A Parametric Study of the SASI Comparing General Relativistic and Non-Relativistic Treatments

Samuel J. Dunham

August 30, 2023

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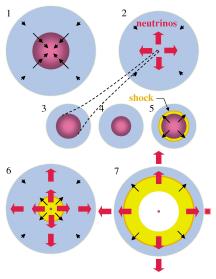
Samuel J. Dunham , ¹ Eirik Endeve ¹ Anthony Mezzacappa ¹ John M. Blondin ¹ Jesse Buffaloe ¹ Para ¹ John M. Blondin ¹ Jesse Buffaloe ¹ Para ¹ John M. Blondin ¹ Jesse Buffaloe ¹ Para ¹ John M. Blondin ¹ Jesse Buffaloe ¹ Para ¹ John M. Blondin ¹ Jesse Buffaloe ¹ Para ¹ John M. Blondin ¹ Para ¹ Par AND KELLY HOLLEY-BOCKELMANN 101,5

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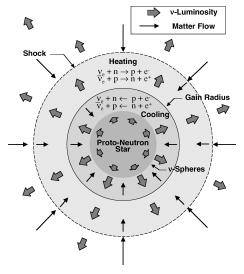
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Overview



Mezzacappa A. 2005. Annu. Rev. Nucl. Part. Sci. 55:467–515



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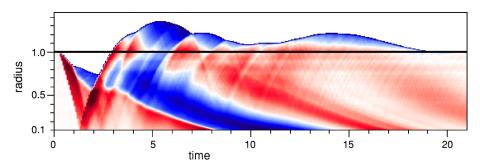


Figure: Equilibrium (white), under- (blue), and over- (red) pressure (?).

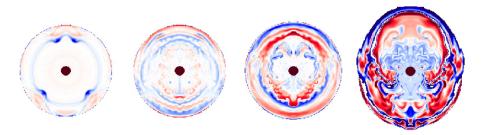
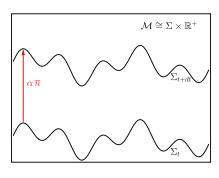


Figure: Equilibrium (white), under- (blue), and over- (red) entropy (?).

$d{+}1$ Decomposition

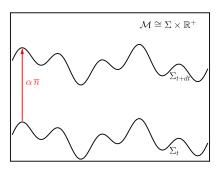
$$ds^2 = g_{\mu\nu}\, dx^\mu\, dx^\nu$$



$d{+}1$ Decomposition

$$ds^2 = g_{\mu\nu} \, dx^\mu \, dx^\nu$$

 $\underline{\underline{g}} \colon \operatorname{spacetime} \operatorname{metric} \operatorname{on} \mathcal{M}$

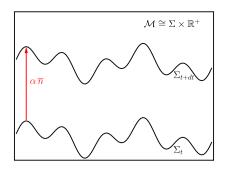


d+1 Decomposition

$$ds^2 = g_{\mu\nu} \, dx^\mu \, dx^\nu$$

 $\underline{\underline{g}}$: spacetime metric on $\mathcal M$

$$g_{\mu\nu} = \begin{pmatrix} -\alpha^2 + \beta^k \beta_k & \beta_j \\ \beta_i & \gamma_{ij} \end{pmatrix}$$



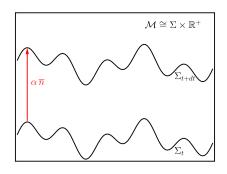
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d+1 Decomposition

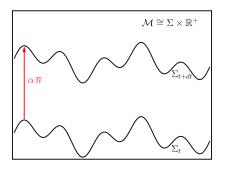
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 $0 < \alpha \le 1$: Lapse Function



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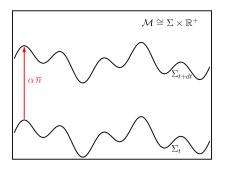
$$g_{\mu\nu} = \begin{pmatrix} -\alpha^2 + \beta^k \beta_k & \beta_j \\ \beta_i & \gamma_{ij} \end{pmatrix}$$

