Abstract

This work investigates the effect of anisotropy in interface energy on orientation selection during repeated nucleation in the polycrystalline film growth and the implications for the formation of crystallographic texture. For this purpose, a methodology was developed, which allowed to obtain the nucleation probability as function of the substrate and nucleus orientation. For every pair of these orientations, the nucleation probability was derived from the equilibrium stable shape of the nucleus at those conditions. In order to determine the arbitrarily oriented nucleus shape, a multi-phase field model incorporating anisotropic interface energy was extensively developed.

Multiple model variants were developed and methodology modified several times to solve various difficulties emerging due to inherent limitations of the models. Own solver for the governing equations was written in MATLAB. Two model-independent benchmark problems were developed to quantitatively compare performance of the developed models in terms of how they reproduced a) the anisotropic curvature driving force and b) the anisotropic curvature driving force when combined with strong anisotropy in kinetic coefficient. In both cases the system consisted of a shrinking grain with a well-defined shape (Wulff shape in a) and circle in b)). Even though the models were comparably consistent in simulations of shrinking Wulff shapes and circles, an additional benchmark problem concerning triple junction angles revealed thermodynamic inconsistency in two of the three models, causing them to be reliable only in isotropic or weakly anisotropic systems.

Analytic solution to the geometric problem was found, which allowed to implement a simpler and more efficient algorithm. So-called shape factor-orientation maps were constructed with high resolution and were used as input in a Monte Carlo simulation of growing polycrystal to demonstrate how the nucleation with anisotropic interface energy may affect the course of texture evolution. New modes of texture evolution were identified and the insight was used to qualitatively explain an experimentally observed peculiar texture evolution in electrodeposited nickel.

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Chapter 1

Introduction

1.1 General introduction

Irrespective of the deposition method, it is known that the process parameters often have pronounced effect on crystallographic texture and microstructure of polycrystalline deposits. Evaporation, sputtering, chemical vapour deposition or electrodeposition are just examples of these methods. Each is tied with a unique combination of physical and chemical interactions at the interface between the deposit and parent phase. The as-deposited texture is the result of synergy of many factors. However, the anisotropy on the deposit/parent phase interface during the film growth certainly plays an important role.

Typically, nucleation in film deposition is studied as a phenomenon occurring on substrates, rather than on the deposit itself. However, there seems to be no obvious reason to neglect such an option, given the complicated anisotropy of grain boundary energy and local deposition conditions, varying possibly in both space and time.

Despite the fact, that the anisotropy in interface energy is widely accepted to be one of the fundamental texture-forming factors, its effect on the orientation selection in nucleation remains fundamentally unclear.

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1.2 Textures in coatings and interface energy anisotropy

A widely accepted theory by Thompson [60] suggests that the texture in deposits is controlled by two main factors: minimization of interface energy and minimization of strain energy. The first favors growth of grains with low-energy interfaces facing the parent phase and the second aligns the grains in the deposit to accommodate the stress. The energy densities are explicitly expressed for an abstracted columnar microstructure and compared as functions of the film thickness. He concluded that thin enough deposits have the texture controlled by minimization of interface energy. Above certain thickness, the strain energy overtakes. Later, additional factors contributing to the strain energy were included in the theory [14], allowing sensible interpretation of textures observed in sputter-deposited CdTe and electrodeposited Cu, but the results of some others are in agreement with the theory as well [56, 57].

A more phenomenological work by Rasmussen et al. [53] studied Cu electrodeposits on crystalline and amorphous substrates and proposed that in electrodeposition, the two competing texture-forming factors are the influence of the substrate and that of deposition conditions. Through the film thickness, three zones are defined: A (substrate-controlled texture), B (mixed) and C (deposition-conditions-controlled texture). The effect of the substrate in zone A can be e.g. epitaxy or chemical reactions, in zone C the anisotropy modifications are believed to be due to specific adsorption of different species [4, 48]. A more recent study of Ni electrodeposition [2] investigated the texture evolution through thickness with microscopic detail and showed, that the transition zone B does not seem to always occur as a simple mixture of A and C.

Two different theories by Pangarov [50, 51] and Kozlov [29] suggested that the textures of the electrodeposits are originated in the anisotropy of the nucleation barrier (work of formation), but each used a different reasoning. Pangarov treated 2D nuclei only and made the work of formation depending exclusively on overpotential and lattice parameters. This theory was inconsistent with experimental observation of nuclei character as function of overpotential [9]. Kozlov studied the formation of initial texture on amorphous substrates in fcc metals by means of atomistic nucleation theory. Specifically, he computed the nucleation barriers and sizes of nuclei with three significant orientations. His prediction of initial $\langle 111 \rangle$ texture is in agreement with Thompson. It can thus be deduced, that in the initial stages of growth in fcc metals on amorphous substrates, both the nucleation and growth may lead to the same fiber texture by means of interface energy minimization.

In practice, the interface energy anisotropy or deposition rate anisotropy cannot

be measured directly during the deposition, which complicates the interpretation of the observed textures and microstructures. Additionally, the effect of the two above physical quantities on the polycrystalline deposit growth was shown to be equivalent in phase-field simulations to some extent [64]. It seems analogical to equilibrium Wulff shapes and kinetic Wulff shapes, which are geometrically similar [28]. However, this makes the fundamental discussion about how the growth is affected by interface anisotropy in general even more challenging, because the growth rate and interface energy are very different quantities, taking part in different processes. This highlights the utility of simulation tools, which allow to set the anisotropy in interface energy and kinetic coefficient independently, because in practice the individual effects may be hard or impossible to distinguish. Among such tools is (multi-)phase field method.

1.3 TODO Goals of the thesis

This thesis explores the effect of anisotropic interface energy on nucleation and growth of polycrystalline deposits in various simulation concepts. It is known that the anisotropy in interface energy affects the equilibrium stable shape of a particle (on a plane). Also, it is known that the nucleation probability is proportional to the particle area (in 3d to its volume).

The goals of the thesis are:

- 1. Develop a methodology which would allow comparison of nucleation probability of differently oriented nuclei on an arbitrarily oriented substrate in 2D
- 2. Create a program which will
- 1. Develop a methodology to simulate the equilibrium stable shapes of particles with anisotropic interface energy on a plane using the multiphase field method

2.

A perspective approach for simulation of a growing film including nucleation is multi-phase field method, which allows inclusion of anisotropic interface energy, but also anisotropy of the kinetic coefficient (deposition rate).

The goal

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An established multi-phase field method was extended so that the equilibrium stable shapes could be obtained.

Chapter 2

Particles with anisotropic interface energy

2.1 Isolated particle with anisotropic interface energy

It was already discovered in 1901, how to determine the stable equilibrium shape of a particle with crystalline anisotropy [67]. The chemical potential $\mu(\theta)$ must be constant along the surface of equilibrium shape [5]. Assuming that the surface is only under the influence of anisotropic capillary force, the chemical potential on the curved surface in 3D space depends on the local principal curvatures and interface energy anisotropy as follows from the Herring equation [23, 26]. In 2D the surfaces become curves and the expression for the chemical potential is simplified to

$$\mu(\theta) = \tilde{\sigma}(\theta)K\,,\tag{2.1}$$

with local interface normal angle θ and local curvature K. The interface stiffness $\tilde{\sigma}(\theta)$ is the thermodynamic driving force for the capillary motion and the local curvature κ is the thermodynamic coordinate. The interface stiffness is

$$\tilde{\sigma}(\theta) = \sigma_0[h(\theta) + h''(\theta)], \qquad (2.2)$$

where $h(\theta)$ and $h(\theta)''$ stand for the anisotropy function and its second derivative, respectively. The anisotropy function defines the interface energy $\sigma(\theta)$

$$\sigma(\theta) = \sigma_0 h(\theta) \,, \tag{2.3}$$

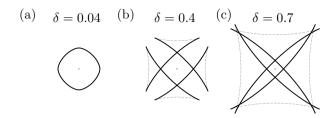


Figure 2.1: Examples of full Wulff plots (i.e. $-\pi \le \theta \le \pi$) of 4-fold symmetry with indicated strengths of anisotropy. The dotted segments are the forbidden orientations, solid lines are the allowed ones. The "ears" in b) do not contribute to the isolated particle shape. However, at even stronger anisotropy in c) another intersection occurs among the branches, which introduces new solutions to the equilibrium shape.

where σ_0 is a constant with physical dimensions of interface energy (i.e. J/m²). In this chapter it is assumed, that the anisotropy function takes the following form

$$f(\theta - \alpha) = 1 + \delta \cos(n(\theta - \alpha)), \qquad (2.4)$$

where $0 \le \delta < 1$ is the strength of anisotropy, n the order of symmetry and α the angle rotating the anisotropy function, representing the 2D crystal orientation. A normalized strength of anisotropy $0 \le \Omega < (n^2 - 1)$ can be defined as $\Omega = \delta(n^2 - 1)$. In this paper, weak anisotropy will specifically refer to $\Omega < 1$ and strong anisotropy to $\Omega \ge 1$.

The interface stiffness then reads

$$\tilde{\sigma}(\theta) = \sigma_0 \{ 1 - \Omega \cos[n(\theta - \alpha)] \}. \tag{2.5}$$

The isolated particle with anisotropic interface energy takes on the well-known Wulff shape, which in 2D can be described as a parametric curve $\boldsymbol{w}(\theta) = [w_x(\theta), w_y(\theta)]^T[10, 17, 28]$

$$w_x(\theta) = R_W[h(\theta)\cos(\theta) - h'(\theta)\sin(\theta)]$$
 (2.6)

$$w_y(\theta) = R_W[h(\theta)\sin(\theta) + h'(\theta)\cos(\theta)]. \tag{2.7}$$

See Figure 2.1 for examples of 4-fold symmetry at various strengths of anisotropy δ .

From equation (2.5), it is easily seen that for $\Omega > 1$, there are such θ s, for which the interface stiffness is negative. These normal angles are unstable [5, 17, 28], hence they do not occur on the stable equilibrium Wulff shape (see the dotted

segments in Figure 2.1) and are called missing or forbidden. Closer inspection of the anisotropy function (2.4) and its derivatives reveals that the forbidden angles correspond to interface energies near maxima of the anisotropy function. In fact, the latter is a general observation, independent on the particular anisotropy function (2.4).

For $\Omega > 1$, the equations (2.6)-(2.7) must be understood in a piece-wise fashion for the intervals of allowed θ s. Each of these allowed-angles interval produces a single branch in the Wulff curve. The discontinuity in allowed angles causes that the Wulff shape is no longer a smooth and continuous closed curve (like in Figure 2.1a). Instead, it is a piece-wise continuous and self-intersecting curve (Figure 2.1b-c), where the individual branches enclose a polygon-like closed curve with corners. In the case of Figure 2.1b, the outer parts of the shape (so-called ears) cannot contribute to the (isolated) crystal shape, but for strong enough anisotropy (like in Figure 2.1c) the ears cross again, giving rise to a multitude of other stable equilibrium solutions. That is because every single of the four spikes can be part of the solution but does not need to be. This repeated crossing of the ears and its implications for the equilibrium shapes have not been reported in any work known to the author, hence it will be elaborated on in detail.

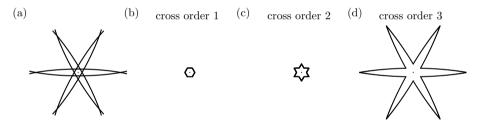


Figure 2.2: Orders of corners, defined by the branches intersections, shown for 6-fold anisotropy and $\delta=0.7$. In a) there are all the allowed angles, in b)-d) are the stable equilibrium shapes defined by the b) first order, c) second order and d) third order of the corners. Any closed curve is a stable equilibrium shape, which connects these corners along the allowed-angles branches and has the interface normals pointing outwards in every point.

To emphasize the general features of this "crossed-ears" type of solutions, the Figure 2.2a provides another example of very strong, but 6-fold anisotropy, where the ears cross even twice behind the polygon-like isolated Wulff shape. Then, Figures 2.2b-d show the different possible shapes, where the outer corners always correspond to branch intersection of the same type. The type of intersection can be characterized by the radial distance from the center, based on which they can

be ordered. Practically, the corner of the first order is a corner on the polygon-like shape, where two branches forming adjacent sides of the polygon-like shape cross (Figure 2.2b). A corner of the second order is farther from the center and is the intersection of two branches forming two polygon sides separated by a single side (Figure 2.2c). A corner of the third order (Figure 2.2d) is the intersection of two branches separated by two sides etc.

For convenience, the shapes in Figures 2.2b-d will be called isolated Wulff shapes of first, second and third order, respectively.

Below, a summary of some key points is given:

- for $n \ge 2$ and $\delta \ge 1/(n^2 1)$ the Wulff shape comprises of n mutually crossing branches and the isolated shape (of first order) exhibits corners,
- for $n \ge 4$ multiple branches intersections occur for sufficiently large δ ,
- the branches intersections (corners) can be characterized by their radial distance from the center,
- the upper limit on the number of intersections of a single branch with others depends on the total number of branches, i.e. on the order of symmetry n,
- as long as the orientation of the curve is respected (i.e. interface normal of all segments on the shape points outwards), any closed curve connecting the above mentioned corners along the branches is a stable equilibrium shape,
- the shape of higher order contains all the lower-order shapes and has larger area than any of them. The first-order shape thus has the smallest area and, consequently the least total interface energy.

	order of corner				
n	1	2	3		
4	0.06667	0.60000	X		
	(1.00000)	(9.00000)			
6	0.02857	0.16422	0.62136		
	(1.00000)	(5.74773)	(21.74773)		

Table 2.1: The approximate minimal strengths of anisotropy at which the corners of indicated order appear on the Wulff shapes in 4-fold and 6-fold anisotropy. The first number is the strength of anisotropy δ and in parenthesis is the corresponding normalized strength of anisotropy $\Omega = \delta(n^2 - 1)$.

In Table 2.1, there are provided the minimal strengths of anisotropy at which the corners of indicated order appear on the Wulff shape in 4-fold and 6-fold anisotropy. Of these, only the first-order corners have an analytic expression, the higher-order corners were obtained numerically and thus are only approximate.

2.2 Particle with anisotropic interface energy on a plane (Winterbottom construction)

The problem of equilibrium shape of a particle with anisotropic interface energy on a planar substrate was treated in various works. It can either be derived from variational principles by total interface energy minimization [33, 66] or by the Cahn-Hoffman vector formalism [11].

The shape is strongly determined by the force balance of the three meeting interfaces: substrate-parent phase (σ_1) , particle-parent phase (σ_2) and particlesubstrate (σ_3) . There are two such contact points in 2D (in 3D it is a triple line) and both must be in force balance in order to sustain a stable equilibrium shape.

Winterbottom construction identifies which segments on the full Wulff shape are (or can be) part of the equilibrium shape of a particle with anisotropic interface energy on a plane.

For tractability of the general solution (of the particle shape), it is essential to assume co-planarity of the particle-substrate and substrate-parent phase interfaces [5, 11, 66].

In 2D, the stability of the shape derives from the stability of the two contact points. In this geometry, all interface energies are known and the orientations of two of the interfaces are known (the former substrate plane). The orientation of the last interface (the particle) in the left and right contact points are expressed by normal angles θ_c^L, θ_c^R , respectively. These are found using the Young's equation with inclination dependence, solved in every contact point

$$\sigma_1 - \sigma_3 = \sigma_2(\theta_2)\sin(\theta_2) + \sigma_2'(\theta_2)\cos(\theta_2), \qquad (2.8)$$

which can be written in a non-dimensional form

$$\Gamma = h(\theta_2)\sin(\theta_2) + h'(\theta_2)\cos(\theta_2), \qquad (2.9)$$

where $\Gamma = [\sigma_1 - \sigma_3]/\sigma_2^0$ will be denoted wetting parameter in this paper. It can be both positive or negative.

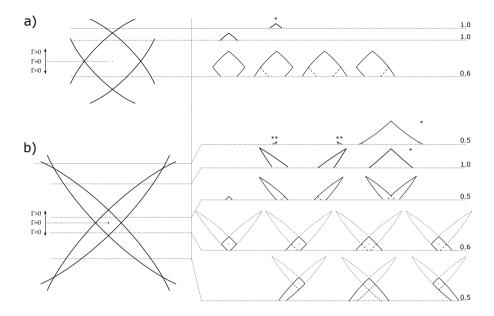


Figure 2.3: Complete list of non-trivial equilibrium shapes of a particle with anisotropic interface energy on a plane (right part of the sketches) as derived from full Wulff shapes (left part of the sketches) using generalized Winterbottom construction. In a) for strength of anisotropy $\delta=0.3$, in b) with $\delta=0.7$. The horizontal dotted lines indicate different truncating lines for the Wulff shape and at the same time the substrate for the equilibrium shapes. The scale of the equilibrium shapes relative to the full Wulff shapes on the left is provided on the right on every respective line. Inverted shape solutions of first order are indicated by asterisk *, of the second order by **. The dashed lines within some of the the equilibrium shapes show the segments on the Wulff shape which were enclosed by the equilibrium shape. The dotted lines above the equilibrium shapes in b) indicate the crossed-ears solutions, which either may be present both, separately or both be absent. The new solutions can be seen in b).

The solutions are from the ranges $3\pi/2 > \theta_c^L > \pi/2$ and $\pi/2 > \theta_c^R \ge -\pi/2$ and must be allowed angles. When (2.9) is compared to (2.7), it is obvious that the solutions θ_c^L, θ_c^R are found as normal angles in points with y coordinate equal to Γ on an unit-radius Wulff shape. Thus, the resulting particle shape is a segment of the isolated shape, where the position of the sectioning plane Γ is determined by the force balance. That is the Winterbottom construction. Originally, it served to find the shape of the least interface energy, but Bao [5] generalized it for any stable equilibrium shape. For illustration of the generalized

Winterbottom construction depending on the position of the truncating line see Figure 2.3.

When the anisotropy is strong, multiple stable equilibrium shapes may exist even when the isolated shape has only a single equilibrium stable shape. All these solutions are illustrated in Figure 2.3a for 4-fold anisotropy, moderately strong [5]. In Figure 2.3b all possible solutions are presented for very strong 4-fold anisotropy, which involves the new solutions derived from the second-order corners.

If multiple solutions existed in the isolated particle in the above-truncating-line part, this multiplicity is present in the possible solutions on the plane too (Figure 2.3b). Compared to the isolated case, additional solutions are present, when the truncating line passes below the center of the isolated shape, with the shape being defined also by the "ears" extending behind the corners (in both Figure 2.3a and 2.3b).

An exceptional class of the solutions is the *inverted shape solution* (indicated by asterisk in Figure 2.3), which is found *below* the sectioning plane, as described by Bao [5] on a shape like in Figure 2.3a. However, due to the inverted shape symmetry, it is not clear, what operation(s) of symmetry the "inversion" refers to according to Bao and how to construct the solution in a general case. More detailed characterization of this solution type was given in the following section 2.3. It should be noted that in systems with corners of higher order, multiple orders of the inversion shape are possible (also indicated in Figure 2.3 by asterisk * and described in more detail in 2.3).

2.3 Inverted solution in Winterbottom construction

The inverted shape in a Winterbottom construction with single plane was described by [5] in a highly symmetric case. The same idea was used much sooner in the Winterbottom construction in a corner (meaning Wulff shape resting on two lines at once, intersecting in a corner) by Zia et al. [71] ("Summertop construction"). Below, the inverted shape construction is described, taking into account the higher-order intersection solutions and arbitrary rotations.

Let the Wulff shape exhibit strong anisotropy, i.e. be there at least corners of first order. Further, let the wetting parameter Γ be such, that the truncating line passes at least above the isolated shape of the first order, i.e. the first-order shape is submerged below the substrate plane.

Then it may happen, that the truncating line intersects two branches, which crossed below it (see Figure 2.4a1). Such situation allows a solution, where the

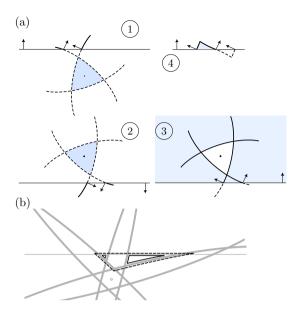


Figure 2.4: Illustration of the inverted Winterbottom construction. In (a) explained in 4 steps and in (b) illustration of inverted solutions of higher orders. In (a) the plane normal points from the substrate to the parent phase, the solid lines of the Wulff plot are thus above the substrate plane and the dashed ones are below. Interface normals in the contact points are indicated as well. In 1 is the initial state, in 2 after center inversion, in 3 after the curve orientation flip and in 4 is the resulting shape next to the original part of the Wulff plot. In (b) the dashed black line indicates inverted shape of the first order, the solid line that of second order.

equilibrium shape is the part below the substrate plane, but above the corner where the two branches intersect. Two steps must be followed to obtain the inverted equilibrium shape:

- 1. point inversion of the whole system (Wulff plot + truncating line) about the center of the Wulff shape (see Figure 2.4a2)
- 2. flip orientation of both the inverted Wulff plot and inverted truncating line, i.e. flip the inverted inward and outwards directions (see Figure 2.4a3)

The first operation places the truncating line below the said corner, where the two branches intersect. At the same time, the whole shape is inverted. The second operation turns the outside of the Wulff plot to be inside - i.e. instead

of having a crystal in the shape of inverted Wulff plot in some medium, we have a medium in the shape of Wulff plot surrounded by the crystal.

