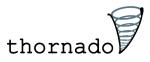
### thornado-Hydro (xCFC)

Samuel J. Dunham

October 21, 2024



#### toolkit for high-order neutrino-radiation hydrodynamics

https://github.com/endeve/thornado

- DG
- SSPRK/IMEX<sup>a</sup>
- GR (xCFC)
- Hydro<sup>b</sup> (Valencia)
- Neutrino transport<sup>c</sup> (M1)
- Interfaces to tabulated EoS/Opacities (weaklib: https://github.com/ starkiller-astro/weaklib)

- Fluid self-gravity via Poseidon: https: //github.com/jrober50/Poseidon
- GPUs via OpenACC or OpenMP pragmas
- MPI parallelism and AMR via AMReX: https: //github.com/AMReX-Codes/amrex

(2019); Pochik et al. (2021)

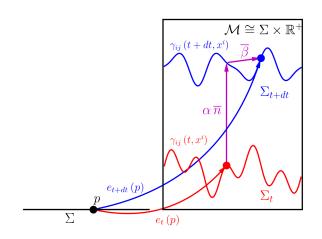
<sup>c</sup>Laiu et al. (2021); Laiu et al. (2025)

<sup>&</sup>lt;sup>a</sup>Chu et al. (2019)

<sup>&</sup>lt;sup>b</sup>Dunham et al. (2020); Endeve et al.

- Capturing diffusion limit for neutrino transport without ad-hoc methods
- High-order accuracy on a compact stencil
- An alternative simulation code than the standard FV/FD methods
- An experiment in applied math: are DG methods effective for CCSN simulations?

#### 3+1 Decomposition



$$ds^{2} = g_{\mu\nu} dx^{\mu} dx^{\nu} = -\alpha^{2} dt^{2} + \gamma_{ij} (dx^{i} + \beta^{i} dt) (dx^{j} + \beta^{j} dt)$$

#### Conformally-Flat Condition

Developed by Wilson et al. (1996), extended by Cordero-Carrión et al. (2009)

$$\gamma_{ij}(x) = \psi^4(x) \, \overline{\gamma}_{ij}(x^i)$$

$$K = 0, \, \partial_t K = 0$$
(Always and everywhere)

- Exact in spherical symmetry!
- Hyperbolic → Elliptic equations
- Good for long-time simulations

Special case: Schwarzchild spacetime in isotropic coordinates (G=c=1)

$$\alpha = \left(1 + \frac{1}{2}\Phi\right) \left(1 - \frac{1}{2}\Phi\right)^{-1}$$