 $\label{eq:spatial_problem} \underline{\underline{\gamma}} \colon \text{spatial three-metric on } \Sigma_t$

 $0 < \alpha \le 1$: Lapse Function

 \overline{n} : Eulerian four-velocity

$$\left(\underline{g}\left(\overline{n},\overline{n}\right) = \overline{n} \cdot \overline{n} = -1\right)$$



d+1 Decomposition

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 $\underline{g} \colon$ spacetime metric on \mathcal{M}

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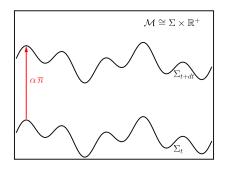
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$$\left(\underline{g}\left(\overline{n},\overline{n}\right) = \overline{n} \cdot \overline{n} = -1\right)$$

(Also, \underline{K} : Extrinsic curvature)



Conformally-Flat Condition

??

$$\gamma_{ij} = \psi^4 \, \bar{\gamma}_{ij}$$

 $1 \le \psi < 2$: Conformal factor

$$\bar{\gamma}_{ij} = \operatorname{diag}\left(1, r^2, r^2 \sin^2 \theta\right)$$

??

$$\gamma_{ij}=\psi^4\, \bar{\gamma}_{ij}$$

$$1\leq \psi <2 \colon {\sf Conformal\ factor}$$
 $\bar{\gamma}_{ij}={\rm diag}\left(1,r^2,r^2\sin^2\theta
ight)$ (Also, $K:={\rm Tr}_{\gamma_{ij}}\left(\underline{K}
ight)=\partial_t\,K=0$)

Isotropic Coordinates (
$$c = G = 1$$
)

$$\alpha\left(r\right) = \frac{1 - r_{\mathrm{Sc}}/r}{1 + r_{\mathrm{Sc}}/r}$$

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$$r > r_{Sc} := M/2 \quad \left(= GM/\left(2c^2\right)\right)$$

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$$\beta^i = 0$$

$$K_{ij} = 0$$

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 $\overline{\nabla} \cdot \overline{\overline{T}} = \overline{0} \hspace{0.5cm} (\overline{\overline{T}}: \hspace{0.1cm} \mathsf{Rank} \hspace{0.1cm} (2,0) \hspace{0.1cm} \mathsf{stress\text{-}energy} \hspace{0.1cm} \mathsf{tensor})$

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$$\overline{J} := \rho \, \overline{u} \quad (\rho: \text{ comoving baryon mass density})$$

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$$\overline{J}:=
ho\,\overline{u} \quad \left(
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$$\overline{\overline{T}}:=
ho\,h\,\overline{u}\otimes\overline{u}+p\,\overline{\overline{g}}$$
 (p : comoving thermal pressure,

 $h:=1+\left(e+p\right) /\rho$: specific enthalpy, e: comoving internal energy density)

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Five equations with six unknowns ©

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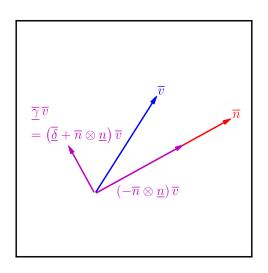
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Five equations with six unknowns ©

Close with an equation of state: $p = p\left(e\right) := \left(\Gamma - 1\right)e$, $\Gamma = 4/3$



Extensible to higher-rank tensors!

$$E:=n_{\mu'}\,n_{\nu'}\,T^{\mu'\nu'}$$

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$$S^{\mu} := - \gamma^{\mu}_{\ \mu'} \, n_{\nu'} \, T^{\mu'\nu'}$$

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Math...