Then, the resulting inverted shape as in Figure 2.4a4 is obtained as an enclosed intersection of the anisotropic Wulff plot and a half-space above the line. All the interface normal orientations on the surface of the inverted shape are the same as on the non-inverted part of the Wulff plot. Importantly, this includes the normal angles in the contact points, as can be seen in Figure 2.4a4.

The order of the intersection in which the two Wulff branches intersect, determines the order of the inverted shape. Figure 2.4b shows an example of strong 6-fold anisotropy with such truncation, that it allows two orders of inverted shapes at once.

2.4 Contact point stability

The contact points mentioned above are essentially triple junctions. The Young's equation with inclination dependence (2.8) represents equilibrium of the tractions due to the interfaces on the junction in the x direction. It concerns a special case, where the position of two of the three interfaces are fixed to lay on the x axis and the third has inclination-dependent interface energy. However, the *stability* of the triple junction configuration is not guranteed by Young's equation alone [34]. In the present case, the configuration means a solution θ_2 to (2.8). Technically, the solutions to (2.8) are stationary points of the triple junction energy as function of virtual translations, but not necessarily its minimum (i.e. the stable configuration). Stability of the found solutions should be thus tested using additional conditions derived from second-order derivatives of the triple junction energy with respect to the virtual translations [34]:

$$0 < \tilde{\sigma}_1 s_1 + \tilde{\sigma}_2 s_2 + \tilde{\sigma}_3 s_3 \tag{2.10}$$

$$0 < \tilde{\sigma}_1 \tilde{\sigma}_2 S_{12} + \tilde{\sigma}_2 \tilde{\sigma}_3 S_{23} + \tilde{\sigma}_1 \tilde{\sigma}_3 S_{13}, \qquad (2.11)$$

where $\tilde{\sigma}_i$ are the respective interface stiffnesses, $s_i = \cos^2(\theta_i)/d_i$, $S_{ij} = \sin^2(\theta_i - \theta_j)/d_id_j$. In the present geometry, interfaces 1 and 3 are fixed with $\theta_1 = \pm \pi/2$ and $\theta_3 = \mp \pi/2$, hence it follows that $s_1 = s_3 = 0$ and also that $S_{13} = 0$. Note also that $s_2 > 0$ and $S_{12} = S_{23} = S > 0$. By the geometry alone, the conditions are thus simplified to

$$0 < \tilde{\sigma}_2 \tag{2.12}$$

$$0 < \tilde{\sigma}_1 \tilde{\sigma}_2 + \tilde{\sigma}_2 \tilde{\sigma}_3 \,, \tag{2.13}$$

where only the interface stiffnesses matter. Using the first equation, the second one may further be simplified

$$0 < \tilde{\sigma}_1 + \tilde{\sigma}_3. \tag{2.14}$$

The condition (2.12) implies, that the particle-parent phase interface must be oriented under allowed angle in the contact point. That is fulfilled automatically by the truncated Wulff shape solution, because it contains only the allowed angles.

Note that the interfaces 1 and 3 have fixed orientations and are assumed to be planar. The condition (2.14) is surely fulfilled when both interfaces 1 and 3 have positive interface stiffness in their orientations. That holds automatically in weak or no anisotropies. When either or both of the two interfaces exhibit strong anisotropy, it means that there are normal angles where the interface stiffness is negative, hence the condition (2.14) may be violated and must be verified in the given configuration.

2.5 Selected geometric properties of Wulff shape

With $h(\theta) = 1 + \delta \cos(n\theta)$, the minimal distance R_{min} between the Wulff shape center and the contour can be related to R_W as

$$R_{min} = R_W(1 - \delta), \qquad (2.15)$$

which holds for arbitrarily strong anisotropy because the minimal-radius point normal is always inclined under a non-missing angle. This formula was used to find the radius R_W of the phase field contour. Then, the phase field contour was scaled to unit radius and compared to the analytic Wulff shape (by means of Hausdorff distance, see the next subsection).

The radius of curvature $\rho(\theta)$ of Wulff shape is locally

$$\varrho(\theta) = R_W[h''(\theta) + h(\theta)], \qquad (2.16)$$

which may technically get negative for some θ . Because that does not have any geometric interpretation, the Wulff shape does not contain such oriented segments and the angles θ_m fulfilling $\varrho(\theta_m) \leq 0$ are called missing or forbidden.

For a given radius R_W , the area of a Wulff shape A_W differs from the area of a circle $A = \pi R_W^2$ by the anisotropy-dependent factor $C_W(\Omega, n) = A_W/A$. For weak anisotropy $\Omega < 1$ there is an analytic expression

$$C_W = 1 - \frac{\Omega^2}{n^2 - 1} \,, \tag{2.17}$$

but there is none for strong anisotropy $\Omega > 1$. The factor $C_W(\Omega, n)$ was obtained numerically for n=4 in the range $0 \le \Omega \le 7.5$ (the Wulff shape was plotted, its area numerically computed and divided by πR^2). There, the Ω -dependence was fitted very well by a polynomial of 4-th order $C_W(\Omega, 4) = \sum_{i=0}^4 a_i \Omega^{N-i}$ with $a_0 = -0.00032, a_1 = 0.00639, a_2 = -0.04219, a_3 = 0.00034, a_4 = 1.00000$.

2.6 Shrinkage rate of a Wulff shape

The normal velocity of anisotropic curvature-driven interface in 2D is [1]

$$v_n(\theta, t) = -\frac{\mu}{\varrho(\theta, t)} \sigma_0[h''(\theta) + h(\theta)]$$
 (2.18)

where μ is the interface mobility and $\sigma_0[h''(\theta) + h(\theta)]$ is interface stiffness. When it is the Wulff shape shrinking, the radius of curvature $\varrho(\theta, t)$ can be expressed as in equation (2.16) and the inclination-dependent part of the interface stiffness is cancelled out. The normal velocity is then constant along the perimeter

$$v_n(t) = -\frac{\mu \sigma_0}{R_W(t)} \,. \tag{2.19}$$

The sign is such that the interface moves towards the center of curvature.

Because the area of a Wulff shape is $A_W = \pi R_W^2 C_W(\Omega, n)$, the shrinkage rate of a Wulff shape is

$$\frac{\mathrm{d}A_W}{\mathrm{d}t} = \frac{\mathrm{d}}{\mathrm{d}t} [\pi R_W^2(t) C_W(\Omega, n)]$$

$$= 2\pi C_W(\Omega, n) R_W(t) \frac{\mathrm{d}R_W}{\mathrm{d}t}.$$
(2.20)

Because the normal velocity is constant along the perimeter, the point R_{min} from equation (2.15) moves also by $v_n(t)$. Importantly, the normal of interface in R_{min} is aligned with the radial direction of the polar coordinates, hence we can write $v_n = \mathrm{d}R_{min}/\mathrm{d}t$. From equation (2.15) then must be

$$\frac{\mathrm{d}R_W}{\mathrm{d}t} = \frac{1}{1-\delta} \frac{\mathrm{d}R_{min}}{\mathrm{d}t} = \frac{1}{1-\delta} v_n(t), \qquad (2.21)$$

which together with 2.19 implies in 2.20

$$\frac{\mathrm{d}A_W}{\mathrm{d}t} = 2\pi\mu\sigma_0 \frac{C_W(\Omega, n)}{1 - \delta} \tag{2.22}$$

Taylor and Cahn [59] proved that the shrinkage rate for the equilibrium shape must be a constant, which is met by the final expression.

2.7 Summary and the author's contribution

This chapter contained fundamental information about the equilibrium stable shapes of particles with anisotropic interface energy, assuming absence of other driving forces which might affect their shape. Isolated particles as well as those on a plane were assessed in 2D. New stable equilibrium shapes were described in the case of very strong anisotropy, where the Wulff plot may self-intersect multiple times, thus giving rise to the solutions containing parts formed by crossed "ears". The notion of "order of the intersection" was introduced to generalize the description of the emerging solution type.

Regarding the particle on a plane, the inverted solution type was described in detail allowing its construction for arbitrary nucleus orientation and wetting. Inverted solutions of higher orders were described as well.

The author is not aware of another work applying the stability conditions for the triple junction configuration to the nucleus contact points. Later, it is shown that for particles with strongly anisotropic interface energy, the stability condition is a very strong influence on the orientation selection in repeated nucleation.

Some basic geometric properties of the Wulff shapes were reviewed, which were later utilized in the phase-field model benchmarkig. Specifically, the steady-state shrinkage rate of a Wulff shape is an useful formula derived by the author.

Chapter 3

Phase field method

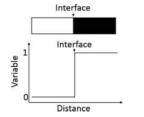
3.1 General introduction

Phase field method is a diffuse-interface approach to solving moving-boundary problems. In materials science it is usually used as meso-scale method. It has numerous applications such as e.g. (multi-component) solidification, grain growth, Ostwald ripening, solid-solid phase transitions (martensitic, precipitation, spinodal...), two-phase flow, nucleation, wetting and much more.

The method is based on the principle of total system energy minimization. The free energy density of the system is a function of all the possibly mutually coupled physical fields, including the fields representing the phases. These are typically called phase fields, order parameters or continuum fields. The other (bulk) physical fields can be e.g. concentration, temperature, electric charge or others. The total energy of the system is then spatial integral of the free energy density over the simulation domain - the free energy functional.

The governing equation(s) for the phase fields are derived using principles of variational calculus, which ensure the fields evolution so that the free energy functional is minimized along the direction of steepest descent, i.e. the most likely way towards the equilibrium. Additionally, the curvature driving force and the possible coupling with the bulk fields are represented. That is the reason for high versatility of the method and its capability to simulate coupled multiphysics.

Due to the thermodynamic nature of the method, the interface energy is naturally included as an input parameter. Multigrain or multiphase systems 18 ______ PHASE FIELD METHOD



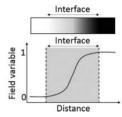


Figure 3.1: Illustration of the diffuse interface in phase field method (on the right) and the comparison to an actual, sharp interface (on the left). [6]

with interfaces of various properties can be simulated with multi phase field models. Some of these allow to include also inclination dependence of interface energy. In fact, it was phase field simulation which revealed the paramount importance of interface energy anisotropy in pattern formation during dendritic solidification [27].

A phase (or phases) in an inhomogeneous system is described by a continuous function (called phase field, order parameter or continuum field), having different constant values in the bulk of the phases (typically 1 and 0 or 1 and -1). At the interface between phases, the phase field varies smoothly between the values and the interface region is thus characterized by a certain width, as sketched in Figure 3.1. This interface width is an important input parameter in phase field models. In quantitative phase field models, the model behavior is in principle not affected by the particular value of interface width chosen, which allows to make quantitative simulations at mesoscale (whereas the real interface width is in the order of units of nm).

The governing equations for the phase fields are typically either conservative (Cahn-Hilliard type of equation, conserving volume) or non-conservative (Allen-Cahn type of equation). The latter is more relevant for this work. In the following, no coupling to other physical fields is assumed.

3.2 Allen-Cahn equation for two-phase system

It was shown that the Alen-Cahn equation is a diffuse-interface approximation of a curvature-driven system. It represents evolution of a non-conserved field

 $\eta(\mathbf{r})$ in a domain $\Omega \in \mathbb{R}^3$ with boundary $\partial \Omega$ and can be written as

$$\frac{\partial \eta}{\partial t} = -L \frac{\delta F}{\delta \eta} \quad \text{in } \Omega \tag{3.1}$$

$$\nabla \eta \cdot \boldsymbol{n} = 0 \quad \text{at } \partial \Omega \,, \tag{3.2}$$

where L is a positive constant and the function $\frac{\delta F}{\delta \eta}$ stands for functional derivative of the free energy functional F. The functional is an integral with free energy density $f(\eta, \nabla \eta) = f_{hom}(\eta) + f_{int}(\nabla \eta)$ comprising of the homogeneous and interfacial free energy densities

$$F = \int_{V} f_{hom}(\eta) + f_{int}(\nabla \eta) dV$$
 (3.3)

$$= \int_{V} m\eta^{2} (1 - \eta)^{2} + \frac{\kappa}{2} |\nabla \eta|^{2} dV$$
(3.4)

with positive constants m, κ being the double barrier height and gradient energy coefficient, respectively. The homogeneous free energy density f_{hom} in the form of the double-well function assures two stable states for the field variable in values 0 and 1. The gradient energy contribution is lessened by the diffuse interface widening, whereas conversely the homogeneous free contribution is lessened by its narrowing. When minimizing the full integral, a balance between the two effects is reached in equilibrium.

As described in more detail in the Appendix A, the functional derivative is computed as

$$\frac{\delta F}{\delta \eta} = \left[\frac{\partial f}{\partial \eta_p} - \nabla \cdot \frac{\partial f}{\partial (\nabla \eta_p)} \right] \tag{3.5}$$

where $\nabla \cdot \partial f / \partial (\nabla \eta_p)$ is divergence of vector field $\partial f / \partial (\nabla \eta_p)$ defined by relation

$$\frac{\partial f}{\partial (\nabla \eta_p)} = \frac{\partial f}{\partial (\partial_x \eta_p)} \mathbf{n}_x + \frac{\partial f}{\partial (\partial_y \eta_p)} \mathbf{n}_y + \frac{\partial f}{\partial (\partial_z \eta_p)} \mathbf{n}_z$$
 (3.6)

with $\partial_x, \partial_y, \partial_z$ being operators for unidirectional derivatives in the corresponding directions and $\mathbf{n}_x, \mathbf{n}_y, \mathbf{n}_z$ coordinate base vectors.

When applied applied to the above functional, it will result in the Allen-Cahn equation

$$\frac{\partial \eta}{\partial t} = -L[f'_{hom}(\eta) - \kappa \nabla^2 \eta] \tag{3.7}$$

$$= -L[2m\eta(1-\eta)(1-2\eta) - \kappa \nabla^2 \eta].$$
 (3.8)

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In equilibrium, the interface is static, $\partial \eta/\partial t = 0$, which means that the forces narrowing and widening the interface are in balance, i.e. $f'_{hom}(\eta) = \kappa \nabla^2 \eta$, or equivalently (in 1D)

$$2\eta(1-\eta)(1-2\eta) = \frac{\kappa}{m} \frac{\partial^2 \eta}{\partial x^2}.$$
 (3.9)

Solution to this equation provides the equilibrium profile of the phase field (centered in x = 0), which in this case is the following sigmoid function

$$\eta(x) = \frac{1}{2} \tanh\left(\sqrt{\frac{m}{2\kappa}}x\right).$$
(3.10)

The equilibrium also implies, that the homogeneous and interfacial energies are equal except for a constant, i.e. $f_{hom} = f_{int} + C$, which can be shown to be zero C = 0 (see e.g. [42]).

Interface width l_{int} (unit m) can be defined using the reciprocal maximal slope in the interface (present in $\eta = 0.5$)

$$l_{int} = \frac{1}{\mathrm{d}\eta/\mathrm{d}x|_{max}} = \left. \frac{1}{\sqrt{2f_{hom}/\kappa}} \right|_{n=0.5} = 2\sqrt{2}\sqrt{\frac{\kappa}{m}}, \quad (3.11)$$

where the relation $d\eta/dx = \sqrt{2f_{hom}/\kappa}$ is derived from $f_{hom} = f_{int}$.

The excess free energy σ (the *interface energy*, unit J/m²) at a flat interface can be computed as an 1D integral of the free energy density along perpendicular line through the interface

$$\sigma = \int_{-\infty}^{+\infty} f_{hom} + \frac{\kappa}{2} |\nabla \eta|^2 dx.$$
 (3.12)

From $f_{hom} = f_{int}$ it can be written $dx = \sqrt{\kappa/2f_{hom}}d\eta$, which allows to solve (3.12) by substitution as follows

$$\sigma = \int_{-\infty}^{+\infty} 2f_{hom} \mathrm{d}x \tag{3.13}$$

$$= \int_0^1 \sqrt{2f_{hom}\kappa} d\eta \tag{3.14}$$

$$=\frac{\sqrt{2}}{6}\sqrt{m\kappa}\,. (3.15)$$

In practice, the model is parameterized by the interface width l_{int} and interface energy σ , from which the model parameters m, κ must be computed. The

necessary formulas are obtained from (3.11) and (3.15) as

$$\kappa = \frac{3}{2}\sigma l_{int} \tag{3.16}$$

$$m = 12 \frac{\sigma}{l_{int}} \tag{3.17}$$

The l_{int} is chosen depending on the scale of the problem, disregarding the real physical interface width.

As shown in [42], the following equality between the model parameters and the interface mobility μ holds

$$\mu\sigma = L\kappa \,, \tag{3.18}$$

which allows to express the phase field kinetic coefficient as

$$L = \frac{2}{3} \frac{\mu}{l_{int}} \,, \tag{3.19}$$

where the equation (3.16) was used.

3.3 Allen-Cahn equation for two-phase system including anisotropic interface energy

3.3.1 Fundamentals

The most straightforward way to introduce anisotropy in the interface energy in the model is to make either of the model parameters κ, m a function of the inclination angle respecting both the anisotropy function and the defining parameter relations.

Tschukin [63] proposes the terminology of classical and natural models (corresponding to the notation VIW and VIE, respectively, as used by Fleck [18], standing for Variable Interface Width and Variable Interface Energy). The classical models [17, 27, 35, 59, 65], introduce anisotropy in the gradient energy coefficient κ , whereas the natural models [18, 32, 42, 61] in both the gradient energy coefficient κ and homogeneous energy density barrier m. The main difference between the classical and natural formulations is that in the first case the diffuse interface width varies proportionally to the local interface energy (see equation (3.11)), whereas in the latter the interface width can be decoupled from the interface energy and the interface width maintained constant (see (3.11) and (3.15)). Below, more details are provided on both model formulations.

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Let's assume that the interface energy is inclination-dependent in the following form (in 2D)

$$\sigma(\theta) = \sigma_0 h(\theta) \,, \tag{3.20}$$

with the inclination angle θ of the interface normal $\mathbf{n} = [n_x, n_y]^{\mathrm{T}}$ being defined as $\theta = \mathrm{atan2d}(n_y/n_x)$ and σ_0 constant with physical dimensions J/m². The interface normal is defined from phase field as

$$\boldsymbol{n} = \frac{\nabla \eta}{|\nabla \eta|} \,. \tag{3.21}$$

First, a classical model with anisotropy only in κ will be developed. The equations (3.15) and (3.20) are combined

$$\sigma(\theta) = \sigma_0 h(\theta) = \frac{\sqrt{2}}{6} \sqrt{m\kappa(\theta)}$$
 (3.22)

and allow to express the anisotropic $\kappa(\theta)$

$$\kappa(\theta) = \kappa_0 [h(\theta)]^2 \quad \text{and} \quad \sigma_0 = (\sqrt{2}/6) \sqrt{m\kappa_0},$$
(3.23)

which recovers the interface energy anisotropy (3.20).

Recall that $\theta = \theta(\mathbf{n}(\nabla \eta))$, hence the θ dependence in (any) model parameter implies that in the governing equations new terms arise from the divergence part of the functional derivative $\nabla \cdot \partial f/\partial (\nabla \eta_p)$, where the free energy density is taken derivative of with respect to the gradient components.

This anisotropy implementation implies that the interface width l_{int} varies proportionally to the interface energy (see (3.11)) and this in turn introduces nonphysical anisotropy in the kinetic coefficient L (see (3.19)). The latter is should be compensated, the first must be taken into account im practical simulations so that there are no excessively large or narrow interfaces in the domain.

In strongly anisotropic systems, the variable interface width poses a practical problem. However, it can be prevented by including the anisotropy also in the parameter m. When the anisotropy is assigned in the natural model formulation as

$$\kappa(\theta) = \kappa_0 h(\theta) \tag{3.24}$$

$$m(\theta) = m_0 h(\theta) \,, \tag{3.25}$$

the interface width is constant $l_{int}=2\sqrt{2}\sqrt{\frac{\kappa_0}{m_0}}$ (see (3.11)) and the following relation holds as well

$$\sigma_0 h(\theta) = \frac{\sqrt{2}}{6} \sqrt{m(\theta)\kappa(\theta)}. \tag{3.26}$$

Because both κ , m depend on θ , even more terms derive from the divergence in the functional derivative, as described earlier.

Another notable approach uses Finsler geometry to replace the Euclidean metric in the simulation domain by an anisotropic one [7, 8]. Formally, the anisotropic Allen-Cahn equation is then not much more complicated then the isotropic one, even though new terms arise as well. The interface width is constant, though.

3.3.2 Challenges and limitations

The main challenges come from the type of anisotropy function and from the strength of anisotropy. So far, it was implicitly assumed that the anisotropy function is continuous and twice-differentiable homogeneous function of order 1. In chapter 2 the problem arising from forbidden angles was worked in some detail, but here it will be put into a context specific for phase field method.

Generally, two types of singularities may occur on the Wulff shapes of the first order (for the terminology see section 2.1): corners and facets. These two depend on the following properties of the particular anisotropy function

- 1. (non-)convexity of the Frank's plot, i.e. $1/h(\theta)$ and
- 2. (non-)differentiability with respect to the polar angle (typically presence of a cusp).

The first produces corners (when non-convex) and the second facets (when non-differentiable).

The easiest solution to the non-differentiability is to replace the anisotropy function by such one, in which the cusps are a smooth tips. These functions are typically formulated in such a way, that the sharpness of the tip is scalable, but formally the function is always smooth. This allows to approximate the former anisotropy close enough while assuring validity of the formulas relying on the differentiability. This was employed for example in [16, 31, 40].