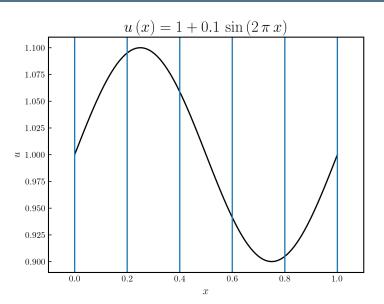
$$\psi = 1 - \frac{1}{2}\Phi$$

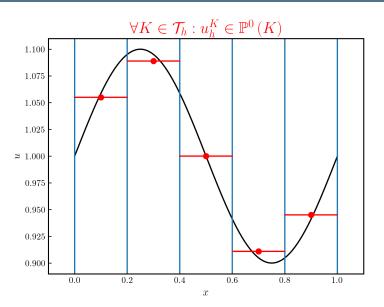
$$\beta^{i} = 0,$$

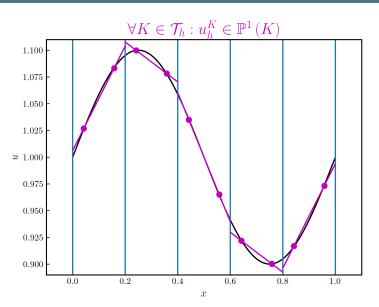
with

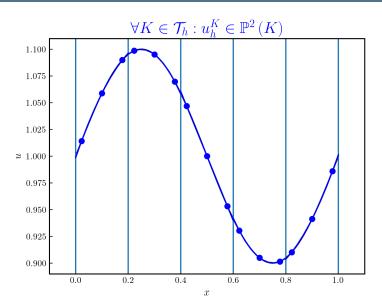
$$\Phi\left(r\right) := -\frac{M}{r}$$

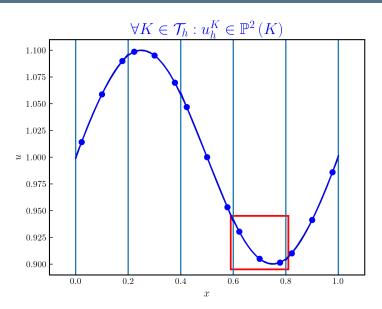
# Discontinuous Galerkin (DG) Method





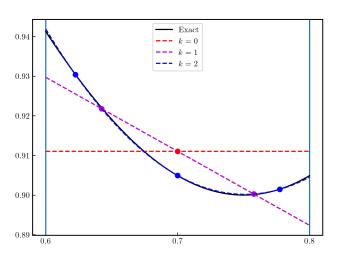






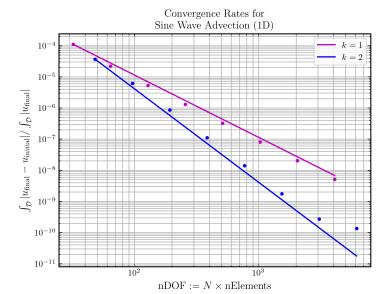
## Discontinuous Galerkin (DG) Method

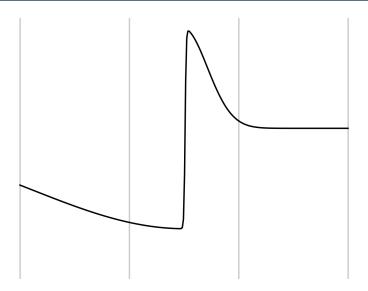
$$u_{h}\left(x,t\right):=\sum_{i=1}^{k+1}u_{i}\left(t\right)\,\ell_{i}\left(x\right)$$



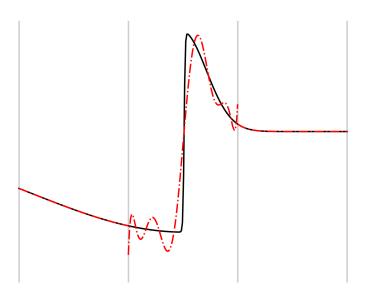
#### Time Stepping

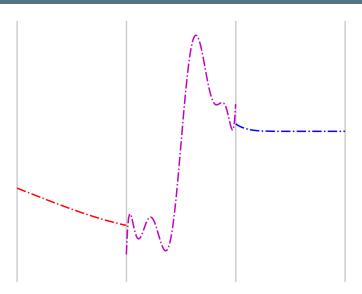
- For hydro-only, strong-stability-preserving Runge-Kutta (SSPRK) methods
  - Convex combinations of forward-Euler timesteps
- For hydro+neutrinos, realizability-preserving ImEx methods (Chu et al. (2019))



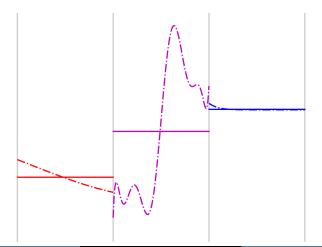


$$k = 9$$



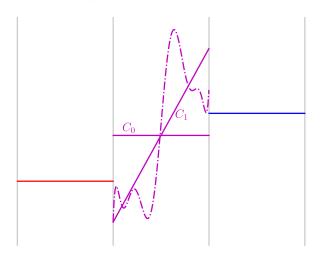


$$\bar{u}\left(t\right) := \frac{1}{V_h\left(t\right)} \int_K u_h\left(x,t\right) \sqrt{\gamma_h}\left(x,t\right) dx, \ V_h\left(t\right) := \int_K \sqrt{\gamma_h}\left(x,t\right) dx$$

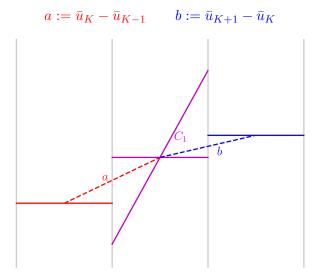


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$$u_{h}\left(x,t\right)=\sum_{i=1}^{k+1}u_{i}\left(t\right)\,\ell_{i}\left(x\right)=\sum_{n=0}^{k}C_{n}\left(t\right)\,P_{n}\left(x\right)$$

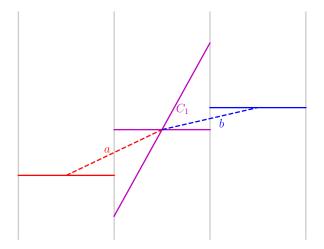


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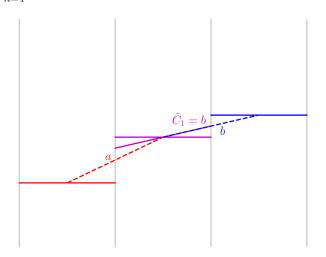
$$\tilde{C}_1 := \min \operatorname{mod}(a, b, C_1),$$

$$\operatorname{minmod}\left(a,b,C_{1}\right):=\begin{cases}\operatorname{sgn}\left(a\right)\times\operatorname{min}\left(\left|a\right|,\left|b\right|,\left|C_{1}\right|\right), & \operatorname{sgn}\left(a\right)=\operatorname{sgn}\left(b\right)=\operatorname{sgn}\left(C_{1}\right)\\ 0, & \operatorname{else}\end{cases}$$