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_i \left[\alpha \sqrt{\gamma} \mathbf{F}^i (\mathbf{U}) \right] = \mathbf{S} (\mathbf{U})$$

GR

$$\partial_{t} \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_{i} \left[\alpha \sqrt{\gamma} \mathbf{F}^{i} \left(\mathbf{U} \right) \right] = \mathbf{S} \left(\mathbf{U} \right)$$

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$$\boldsymbol{U} := (D, S_j, \tau)^\top$$

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$$D := \rho W$$

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GR NR

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$$U := \left(D_{\mathrm{NR}}, S_j^{\mathrm{NR}}, \tau_{\mathrm{NR}}\right)^{\mathsf{T}}$$

$$\partial_{t} \boldsymbol{U} + \frac{1}{\sqrt{\gamma}} \partial_{i} \left[\alpha \sqrt{\gamma} \, \boldsymbol{F}^{i} \left(\boldsymbol{U} \right) \right] = \boldsymbol{S} \left(\boldsymbol{U} \right)$$

$$\begin{aligned} \boldsymbol{U} &:= (D, S_{j}, \tau)^{\top} \\ D &:= \rho W \\ S_{j} &:= \rho h W^{2} v_{j} \\ \tau &:= E - D = \rho h W^{2} - p - \rho W \\ \alpha &= (1 - r_{Sc}/r) / (1 + r_{Sc}/r) \end{aligned}$$

NR

$$U := \left(D_{NR}, S_j^{NR}, \tau_{NR}\right)^{\top}$$
$$D_{NR} := \rho$$

 $\sqrt{\gamma} = \psi^6 \sqrt{\bar{\gamma}}$

$$\partial_{t} \boldsymbol{U} + \frac{1}{\sqrt{\gamma}} \partial_{i} \left[\alpha \sqrt{\gamma} \boldsymbol{F}^{i} \left(\boldsymbol{U} \right) \right] = \boldsymbol{S} \left(\boldsymbol{U} \right)$$

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$$\begin{split} & \text{NR} \\ & \boldsymbol{U} := \left(D_{\text{NR}}, S_j^{\text{NR}}, \tau_{\text{NR}}\right)^\top \\ & D_{\text{NR}} := \rho \\ & S_j^{\text{NR}} := \rho \, v_j \\ & \tau_{\text{NR}} = e + \frac{1}{2} \, \rho \, v^i \, v_i \end{split}$$

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$$\begin{aligned} & \boldsymbol{U} := (D, S_j, \tau)^\top \\ & \boldsymbol{D} := \rho \, W \\ & S_j := \rho \, h \, W^2 \, v_j \\ & \tau := E - D = \rho \, h \, W^2 - p - \rho \, W \\ & \alpha = \left(1 - r_{\mathrm{Sc}}/r\right) / \left(1 + r_{\mathrm{Sc}}/r\right) \\ & \sqrt{\gamma} = \psi^6 \, \sqrt{\bar{\gamma}} \end{aligned}$$

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$$\partial_{t} \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_{i} \left[\alpha \sqrt{\gamma} \mathbf{F}^{i} \left(\mathbf{U} \right) \right] = \mathbf{S} \left(\mathbf{U} \right)$$

GR
$$F^{i}(U) = \begin{pmatrix} \rho W v^{i} \\ \rho h W^{2} v^{i} v_{j} + p \delta^{i}_{j} \\ (\rho h W^{2} - \rho W) v^{i} \end{pmatrix}$$

NR
$$\mathbf{F}^{i}(\mathbf{U}) = \begin{pmatrix} \rho v^{i} \\ \rho v^{i} v_{j} + p \delta^{i}_{j} \\ (\rho h_{NR} + \frac{1}{2} v^{j} v_{j}) v^{i} \end{pmatrix}$$

$$\partial_t \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_i \left[\alpha \sqrt{\gamma} \mathbf{F}^i \left(\mathbf{U} \right) \right] = \mathbf{S} \left(\mathbf{U} \right)$$