The non-convexity of the Frank's plot does not have so easy solution. It causes the governing equations to be ill-posed, which means that the solution to is either not existing, unique or does not depend continuously on the initial/boundary conditions. In this case the uniqueness is the problem, because it is not clear, whether a certain segment on the surface should align with the "ear" orientation or some on the Wulff plot of the first order. In practice, it causes corrugations of the phase field contour near the corners, see e.g. [63]. The the ill-posedness can be removed by regularization (and thus the match in corners improved), the following approaches are known to the author

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- 1. regularization of the anisotropy function [17, 28, 59],
- 2. inclusion of higher order terms in expansion of interface energy in the free energy functional [61, 63, 65],
- 3. models incorporating the Finsler geometry do not need regularization [7, 8],
- 4. one study reports numerical regularization which occurred after using spectral solver [62].

All of the above approaches were successful in delivering cornered or faceted Wulff shapes of the first order and anisotropic crystals grown in either temperature or concentration field. These were the main applications of these models so far. Very likely, each of these will have its own limitations in the future applications, but currently the author is unaware of any systematic comparison of these in the literature and himself only implemented the first approach.

3.4 Allen-Cahn equations for multi-phase systems

3.4.1 Historical context

Multi-phase field models emerged from the need to simulate systems with more than two phases. At first, these models were developed to simulate ideal grain growth in systems with isotropic grain boundary energy. That is a natural choice, because the grain coarsening is curvature driven and the Allen-Cahn equation approximates the mean curvature flow. The main characteristic of the multi-phase field models is that every grain or phase is characterized by it own field variable, which implies that these systems may contain many governing equations, even tens of thousands in large scale grain growth simulations [69].

Even though there were published many model formulations, all those known to the author can be attributed to one of two main branches of development: one initiated by the work of Long-Qing Chen in 1994 [13] and the other by Steinbach in 1996 [58]. The most fundamental difference is that in Steinbach's branch, the field variables are interpreted as local phase fractions, whereas in the Chen's there is no direct physical interpretation of the field variables. For this reason, it is not accurate to call the latter multi-phase-field models (MPF), but the difference is rather technical and not obvious to broader scientific community. When there is the need to differentiate the two groups of models, the multi-continuum-field (CF) or multi-order-parameter model are the options to call the models in Chen's legacy.

Both multi-phase-field and multi-continuum-field models were extensively developed beyond the ideal grain growth, and nowadays the two model families are comparable in terms of maturity and applicability. The models offer anisotropy in interface energy and kinetic coefficient (pair-wise isotropic as well as the inclination dependence), coupling to multi-component systems using the Onsager's theory, possibly employing the thermodynamic databases to express the concentration dependence of the free energy density.

Real interfaces (including the grain boundaries) exhibit anisotropic (inclination-dependent) interface energies [9, 49], hence a significant effort was made to introduce this feature in some multi-phase field [19, 55, 62, 64] and continuum-field models [25, 42].

Toth provided structured and rational fundamental criteria to assess thermodynamic consistency of a multi-phase field model [62], most of which are applicable to the multi-continuum-field models as well. They include e.g. the requirement that the

- results should not depend on the fields numbering
- model should be scalable in the number of simulated fields
- model should allow assignment of binary interface properties (interface energy and kinetic coefficient) independently to every possible pair of fields
- total system energy decreases monotonically (second law of thermodynamics)
- ghost phases do not appear

and some others. Both model families have models which fulfill most of the criteria or even all of them.

3.4.2 Challenges and limitations

One of pragmatic requirements on these models not listed in [62] is that they must correctly reproduce force balance in triple and higher-order junctions. Despite high maturity of the models and great relevance of triple junctions for the grain coarsening, only very recently there was a structured attempt to quantify the performance of different models in how they reproduced the triple junction behavior in well defined scenarios [15].

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This study emphasized, that probably the largest challenge for both CF and MPF models are the ghost phases. These manifest as a finite amount of a third phase within a binary interface between two other fields. Ghost phases are unphysical and their amount was correlated with the (poorer) quality of match of dihedral angles in the triple junctions benchmark problems [15].

The ghost phases mitigation strategies include in the first place reformulation of either homogeneous or gradient energy contributions to avoid problematic driving force terms emerging from the functional derivative, then additions of ternary terms, or even neglecting the problematic driving force terms. In continuum-field models, the ghost phases were eliminated by specific way the model parameters were made anisotropic, but at the cost of having variable interface width.

The community has just acquired for itself the tools to quantify the models performance in triple junctions of interfaces with isotropic interface energies [15]. Certainly, a challenge ahead are the triple junctions of interfaces with inclination-dependent interface energy. First, a benchmark problem must be developed a and then used to compare the models. However, before such study, no simulation of grain growth with inclination dependent interfaces including the triple junctions is reliable. In the multi-phase field models including the inclination dependent interface energy there are many more terms in the governing equations, which makes the ghost phases prevention by model reformulation significantly more complicated than in the pair-wise isotropic case.

3.5 Benchmarking of phase field models

The community produces many applications of the phase field method, often involving complicated multi-physics, to deliver solutions to complex engineering and scientific problems. However, the know-how related to the individual modeling approaches seems to be rather fragmented into the individual groups and works comparing different models are rather exceptional. Different models for the same purpose may have different formulations, but the differences in their behavior are not obvious, especially because there is no unified, generally accepted validation procedure for quantitative assessment of performance in terms of represented physics.

Another source of confusion about the limits of phase field method in general is that some works claim to bring "quantitative" or "thermodynamically consistent" phase field models of complicated multiphysics problem (e.g. electrodeposition of lithium [12]) after doing some lower-dimensional check on a not coupled problem. This way of presentation makes the impression that everything is

possible, when in fact the more equations are coupled, the easier it is to hit reasonable limits of applicability for the model.

Where exactly lay the limits of the current models is also partially obscured by the fact, that it is not a norm to publish the codes used to produce published results, which significantly affects their reproducibility.

Systematic and reproducible parametric studies in well-defined problems are needed for quantitative assessment of phase field models reliability. The lack of benchmarking is present in terms of physics and also of models comparison. It is critical especially in models coupling many physical driving forces. In order to confidently and honestly interpret results of quantitative phase field simulations, both physical models and numerical implementations must be validated and verified [24]. Recent initiative PFHub addresses the need for benchmarking of the multitude of software and numeric approaches for solving the phase field governing equations.

Usually, the validation of models with inclination-dependent interface energy was carried out by visual comparison of the phase field contours to corresponding Wulff shapes for single or several values of strength of anisotropy [17–19, 32, 63]. Such approach does not reveal the limits of reliability of these models, though.

The cited study comparing the MPF and MCF models in the triple junctions benchmarks [15] was initiated by the author and he believes that using the presented benchmarks for assessment of the new models or implementatinos performance adds more reliability whenever the tested multi-phase field models will be used.

Chapter 4

Extended multi-phase field model

The multi-phase field model by Moelans [42] is an established [41] quantitative phase field model of grain growth with anisotropic grain boundary properties. Using asymptotic analysis, Moelans derived that in the original model [42], the local interface energy and width are related to three model parameters. Because these are two equations of three variables, the system is undetermined and one of the model parameters is free. This degree of freedom introduces the possibility of many different parameters assignment strategies, all of which represent the same physical input. The effect of different but equivalent parameters choices can thus be investigated in this model. Moelans proposed such parameters assignment strategy [42], which assured constant interface width irrespective of the strength of anisotropy in interface energy (a kind of *natural* formulation). Two of the three model parameters (γ and κ) were made anisotropic in order to achieve such behavior. However, this approach (here denoted IWc - Interface Width constant) does not reproduce well the angles between interfaces in triple junctions for stronger anisotropies. This was already first noted by Moelans in [43]. An alternative parameters assignment strategy with only parameter γ anisotropic was used in [41, 54], but no systematic comparison was made. In this approach, the interface width is not constant in anisotropic systems, but it is not simply classifiable as *classical* anisotropy formulation, because the gradient energy coefficient is constant. It will be denoted IWvG (Interface Width variable and Gamma anisotropic). The third compared parameters assignment strategy is the classical formulation, varying only gradient energy coefficient κ to achieve the desired interface energy anisotropy (denoted IWvK - Interface

Width variable and Kappa anisotropic). Additionally, the inclination dependence of interface energy in IWvG and IWvK have not yet been addressed in the framework of Moelans' model.

4.1 Introduction

The system consists of n non-conserved continuous-field variables (further denoted phase fields) $\eta_1(\mathbf{r},t), \eta_2(\mathbf{r},t), \dots, \eta_n(\mathbf{r},t)$, which are functions of space and time. The total free energy of the system is expressed as functional of the phase fields and their gradients $\nabla \eta_1(\mathbf{r},t), \nabla \eta_2(\mathbf{r},t), \dots, \nabla \eta_n(\mathbf{r},t)$

$$F = \int_{V} \left\{ m f_0(\vec{\eta}) + \frac{\kappa}{2} \sum_{i=1}^{n} (\nabla \eta_i)^2 \right\} dV, \qquad (4.1)$$

where the homogeneous free energy density $f_0(\vec{\eta}) = f_0(\eta_1, \eta_2, \dots, \eta_n)$ is expressed as

$$f_0(\vec{\eta}) = \sum_{i=1}^n \left(\frac{\eta_i^4}{4} - \frac{\eta_i^2}{2} \right) + \gamma \sum_{i=1}^n \sum_{i>j} \eta_i^2 \eta_j^2 + \frac{1}{4}.$$
 (4.2)

The parameters m, κ, γ are model parameters, which together define interface energy and interface width (see the following section for more details).

The governing equations for each phase field η_p are obtained based on the functional derivative of the free energy functional with respect to η_p , assuming that the phase-fields are non-conserved, i.e.

$$\frac{\partial \eta_p}{\partial t} = -L \frac{\delta F}{\delta \eta_p} = -L \left[\frac{\partial f}{\partial \eta_p} - \nabla \cdot \frac{\partial f}{\partial (\nabla \eta_p)} \right], \tag{4.3}$$

where L is the kinetic coefficient (also dependent on the model parameters), f is the full integrand in (4.1) and $\nabla \cdot \partial f/\partial(\nabla \eta_p)$ is divergence of vector field $\partial f/\partial(\nabla \eta_p)$ defined by relation

$$\frac{\partial f}{\partial (\nabla \eta_p)} = \frac{\partial f}{\partial (\partial_x \eta_p)} \mathbf{n}_x + \frac{\partial f}{\partial (\partial_y \eta_p)} \mathbf{n}_y + \frac{\partial f}{\partial (\partial_z \eta_p)} \mathbf{n}_z \tag{4.4}$$

with $\partial_x, \partial_y, \partial_z$ being operators for unidirectional derivatives in the corresponding directions and $\mathbf{n}_x, \mathbf{n}_y, \mathbf{n}_z$ coordinate base vectors.

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4.2 Isotropic model

In a system with uniform grain boundary properties, the interface energy is equal for all interfaces and hence the phase-field model parameters m, κ, γ (and interface width l) are constant in the system.

Then, using the expression (4.3), the governing equation for each phase field η_p takes the following form

$$\frac{\partial \eta_p}{\partial t} = -L \left[m \left(\eta_p^3 - \eta_p + 2\gamma \eta_p \sum_{j \neq p} \eta_j^2 \right) - \kappa \nabla^2 \eta_p \right]$$
 (4.5)

The interface energy σ of the system is related to the model parameters via

$$\sigma = g(\gamma)\sqrt{m\kappa}\,,\tag{4.6}$$

where $g(\gamma)$ is a non-analytic function of parameter γ . The interface width l is expressed as

$$l = \sqrt{\frac{\kappa}{m f_{0c}(\gamma)}}, \qquad (4.7)$$

where $f_{0c}(\gamma)$ is the value of $f_0(\eta_{i,cross}, \eta_{j,cross})$ in the points where the two phase fields η_i, η_j cross. $f_{0c}(\gamma)$ is a non-analytic function too. Values of both $g(\gamma)$ and $f_{0c}(\gamma)$ were tabulated and are available in [54]. Both functions are positive and monotonously rising.

Usually, the interface energy σ is known as material property and l is chosen for computational convenience, together with γ . Then, the parameter values are assigned from the following formulae

$$\kappa = \sigma l \frac{\sqrt{f_{0c}(\gamma)}}{g(\gamma)} \approx \frac{3}{4} \sigma l \tag{4.8}$$

$$m = \frac{\sigma}{l} \frac{1}{g(\gamma)\sqrt{f_{0c}(\gamma)}} \approx 6\frac{\sigma}{l}$$
 (4.9)

$$L = \frac{\mu}{l} \frac{g(\gamma)}{\sqrt{f_{0c}(\gamma)}} \approx \frac{4}{3} \frac{\mu}{l}$$
 (4.10)

The symbol μ stands for interface mobility. The approximate relations above hold exactly when $\gamma = 1.5$ and are well applicable when $0.9 \le \gamma \le 2.65$ [42].

4.3 Anisotropic model and parameters assignment strategies

Two cases of interface energy anisotropy may occur, together or separately. Firstly, in the system there may be multiple interfaces with different interface energies (termed $misorientation\ dependence$ in [42], here $pair\ wise\ isotropy$ for greater generality). Secondly, there may be an interface with inclination-dependent interface energy. Additionally, the kinetic coefficient L can be inclination dependent.

In both cases of anisotropy in interface energy, some of the model parameters m, κ, γ must become spatially dependent in order to assure correct local representation of the interface energy and width. In other words, the equations (4.6) and (4.7) must hold in every point of the anisotropic system. These two equations locally form an undetermined system of three variables, hence one of the model parameters is free and many different parameters assignment strategies are possible.

In this paper, the parameter m is always a constant, because when m was spatially varied [42], the model behavior in multijunctions was reported to be strongly affected by the interface width. Such model would be non-quantitative and thus will not be further regarded in this paper.

Three different parameters assignment strategies are considered, which differ in value of m and further in which of parameters κ, γ is constant and which varies in space to keep equation (4.6) valid. The three strategies are denoted: IWc (variable γ, κ so that interface width is constant [42]), IWvG (variable interface width and γ [54]) and IWvK (variable interface width and κ). Table 4.1 summarizes, which parameters are kept constant and which vary to capture the anisotropy in the different strategies. The detailed procedure of the parameters assignment and ways to control the width of the narrowest interface are described in Appendix B. Note that for the IWc model we propose a single-step parameters determination procedure, which is more predictable and simpler than the original iterative one [42]. The two are equivalent, though.

Below follow details about the incorporation of pair-wise isotropy and inclination-dependence in the model.

4.3.1 Systems with pair-wise isotropic IE

In the system with n phase fields, there are n(n-1)/2 possible pair-wise interfaces, each of which may have different (mean) interface energy $\sigma_{i,j}$. The indices i, j denote interface between phase fields η_i, η_j . A set of parameters

Table 4.1: Characterization of the three parameter assignment strategies: the one with constant interface width (IWc), with variable interface width and all anisotropy in γ (IWvG) and with variable interface width and all anisotropy in κ (IWvK). In the latter, it is inconvenient to choose other value of γ than $\gamma=1.5$. IW stands for interface width, other symbols have meaning as in the text.

 $m, \kappa_{i,j}, \gamma_{i,j}, L_{i,j}$ (all scalars) is obtained by appropriate procedure (depending on the strategy, see Appendix B) so that the relations (4.6) and (4.7) are valid for each interface independently (equations (4.8)-(4.10) hold for each interface (i-j)). Then, these are combined together to produce the model parameter fields $\kappa(\mathbf{r}), \gamma(\mathbf{r}), L(\mathbf{r})$:

$$\kappa(\mathbf{r}) = \frac{\sum_{i=1}^{n} \sum_{j>i}^{n} \kappa_{i,j} \eta_i^2 \eta_j^2}{\sum_{i=1}^{n} \sum_{j>i}^{n} \eta_i^2 \eta_j^2}$$
(4.11)

$$\gamma(\mathbf{r}) = \frac{\sum_{i=1}^{n} \sum_{j>i}^{n} \gamma_{i,j} \eta_i^2 \eta_j^2}{\sum_{i=1}^{n} \sum_{j>i}^{n} \eta_i^2 \eta_j^2},$$
(4.12)

$$L(\mathbf{r}) = \frac{\sum_{i=1}^{n} \sum_{j>i}^{n} L_{i,j} \eta_i^2 \eta_j^2}{\sum_{i=1}^{n} \sum_{j>i}^{n} \eta_i^2 \eta_j^2}$$
(4.13)

which stand in place of κ, γ in the functional in equation (4.1) and in place of L in the governing equation (4.3).

Notice that from Table 4.1 stems that in IWvG is $\kappa(\mathbf{r}) = \text{const}$ by definition (i.e. all the $\kappa_{i,j}$ s are equal) and similarly in IWvK $\gamma(\mathbf{r}) = \text{const} = 1.5$ (i.e. all $\gamma_{i,j}$ s are equal).

The free energy functional for IWc is then (those for IWvG and IWvK are equal, only with either $\kappa(\mathbf{r})$ or $\gamma(\mathbf{r})$ being constants, respectively):

$$F = \int_{V} \left\{ m f_0(\vec{\eta}) + \frac{\kappa(\mathbf{r})}{2} \sum_{i=1}^{n} (\nabla \eta_i)^2 \right\} dV, \qquad (4.14)$$

$$f_0(\vec{\eta}) = \sum_{i=1}^n \left(\frac{\eta_i^4}{4} - \frac{\eta_i^2}{2} \right) + \sum_{i=1}^n \sum_{i>j} \gamma_{i,j} \eta_i^2 \eta_j^2 + \frac{1}{4}.$$
 (4.15)

Both parameter fields $\kappa(\mathbf{r}), \gamma(\mathbf{r})$ are functions of phase fields $\vec{\eta}$. This dependence should produce new terms in the governing equations (from $\partial f/\partial \eta_p$ in

equation (4.3)). However, because the denominator of $\gamma(\mathbf{r})$ cancels out in the functional, the new terms only arise from $\partial \kappa / \partial \eta_p$. The governing equations then are

$$\frac{\partial \eta_p}{\partial t} = -L(\mathbf{r}) \left[m \left(\eta_p^3 - \eta_p + 2\eta_p \sum_{j \neq p} \gamma_{p,j} \eta_j^2 \right) + \frac{1}{2} \frac{\partial \kappa}{\partial \eta_p} \sum_{i=1}^n (\nabla \eta_i)^2 - \kappa(\mathbf{r}) \nabla^2 \eta_p \right]$$
(4.16)

The above procedure is fully variational, nevertheless inclusion of the term proportional to $\partial \kappa/\partial \eta_p$ enables the model to reduce the total energy of the system by introduction of so called third phase contributions (also ghost or spurious phases) at diffuse interfaces. [42] That is a common problem in multiphase field models [62], where a third phase field attains non-zero value within an interface of two other phase fields. This mathematical artefact affects triple junction angles and in general is not physically justified. Several ways of elimination or suppression of ghost phases were described in [62] and the references therein.

In this work, the ghost phases were eliminated by neglecting the term proportional to $\partial \kappa/\partial \eta_p$. However, because such model is not fully variational, the thermodynamic consistency can no longer be guaranteed in IWc and IWvK. This does not affect IWvG, because there is $\partial \kappa/\partial \eta_p = 0$ anyway. That accounts for a clear advantage of the IWvG model, as no ghost phases appear even when fully variational.

4.3.2 Systems with inclination-dependent interface energy

The orientation of an interface in 2D system is given by interface normal, inclined under the angle θ . Local value of interface energy may be a function of local interface inclination, i.e. $\sigma = \sigma(\theta)$. In Moelans' model [42], the normal at interface between η_i, η_j , denoted $\hat{n}_{i,j}$, is defined as

$$\hat{\boldsymbol{n}}_{i,j} = \frac{\nabla \eta_i - \nabla \eta_j}{|\nabla \eta_i - \nabla \eta_j|} = \begin{bmatrix} (\hat{n}_{i,j})_x \\ (\hat{n}_{i,j})_y \end{bmatrix}$$
(4.17)

and the definite inclination of that normal

$$\theta_{i,j} = \text{atan2}[(\hat{n}_{i,j})_y, (\hat{n}_{i,j})_x],$$
(4.18)

 $\sigma_{i,j}(\theta_{i,j}) = \sigma_{i,j}^0 h_{i,j}(\theta_{i,j})$. Symbols $\kappa_{i,j}^0$, $\gamma_{i,j}^0$ stand for scalar values of the parameters determined from $\sigma_{i,j}^0$ (see B in [39]). Expressions for $\gamma_{i,j}(\theta_{i,j})$ follow the so called weak anisotropy approximation [42], i.e. they assume that the values of Table 4.2: Inclination dependence of the variable parameters in the respective models. The interface energy is $\gamma_{i,j}(\theta_{i,j})$ do not diverge far from 1.5, so that the approximation $g^2[\gamma_{i,j}(\theta_{i,j})] \approx 16[2\gamma_{i,j}(\theta_{i,j}) - 1]/9[2\gamma_{i,j}(\theta_{i,j}) + 1]$ is applicable (see [42] for details). Second row contains expressions used in equations 4.23 and 4.24.

