = -\mu_0 \mathbf{J} \quad (\text{A.6})$$

$$\nabla^2 \phi - \mu_0 \varepsilon_0 \frac{\partial \phi}{\partial t} = -\frac{\rho}{\varepsilon_0} \quad (\text{A.7})$$

A.2 Gauges

Coloumb Gauge

$$\nabla \cdot \mathbf{A} = 0 \quad (\text{A.8})$$

Lorenz Gauge

$$\nabla \cdot \mathbf{A} + \frac{1}{c^2} \frac{\partial \phi}{\partial t} = 0 \quad (\text{A.9})$$

A.3 Vector Identities

Laplacian of a Vector

$$\nabla \times \nabla \times \mathbf{A} = \nabla (\nabla \cdot \mathbf{A}) - \nabla^2 \mathbf{A} \quad (\text{A.10})$$

A.4 EM Waves

Relation between \mathbf{E} and \mathbf{B}

$$\mathbf{B} = \frac{1}{c} \hat{\mathbf{k}} \times \mathbf{E}. \quad (\text{A.11})$$