GR

$$S(U) = \begin{pmatrix} 0 \\ \frac{1}{2} \alpha P^{ik} \partial_j \gamma_{ik} - E \partial_j \alpha \\ -S^j \partial_j \alpha \end{pmatrix}$$

$$S\left(U\right) = \begin{pmatrix} 0 \\ \frac{1}{2} P^{ik} \partial_{j} \bar{\gamma}_{ik} - \rho \partial_{j} \Phi \\ -S^{j} \partial_{j} \Phi \end{pmatrix}$$

$$\Phi\left(r\right) := -M/r$$

$$\partial_{t} \mathbf{U} + \frac{1}{\sqrt{\gamma}} \partial_{r} \left[\alpha \sqrt{\gamma} \mathbf{F}^{r} \left(\mathbf{U} \right) \right] = \mathbf{S} \left(\mathbf{U} \right)$$

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$$\alpha\,\psi^6\,W\times 4\pi\,r^2\,\rho\,v=-\dot{M}$$

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$$\alpha \, \psi^6 \, W \times 4\pi \, r^2 \, \rho \, v = -\dot{M}$$
$$\alpha \, h \, W = \mathcal{B}$$

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$$p = K_1 \rho^{\Gamma}$$

$$\partial_{t} \boldsymbol{\mathcal{U}} + \frac{1}{\sqrt{\gamma}} \partial_{r} \left[\alpha \sqrt{\gamma} \, \boldsymbol{F}^{r} \left(\boldsymbol{U} \right) \right] = \boldsymbol{S} \left(\boldsymbol{U} \right)$$

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$$4\pi r^2 \rho v = -\dot{M}$$

$$\alpha \, h \, W = \mathcal{B}$$

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$$\alpha \psi^{6} W \times 4\pi r^{2} \rho v = -\dot{M}$$

$$\alpha h W = \mathcal{B}$$

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$$4\pi r^2 \rho v = -\dot{M}$$

$$\frac{1}{2}v^2 + h_{NR} + \Phi = \mathcal{B}_{NR}$$

$$\partial_{t} \boldsymbol{U} + \frac{1}{\sqrt{\gamma}} \partial_{r} \left[\alpha \sqrt{\gamma} \, \boldsymbol{F}^{r} \left(\boldsymbol{U} \right) \right] = \boldsymbol{S} \left(\boldsymbol{U} \right)$$

$$\alpha \psi^6 W \times 4\pi r^2 \rho v = -\dot{M}$$

$$\alpha h W = \mathcal{B}$$

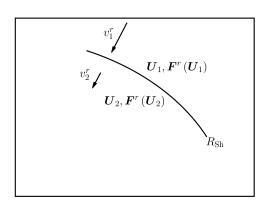
$$p = K_1 \rho^{\Gamma}$$

$$4\pi r^2 \rho v = -\dot{M}$$

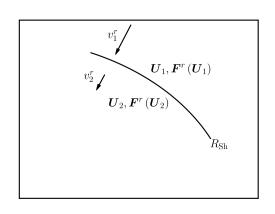
$$\frac{1}{2}v^2 + h_{\rm NR} + \Phi = \mathcal{B}_{\rm NR}$$

$$p = K_1 \rho^{\Gamma}$$

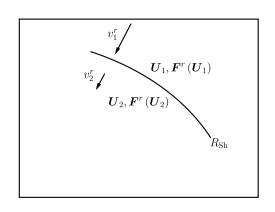
$$U_1 \neq U_2$$



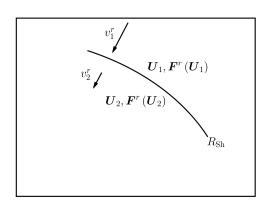
$$egin{aligned} oldsymbol{U}_1
eq oldsymbol{U}_2 \\ oldsymbol{F}^r\left(oldsymbol{U}_1
ight) &= oldsymbol{F}^r\left(oldsymbol{U}_2
ight) \end{aligned}$$



$$egin{aligned} oldsymbol{U}_1
eq oldsymbol{U}_2 \ oldsymbol{F}^r\left(oldsymbol{U}_1
ight) = oldsymbol{F}^r\left(oldsymbol{U}_2
ight) \ ext{Yields:} \end{aligned}$$