IWvK	$\kappa_{i,j}(\theta_{i,j}) = \kappa_{i,j}^0 [h_{i,j}(\theta_{i,j})]^2$	$\partial \kappa_{i,j}/\partial h_{i,j} = 2\kappa_{i,j}^0 h_{i,j}(\theta_{i,j})$
IWvG	$\gamma_{i,j}(\theta_{i,j}) = -\frac{\frac{9}{4}[g(\gamma_{i,j}^0)h_{i,j}(\theta_{i,j})]^2 + 1}{\frac{9}{2}[g(\gamma_{i,j}^0)h_{i,j}(\theta_{i,j})]^2 - 2}$	$\frac{\partial \gamma_{i,j}/\partial h_{i,j} =}{18g^2(\gamma_{i,j}^0)h_{i,j}(\theta_{i,j})} \\ \frac{1}{2} [g(\gamma_{i,j}^0)h_{i,j}(\theta_{i,j})]^2 - 2\}^2$
IWc	$\kappa_{i,j}(\theta_{i,j}) = \kappa_{i,j}^{0} h_{i,j}(\theta_{i,j}) $ $\gamma_{i,j}(\theta_{i,j}) = -\frac{1}{2} \frac{1}{2} \sigma^{2}(\gamma_{i,j}) h_{i,j}(\theta_{i,j}) + 1}{\frac{1}{2} g^{2}(\gamma_{i,j}) h_{i,j}(\theta_{i,j}) - 2} $ $\gamma_{i,j}(\theta_{i,j}) = -\frac{1}{2} \frac{1}{2} \sigma^{2}(\gamma_{i,j}) h_{i,j}(\theta_{i,j}) - 2}$	$\partial \kappa_{i,j}/\partial h_{i,j} = \kappa_{i,j}^0$ $\partial \gamma_{i,j}/\partial h_{i,j} = \frac{9g^2(\gamma_{i,j}^0)}{\left[\frac{9}{2}g^2(\gamma_{i,j}^0)h_{i,j}(\theta_{i,j}) - 2\right]^2}$
model	variable pa- rameter(s)	$rac{\partial \kappa_{i,j}/\partial h_{i,j}}{\mathrm{and}}$

which is the standard 2-argument arctangent function.

In 2D, the inclination-dependence of interface energy can be expressed as

$$\sigma_{i,j}(\theta_{i,j}) = \sigma_{i,j}^0 h_{i,j}(\theta_{i,j}) \tag{4.19}$$

where $\sigma_{i,j}^0$ is a scalar and $h_{i,j}(\theta_{i,j})$ is anisotropy function. The used anisotropy function was

$$h_{i,j}(\theta_{i,j}) = 1 + \delta \cos(n\theta_{i,j}), \qquad (4.20)$$

with δ being strength of anisotropy and n the order of symmetry. Some properties of this anisotropy function and the resulting Wulff shapes are given in 2.5.

The inclination dependence of $\sigma_{i,j}$ implies that some of the model parameters $\gamma_{i,j}$, $\kappa_{i,j}$ must be taken inclination-dependent too. Depending on the model used (IWc, IWvG or IWvK), the local validity of equation (4.6) is achieved using different inclination dependence of the variable parameters (see Table 4.2 for details).

Because the inclination-dependent $\kappa_{i,j}, \gamma_{i,j}$ are functions of components of gradients $\nabla \eta_i, \nabla \eta_j$, the divergence term in the functional derivative (equation (4.3)) produces additional driving force terms. In the general case with multiple inclination-dependent interfaces, the divergence term equals

$$\nabla \cdot \frac{\partial f}{\partial (\nabla \eta_p)} = 2m\eta_p \nabla \eta_p \cdot \left[\sum_{j \neq p} \eta_j^2 \frac{\partial \gamma_{p,j}}{\partial (\nabla \eta_p)} \right]$$

$$+ 2m\eta_p^2 \sum_{j \neq p} \left[\eta_j \nabla \eta_j \cdot \frac{\partial \gamma_{p,j}}{\partial (\nabla \eta_p)} \right]$$

$$+ m\eta_p^2 \sum_{j \neq p} \eta_j^2 \left[\nabla \cdot \frac{\partial \gamma_{p,j}}{\partial (\nabla \eta_p)} \right]$$

$$+ \frac{1}{2} \left[\nabla \cdot \frac{\partial \kappa}{\partial (\nabla \eta_p)} \right] \sum_{i=1}^n (\nabla \eta_i)^2$$

$$+ \frac{1}{2} \frac{\partial \kappa}{\partial (\nabla \eta_p)} \cdot \left[\nabla \sum_{i=1}^n (\nabla \eta_i)^2 \right]$$

$$+ \nabla \kappa(\mathbf{r}) \cdot \nabla \eta_p + \kappa(\mathbf{r}) \nabla^2 \eta_p .$$

$$(4.21)$$

The vector field $\partial \kappa / \partial (\nabla \eta_p)$ is

$$\frac{\partial \kappa}{\partial (\nabla \eta_p)} = \frac{\sum_{j \neq p}^n \left(\frac{\partial \kappa_{p,j}}{\partial (\nabla \eta_p)}\right) \eta_p^2 \eta_j^2}{\sum_{k=1}^n \sum_{l > k} \eta_k^2 \eta_l^2}$$
(4.22)

where the sum in the numerator goes through all pair-wise interfaces of $\eta_p(\mathbf{r})$. The vector fields $\partial \kappa_{p,j}/\partial (\nabla \eta_p)$ are expressed

$$\frac{\partial \kappa_{p,j}}{\partial (\nabla \eta_p)} = \frac{1}{|\nabla \eta_i - \nabla \eta_j|} \frac{\partial \kappa_{p,j}}{\partial h_{p,j}} \frac{\partial h_{p,j}}{\partial \theta_{p,j}} \begin{bmatrix} -(\hat{n}_{i,j})_y \\ (\hat{n}_{i,j})_x \end{bmatrix}. \tag{4.23}$$

Note, that the above vector field is nonzero only in IWc and IWvK models at the interfaces (p-j) with inclination-dependent IE. Likewise, the below vector field is nonzero only in IWc and IWvG

$$\frac{\partial \gamma_{p,j}}{\partial (\nabla \eta_p)} = \frac{1}{|\nabla \eta_i - \nabla \eta_j|} \frac{\partial \gamma_{p,j}}{\partial h_{p,j}} \frac{\partial h_{p,j}}{\partial \theta_{p,j}} \begin{bmatrix} -(\hat{n}_{i,j})_y \\ (\hat{n}_{i,j})_x \end{bmatrix}. \tag{4.24}$$

The multipliers $\partial \kappa_{p,j}/\partial h_{p,j}$ and $\partial \gamma_{p,j}/\partial h_{p,j}$ differ in individual models and are also provided in Table 4.2. The term $\partial h_{p,j}/\partial \theta_{p,j}$ is defined by the inclination-dependence at the interface (p-j).

The governing equation then is

$$\frac{\partial \eta_p}{\partial t} = -L(\mathbf{r}) \left[m \left(\eta_p^3 - \eta_p + 2\eta_p \sum_{j \neq p} \gamma_{p,j} (\theta_{p,j}) \eta_j^2 \right) -\nabla \cdot \frac{\partial f}{\partial (\nabla \eta_p)} \right].$$
(4.25)

Note, that the term proportional to $\partial \kappa / \partial \eta_p$ was neglected here.

In models with variable interface width (IWvG, IWvK), at the interfaces with inclination-dependent interface energy, the interface width is a function of the inclination, i.e. $l_{i,j} = l_{i,j}(\theta_{i,j})$. Because the kinetic coefficient $L_{i,j}$ is inversely proportional to the interface width $l_{i,j}$ (see equation (4.10)), the kinetic coefficient is inclination-dependent as well (even for *constant* grain boundary mobility $\mu_{i,j}$). The inclination dependence of $L_{i,j}(\theta_{i,j})$ due to interface width variation is in the IWvG model

$$L_{i,j}(\theta_{i,j}) = L_{i,j}h_{i,j}(\theta_{i,j}) \tag{4.26}$$

and in the IWvK model

$$L_{i,j}(\theta_{i,j}) = L_{i,j}/h_{i,j}(\theta_{i,j}),$$
 (4.27)

where $h_{i,j}(\theta_{i,j})$ is the anisotropy function in interface energy (4.19).

The equations (4.26),(4.27) were derived from an alternative expression for the kinetic coefficient $L_{i,j}$

$$L_{i,j} = \frac{\mu_{i,j}\sigma_{i,j}(\theta_{i,j})}{\kappa_{i,j}(\theta_{i,j})}, \qquad (4.28)$$

where the inclination dependencies of the right-hand side were expressed correspondingly to the model (see Table 4.2 for $\kappa_{i,j}(\theta_{i,j})$).

Due to varying number of driving force terms in the three parameter assignment strategies, the governing equations are different in each and hence it is justified to call them different models.

Systems with inclination-dependent mobility

Let the interface (i-j) have isotropic interface energy and inclination-dependent grain boundary mobility with anisotropy function $h_{i,j}^{\mu}(\theta_{i,j})$, i.e. $\mu_{i,j} = \mu_{i,j}(\theta_{i,j}) = \mu_{i,j}^0 h_{i,j}^{\mu}(\theta_{i,j})$. From equation (4.10) we can see that the kinetic coefficient must have the same anisotropy, i.e. $L_{i,j}(\theta_{i,j}) = L_{i,j}^0 h_{i,j}^{\mu}(\theta_{i,j})$, where $L_{i,j}^0 = \mu_{i,j}^0 g(\gamma_{i,j})/l_{i,j} f_{0c}(\gamma_{i,j})$.

If the interface energy is inclination-dependent as well and a model with variable interface width is used (either IWvG or IWvK), the inclination dependence in $L_{i,j}(\theta_{i,j})$ due to the interface width variation must be included similarly like in (4.26) and (4.27). The physical inclination-dependence is independent from the one due to interface width variation, implying the following expression for IWvG model

$$L_{i,j}(\theta_{i,j}) = L_{i,j}^0 h_{i,j}(\theta_{i,j}) h_{i,j}^{\mu}(\theta_{i,j})$$
(4.29)

and for the IWvK model analogically

$$L_{i,j}(\theta_{i,j}) = L_{i,j}^0 \frac{h_{i,j}^{\mu}(\theta_{i,j})}{h_{i,j}(\theta_{i,j})},$$
(4.30)

where $h_{i,j}(\theta_{i,j})$ is the interface energy anisotropy function.

4.3.3 Interface profiles in different models

The main difference in the model modifications is how the interface width varies as function of local interface energy. Obviously, in IWc the width is constant.

In IWvK (with $\gamma = 1.5$) the width of interface *i-j* can be computed as

$$l_{i,j} = 6\frac{\sigma_{i,j}}{m} \,. \tag{4.31}$$

Apparently, in IWvK model the interface width is proportional to the interface energy, i.e. the larger the interface energy, the larger the interface width.

In IWvG the width can be expressed from (4.7) and (4.6) assuming $\frac{4}{3}\sqrt{f_{0,c}}(\gamma_{i,j}) = g(\gamma_{i,j})$ (which holds for small values of $\gamma_{i,j}$). Then, it goes approximately

$$l_{i,j} \approx \frac{\kappa}{\sigma_{i,j}}$$
, (4.32)

and apparently the larger interface energies are associated with lower interface widths in IWvG.

4.4 General Neumann boundary conditions to control interface inclination at domain boundary

This model extension was inspired by [20]. It allows to control the interface inclination angle ϕ at the boundary, which thus becomes an input parameter. The principle is explained using a single phase field $\eta(\mathbf{r})$. First, we remark that the Neumann BCs $\nabla \eta \cdot \mathbf{n}_D = 0$ must be a special case of the general BC (\mathbf{n}_D is the domain boundary normal). The Neumann BC implies perpendicularity of the PF gradient and domain boundary normal. In other words, the interface is perpendicular to the domain boundary. Generally, we can write $\mathbf{n}_D \cdot \nabla \eta = |\nabla \eta| \cos(\phi)$. In the model by Granasy [20], the local magnitude of phase field gradient may be expressed using the local phase field value, giving rise to expression of the boundary condition as

$$\boldsymbol{n_D} \cdot \nabla \eta = \frac{\cos(\phi)}{\delta \sqrt{2}} \eta (1 - \eta),$$
 (4.33)

where δ is the interface width. At the domain boundary, the gradient $\nabla \eta$ is expressed as a polynomial, which allows straightforward implementation of the condition, especially in rectangular simulation domains.

In Moelan's model the interface normal between the phase fields η_i , η_j denoted $n_{i,j}$ is defined as local difference in neighboring PF gradients (see equation 4.17). In the spirit of the introductory example we can write

$$\boldsymbol{n_D} \cdot \boldsymbol{n}_{i,j} = \cos(\phi) \implies \boldsymbol{n} \cdot (\nabla \eta_i - \nabla \eta_j) = |\nabla \eta_i - \nabla \eta_j| \cos(\phi). \quad (4.34)$$

Now we have single BC with the correct physical interpretation (fixed inclination angle at the boundary) but because the governing equations are solved for the phase fields, a set of equivalent boundary conditions (one for each phase field) must be derivable from the above. In these independent BC, it must be possible to express the gradient magnitude (the introductory example finds a polynomial of local PF values equal to the gradient magnitude).

The first problem (coupling in the BC) would be solved if one gradient could be written as a function of the other and vice versa. However, this dependence is non-analytical for all $\gamma_{i,j} \neq 1.5$. For $\gamma_{i,j} = 1.5$ it is simply $\eta_i = 1 - \eta_j$ and therefore also $\nabla \eta_i = -\nabla \eta_j$. [42]

The second problem (the expression for gradient magnitude) is also enabled by the choice $\gamma = 1.5$, as in such case there were derived analytic expressions for the gradients $\nabla \eta_i$, $\nabla \eta_i$. [42] In 1D system with 2 phase fields holds

$$\frac{\mathrm{d}\eta_{i}}{\mathrm{d}x} = \sqrt{\frac{2m}{\kappa_{i,j}}} \eta_{i} (1 - \eta_{i}) \stackrel{\mathrm{2D,3D}}{=} |\nabla \eta_{i}|$$

$$\frac{\mathrm{d}\eta_{j}}{\mathrm{d}x} = \sqrt{\frac{2m}{\kappa_{i,j}}} \eta_{j} (1 - \eta_{j}) \stackrel{\mathrm{2D,3D}}{=} |\nabla \eta_{j}|.$$
(4.35)

Validity of the last equation sign in 2D and 3D (in systems with 2 phase fields) can be checked using the expressions in [42] as follows. Write the equation 7 from [42] for 2D. From it one gets 2D equivalents of equations 8a and 8b, where $\frac{d\eta}{dx}$ is replaced by $|\nabla\eta|$. Further, because the equation 19 does not depend on problem dimensionality, we obtain here the expressions in equation 4.35. In systems with more than two phase fields the general expression for the gradient magnitude is more complex. However, the same relations are satisfied at a pair-wise interface not close to the triple junction.

With the above and assuming that there is no triple junction near the domain boundary, we can thus write

$$\mathbf{n} \cdot (\nabla \eta_i - \nabla \eta_j) = 2\mathbf{n} \cdot \nabla \eta_i = 2|\nabla \eta_i|\cos(\phi)$$

$$= -2\mathbf{n} \cdot \nabla \eta_j = -2|\nabla \eta_j|\cos(\phi),$$
(4.36)

which provides very similar BC like in the introductory example for each of the two phase fields.

The only way how to use these relations is to have $\gamma_{i,j} = 1.5$. In the case of inclination-dependent interface energy $\sigma_{i,j}(\theta_{i,j}) = \sigma^0_{i,j} h^{\sigma}_{i,j}(\theta_{i,j})$, it is necessary to express the anisotropy fully by the parameter $\kappa_{i,j}(\theta_{i,j})$.

4.5 Volume conserving multi phase field models

There are several approaches which accomplish the volume conservation of some species or phases. Each has its drawbacks though. It was not clear which of the possible approaches was the most convenient for coupling with inclination-dependent interface energy in curvature-driven systems.

Three conceptual solutions can be used in a multi-phase field model, which are known to the author: Cahn-Hilliard equation, Lagrange multipliers and fictitious concentrations field.

Given the intended application with anisotropic interface energy and the complicated driving force, a less computationally demanding option than the Cahn-Hilliard equation was seeked, because that one is a partial differential equation of 4-th order in space.

The approach using Lagrange multipliers to conserve volume was thoroughly considered, but eventually it was found out that in Moelan's model is not a suitable solution. The reason is that the volume of a single phase/grain is defined using all other phase fields, which introduces excessive computational complexity in application of the principles leading to volume conservation. These findings were documented in Appendix D.

The approach using fictitious concentration field is not new in combination with Moelans' multi-phase field model (see e.g. [68, 70]), but it will be reviewed.

4.5.1 Fictitious concentration field

The idea is to couple the AC equation to conserved concentration field in such a way that the change in phase fraction is only possible with exchange of species between phases. If the concentration of the independent species is set to equilibrium value no phase transformation occurs and the volume of the phase fraction should be conserved.

In the description below, it is assumed that there are two phases, one solid and the other liquid.

The free energy functional is then

$$F_{cons} = \int_{V} \left[f_{hom}(\vec{\eta}, \nabla \vec{\eta}, c) + f_{grad}(\vec{\eta}, \nabla \vec{\eta}) \right] dV, \qquad (4.37)$$

with

$$f_{hom}(\vec{\eta}, \nabla \vec{\eta}, c) = f_0(\vec{\eta}, \nabla \vec{\eta}) + f_{chem}(\vec{\eta}, c)$$
(4.38)

$$f_{chem}(\vec{\eta}, c) = h_S(\vec{\eta}) f_S(c) + h_L(\vec{\eta}) f_L(c)$$

$$(4.39)$$

 $h_S(\vec{\eta}), h_L(\vec{\eta})$ are solid and liquid interpolation functions. Assuming that the solid phase is composed of s phase fields $\eta_{S1}, \eta_{S2}, \ldots, \eta_{Ss}$ and the liquid phase of l phase fields $\eta_{L1}, \eta_{L2}, \ldots, \eta_{Ll}$ we can write

$$h_S(\vec{\eta}) = \frac{\sum_i^s \eta_{Si}^2}{\sum_i^s \eta_{Si}^2 + \sum_j^l \eta_{Lj}^2}$$
(4.40)

$$h_L(\vec{\eta}) = \frac{\sum_j^l \eta_{Lj}^2}{\sum_i^s \eta_{Si}^2 + \sum_j^l \eta_{Lj}^2} \quad . \tag{4.41}$$

Parabolic energy approximation will be used for the concentration dependence of the homogeneous phase free energy densities $f_S(c)$, $f_L(c)$ as

$$f_S(c_S) = A(c_S - c_{S,eq})^2 (4.42)$$

$$f_L(c_L) = B(c_L - c_{L,eq})^2$$
 (4.43)

An ideal solution approximation is adopted.

The governing equations for two phase fields are then

$$\frac{\partial \eta_1}{\partial t} = -L \left[\frac{\partial f_0}{\partial \eta_1} - \nabla \cdot \frac{\partial F}{\partial (\nabla \eta_1)} + \frac{\partial h_S}{\partial \eta_1} [f_S(c_S) - f_L(c_L) - (c_S - c_L)\mu] \right]$$
(4.44)

$$\frac{\partial \eta_2}{\partial t} = -L \left[\frac{\partial f_0}{\partial \eta_2} - \nabla \cdot \frac{\partial F}{\partial (\nabla \eta_2)} + \frac{\partial h_S}{\partial \eta_2} [f_S(c_S) - f_L(c_L) - (c_S - c_L)\mu] \right]$$
(4.45)

$$\mu = 2A(c_L - c_{L,eq}) \tag{4.46}$$

$$\frac{\partial c}{\partial t} = \frac{D}{A} \Delta \mu = 2D\Delta (c_L - c_{L,eq}) \tag{4.47}$$

where the diffusion coefficient D is a constant.

The interpolation functions have the following property

$$\frac{\partial h_S}{\partial \eta_p} = -\frac{\partial h_L}{\partial \eta_p} \tag{4.48}$$

4.6 Summary and the author's contribution

The author extended the multi-phase field model by Moelans [42]. The work [42] promoted usage of the model variant with constant interface width (here denoted IWc) with not-fully-variational formulation and described it in great detail.

At the beginning of the author's work, it was already known that IWc was not reliably reproducing the angles in triple junctions of interfaces with different interface energies [44] (i.e. in pair-wise isotropic systems). The model variant with variable interface width and all anisotropy in the parameter γ (here IWvG) was used in one grain growth study [54], but its better reliability in triple junction angles had not been documented yet.

As provided below with all details, the model extensions included

- 1. Description of another model variant with all anisotropy in the parameter κ and variable interface width (IWvK)
- 2. Derivation of the governing equation for all the three model variants in both 2D and 3D for the case of inclination-dependent interface energy
- 3. The model behavior depends heavily on proper parametrization of the equations. Best practices in parameters determination were developed and described (different in each model variant) in order to assure control over both the local interface width and local interface energy.
- 4. In the IWvK model variant such boundary conditions were incorporated, which allowed to control the angle under which the interface intersecting the domain boundary will align.

Besides the listed model extensions, the author also implemented two approaches adding the volume conservation to the system based on Allen-Cahn type of equations, namely one using Lagrange multipliers to conserve the volume and pne using fictitious concentration field. Volume conservation would be necessary in simulations, where the nucleus interacted dynamically with its surroundings, but these simulations were not needed eventually. Regardless, the model modifications including volume conservation were described in Appendix.

