

Background

I'm using the wind scalings from Muratov et al. [2015], with modifications from Davé et al. [2016] (the MUFASA simulations). I need a wind velocity (v_w) and mass-loading factor (η) for the parameterized model.

The FIRE wind velocity scalings follow,

$$v_w = 0.854v_c^{1.12}, \quad (1)$$

where v_c is the circular velocity of the galaxy. We use the results of Mo et al. [1998] to calculate v_c ,

$$v_c = (M_b/102.329M_\odot)^{0.26178}(H(z)/H_0)^{1/3}, \quad (2)$$

where M_b is the baryonic mass of the galaxy, and $H(z)$ is the Hubble function. I end up with an average circular velocity of $v_c \approx 327$ km/s. Using the average v_c , I calculate an average wind velocity of $v_w \approx 538$ km/s.

Next, I need a mass loading factor, η . The FIRE scalings give,

$$\eta = 3.55 \left(\frac{M_*}{10^{10}M_\odot} \right)^{-0.351}, \quad (3)$$

where M_* is the stellar mass of the galaxy. I calculate an average mass-loading factor of $\eta \approx 3.13$ using my estimated stellar masses.

For the recoupling time, I follow Davé et al. [2016] and use 2% of the Hubble time at launch. This works out to $t_{\text{recouple}} \approx 4.21$ Myr.

Summarized Results

$$\eta = 3.13$$

$$v_w = 538 \text{ km/s}$$

$$t_{\text{recouple}} = 4.21 \text{ Myr}$$

References

- Alexander L Muratov, Dušan Kereš, Claude-André Faucher-Giguère, Philip F Hopkins, Eliot Quataert, and Norman Murray. Gusty, gaseous flows of fire: galactic winds in cosmological simulations with explicit stellar feedback. *Monthly Notices of the Royal Astronomical Society*, 454(3):2691–2713, 2015.
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- HJ Mo, Shude Mao, and Simon DM White. The formation of galactic discs. *Monthly Notices of the Royal Astronomical Society*, 295(2):319–336, 1998.