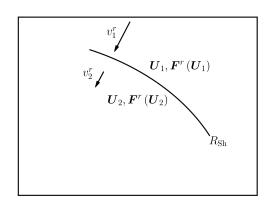


$$egin{aligned} oldsymbol{U}_1
eq oldsymbol{U}_2 \ oldsymbol{F}^r\left(oldsymbol{U}_1
ight) = oldsymbol{F}^r\left(oldsymbol{U}_2
ight) \ ext{Yields:} \
ho_2 >
ho_1 \end{aligned}$$

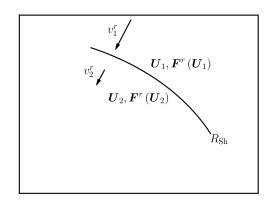


$$egin{aligned} oldsymbol{U}_1
eq oldsymbol{U}_2 \ oldsymbol{F}^r\left(oldsymbol{U}_1
ight) = oldsymbol{F}^r\left(oldsymbol{U}_2
ight) \ ext{Yields:} \
ho_2 >
ho_1 \end{aligned}$$

 $e_2(p_2) > e_1(p_1)$

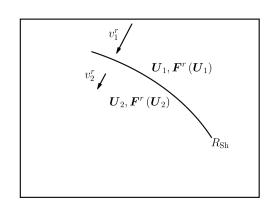


$$egin{aligned} oldsymbol{U}_1
eq oldsymbol{U}_2 \ oldsymbol{F}^r\left(oldsymbol{U}_1
ight) = oldsymbol{F}^r\left(oldsymbol{U}_2
ight) \ egin{aligned} \operatorname{Yields} : \ & \rho_2 > \rho_1 \ & e_2\left(p_2\right) > e_1\left(p_1\right) \end{aligned}$$



 $K_2 > K_1$

$$egin{aligned} m{U}_1 &
eq m{U}_2 \\ m{F}^r \left(m{U}_1
ight) &
eq m{F}^r \left(m{U}_2
ight) \\ ext{Yields:} \\ \rho_2 & > \rho_1 \\ e_2 \left(p_2
ight) & > e_1 \left(p_1
ight) \\ ext{} K_2 & > K_1 \\ |v_2^r| & < |v_1^r| \end{aligned}$$



$$\partial_{t} \boldsymbol{U} + \frac{1}{\sqrt{\gamma}} \partial_{r} \left[\alpha \sqrt{\gamma} \, \boldsymbol{F}^{r} \left(\boldsymbol{U} \right) \right] = \boldsymbol{S} \left(\boldsymbol{U} \right)$$

$$\alpha \psi^{6} W \times 4\pi r^{2} \rho v = -\dot{M}$$

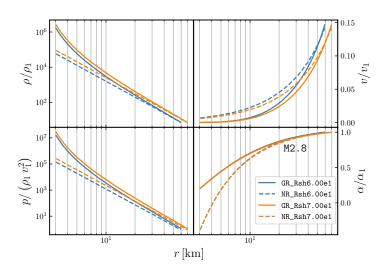
$$\alpha h W = \mathcal{B}$$

$$p = K_{2} \rho^{\Gamma}$$

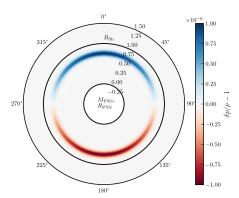
$$4\pi r^2 \rho v = -\dot{M}$$

$$\frac{1}{2}v^2 + h_{NR} + \Phi = \mathcal{B}_{NR}$$

$$p = K_2 \rho^{\Gamma}$$



$$\eta\left(r
ight):=rac{r-R_{\mathrm{PNS}}}{R_{\mathrm{Sh}}-R_{\mathrm{PNS}}}$$

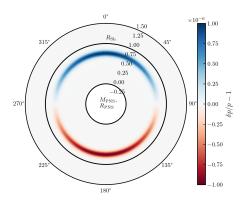


$$\eta(r) := \frac{r - R_{\text{PNS}}}{R_{\text{Sh}} - R_{\text{PNS}}}$$

$$\frac{\delta p(\eta, \theta)}{p(\eta_c)} = 10^{-6} \times \exp\left[\frac{-(\eta - \eta_c)^2}{2\sigma^2}\right] \cos \theta$$