Chapter 5

Phase filed model benchmarking

5.1 Introduction

This chapter proposes two benchmarks for quantitative assessment of a) the anisotropic curvature driving force, and b) the anisotropic curvature driving force in combination with anisotropic kinetic coefficient. Both are 2-phase systems, hence could be simulated using a single phase field model. Nevertheless, because the multi-phase field models allow simulations of multiple phases, they have wider application potential than single phase models. For this reason, the validations are demonstrated on a multi-phase field model from chapter 4. However, the nature of these benchmarks is independent on the model formulation.

In order to provide more complete comparison of the model variants, another supplementary benchmark was carried out, which determined equilibrium triple junction angles.

Throughout this chapter the term "Wulff shape" is used in place of "Wulff shape of the first order" (see section 2.1), i.e. the smallest of the possible equilibrium stable shapes.

This chapter first provides details on the numerical implementation of the model(s) in MATLAB, then the methodology is explained and the models quantitatively compared. The first two benchmarks involve quantitative matching of the shrinking shape to the analytic one and also the determination of shrinkage rate. In the third benchmark, the triple junction angles are determined

from the phase field contour using two different approaches. For all simulations the effect of interface width and number of grid points through the interface are investigated by re-running the simulations using different numerical settings.

5.2 Numerical implementation

All models were implemented in a single MATLAB function, where the governing equations were solved by centered finite differences of second order, explicit Euler time stepping and boundary conditions implemented using ghost nodes. For more details on the used finite difference schemes, refer to the Appendix C. The minimal code to run the simulations is available in the dataset [36]. During the parameters assignment in models with variable interface width (IWvG, IWvK), there was assured control over the minimal interface width, i.e. that there would be no interface narrower than the user-specified one. That is to prevent unphysical behavior of the interface due to too small grid resolution. Different strategies had to be adopted in IWvK and IWvG, respectively. They are described in Appendix B together with other best practices in parameters determination for the respective models. The MATLAB functions which were used for parameters determination were also included in the dataset [36].

In the simulations with inclination-dependent interface energy, the vector field $\partial \kappa / \partial (\nabla \eta_p)$ was computed as in equations (4.22) and (4.23), the fields $\partial \gamma_{p,j} / \partial (\nabla \eta_p)$ as in equation (4.24) and their divergences (see equation (4.21)) were computed numerically (by centered differences), as well as the gradient $\nabla \kappa(\mathbf{r})$. All the above terms were computed in the IWc model, whereas in the models with variable interface width some of them could be omitted (as explained in section 4.3.2).

When the anisotropy in inclination-dependent interface energy was strong (i.e. $\delta > 1/(n^2 - 1)$ or $\Omega > 1$), the anisotropy function had to be regularized as described in [17] in order to avoid ill-posedness of the governing equations for interfaces with missing inclination.

In finite difference method, the ratio of the time step Δt to the grid spacing Δx determines the numeric stability of the simulation. Specifically, the so-called Courant number C must be 0 < C < 0.5, where in this model

$$C = \min[\kappa(\mathbf{r})L]\Delta t/(\Delta x)^{2}. \tag{5.1}$$

In practice, the time step was computed using equation (5.1) assuming the value of $C \leq 0.3$. For stronger anisotropies C had to be reduced farther. When $0.129 \leq \sigma_{min}/\sigma_{max} \leq 0.2$, the value C = 0.12 was used, when $0.2 < \sigma_{min}/\sigma_{max} \leq 0.3$, C = 0.17 and for $\sigma_{min}/\sigma_{max} > 0.3$ the used Courant number

Table 5.1: Parameters from equation (5.2) describing anisotropic dependence of the Courant number C in simulations with inclination-dependent interface energy and strong anisotropy $\Omega > 1$.

was C=0.3. The latter value of Courant number was also applied in simulations with inclination-dependent interface energy and smooth Wulff shapes $(\Omega \leq 1)$. For cornered Wulff shapes $(\Omega > 1)$, the following dependence was obtained from visual inspection of the interface normal inclination field $\theta(\boldsymbol{r})$ defined in equation (18) $(\theta(\boldsymbol{r})$ was sensitive to the numeric instabilities) during the simulation

$$C(\Omega) = \frac{1}{a\Omega + b}. (5.2)$$

The Table 5.1 contains the values of a, b in the respective model modifications.

The following 9-point stencil was used for computation of the laplacian in grid point (k, l)

$$(\nabla^{2}\eta)_{k,l} = \frac{1}{6h^{2}} [4(\eta_{k-1,l} + \eta_{k+1,l}\eta_{k,l-1} + \eta_{k,l+1}) + \eta_{k+1,l+1} + \eta_{k-1,l-1} + \eta_{k-1,l+1} + \eta_{k+1,l-1} - 20\eta_{k,l}].$$

$$(5.3)$$

When the interface energy is inclination-dependent, it can be seen in equations (4.23), (4.24) that some driving force terms are proportional to $1/|\nabla \eta_i - \nabla \eta_j|$. In fact, some are even proportional to $1/|\nabla \eta_i - \nabla \eta_j|^3$ when worked out analytically. Far enough from the interface, the magnitude $|\nabla \eta_i - \nabla \eta_i|$ inevitably gets smaller than the numeric resolution, which causes severe numeric instabilities due to locally very large driving forces. For this reason, all the driving force terms from equation (4.21) were only computed in a region close enough to the interface. Two such regions Γ_{out} , Γ_{inn} (outer and inner) were selected based on local value of $|\nabla \eta_i - \nabla \eta_j|$ (see Figure 5.1). The region Γ_{inn} was a subset of Γ_{out} . The interface inclination field $\theta_{i,j}$ was computed in Γ_{out} , where $|\nabla \eta_i - \nabla \eta_j|^3 (\Delta x)^3 > C_{out}^3$ and the driving force terms $\nabla \cdot \partial f / \partial (\nabla \eta_p)$ were computed in region Γ_{inn} where $|\nabla \eta_i - \nabla \eta_j|^3 (\Delta x)^3 > C_{inn}^3$. The constants $C_{out} = 10^{-6}$ and $C_{inn} = 0.005$ have the meaning of minimal local change in value of phase field which is acceptable to the respective near-interface region (this is exactly true for $\gamma = 1.5$ and approximate otherwise). Note, that the driving force from homogeneous free energy density (i.e. $\partial f/\partial \eta_p$) and the

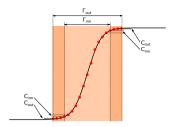


Figure 5.1: Illustration of the near-interface regions Γ_{out} , Γ_{inn} , defined by absolute values of difference in phase fields between neighbouring points.

laplacian term $\kappa(\mathbf{r})\nabla^2\eta_p$ were computed and applied in the whole simulation domain.

In the Wulff shape simulations, the accuracy of match to the analytical shape with IWvG phase field contour turned out to be sensitive to alignment of the shape within the numerical grid. For strong anisotorpies $\Omega>1$, when the corner laid exactly on some grid line, the IWvG contours were slightly deformed, making the model perform seemingly worse than IWc and IWvK (which were not affected by this). This problem was solved by shifting the center of the initial shape by half a pixel in both x and y direction. No corner than laid on any grid line and that solved the problem. Rotating the Wulff shape relative to the grid had similar effect. All the results in the main text were simulated with the shifted initial shape. This is not likely to play significant role in grain growth simulations as the corner motion along grid line is very unlikely with arbitrary grain structure and local grain orientation.

5.3 Methodology

Three different simulation experiments were performed: a shrinking Wulff shape, kinetically compensated anisotropic curvature-driven circle shrinkage and triple junction angles. The initial-state geometries and grid dimensions are in Figure 5.2. Except for the shrinking circles simulation, a parametric study was carried out in every experiment in order to validate the model behavior. Table 5.2 summarizes the variable parameters in every experiment.

Each of the experiments was simulated using all the three model modifications (IWc, IWvG and IWvK) successively and the results were quantitatively compared.

In order to distinguish the behavior of the model from the artefacts of numeric

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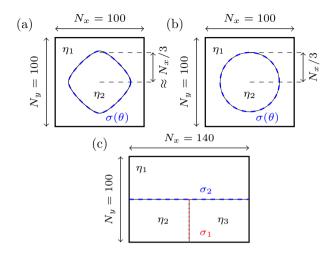


Figure 5.2: Initial conditions in the different numeric experiments with indicated interface energies. In a) Wulff shape shrinkage, b) kinetically compensated anisotropic curvature-driven circle shrinkage and c) measurement of triple junction angle. Grid dimensions correspond to the base run and 1000IW run (see text for details). All interfaces have equal mobilities.

Table 5.2: Overview of the simulations carried out in every simulation experiment. Note that these were carried out in every model modification (i.e. IWc, IWvG and IWvK) and simulation run (see Table 5.3). Number of simulations in every experiment is provided in the column Count. See text for more details.

Experiment	Varied par.	Values	Count
Wulff shape	Ω (-)	0.2, 0.4, 0.6, 0.8, 1.0,	10
		2.3, 3.6, 4.9, 6.2, 7.5	
Kin.comp.aniso. circle	Ω (-)	0.1, 0.3, 0.5, 0.7, 0.9	5
Triple junction	σ_1/σ_2 (-)	0.13, 0.2, 0.3, 0.4, 0.5,	20
		0.6, 0.7, 0.8, 0.9, 1.0,	
		1.1, 1.2, 1.3, 1.4, 1.5,	
		1.6, 1.7, 1.8, 1.9, 2.0	

Table 5.3: Numeric settings of different simulation runs (for each simulation experiment and model). l_{min} is the minimal interface width. N_x and N_y are grid dimensions in the base run as shown in Figure 5.2

	base run	IW/2	14IWpts	1000IW
grid dimensions	$N_x \times N_y$	$2N_x \times 2N_y$	$2N_x \times 2N_y$	$N_x \times N_y$
$l_{min} \text{ (nm)}$	1	0.5	1	1000
points in l_{min}	7	7	14	7

discretization, the above described series of simulations was re-run for several numeric settings. Throughout the paper, these large series are called 'runs'.

Results of four runs are presented in this paper, the basic settings of which are summarized in Table 5.3. The difference between the runs were: a) the minimal set interface width l_{min} and b) the number of points in the interface. The runs are denoted base run, IW/2, 14IWpts and 1000IW. The base run had rather coarse grid and 7 points in the interface, which should in general be reliable yet not very computationally heavy settings. The minimal interface width was $l_{min} = 1$ nm. In the run IW/2 the interface width was halved $(l_{min} = 0.5 \text{ nm})$ and the number of points in the interface was 7 again. In the 14IWpts run, the interface width was as in the base run but the points in the interface were doubled (to be 14). Physical dimensions of the simulated domain were equal in the described runs but the grid spacing was halved in IW/2 and 14IWpts runs. In the last 1000IW run, the grid was equal as in the base run, but the physical dimensions were scaled by a factor of 1000, meaning the minimal interface width $l_{min} = 1 \,\mu\text{m}$. This is the practical settings for the actual grain growth simulations. The latter run validated that the meso-scale behavior of the model is equal to that one at the nanometer scale.

In the simulation experiments, position of the interface in the respective geometry was compared to the expected shape. Depending on the parameters assignment strategy, the phase field profiles may not be symmetric about the point 0.5 in Moelans' model. For this reason, it was considered that the position of the interface i-j was in points $\mathbf{r} = (x, y)$ of the domain, where $\eta_i(\mathbf{r}) = \eta_j(\mathbf{r})$, i.e. where the two profiles crossed. This way the contours were always well defined and could be quantitatively compared to appropriate analytical models even for Wulff shape simulations, where the profile shape varies along the interface.

The below subsections describe in detail the analytic solutions of the problems and results processing in the individual simulation experiments.

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5.3.1 Used anisotropy function and its Wulff shape

The anisotropy function $h(\theta) = 1 + \delta \cos(n\theta)$ was used, together with the normalized strength of anisotropy $\Omega = \delta(n^2 - 1)$. Ω easily distinguishes weak from strong anisotropy, because for $0 < \Omega < 1$ the Wulff shape is smooth, whereas for $1 \le \Omega < n^2 - 1$ it has corners. Section 2.5 provides more details about this anisotropy function. Fourfold symmetry was assumed (i.e. n = 4).

Ill-posedness of the governing equations for forbidden orientations on Wulff shapes for strong anisotropies $(\Omega > 1)$ was treated by regularization of the anisotropy function as in [17].

For validation of the kinetics of shrinkage, an analytic expression for Wulff shape shrinkage rate was derived in 2.6. Measurement of area/volume occupied by a grain/phase is trivial in phase field method, hence the rate of its change (i.e. the shrinkage rate) can be easily used for validation or benchmark. The derivation in [39] delivers the expression

$$\frac{\mathrm{d}A_W}{\mathrm{d}t} = -2\pi\mu\sigma_0 \frac{C_W(\Omega, n)}{1 - \delta} \tag{5.4}$$

where $C_W(\Omega,n) = A_W/A_{circle}$ is an anisotropic factor relating the area of a Wulff shape and a circle of equal radius. For fourfold symmetry it was numerically computed and fitted by polynomial $C_W(\Omega,4) = \sum_{i=0}^4 a_i \Omega^{N-i}$ with $a_0 = -0.00032, a_1 = 0.00639, a_2 = -0.04219, a_3 = 0.00034, a_4 = 1.00000$ (rounded to the relevant decimal digits). As can be seen, the analytic shrinkage rate is a constant, which is consistent with [59].

With isotropic interface energy ($\delta = \Omega = 0$) the steady state shape is a circle and the anisotropic factor is $C_W/(1-\delta) = 1$, hence the isotropic curvature-driven shrinkage rate of a circle is (as also e.g. in [46])

$$\frac{\mathrm{d}A}{\mathrm{d}t} = -2\pi\mu\sigma_0. \tag{5.5}$$

5.3.2 Quantifying the match in shape

The Hausdorff distance was used for quantification of the match in shape. Let the 2D curves $\boldsymbol{w}, \boldsymbol{w}_{PF}$ be the analytic shape and the phase-field contour, respectively. The Hausdorff distance $d_H(\boldsymbol{w}, \boldsymbol{w}_{PF}) = \lambda$ between them implies, that λ is the smallest number such that \boldsymbol{w} is completely contained in λ -neighborhood of \boldsymbol{w}_{PF} and vice-versa [3]). Formally, it is defined between sets P and Q as

$$d_H(Q, P) = \max(\tilde{d}_H(P, Q), \tilde{d}_H(Q, P)), \qquad (5.6)$$

where

$$\tilde{d}_H(P,Q) = \max_{x \in P} (\min_{y \in Q} ||x - y||)$$
 (5.7)

is directed Hausdorff distance. It is always $d_H(\cdot,\cdot) \geq 0$, and the closer to zero, the more alike the compared sets are. It has been extensively used for image matching and pattern recognition [30].

For comparability in the two validation experiments with inclination-dependent interface energy, it is essential that the two curves w, w_{PF} are co-centric and scaled to unit radius.

Note tat the data set [36] includes also MATLAB functions for the contour shape matching.

5.3.3 Quantifying match in shrinkage rate

The shrinkage rate was obtained as mean value of shrinkage rates in simulation time interval where the area of the shape was in between 0.95-0.6 fraction of the initial area. This choice should prevent the diffuse interface from being too large compared to the shape itself, in which case it would affect the kinetics. Additionally, this approach turned out to be rather insensitive to the particular numeric settings, which is convenient for validations.

The results are presented as relative error δx , defined in the following convention

$$\delta x = 100 \frac{x_0 - x}{x} \%, (5.8)$$

where x_0 is the measured value and x is the expected one. In this convention, the shrinkage was slower than expected when $\delta x < 0$, and faster when $\delta x > 0$.

5.3.4 Wulff shape

Shrinking Wulff shapes with different strengths of anisotropy were simulated. The match to the analytic shape was measured in Hausdorff distance. The shrinkage rate was expressed analytically and used for validation as well.

The Neumann boundary conditions were applied to all boundaries in a system with initial condition like in Figure 5.2a. The initial condition in every simulation was the analytic Wulff shape of the corresponding strength of anisotropy Ω as discretized by the grid. The radius was taken such that the initial shape occupied the area fraction in the domain of at least 0.25. Because the initial Wulff shape already minimized the interface energy, the shrinkage with constant

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rate as in (5.4) was expected and any change in the shape was a departure from the analytic solution.

5.3.5 Kinetically compensated anisotropic circle shrinkage

This simulation experiment validates the inclination-dependence of the kinetic coefficient in combination with inclination-dependent interface energy. Specifically, the anisotropy of kinetic coefficient was chosen such that it compensated the anisotropic driving force so that the resulting interface motion was isotropic.

Again, the match to the steady-state shape (a circle) was quantified by Hausdorff distance and the mean shrinkage rate was measured when the shape area was a fraction 0.95-0.6 relative to the initial condition. Parametric study in Ω were carried out to validate the model.

The initial condition for the simulation experiment was as in Figure 5.2b, i.e. a two-phase-field system of a circular grain in a matrix.

Normal velocity v_n of a curvature-driven interface with inclination-dependent interface energy is [1]

$$v_n(\theta) = \frac{\mu}{\rho} \sigma_0[h(\theta) + h''(\theta)], \qquad (5.9)$$

where μ is interface mobility, ϱ is local radius of curvature and $\sigma_0[h(\theta) + h''(\theta)]$ is the interface stiffness. With $h(\theta) = 1 + \delta \cos(n\theta)$ the inclination-dependent factor in (5.9) is $[h(\theta) + h''(\theta)] = 1 - \delta(n^2 - 1)\cos(n\theta) = 1 - \Omega\cos(n\theta)$.

When the interface mobility is set anisotropic as

$$\mu(\theta) = \frac{\mu_0}{1 - \Omega \cos(n\theta)}, \qquad (5.10)$$

the resulting interface normal velocity v_n does not depend on interface inclination θ anymore, i.e. it is isotropic. The shrinkage rate is then (5.5).

The below presented simulations were all carried out with $\Omega < 1$, because the kinetic coefficient as in equation (5.10) is then positive for all interface inclinations.

The ratio of maximal to minimal interface velocity due to the anisotropic interface energy is $(1 + \Omega)/(1 - \Omega)$, which indicates rather strong kinetic anisotropy when Ω is close to 1. E.g. with $\Omega = 0.9$ the ratio of maximal to minimal interface velocity is 19 (assuming constant ϱ for all inclinations, which holds for a circle). Note that the corresponding ratio of maximal to minimal interface energy is only 1.0664 (with four-fold symmetry).

5.3.6 Triple junction angles

In this experiment triple junction angles are measured in systems with different combinations of pair-wise isotropic interface energies. This way it is validated how well the triple junction force balance is reproduced by the model.

The initial geometry was like in Figure 5.2b with periodic left and right boundaries, and Neumann boundary conditions on the top and bottom ones. The individual interfaces are isotropic but have different interface energies. The initially straight interface segments (1-2) and (1-3) with grain boundary energy σ_2 turn into circular arcs, which then move towards the center of curvature, i.e. downwards. The two grains η_2 and η_3 will shrink and in the steady state the angle α between the arcs (see Figure 5.7a) in the triple junction is described by Young's law [52] (section 3.3.3):

$$\alpha = 2a\cos(\sigma_1/2\sigma_2). \tag{5.11}$$

The ratio σ_1/σ_2 was varied in the parametric study to validate the model (see Table 5.2). It was always $\sigma_1 = 0.3 \,\mathrm{J/m^2}$ and σ_2 was computed from the ratio.

The phase field contours of interfaces (1-2) and (1-3) were analyzed by two methods in order to determine the triple junction angle α . First, the points on both the arcs nearest to the triple junction were fitted by a straight line (indicated by red segments in Figure 5.7a) and second, the remaining arc points were fitted by a circular arc (see green segments in Figure 5.7a). Simple geometric construction allows to determine the angle α from the fitted parameters in the latter case [46] as

$$\alpha = 2a\cos(x/R), \tag{5.12}$$

where R is the fitted circle arc radius and x is the horizontal distance from the triple junction (see the scheme in Figure 5.7a). Width of the interval in which the arcs were fitted by straight lines (width of pink rectangles in Figure 5.7a) was set to width of the interface (2-3) (i.e. half of the width on each side). Accuracy of the lines fitting is affected especially by the width of the above interval and that of the circle arc fitting is mostly affected by the simulated arc shape.

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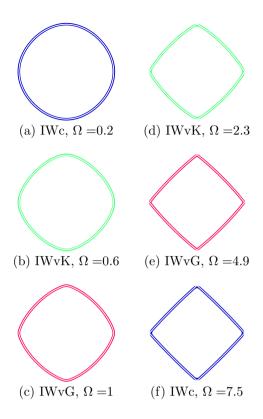


Figure 5.3: Demonstration of the simulated Wulff shapes for strengths of anisotropy Ω with the different models (base run). The white line is the analytic Wulff shape and the colored ones are the extracted phase field contours.

5.4 Results

5.4.1 Wulff shapes

The Figures 5.3a-5.3f visually compare the Wulff shapes obtained from simulation by the different models (in base run) to the analytic ones. As can be seen, the overall match is very good in all the three models, although a rather round contour near the corners in strong-anisotorpy Wulff shapes are observed (this is more apparent in Figure 5.4b, which shows detail of the contours near a corner for $\Omega = 7.5$). That was expected, as no special finite difference scheme was used near corners (in [17] a one-sided finite difference scheme was proposed to avoid

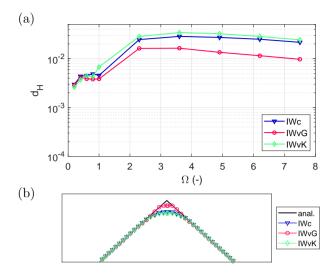


Figure 5.4: In (a) the match to Wulff shape for the model modifications in the base run as function of normalized strength of anisotropy Ω . In (b) a detail of the phase field contours near the Wulff shape corner for simulation with $\Omega = 7.5$.

corners rounding).