$$\eta_c = 0.75$$

$$\sigma = 0.05$$



Parameter Space

Model parameters:

$$M_{\mathrm{PNS}}$$
, R_{PNS} , R_{Sh} $(t=0)$, \dot{M} , K_{1} , $\mathcal{B}_{\mathrm{(NR)}}$, Γ

Parameter Space

Model parameters:

$$M_{\rm PNS}$$
, $R_{\rm PNS}$, $R_{\rm Sh}$ $(t=0)$, \dot{M} , $K_{\rm 1}$, $\mathcal{B}_{\rm (NR)}$, Γ

Parameters we varied:

$$\xi = (M_{\rm PNS}/M_{\odot}) / (R_{\rm PNS}/20 \, {\rm km})$$
 (?), $R_{\rm Sh} \, (t=0)$

Parameter Space

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Parameters we varied:

$$\xi = (M_{\rm PNS}/M_{\odot}) / (R_{\rm PNS}/20 \, {\rm km})$$
(?), $R_{\rm Sh} (t=0)$

Table 1. Model Parameters

Model	$M_{ ext{PNS}}\left[M_{\odot} ight]$	$R_{ ext{PNS}}\left[ext{km} ight]$	$R_{ m sh}[{ m km}]$	ξ
M1.4_Rpns040_Rsh1.20e2	1.4	40	120	0.7
M1.4_Rpns040_Rsh1.50e2	1.4	40	150	0.7
M1.4_Rpns040_Rsh1.75e2	1.4	40	175	0.7
M2.8_Rpns020_Rsh6.00e1	2.8	20	60	2.8
M2.8_Rpns020_Rsh7.00e1	2.8	20	70	2.8

NOTE—Model parameters chosen for the 5 models. All models were run with both GR and NR. The first three rows correspond to the low-compactness models and the last two rows correspond to the high-compactness models.

$$A := \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left(v^{\theta} \sin \theta \right)$$
 (?)

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 (?)

$$A(r, \theta, t) = \sum_{\ell'=0}^{\infty} G_{\ell'}(r, t) P_{\ell'}(\cos \theta)$$

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 (?)

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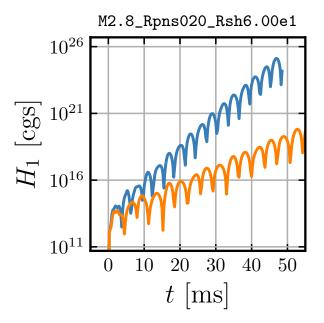
$$\implies G_{\ell}(r,t) := \frac{1}{N_{\ell}} \int_{0}^{\pi} A(r,\theta,t) P_{\ell}(\cos\theta) \sin\theta d\theta$$

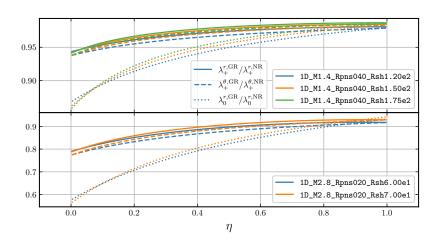
$$A := \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \left(v^{\theta} \sin \theta \right)$$
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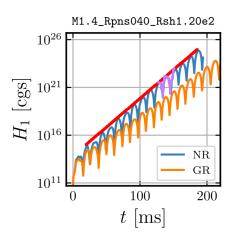
$$H_{\ell}\left(t
ight):=4\pi\int\limits_{r_{a}}^{r_{b}}\left[G_{\ell}\left(r,t
ight)\right]^{2}\left[\psi\left(r
ight)\right]^{6}r^{2}\,dr$$





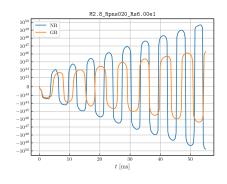
$$F(t) = F(0) e^{2\omega t} \sin^2 \left(\frac{2\pi t}{T} + \delta\right)$$

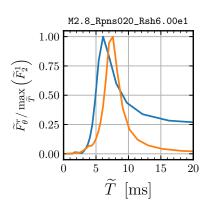
(?)