Figure 5.4a, shows the match to Wulff shape as function of normalized strength of anisotropy Ω in the base run. As can be seen, the IWvG model slightly outperformed the other two in strong anisotropies because it was able to resolve the corners the best (see Figure 5.4b). However, when the interface width was halved in the IW/2 run, all the models performed nearly equally well because with smaller interface width the rounding near the corners was reduced. Interestingly, the IWvG model performed comparably in the IW/2 and base runs, which is in contrast to IWc and IWvK, which improved markedly with narrower interface width.

The best results in match to Wulff shape were obtained in the 14IWpts run with IWvG model. IWc and IWvK models performed comparably in the 14IWpts and IW/2 runs.

The mean shrinkage rates of the Wulff shapes as function of strength of anisotropy are in Figure 5.5a for the base run. At first sight, all the models perform comparably well, following the analytical prediction within an absolute error of 3 %. The Figures 5.5b and 5.5c show the time evolution of shrinkage rate with $\Omega = 7.5$ in the base and IW/2 runs, respectively. It can be seen that

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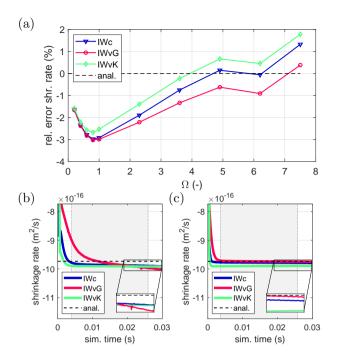


Figure 5.5: Wulff shape shrinkage rate results. In (a) the mean shrinkage rate in the base run as function of normalized strength of anisotropy Ω , in (b) and (c) there is time evolution of shrinkage rate for $\Omega=7.5$ in the base and IW/2 runs, respectively. The shaded areas in b) and c) indicate the time interval from which the mean shrinkage rate was computed.

despite the mean shrinkage rates being near the prediction, in the base run, the shrinkage rate of IWvG model did not converge to the analytic prediction (see the inset of Figure 5.5b, where the IWvG curve clearly declines). In the IWv2 run the IWvG model did converge close to the analytic shrinkage rate (see also the inset of Figure 5.5c) and all the mean values are a little closer than in the base run. Note that in the IWvG model the lowest-energy interface has the widest interface width and that the cornered Wulff shape contains only interface orientations with lower energy. For this reason it required narrower interface width to reach constant shrinkage rate.

The relative error in mean shrinkage rate obtained with this methodology was nearly the same in the base and IW/2 runs though (both within $\pm 3\%$). Apparently, the difference is that with narrower interface the attained shrinkage rate is more steady. Additional run with even finer grid was carried out and no

improvement in the mean values of shrinkage rates was observed. It can thus be concluded that the methodology is robust enough to assess the shrinkage rate even in the domain 100x100 (i.e. base run).

In addition to the four simulation runs discussed so far, the Wulff shapes simulations were re-run also with 4 and 5 points in the interface. The match to Wulff shape was worse than with 7 or 14 points, but the shapes were resolved qualitatively well regardless. The shrinkage rates were smaller than expected though, due to grid pinning. Seven points in the interface were thus confirmed as a reasonable value for the validations and practical simulations.

No significant effect of interface width scaling in the run 1000IW was found.

5.4.2 Kinetically compensated anisotropic circle shrinkage

The quantified match to the circle and the mean shrinkage rate as functions of strength of anisotropy Ω are in Figures 5.6a and 5.6b, respectively (results of the base run showed). All models and runs retained the initial circle well or up to excellent geometrical match. Nevertheless, the IWvG model gave the best results, except for the strongest anisotropy, where the IWvK model was better. Only minor improvement in the match was achieved in the IW/2 run when compared to the base run. As with Wulff shapes simulations, the best match was obtained in the 14IWpts run, which also exhibited the best mean shrinkage rates (for all the three models).

The relative error in shrinkage rates in Figure 5.6b shows slightly decreasing trend with Ω for IWc and IWvG models (i.e. slowing down). The values of IWvK model were not affected and were constant. Apparently, the symmetric profiles of IWvK model provide an advantage for preserving the expected kinetics in simulations with strong kinetic anisotropy.

Time evolution of shrinkage rate for the strongest considered anisotropy (i.e. $\Omega=0.9$) in Figures 5.6c and 5.6d (base and 14IWpts runs, respectively), shows that in both runs the lines slightly diverge (this being applicable to all Ω s). The convergence was better in IW/2 run, but the mean shrinkage rates were worse than in the 14IWpts run.

Apparently, optimal results would be obtained here with more points than 7 in the interface and with smaller interface-width-to-circle ratio than in the base run and 14IWpts runs. However, as noted earlier, the kinetic anisotropy is very strong in the case of $\Omega=0.9$. For weaker anisotropy the discussed effects are less pronounced and there is little difference in the shrinkage rates among the models.

There was no significant difference between the results of base and 1000IW runs.

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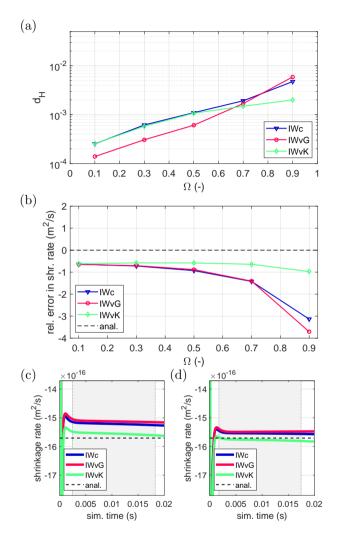


Figure 5.6: Results for kinetically compensated anisotropic shrinkage. In (a) the match to circle for the base run, in (b) the mean shrinkage rate for 14IWpts run, in (c) and (d) the shrinkage rate time evolution in simulation with $\Omega=0.9$ in the base run and 14IWpts runs, respectively.

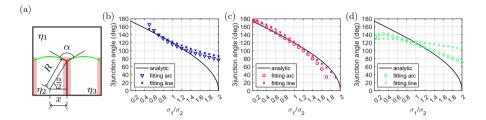


Figure 5.7: Triple junction angles. In a) the two methods for angles determination are illustrated (points fitted by straight lines in red and those fitted by circular arc in green). Subfigures b)-d) show the simulation results for different models, these being: in b) IWc, in c) IWvG and in d) IWvK. The hollow symbols correspond to the the angles determined by arc fitting and crosses to the lines fitting.

5.4.3 Equilibrium triple junction angles

Figures 5.7b-5.7d show the simulation results from the triple junction simulations as function of σ_1/σ_2 . Neither of IWc or IWvK models show good agreement to Young's law when deviating farther from an isotropic system (which has ratio $\sigma_1/\sigma_2 = 1$). The model IWvG, on the other hand, always shows very good agreement in at least one of the fitting methods along the whole range of probed ratios of interface energies. With σ_1/σ_2 closer to 2 in IWvG modification, the grains shape was slightly elongated in the vertical direction, resulting in too small radius of the fitted arcs to cross in the triple junction. The angles could not be determined this way then and the linear fit is more reliable.

For the ratios approximately $\sigma_1/\sigma_2 \leq 0.45$ the IWc model behaves non-physically. The triple junction was observed to move in the opposite than expected direction (i.e. upwards, as if the triple junction angle was larger then 180°). The overall shape of contours was not as in Figure 5.7a because the green arcs curved in the other way. In IWvK this behavior was not observed, but the Young's law is not followed. Qualitative explanation is that the IWc and IWvK models are not fully variational.

No significant change was observed in the results of the triple junction angles in runs IW/2, 14IWpts or 1000IW when compared to the base run. Quality of the results is thus not improved when more points in the interface than 7 are used. Also it implies, that the model behavior (for all parameters assignment strategies) is not affected by reducing the interface-width-to-feature ratio or the interface width scaling. The latter confirms that the model is quantitative.

CONCLUSIONS ______61

5.5 Conclusions

This chapter presented a quantitative methodology for assessment of the anisotropic curvature driving force in phase field method. It was demonstrated in comparison of three different modifications of a multi-phase field model. The match to the expected shape and the shrinkage rate were quantified in two different benchmark problems. The methodology was sensitive enough to capture differences between the model modifications and is suitable for validation and benchmarking of different models and numerical solvers.

The overall performance of the three model modifications in the benchmarks was comparable. Both match to the steady-state shapes and the shrinkage rates followed the expected results as in the anisotropic mean curvature flow. However, a significant difference was noted in a supplementary benchmark simulation where triple junction angles were measured. It was observed that only the IWvG model modification (with only the parameter γ anisotropic) reproduced the triple junction angles in the full interval 0.13-2 of σ_1/σ_2 , whereas IWc and IWvK modifications failed for ratios farther from 1. It is noted that these two modifications were not fully variational in order to avoid ghost phases, whereas IWvG is fully variational and yet the ghost phases do not appear.

Even though the triple junction benchmark did not involve interfaces with inclination-dependent interface energies, one conclusion can be made about that case regardless. As the IWc and IWvK were shown unreliable in the simpler pair-wise isotropic case, there is no reason why they should be reliable in the more complicated one. Further development and validations should thus focus on the IWvG model modification.

No results were significantly affected by the interface width scaling. Also, it was confirmed that 7 points in the interface were sufficient for retaining the expected kinetics in most cases, unless the inclination dependence of the kinetic coefficient was very strong.

The results in this chapter are reproducible with codes provided in the data set [36].

Appendix A

Functional derivative

The Alen-Cahn equation describing evolution of a non-conserved field $\xi(\mathbf{r})$ in a domain $\Omega \in \mathbb{R}^3$ with boundary $\partial \Omega$ can be written as

$$\frac{\partial \xi}{\partial t} = -L \frac{\delta F}{\delta \xi} \quad \text{in } \Omega \tag{A.1}$$

$$\nabla \xi \cdot \boldsymbol{n} = 0 \quad \text{at } \partial \Omega \,, \tag{A.2}$$

where L is a positive constant and the function $\frac{\delta F}{\delta \xi}$ stands for functional derivative of the free energy functional F. The latter expresses how much the energy changes when the function $\xi(\mathbf{r})$ is infinitesimally varied by addition of some test function multiplied by small number, i.e. multiplied by $\varepsilon g(\mathbf{r})$, where $0 < \varepsilon < 1$.

The functional is a map from some function space to real numbers and in our case $F[\xi(\mathbf{r})] = \int f(\mathbf{r}, \xi(\mathbf{r}), \nabla \xi(\mathbf{r})) dV$ with the integrand being energy density. The change in functional value depending on ε is called functional variation and is defined $\delta F = F[\xi(\mathbf{r}) + \varepsilon g(\mathbf{r})] - F[\xi(\mathbf{r})]$. The term $F[\xi(\mathbf{r}) + \varepsilon g(\mathbf{r})]$ can be seen as a function of ε and the functional derivative is a functional defined as

$$\frac{\delta F}{\delta \xi}[\xi, g] := \left[\frac{\mathrm{d}}{\mathrm{d}\varepsilon} F[\xi + \varepsilon g] \right]_{\varepsilon = 0} = \lim_{\varepsilon \to 0} \frac{F[\xi(\mathbf{r}) + \varepsilon g(\mathbf{r})] - F[\xi(\mathbf{r})]}{\varepsilon} \tag{A.3}$$

The functional F is differentiable when the derivative A.3 exists and can be written in form of integral

$$\left[\frac{\mathrm{d}}{\mathrm{d}\varepsilon}F[\xi+\varepsilon g]\right]_{\varepsilon=0} = \int_{\Omega} \frac{\delta F}{\delta \xi}(\mathbf{r})g(\mathbf{r})dV \tag{A.4}$$

the kernel of which $\frac{\delta F}{\delta \xi}(\mathbf{r})$ is the function appearing in the Allen-Cahn equation. The expression for $\frac{\delta F}{\delta \xi}(\mathbf{r})$ can be obtained using Taylor expansion of $F[\xi(\mathbf{r}) + \varepsilon g(\mathbf{r})]$ in $\varepsilon = 0$, which will be shown first in the case of an 1D system to illustrate the procedure.

The functional is $F[\xi(x)] = \int_a^b f(x,\xi,\xi') dx$, where the integration limits a,b are constant and $\xi = \xi(x), \xi' = \frac{d\xi}{dx}(x)$. The functional derivative is then

$$\frac{\delta F}{\delta \xi}[\xi, g] = \lim_{\varepsilon \to 0} \frac{1}{\varepsilon} \int_{a}^{b} [f(x, \xi + \varepsilon g, \xi' + \varepsilon g') - f(x, \xi, \xi')] dx =
= \lim_{\varepsilon \to 0} \frac{1}{\varepsilon} \int_{a}^{b} \left[\frac{\partial f}{\partial \xi} \varepsilon g + \frac{\partial f}{\partial \xi'} \varepsilon g' + \mathcal{O}(\varepsilon^{2}) \right] dx =
= \int_{a}^{b} \left(\frac{\partial f}{\partial \xi} g + \frac{\partial f}{\partial \xi'} g' \right) dx =
= \int_{a}^{b} \left(\frac{\partial f}{\partial \xi} - \frac{\partial}{\partial x} \frac{\partial f}{\partial \xi'} \right) g dx + \left[\frac{\partial f}{\partial \xi'} g \right]_{a}^{b}$$
(A.5)

In the last equation sign per partes technique was used.

So far there was no assumption on test function g(x), but in the simplest case the end points of ξ are fixed, i.e. g(a) = g(b) = 0, which removes the second term and allows to write

$$\frac{\delta F}{\delta \xi}(x) = \frac{\partial f}{\partial \xi} - \frac{\partial}{\partial x} \frac{\partial f}{\partial \xi'}.$$
 (A.6)

If the left-hand side is replaced by zero, we obtain *Euler-Lagrange* equation, solution to which minimizes the functional F.

Now, in the case of functional in 3D space $F[\xi(\mathbf{r})] = \int_{\Omega} f(\mathbf{r}, \xi(\mathbf{r}), \nabla \xi(\mathbf{r})) \mathrm{d}^3 r$ we must realize that dependence of the integrand on a vector variable means dependence on its components, i.e. $f(\mathbf{r}, \xi(\mathbf{r}), \nabla \xi(\mathbf{r})) = f(x_1, x_2, x_3, \xi, \partial_1 \xi, \partial_2 \xi, \partial_3 \xi)$, where $\partial_i = \frac{\partial}{\partial x_i}$. The functional derivative is then derived in a way analogous to the previous case, but now there is one term in the Taylor expansion per each gradient component. Before the derivation we define the following vector field

$$\frac{\partial f}{\partial(\nabla \xi)} = \frac{\partial f}{\partial(\partial_1 \xi)} \mathbf{i} + \frac{\partial f}{\partial(\partial_2 \xi)} \mathbf{j} + \frac{\partial f}{\partial(\partial_3 \xi)} \mathbf{k} =
= \frac{\partial f}{\partial(\frac{\partial \xi_p}{\partial x})} \mathbf{i} + \frac{\partial f}{\partial(\frac{\partial \xi_p}{\partial y})} \mathbf{j} + \frac{\partial f}{\partial(\frac{\partial \xi_p}{\partial z})} \mathbf{k}$$
(A.7)

And we proceed to the functional derivative

$$\frac{\delta F}{\delta \xi}[\xi, g] = \lim_{\varepsilon \to 0} \frac{1}{\varepsilon} \int_{\Omega} [f(x_i, \xi + \varepsilon g, \partial_i(\xi + \varepsilon g)) - f(x_i, \xi, \partial_i \xi)] d^3 r =
= \lim_{\varepsilon \to 0} \frac{1}{\varepsilon} \int_{\Omega} \left[\frac{\partial f}{\partial \xi} \varepsilon g + \sum_{i=1}^{3} \frac{\partial f}{\partial (\partial_i \xi)} \varepsilon \partial_i g + \mathcal{O}(\varepsilon^2) \right] d^3 r =
= \int_{\Omega} \left[\frac{\partial f}{\partial \xi} g + \frac{\partial f}{\partial (\nabla \xi)} \cdot \nabla g \right] d^3 r
= \int_{\Omega} \left[\frac{\partial f}{\partial \xi} - \nabla \cdot \frac{\partial f}{\partial (\nabla \xi)} \right] g d^3 r + \int_{\partial \Omega} g \frac{\partial f}{\partial (\nabla \xi)} \cdot d \mathbf{n}$$
(A.8)

In the last equation there was used Stokes' theorem¹ in combination with the rule for divergence of product of vector and scalar field². Hence, in order to null the surface integral in equation A.8, either the boundary values must be fixed (Dirichlet BCs) or the following boundary condition (so called *natural BC*) must be satisfied

$$\frac{\partial f}{\partial (\nabla \xi_p)} \cdot \boldsymbol{n} = 0. \tag{A.9}$$

Different choice of the boundary conditions will interfere with minimization of the functional. The safest option is when the interface does not come near the boundary.

with scalar field f(r) and vector field F(r) the Stokes theorem gives: $\int_{\Omega} \nabla \cdot (Ff) = \int_{\partial \Omega} f F \cdot dn$, where n is a vector normal to the domain boundary $\nabla \cdot (Ff) = (\nabla \cdot F)f + F \cdot (\nabla f)$

Appendix B

Model parameters determination - best practices

The multi-phase field model by Moelans [42] is sensitive to how exactly the values of the parameters m, κ, γ are assigned in systems with anisotropic interface energy. In the model variants with variable interface width (IWvG and IWvK), it is important to assure that no interface becomes too narrow, so that grid pinning is prevented. Otherwise, the kinetics is no longer accurately reproduced by the implementation. The best practices in parameters determination described in this section were implemented in MATLAB functions, which were published in dataset [36], distributed under GPLv3 license.

B.1 Polynomial expressions for determination of γ

Interface energy and width are non-analytic functions of the parameter γ (equations (4.6) and (4.7) in the main text). In models IWc and IWvG with non-uniform interface energy, γ varies in space and depends on physical input (local interface energy σ) and chosen interface width. Assuming that κ, m and σ are known, $g(\gamma)$ is computed using equation (4.6) and must be inverted in order to obtain γ . For numeric convenience, functions $g^2 \to 1/\gamma$ and $1/\gamma \to f_{0,c}$ were fitted by polynomials using the tabulated values from [54]. The same approach was taken in [42]. Note that in one of the following subsections an alternative algorithm for determination of parameters in IWc is proposed.

The values of $f_{0,c}(\gamma), g(\gamma)$ were tabulated and published in [54] in the range

 $0.52 \le \gamma \le 40$. However, the polynomials used in both [42, 54] did not describe the functions along the whole tabulated range. For inconvenient input, that resulted in a loss of control over the interface properties (even negative values of γ were obtained).

To improve the parameters determination, the functions $g^2 \to 1/\gamma$ and $1/\gamma \to f_{0,c}$ were fitted again so that all γ values from the interval $0.52 \le \gamma \le 40$ were reliably available. The new polynomials show good agreement along the given interval and their coefficients are listed in Table B.1.

B.2 IWvG

This parameters determination algorithm [54] requires input of initializing interface energy σ_{init} , and width l_{init} which are assigned to interface with value $\gamma = 1.5$. There, the analytic relations between the model parameters are valid and the (constant) parameters κ , m can be computed:

$$\kappa = \frac{3}{4}\sigma_{init}l_{init} \quad , \quad m = 6\frac{\sigma_{init}}{l_{init}} \,. \tag{B.1}$$

These are then used in equation (4.6) with individual interface energies $\sigma_{i,j}$ to determine $g^2(\gamma_{i,j})$ and further $\gamma_{i,j}$ (using the first polynomial from Table B.1). However, because the said polynomial was only fitted in the interval $0.52 \le \gamma \le 40$ (values tabulated in [54]), it is important to choose the value of σ_{init} such, that none of the interface energies $\sigma_{i,j}$ should be described by $\gamma_{i,j}$ out of the interval. From equation (4.6) with substituted $\sqrt{\kappa m}$ from B.1 we show that the values do not leave the above interval when σ_{init} complies with the following two inequalities

$$\frac{\sigma_{max}}{g(40)}\sqrt{\frac{2}{9}} \le \sigma_{init} \le \frac{\sigma_{min}}{g(0.52)}\sqrt{\frac{2}{9}},$$
(B.2)

where σ_{min} , σ_{max} are minimal and maximal interface energy present in the system. Additionally, in order to fulfill these conditions, it must be satisfied that

$$\frac{\sigma_{min}}{\sigma_{max}} \ge \frac{g(0.52)}{g(40)} = \frac{0.1}{0.76} = 0.129,$$
(B.3)

which sets the maximal anisotropy in interface energy for which all $\gamma_{i,j}$ s may be determined in IWvG based on the values tabulated in [54].