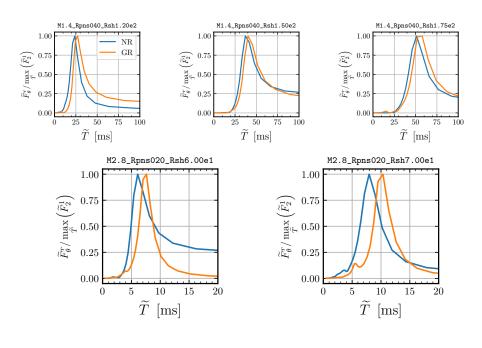
$$F_{\theta}^{r} := \alpha \, \psi^{6} \, h \, W^{2} \times \sqrt{\bar{\gamma}} \, \rho \, v^{r} \, v_{\theta}$$

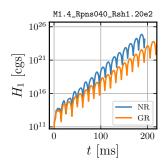
$$\widetilde{F}_{\theta}^{r} = \text{FFT}\left\{F_{\theta}^{r}\right\}$$

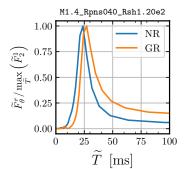


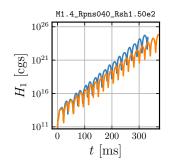


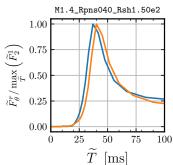
T defined as the unique \widetilde{T} such that $\widetilde{F}_{\theta}^{r}\left(\widetilde{T}\right)=1$

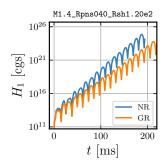


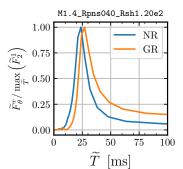


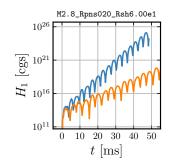


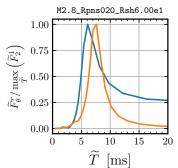


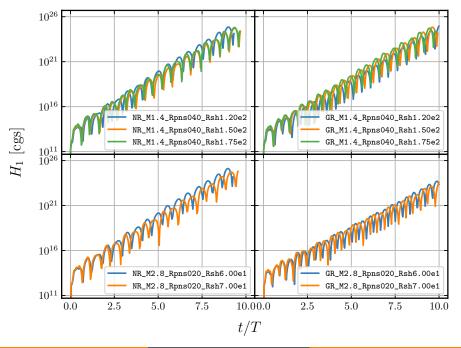


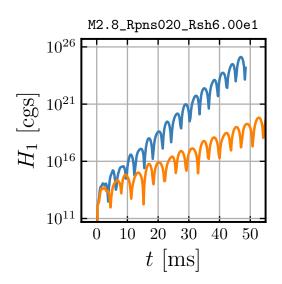


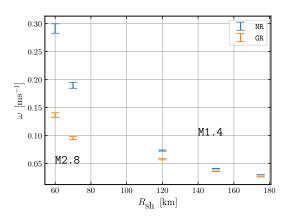












Bibliography

Summary

- Extended study of ? to include GR
- Showed that GR leads to longer SASI oscillation period than NR
- Showed that GR leads to smaller SASI growth rate than NR
- ullet Found that growth rate is such that $\omega\,T$ is roughly constant for some parameter sets: implications for growth rate mechanism
- Future Work
 - Extend study to 3D
 - Include GR monopole (?)