Control over the width of the narrowest interface (with user-defined width l_{min}) is obtained when the parameters are determined in two steps as follows:

1. values of κ , m are calculated from equations (B.1) using appropriate σ_{init} together with l_{min}

IWVG _

gether Table B.1: Coefficients of the new polynomials used in the parameters determination algorithm (IWc and IWvG mode with

ang of the nume $p_N(x) = \sum_{i=0}^{N}$ order $\frac{a_0}{-5.730}$ $\frac{5}{6}$ -0.0729	odels). Based on fitting of the numeric values given in [54]. The polynomial of order N $p_N(x)$ is defined here toger. In its coefficients as $p_N(x) = \sum_{i=0}^N a_i x^{N-i}$. See text for details of use in the respective models.	a_1 a_2 a_3 a_4 a_5 a_6	08 18.8615 -23.0557 7.47952 8.33568 -8.01224 2.00013	-0.072966 0.35784 -0.68325 0.63578 -0.48566 0.53703	103.397 -165.393 105.3469 -44.55661 24.7348 -11.25718 1.999642
dels). Based on fitting of the numeric values given in [54]. The polynomial of order N $p_N(x)$ is defined here toger that its coefficients as $p_N(x) = \sum_{i=0}^N a_i x^{N-i}$. See text for details of use in the respective models. order a_0	rnomial of cuse in the 1	a_3	17952 8.	3578 -0	55661 24
ang of the numeric values given in [\bar{t} $p_N(x) = \sum_{i=0}^N a_i x^{N-i}$. See text fo order a_0 a_1 -5.73008 18.8615 -2.5 -0.072966 0.35784 -0.0 6 103.397 -165.393 105	 The poly of the details of τ 		3.0557 7.4	68325 0.6	5.3469 -44.
ang of the numeric value $p_N(x) = \sum_{i=0}^{N} a_i x^{N-1}$ order a_0	les given in [5 ' i '. See text fo	a_1	18.8615 -23	0.35784 -0.0	165.393 105
ing of th $p_N(x) = \frac{1}{2}$ order $\frac{6}{5}$	e numeric valu $=\sum_{i=0}^{N} a_i x^{N-1}$	a_0	-5.73008	-0.072966	103.397
	sing of th $p_N(x) =$	order	9	5	$q_1/f_{0.c} \to \frac{1}{1}$ 6

2. values of κ , m are recalculated from equations (B.1) with the same σ_{init} and interface width $l_{calib} = l_{min} \sqrt{8f_{0,c}(\gamma_{min})}$, where γ_{min} corresponds to the lowest value of all the $\gamma_{i,j}$ obtained in the previous step, i.e. to the interface with the largest interface energy σ_{max} .

After the second step, the interface with maximal interface energy will have the width l_{min} and all the other interfaces will have it equal or larger.

The other interface widths can be expressed

$$l_{i,j} = \frac{\sigma_{i,j}}{m} \frac{1}{g(\gamma_{i,j})\sqrt{f_{0,c}}(\gamma_{i,j})}$$
(B.4)

B.3 IWc

The iterative parameters determining procedure assuring the constant interface width is described in [42]. The procedure requires initializing value of interface energy σ_{init} , which has decisive role for the parameters values. For inconvenient choice of σ_{init} , some of the parameters $\gamma_{i,j}$ may get out of range of applicability of the polynomials in Table B.1. Systematic investigation revealed that the value $\sigma_{init} = (\sigma_{max} + \sigma_{min})/2$ prevents occurrence of such effects for anisotropies $\sigma_{min}/\sigma_{max} \geq 0.03$.

Here we propose a simplified, yet equivalent single-step procedure. We assume constant interface width l and an initializing interface energy σ_{init} , for which $\gamma=1.5$ and thus the analytic relations between the model parameters hold. As in the original iterative algorithm, the barrier height is computed

$$m = 6 \frac{\sigma_{init}}{l}.$$
 (B.5)

Then, the gradient energy coefficient $\kappa_{i,j}$ can be expressed from both equations (4.7) and (4.8), hence we can write

$$\kappa_{i,j} = f_{0,c}(\gamma_{i,j})ml^2 = \sigma_{i,j}l \frac{\sqrt{f_{0,c}(\gamma_{i,j})}}{g(\gamma_{i,j})},$$
(B.6)

from where we obtain

$$g(\gamma_{i,j})\sqrt{f_{0,c}(\gamma_{i,j})} = \frac{\sigma_{i,j}}{ml} = \frac{\sigma_{i,j}}{6\sigma_{init}}.$$
 (B.7)

In the last equation, m was substituted from (B.5). We denote $G(\gamma_{i,j}) = g(\gamma_{i,j}) \sqrt{f_{0,c}(\gamma_{i,j})}$. This function $\gamma \to G$ can be evaluated from the data tabulated in [54] and its inverse function $G \to \frac{1}{\gamma}$ was fitted by 6-th order

polynomial producing the coefficients in the last row of Table B.1. Apparently, evaluation of this polynomial in points $\sigma_{i,j}/6\sigma_{init}$ produces the searched $1/\gamma_{i,j}$. The gradient energy coefficients are then best computed as in B.6, where m was substituted from B.5:

$$\kappa_{i,j} = 6f_{0,c}(\gamma_{i,j})\sigma_{init}l. \tag{B.8}$$

Here, the fitted polynomial $\frac{1}{\gamma} \to \sqrt{f_{0,c}}$ is used to obtain the values of $f_{0,c}(\gamma_{i,j})$.

Similarly like in IWcG model, there are conditions limiting suitable values of σ_{init}

$$\frac{\sigma_{max}}{6G(40)} \le \sigma_{init} \le \frac{\sigma_{min}}{6G(0.52)}, \tag{B.9}$$

where $G(0.52) \approx 0.0069$, $G(40) \approx 0.4065$. Also must be satisfied

$$\frac{\sigma_{min}}{\sigma_{max}} \ge \frac{G(0.52)}{G(40)} = 0.017.$$
 (B.10)

The main advantage of the proposed algorithm is that no iterations are needed for parameters determination. In the original algorithm the number of iteration steps is uncertain (tens of repetitions are not exceptional though) and sometimes the results may not converge due to reasons tedious to debug. Obtaining each parameter in a single step thus provides better control over the process and simplifies the implementation. Moreover, the permissible values of σ_{init} and the maximal anisotropy $\sigma_{min}/\sigma_{max}$ are clearly defined.

Also, it should be noted that when σ_{init} is chosen close to the upper limit in the inequality B.9, the determined $\gamma_{i,j}$ s are rather closer to value 0.52 (corresponding to long-tailed interfaces). When σ_{init} is near the bottom limit, the values of γ_{max} are close to the value 40 and γ_{min} are much larger than in the previous case. Near the limiting anisotropy $\sigma_{min}/\sigma_{max}\approx 0.017$, the values always are $\gamma_{max}\approx 40$ and $\gamma_{min}\approx 0.52$. It is especially near the limits of admissible σ_{init} , where the original iterative parameters assignment strategy is not always reliable when compared to the proposed one.

B.4 IWvK

When $\gamma = const = 1.5$, the analytic relations between the model parameters hold (i.e. equations (4.8)-(4.10)). Then, the minimal interface width l_{min} is controlled when

$$m = 6 \frac{\sigma_{min}}{l_{min}} \tag{B.11}$$

and

$$\kappa_{i,j} = \frac{9}{2} \frac{\sigma_{i,j}^2}{m} \,. \tag{B.12}$$

Interface width of an interface i-j can then be computed as

$$l_{i,j} = 6 \frac{\sigma_{i,j}}{m} \,. \tag{B.13}$$

Appendix C

Implementation of finite difference method

C.1 General information

The 2D system was discretized into a rectangular equidistant grid and rearranged into a single column vector as indicated in figure C.1. The differential operators were represented by finite-difference matrices left-multiplying the column vectors standing for field variables present on the grid. The finite-difference matrices are block matrices, where each block is a square matrix acting on a single column in the rearranged system.

Let the system be a matrix with n rows and m columns. All blocks are matrices with dimensions $n \times n$ and they are ordered in $m \times m$ grid giving rise to an $mn \times mn$ sparse finite-difference matrix.

The finite-difference matrix can perform linear combination of arbitrary points in the system in each grid point. It takes various form depending on the boundary conditions. However, it is always a toeplitz block matrix.

Second-order accurate centered finite-difference scheme in space and first-order in time (i.e. the Euler method) were used.

In the following sections, the symbol \mathbb{O} stands for $n \times n$ block matrix of zeros and \mathbb{I} for $n \times n$ identity matrix (i.e. matrix with ones in the diagonal and zeros otherwise). The finite difference method is abbreviated as FD in the following. In the finite-difference matrices below, some elements may be highlighted in red. These are the elements, which are involved in the boundary conditions.

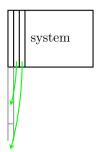


Figure C.1: Rearrangement of the matrix representing the system before application of the discrete deferential operators.

C.2 Ghost nodes in Neumann boundary condition

C.2.1 Zero Neumann BCs

Neumann boundary conditions were implemented using ghost nodes. Ghost nodes are hypothetical nodes adjacent to the boundary nodes, in which the values can be specified using the given boundary condition. In order to have the derivatives in the boundary nodes second-order-accurate, centered difference can be used using the known values at ghost nodes. In the case $\nabla \eta \cdot \mathbf{n} = 0$ we can write at the left boundary edge (i.e. x = 0), where (i, j) = (0, j)

$$\frac{\partial \eta}{\partial x} = 0 \quad \text{at } x = 0 \quad \Rightarrow \quad \forall j \in m + 1 \ (\Delta_x \eta)_{0j} = \frac{\eta_{1j} - \eta_{-1j}}{h} = 0 \quad \Longrightarrow \quad \eta_{1j} = \eta_{-1j}$$

$$\frac{\partial \eta}{\partial y} = 0 \quad \text{at } x = 0 \quad \Rightarrow \quad \forall i \in n + 1 \ (\Delta_y \eta)_{0j} = \frac{\eta_{ij+1} - \eta_{ij-1}}{h}$$
(C.1)

The first equation provides the value at the ghost node. In order to have the centered difference in the respective direction zero at the boundary, apparently the value must be equal to the value in the neighbour on the opposite side.

C.2.2 General Neumann BCs

The boundary condition defining the inclination angle θ at the boundary is $\nabla \eta \cdot \boldsymbol{n} = |\nabla \eta| \cos(\theta)$. This expression implies at the left boundary (i, j) = (0, j) that $\frac{\partial \eta}{\partial x} = |\nabla \eta| \cos(\theta)$ and $\frac{\partial \eta}{\partial y} = |\nabla \eta| \sin(\theta)$. By replacing the partial derivatives $\frac{\partial}{\partial x}$, $\frac{\partial}{\partial y}$ with finite differences Δ_x , Δ_y the values in ghost node (-1, j) can be

expressed as

$$\forall j \in m+1 \quad (\Delta_x \eta)_{0j} = \frac{\eta_{1j} - \eta_{-1j}}{2h} = |\nabla \eta|_{0,j} \cos(\theta)$$
 (C.2)

$$\implies \eta_{-1i} = \eta_{1i} - 2h|\nabla \eta|_{0i}\cos(\theta). \tag{C.3}$$

At the right boundary (i, j) = (n, j) the ghost nodes (n + 1, j) are

$$\forall j \in m+1 \quad (\Delta_x \eta)_{nj} = \frac{\eta_{n+1j} - \eta_{n-1j}}{2h} = |\nabla \eta|_{nj} \cos(\theta)$$
 (C.4)

$$\implies \eta_{n+1j} = \eta_{n-1j} + 2h|\nabla \eta|_{nj}\cos(\theta). \tag{C.5}$$

And at the bottom and top, respectively, using the same principles we obtain

$$\forall i \in n+1 \quad \eta_{i,-1} = \eta_{i,1} - 2h|\nabla \eta|_{i,0}\cos(\theta) \tag{C.6}$$

$$\forall i \in n+1 \quad \eta_{i,m+1} = \eta_{i,m-1} + 2h|\nabla \eta|_{i,m}\cos(\theta). \tag{C.7}$$

Important observation is that the general Neumann BC indeed reduces to the zero condition for $\theta = 90^{\circ}$ and that the values at the ghost nodes are always derived from the neighbour on the opposite side.

C.3 First-order derivatives in FD

The centered finite difference was used with stencil $\begin{bmatrix} -1 & 0 & 1 \end{bmatrix}$ for x-direction derivative and $\begin{bmatrix} 1 & 0 & 1 & 1 \end{bmatrix}$ for y-direction derivative.

C.3.1 Periodic boundary conditions

For the y-direction derivative the FD matrix is

$$\frac{\partial}{\partial y} \approx \Delta_y = \frac{1}{2h} \begin{pmatrix}
\mathbb{Y} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} \\
\mathbb{O} & \mathbb{Y} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} \\
\mathbb{O} & \mathbb{O} & \mathbb{Y} & \dots & \mathbb{O} & \mathbb{O} \\
\vdots & \vdots & \vdots & \ddots & \vdots & \vdots \\
\mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{Y} & \mathbb{O} \\
\mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{Y}
\end{pmatrix}$$
(C.8)

with

$$\mathbb{Y} = \begin{pmatrix}
0 & 1 & 0 & 0 & \dots & 0 & 0 & -1 \\
-1 & 0 & 1 & 0 & \dots & 0 & 0 & 0 \\
0 & -1 & 0 & 1 & \dots & 0 & 0 & 0 \\
\vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots & \vdots \\
0 & 0 & 0 & 0 & \dots & -1 & 0 & 1 \\
1 & 0 & 0 & 0 & \dots & 0 & -1 & 0
\end{pmatrix}$$
(C.9)

In the x-direction, the finite difference matrix is

$$\frac{\partial}{\partial x} \approx \Delta_x = \frac{1}{2h} \begin{pmatrix} \mathbb{O} & \mathbb{I} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} & -\mathbb{I} \\ -\mathbb{I} & \mathbb{O} & \mathbb{I} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \mathbb{O} & -\mathbb{I} & \mathbb{O} & \mathbb{I} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots & \vdots \\ \mathbb{O} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & -\mathbb{I} & \mathbb{O} & \mathbb{I} \\ \mathbb{I} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & -\mathbb{I} & \mathbb{O} \end{pmatrix}. \tag{C.10}$$

C.3.2 Neumann boundary conditions

The FD matrices are equal in both cases of the zero and general NBC, but in the case of the gerneral NBC the following vector must be added at the left and right boundary $|\nabla \eta| \begin{bmatrix} \cos(\theta) \\ \sin(\theta) \end{bmatrix}$ and the following one at the bottom and top boundary $|\nabla \eta| \begin{bmatrix} -\sin(\theta) \\ \cos(\theta) \end{bmatrix}$.

For the y-direction the FD matrix is the same as for the periodic boundary conditions in the block structure (equation C.32), and the block matrix \mathbb{Y} is only modified in its first and last row as

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For the x-direction with Neoumann boundary conditions

$$\frac{\partial}{\partial x} \approx \Delta_x = \frac{1}{2h} \begin{pmatrix} \bigcirc & \bigcirc & \bigcirc & \bigcirc & \bigcirc & \cdots & \bigcirc & \bigcirc & \bigcirc & \bigcirc \\ -\mathbb{I} & \bigcirc & \mathbb{I} & \bigcirc & \cdots & \bigcirc & \bigcirc & \bigcirc & \bigcirc \\ \bigcirc & -\mathbb{I} & \bigcirc & \mathbb{I} & \cdots & \bigcirc & \bigcirc & \bigcirc & \bigcirc \\ \bigcirc & -\mathbb{I} & \bigcirc & \mathbb{I} & \cdots & \bigcirc & \bigcirc & \bigcirc & \bigcirc \\ \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots & \vdots \\ \bigcirc & \bigcirc & \bigcirc & \bigcirc & \bigcirc & \cdots & -\mathbb{I} & \bigcirc & \mathbb{I} \\ \bigcirc & \bigcirc & \bigcirc & \bigcirc & \bigcirc & \cdots & \bigcirc & \bigcirc & \bigcirc \end{pmatrix}$$
(C.12)

C.3.3 Mixed boundary conditions

In 2D system there are only two options when combining periodic and Neumann BCs: the periodic boundary is either parallel to x or y axis. The periodic boundary must always lay opposite to another periodic boundary. In a system with mixed boundaries the finite difference matrices in the x and y direction are combined from the above options so that the BCs are met at the corresponding boundary.

C.4 Laplacian in 2D

bi-equidistant grid with spacing h in a grid point (i, j) as

$$(\nabla^2 u)_{ij}^{5p} = u_{ij,xx} + u_{ij,yy} = \frac{1}{h^2} (-4u_{ij} + u_{i-1j} + u_{i+1j} + u_{ij-1} + u_{ij+1}) + \mathcal{O}(h^4)$$
(C.13)

$$(\nabla^2 u)_{ij}^{9p8} = \frac{1}{3h^2} (-8u_{ij} + u_{i-1j} + u_{i+1j} + u_{ij-1} + u_{ij+1} + u_{ij+1} + u_{i+1j+1} + u_{i-1j-1} + u_{i-1j+1} + u_{i+1j-1}) + \mathcal{O}(h^4)$$
(C.14)

9-point stencil (denoted 9p20) $\frac{\begin{array}{c|ccccc} 1 & 4 & 1 \\ \hline 4 & -20 & 4 \\ \hline 1 & 4 & 1 \end{array} } \text{is expressed in a grid point } (i,j)$ as

$$(\nabla^{2}u)_{ij}^{9p20} = \frac{1}{6h^{2}} [-20u_{ij} + 4(u_{i-1j} + u_{i+1j} + u_{ij-1} + u_{ij+1}) + u_{i+1j+1} + u_{i-1j-1} + u_{i-1j+1} + u_{i+1j-1}] + \mathcal{O}(h^{4})$$
(C.15)

The above formulas are derived from Taylor expansions of the values of function u at the points adjacent to (i,j) assuming the small displacement h in both x and y direction.

The above formulas are represented by matrix, which carries the weighted summation of the function values for each point similarly to the first-order derivatives. Below, the matrices are shown in detail depending on the BCs.

C.4.1 Periodic boundary conditions

For 5p stencil the blocks are unit matrix \mathbb{I} , zero matrix \mathbb{O} and the following matrix \mathbb{D}^{5p} (the red elements are the top-bottom periodic condition)

$$\mathbb{D}_{P}^{5p} = \begin{pmatrix}
-4 & 1 & 0 & 0 & \dots & 0 & 0 & \mathbf{1} \\
1 & -4 & 1 & 0 & \dots & 0 & 0 & 0 \\
0 & 1 & -4 & 1 & \dots & 0 & 0 & 0 \\
\vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots \\
0 & 0 & 0 & 0 & \dots & 1 & -4 & 1 \\
\mathbf{1} & 0 & 0 & 0 & \dots & 0 & 1 & -4
\end{pmatrix}.$$
(C.16)

and the finite-difference matrix \mathbb{L}^{5p} is ordered as follows (red elements are the left-right periodic boundary condition)

$$\mathbb{L}_{P}^{5p} = \frac{1}{h^{2}} \begin{pmatrix} \mathbb{D}_{P}^{5p} & \mathbb{I} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{I} \\ \mathbb{I} & \mathbb{D}_{P}^{5p} & \mathbb{I} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \mathbb{O} & \mathbb{I} & \mathbb{D}_{P}^{5p} & \mathbb{I} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots \\ \mathbb{O} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{I} & \mathbb{D}_{P}^{5p} & \mathbb{I} \\ \mathbb{I} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{I} & \mathbb{D}_{P}^{5p} \end{pmatrix}$$
(C.17)

For 9p20 stencil the blocks are a Toeplitz matrix \mathbb{D}_{P}^{9p20} , a Toeplitz matrix \mathbb{A} and a zero matrix \mathbb{O} . These are defined

$$\mathbb{D}_{P}^{9p20} = \begin{pmatrix} -20 & 4 & 0 & 0 & \dots & 0 & 0 & 4\\ 4 & -20 & 4 & 0 & \dots & 0 & 0 & 0\\ 0 & 4 & -20 & 4 & \dots & 0 & 0 & 0\\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots\\ 0 & 0 & 0 & 0 & \dots & 4 & -20 & 4\\ 4 & 0 & 0 & 0 & \dots & 0 & 4 & -20 \end{pmatrix}$$
 (C.18)

$$\mathbb{A} = \begin{pmatrix} 4 & 1 & 0 & 0 & \dots & 0 & 0 & 1 \\ 1 & 4 & 1 & 0 & \dots & 0 & 0 & 0 \\ 0 & 1 & 4 & 1 & \dots & 0 & 0 & 0 \\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots \\ 0 & 0 & 0 & 0 & \dots & 1 & 4 & 1 \\ 1 & 0 & 0 & 0 & \dots & 0 & 1 & 4 \end{pmatrix}$$
(C.19)

and arranged

$$\mathbb{L}_{P}^{9\text{p}20} = \frac{1}{6h^{2}} \begin{pmatrix} \mathbb{D}_{P}^{9\text{p}20} & \mathbb{A} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{A} \\ \mathbb{A} & \mathbb{D}_{P}^{9\text{p}20} & \mathbb{A} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \mathbb{O} & \mathbb{A} & \mathbb{D}_{P}^{9\text{p}20} & \mathbb{A} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \vdots & \vdots \\ \mathbb{O} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{A} & \mathbb{D}_{P}^{9\text{p}20} & \mathbb{A} \\ \mathbb{A} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{A} & \mathbb{D}_{P}^{9\text{p}20} \end{pmatrix}$$
(C.20)

C.4.2 Neuman boundary conditions

For both zero and general NBCs the FD matrices are equal. Because the value at ghost nodes is either equal or derived from value of neighbour on the opposite side, the laplacian stencil behaves at the Neumann boundary as in Figure C.2.

The 5p stencil finite-difference matrix \mathbb{L}_N^{5p} with Neuman BC is composed of 4 different blocks: tridiagonal \mathbb{D}_N^{5p} , unit matrix \mathbb{I} , $2\mathbb{I}$ and zero matrix \mathbb{O} . These

a)										b)								
/		1					1		L	- /								
	1	-4	1			1	-4	1			-4	2			1	-4	1	
		1					1				2					2		
					1									1				
		1		1	-4	1					1		1	-4	1			
	1	-4	1		1						-4	2		1				
		1									1							

Figure C.2: Effect of Neuman BC on the 5pt laplacian stencil on the boundary points. Ghost points are outside the simulation domain but they can be accounted for thanks to knowing their value from equation C.1.

are:

$$\mathbb{D}_{N}^{5p} = \begin{pmatrix}
-4 & 2 & 0 & 0 & \dots & 0 & 0 & 0 \\
1 & -4 & 1 & 0 & \dots & 0 & 0 & 0 & 0 \\
0 & 1 & -4 & 1 & \dots & 0 & 0 & 0 & 0 \\
\vdots & \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots & \vdots \\
0 & 0 & 0 & 0 & \dots & 1 & -4 & 1 & 1 \\
0 & 0 & 0 & 0 & \dots & 0 & 2 & 4
\end{pmatrix}$$
(C.21)

and are ordered as follows

$$\mathbb{L}_{N}^{5p} = \frac{1}{h^{2}} \begin{pmatrix} \mathbb{D}_{N}^{5p} & 2\mathbb{I} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \mathbb{I} & \mathbb{D}_{N}^{5p} & \mathbb{I} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \mathbb{O} & \mathbb{I} & \mathbb{D}_{N}^{5p} & \mathbb{I} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots \\ \mathbb{O} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{I} & \mathbb{D}_{N}^{5p} & \mathbb{I} \\ \mathbb{O} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & 2\mathbb{I} & \mathbb{D}_{N}^{5p} \end{pmatrix}$$
(C.23)

In case of general NBC in 5-point stencil there must always be added or subtracted the term $2h|\nabla \eta|_{ij}\cos(\theta)$, where (i,j) are the boundary points. At the left and bottom boundary it is subtracted, at the right and top it is added.

The 9p20 stencil finite-difference matrix \mathbb{L}_N^{9p20} with Neuman BC is composed of 4 differnt blocks: tridiagonal \mathbb{D}_N^{9p20} , tridiagonal \mathbb{B} , $2\mathbb{B}$ and zero matrix \mathbb{O} . These

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are

$$\mathbb{D}_{N}^{9p20} = \begin{pmatrix} -20 & 8 & 0 & 0 & \dots & 0 & 0 & 0\\ 4 & -20 & 4 & 0 & \dots & 0 & 0 & 0\\ 0 & 4 & -20 & 4 & \dots & 0 & 0 & 0\\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots\\ 0 & 0 & 0 & 0 & \dots & 4 & -20 & 4\\ 0 & 0 & 0 & 0 & \dots & 0 & 8 & -20 \end{pmatrix}$$
 (C.24)

and are ordered as follows

$$\mathbb{L}_{N}^{9\text{p}20} = \frac{1}{6h^{2}} \begin{pmatrix} \mathbb{D}_{N}^{9\text{p}20} & \frac{2\mathbb{B}}{\mathbb{B}} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \mathbb{B} & \mathbb{D}_{N}^{9\text{p}20} & \mathbb{B} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \mathbb{O} & \mathbb{B} & \mathbb{D}_{N}^{9\text{p}20} & \mathbb{B} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots \\ \mathbb{O} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{B} & \mathbb{D}_{N}^{9\text{p}20} & \mathbb{B} \\ \mathbb{O} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & 2\mathbb{B} & \mathbb{D}_{N}^{9\text{p}20} \end{pmatrix}$$
(C.26)

In case of general NBC in 9-point stencil the term $2h|\nabla \eta|_{ij}\cos(\theta)$ must be added or subtracted (depending on the boundary) correspondingly to the stencil, e.g. at the left boundary

$$\frac{1}{6h^{2}} \frac{\begin{array}{c|c|c} 1 & 4 & 1 \\ \hline 4 & -20 & 4 \\ \hline 1 & 4 & 1 \end{array}}{\begin{array}{c|c|c} 1 & 6h^{2} \end{array}} \frac{1}{6h^{2}} \begin{bmatrix} \begin{array}{c|c} 0 & 4 & 2 \\ \hline 0 & -20 & 8 \\ \hline 0 & 4 & 2 \end{array} - 2h\cos(\theta)(|\nabla \eta|_{0,j+1} + |\nabla \eta|_{0,j-1} + 4|\nabla \eta|_{0,j}) \\ \hline \\ (C.27)$$

or e.g. at the top boundary

$$\frac{1}{6h^2} \frac{\frac{1}{4} \left| \frac{4}{-20} \right| \frac{1}{4}}{1 \left| \frac{4}{4} \right| 1} \to \frac{1}{6h^2} \left[\frac{0 \left| 0 \right| 0}{\frac{4}{-20} \left| \frac{4}{4} \right|} + 2h\cos(\theta)(|\nabla \eta|_{i+1,M} + |\nabla \eta|_{i-1,M} + 4|\nabla \eta|_{i,M}) \right]$$
(C.28)

C.4.3 Mixed boundary conditions

Imagine the system with BCs as indicated in figure C.3. The action of the finite

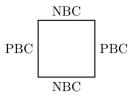


Figure C.3: Layout of the mixed boundaries. NBC and PBC stand for Neumann and periodic boundary conditions, respectively.

difference matrix in the y-direction is given by the structure of the individual block matrices and the x-direction by the arrangement of the blocks in the large matrix.

Specifically, with the periodic BCs present in the x-direction, the structure of the large matrix is as with periodic BCs only the diagonal block matrix. Showing on the example of the 5pt laplacian approximation with mixed BCs as in the figure C.3

$$\mathbb{L}_{PN}^{5p} = \frac{1}{h^{2}} \begin{pmatrix} \mathbb{D}_{N}^{5p} & \mathbb{I} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{I} \\ \mathbb{I} & \mathbb{D}_{N}^{5p} & \mathbb{I} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \mathbb{O} & \mathbb{I} & \mathbb{D}_{N}^{5p} & \mathbb{I} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots \\ \mathbb{O} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{I} & \mathbb{D}_{N}^{5p} & \mathbb{I} \\ \mathbb{I} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{I} & \mathbb{D}_{N}^{5p} \end{pmatrix}$$
(C.29)

In the 9p20 scheme the finite difference matrix is

$$\mathbb{L}_{PN}^{9\text{p}20} = \frac{1}{6h^2} \begin{pmatrix} \mathbb{D}_{N}^{9\text{p}20} & \mathbb{A} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{A} \\ \mathbb{A} & \mathbb{D}_{N}^{9\text{p}20} & \mathbb{A} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \mathbb{O} & \mathbb{A} & \mathbb{D}_{N}^{9\text{p}20} & \mathbb{A} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \vdots & \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots \\ \mathbb{O} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{A} & \mathbb{D}_{N}^{9\text{p}20} & \mathbb{A} \\ \mathbb{A} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{A} & \mathbb{D}_{N}^{9\text{p}20} \end{pmatrix}$$
(C.30)

C.5 Anisotropic driving force terms

In phase field method, usually it is better to keep the interface away from the domain boundary as the boundary conditions often impose specific, non-physical behavior of the interface. In the case of general NBC and the intended use for

the energy assessment, the proximity of interface to the domain boundary is deliberate.

In the implementation of the inclination-dependent model, the anisotropic driving force terms are calculated from numeric derivatives of various fields. Care must be taken to each of the numerical derivative in order to assure the control over the interface inclination at the domain boundary.

When the interface inclination is fixed, the anisotropic driving force terms have fixed value (Dirichlet BC). Some of these fields are numerically differentiated again. However, their values at the ghost nodes are unknown (to be used in centered finite difference), hence one-sided FD scheme must be used at the boundaries. In order to keep second-order accuracy at the boundaries, the

following backward $\begin{bmatrix} 1/2 \\ -2 \end{bmatrix}$ and forward $\begin{bmatrix} -3/2 \\ 2 \\ -1/2 \end{bmatrix}$ FD schemes were used at the

bottom and top boundary, respectively.

For the x-direction derivative the FD matrix is

$$\frac{\partial}{\partial x} \approx \Delta_x = \frac{1}{2h} \begin{pmatrix} -3\mathbb{I} & 4\mathbb{I} & -\mathbb{I} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ -\mathbb{I} & \mathbb{O} & \mathbb{I} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \mathbb{O} & -\mathbb{I} & \mathbb{O} & \mathbb{I} & \dots & \mathbb{O} & \mathbb{O} & \mathbb{O} \\ \vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots & \vdots \\ \mathbb{O} & \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & -\mathbb{I} & \mathbb{O} & \mathbb{I} \\ \mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{I} & -4\mathbb{I} & 3\mathbb{I} \end{pmatrix}. \tag{C.31}$$

For the v-direction derivative the FD matrix is

$$\frac{\partial}{\partial y} \approx \Delta_y = \frac{1}{h} \begin{pmatrix}
\mathbb{Y} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} \\
\mathbb{O} & \mathbb{Y} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{O} \\
\mathbb{O} & \mathbb{O} & \mathbb{Y} & \dots & \mathbb{O} & \mathbb{O} \\
\vdots & \vdots & \vdots & \ddots & \vdots & \vdots \\
\mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{Y} & \mathbb{O} \\
\mathbb{O} & \mathbb{O} & \mathbb{O} & \dots & \mathbb{O} & \mathbb{Y}
\end{pmatrix} \tag{C.32}$$

with

$$\mathbb{Y} = \begin{pmatrix}
-3/2 & 2 & -1/2 & 0 & \dots & 0 & 0 & 0 \\
-1 & 0 & 1 & 0 & \dots & 0 & 0 & 0 \\
0 & -1 & 0 & 1 & \dots & 0 & 0 & 0 \\
\vdots & \vdots & \vdots & \ddots & \vdots & \vdots & \vdots & \vdots \\
0 & 0 & 0 & 0 & \dots & -1 & 0 & 1 \\
0 & 0 & 0 & 0 & \dots & 1/2 & -2 & 3/2
\end{pmatrix}$$
(C.33)

Appendix D

Volume conservation using Lagrange multipliers

D.1 Phase field equations under equality constraintLagrange multipliers

In multivariate calculus, the method of Lagrange multipliers is used to find extrema of a multivariate function under equality constraint, i.e. on a subset of its domain. For example, the extrema of the function f(x,y) over the subset of the domain defined as g(x,y)=0 can be found by analysis of the so-called Lagrange function $\mathcal{L}(x,y,\lambda)=f(x,y)-\lambda g(x,y)$, where λ is the Lagrange multiplier. The critical points are found by solving the set of equations

$$\frac{\partial \mathcal{L}}{\partial x} = 0, \quad \frac{\partial \mathcal{L}}{\partial y} = 0, \quad \frac{\partial \mathcal{L}}{\partial \lambda} = 0.$$
 (D.1)

Similar principles can be generalized in functional calculus. Let's have a functional of free energy $F[\phi] = \int f(\phi, \nabla \phi) dV$ with equality constraint $G[\phi] =$

 $\int g(\phi) dV = 0$. The Lagrangian now takes the form

$$\Lambda[\phi, \lambda] = F[\phi] - \lambda G[\phi] \tag{D.2}$$

$$= \int f(\phi, \nabla \phi) - \lambda g(\phi) dV$$
 (D.3)

$$= \int l(\phi) dV, \qquad (D.4)$$

where $l(\phi)$ is the lagrangian energy density and $\lambda(t)$ is the Lagrange multiplier. It is a function of time, because its value is related to the state of the evolving system. When the phase field $\phi(\mathbf{r})$ evolves due to governing equations

$$\frac{\partial \phi}{\partial t} = \frac{\delta \Lambda}{\delta \phi} \,, \tag{D.5}$$

the equality constraint $G[\phi] = \int g(\phi) dV = 0$ will hold throughout the simulation. The constraint can e.g. impose some conservation law ¹, or enforce local sum of values of phase fields in a multi-phase field model like in all multi-phase field models in the legacy of [58].

The governing equations then read

$$\frac{\delta \Lambda}{\delta \phi} = \frac{\delta F}{\delta \phi} - \lambda(t) \frac{\partial g}{\partial \phi} \,, \tag{D.6}$$

where it was assumed that the constraint is only functional of the phase field, not of its gradient. It can be seen that an additional term appeared, originating from the constraint.

Multiple constraints $G_1[\phi] = 0$, $G_2[\phi] = 0$, ..., $G_p[\phi] = 0$ can be imposed on the system, each of which has its Lagrange multiplier $\lambda_1, \lambda_2, \ldots, \lambda_p$. The Lagrangian is then

$$\Lambda[\phi, \lambda_1, \lambda_2, \dots, \lambda_p] = F[\phi] - \sum_i \lambda_i G_i[\phi], \qquad (D.7)$$

which leads to constrained governing equation

$$\frac{\delta\Lambda}{\delta\phi} = \frac{\delta F}{\delta\phi} - \sum_{i} \lambda_{i}(t) \frac{\partial g_{i}}{\partial\phi} , \qquad (D.8)$$

where $g_i(\phi)$ is the kernel of *i*-th constraint $G_i[\phi] = \int g_i(\phi) dV$.

 $^{^1}$ typically volume, but e.g. in [21, 22] electric charge and behavior of other physical quantities were controlled by Lagrange multipliers

D.2 Lagrange multipliers for volume conservation in multi-phase field models

This approach allows to mathematically enforce certain behavior of the evolving fields. Volume conservation in Allen-Cahn equation is a typical application.

Moelans' model is a multi phase field model with non-conserved fields and with curvature-driven interfaces motion. Inspired by [47], an extension of this model is hereby proposed, which allows to conserve volume of selected phase fields.

The model by Nestler et al. [47] is an extension of a phase field model which differs from the Moelans' one is several aspects. The main differnce is that the sum of all the phase fields locally is equal to 1, which allows to interpret them as order parameters in the strict mathematical sense. However, this comes with a cost - this model property is imposed using a suitable Lagrange multiplier, which introduces new terms in the governing equation(s). Moreover, every model extension must be such that this property is preserved. Nestler's model also differs in the double obstacle potential and in the choice of the non-local gradient term.

D.2.1 Concept - single conserved phase field

The simulated system over the domain Ω consists of n phase fields $\vec{\eta} = (\eta_1, \dots, \eta_n)$ and is described by the free energy functional F with free energy density $f(\vec{\eta}, \nabla \vec{\eta})$ being also function of the phase fields gradients $\nabla \vec{\eta} = (\nabla \eta_1, \dots, \nabla \eta_n)$ as is usual in phase field method

$$F = \int_{\Omega} f(\vec{\eta}, \nabla \vec{\eta}) dV$$
 (D.9)

The volume V_i occupied by i-th grain η_i in the domain Ω can be expressed in Moelans' model as

$$V_i = \int_{\Omega} \Phi_i(\vec{\eta}) dV, \qquad (D.10)$$

where $\Phi_i(\vec{\eta})$ is the *i*-th grain expressed using the special interpolation function [45]

$$\Phi_i(\vec{\eta}) = \frac{\eta_i^2}{\sum_k \eta_k^2} \,. \tag{D.11}$$

The equality constraint G=0 which conserves volume of η_i can be formulated as

$$G = \int \Phi_i(\vec{\eta}) dV - V_i^0, \qquad (D.12)$$

where V_i^0 is the initial volume of the phase field. Because the constant V_i^0 plays no role in the governing equation, it will be disregarded. The volume conservation is ensured by deriving the governing equations from Lagrangian

$$\Lambda[\vec{\eta}] = \int_{\Omega} f(\vec{\eta}, \nabla \vec{\eta}) - \lambda(t) \Phi_i(\vec{\eta}) dV, \qquad (D.13)$$

leading to equation for any p-th phase field

$$\frac{\partial \eta_p}{\partial t} = \frac{\delta \Lambda}{\delta \eta_p} \tag{D.14}$$

$$= \frac{\delta F}{\delta \eta_p} - \lambda(t) \frac{\partial \Phi_i}{\partial \eta_p} \,. \tag{D.15}$$

Expression for the Lagrange multiplier $\lambda(t)$ is obtained from the constraint itself, as the volume conservation of η_i implies

$$\frac{\mathrm{d}V_i}{\mathrm{d}t} = 0 = \frac{\partial}{\partial t} \int \Phi_i(\vec{\eta}) \mathrm{d}V \tag{D.16}$$

$$= \int \frac{\partial \Phi_i}{\partial t} dV \tag{D.17}$$

$$= \int \sum_{k=1}^{n} \frac{\partial \Phi_{i}}{\partial \eta_{k}} \frac{\partial \eta_{k}}{\partial t} dV$$
 (D.18)

$$= -L \int \sum_{k=1}^{n} \frac{\partial \Phi_{i}}{\partial \eta_{k}} \left[\frac{\delta F}{\delta \eta_{k}} - \lambda(t) \frac{\partial \Phi_{i}}{\partial \eta_{k}} \right] dV.$$
 (D.19)

In the last equation, $\lambda(t)$ can already be taken out of the integral as it is not function of space, allowing to write

$$\lambda(t) = \frac{\int \sum_{k=1}^{n} \frac{\partial \Phi_i}{\partial \eta_k} \frac{\delta F}{\delta \eta_p} dV}{\int \sum_{k=1}^{n} \left(\frac{\partial \Phi_i}{\partial \eta_k}\right)^2 dV}.$$
 (D.20)

As can be seen, the expression is notably more complicated than in the original model by Nestler. One reason is that the interpolation function $\Phi(\vec{\eta})$ is a function of all the phase fields, whereas in Nestler the interpolation function $p_k = p_k(\eta_k)$ only. The second reason is that in Nestler's model the volume evolution can be expressed as $\mathrm{d}V_i/\mathrm{d}t = \partial \eta_i/\partial t$. In comparison to Nestler's model, the first complication adds the sum over all phase fields to the integrals and the second adds the term $\partial \Phi_i/\partial \eta_k$.

Hereby the case when all the phase fields are conserved is discussed. With n phase fields that implies n-1 constraints $G_i=\int \Phi_i(\vec{\eta})\mathrm{d}V=0$ as the n-th phase field must be conserved as well when all the others are. Without the loss of generality it is assumed that the n-th phase field does not have its constraint. The Lagrangian is

$$\Lambda[\vec{\eta}] = \int_{\Omega} f(\vec{\eta}, \nabla \vec{\eta}) - \sum_{l=1}^{n-1} \lambda_l(t) \Phi_l(\vec{\eta}) dV, \qquad (D.21)$$

with $\lambda_l(t)$ being the Lagrange multiplier corresponding to l-th phase field.

Each governing equation can then be written in the form

$$\frac{\partial \eta_p}{\partial t} = \frac{\delta \Lambda}{\delta \eta_p} \tag{D.22}$$

$$= \frac{\delta F}{\delta \eta_p} - \sum_{l=1}^{n-1} \lambda_l(t) \frac{\partial \Phi_l}{\partial \eta_p} \,. \tag{D.23}$$

The individual Lagrange multipliers are determined from equation analogical to (D.16), i.e. for arbitrary i

$$\frac{\mathrm{d}V_i}{\mathrm{d}t} = 0 = \int \sum_{k=1}^{n} \frac{\partial \Phi_i}{\partial \eta_k} \frac{\partial \eta_k}{\partial t} \mathrm{d}V$$
 (D.24)

$$= -L \int \sum_{k=1}^{n} \frac{\partial \Phi_{i}}{\partial \eta_{k}} \left[\frac{\delta F}{\delta \eta_{k}} - \sum_{l=1}^{n-1} \lambda_{l}(t) \frac{\partial \Phi_{l}}{\partial \eta_{k}} \right] dV.$$
 (D.25)

This equation can be rearranged to give

$$\int \sum_{k=1}^{n} \frac{\partial \Phi_{i}}{\partial \eta_{k}} \frac{\delta F}{\delta \eta_{k}} dV = \sum_{l=1}^{n-1} \lambda_{l}(t) \int \sum_{k=1}^{n} \frac{\partial \Phi_{i}}{\partial \eta_{k}} \frac{\partial \Phi_{l}}{\partial \eta_{k}} dV, \qquad (D.26)$$

which is a linear equation for n-1 variables $\lambda_l(t)$. Because such equation can be written for every i of the n-1 constrained phase fields, it is a system of linear algebraic equations which can be compactly described using the following notation

$$b_{i} = \int \sum_{k=1}^{n} \frac{\partial \Phi_{i}}{\partial \eta_{k}} \frac{\delta F}{\delta \eta_{k}} dV$$
 (D.27)

$$A_{il} = \int \sum_{k=1}^{n} \frac{\partial \Phi_i}{\partial \eta_k} \frac{\partial \Phi_l}{\partial \eta_k} dV, \qquad (D.28)$$

where b_i is a component of a vector \boldsymbol{b} and A_{il} is a component of the matrix \boldsymbol{A} . Denoting the vector with variables $\boldsymbol{\lambda}$, the system of n-1 equations for n-1 variables can be written

$$A\lambda = b. (D.29)$$

Evaluation of this system is computationally very expensive, because in every component of both \boldsymbol{b} and \boldsymbol{A} there is the sum over all n phase fields. Moreover, it must be evaluated at every time step and it is expected, that all the components of \boldsymbol{A} are non-zero. Even though the matrix is symmetric, the model does not seem to be practical for use in simulations with many conserved grains.

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