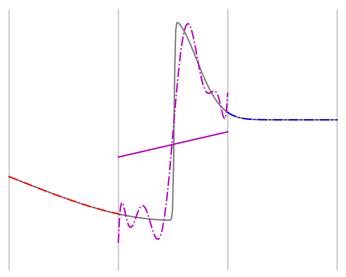


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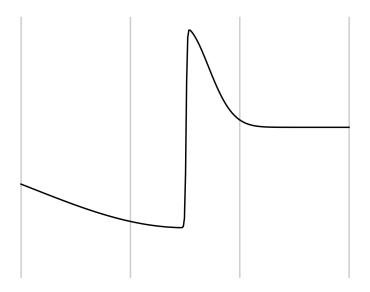
$$\sum_{k=1}^{k} C_n P_n = u_k \to \tilde{u}_k := \Lambda_{SL} (u_k) := C_0 P_0 + \tilde{C}_1 P_1$$



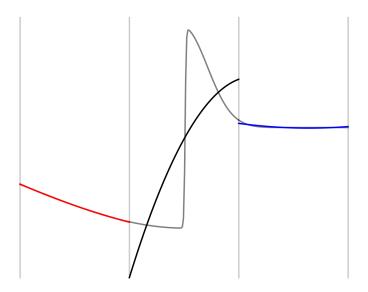
$$k = 9$$
$$\tilde{k} = 1$$

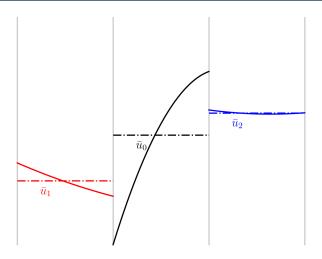


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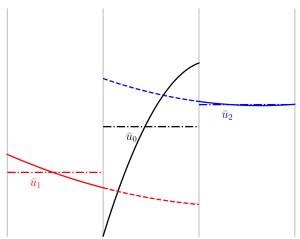


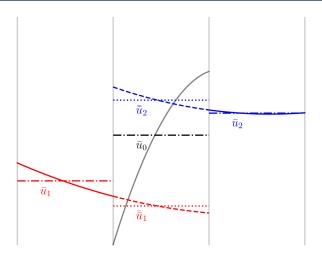




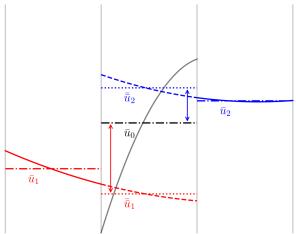


$$u_{h}(x,t) = \sum_{i=1}^{k+1} u_{i} \ell_{i}(x), u_{h}(x,t) = \sum_{i=1}^{k+1} u_{i} \ell_{i}(x)$$



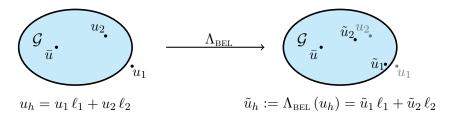


$$I_K := \left( \max_{j \in \{0, \cdots, 2^d\}} (|\bar{u}_j|) \right)^{-1} \sum_{j=1}^{2^d} |\bar{u}_0 - \bar{\bar{u}}_j|$$



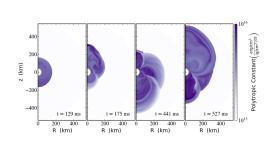
## Bound-Enforcing Limiter $(\Lambda_{ ext{BEL}})$

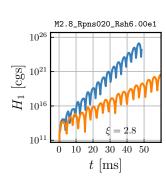
- Proposed in Qin et al. (2016)
- Define  $\mathcal{G}$  as the set of physically-admissible states; i.e., positivite density and energy density, subluminal fluid velocity (for an ideal EoS,  $\mathcal{G}$  is convex)



#### Standing Accretion Shock Instability

Used thornado to investigate the role of GR on the SASI<sup>1</sup>



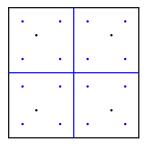


<sup>&</sup>lt;sup>1</sup>Dunham et al. (2020, 2024)



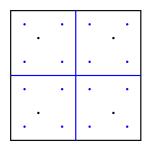
https://github.com/AMReX-Codes/amrex

#### Interpolation



Given 
$$U_h\left(x,t\right) = \sum_{i \in \mathcal{N}} U_i\left(t\right) \, \ell_i\left(x\right)$$
, how to compute  $u_h^{(j)}\left(x,t\right) = \sum_{i \in \mathcal{N}} u_i^{(j)}\left(t\right) \, \ell_i^{(j)}\left(x\right)$ , or vice-versa? ( $\mathcal{N}$  is the set of all grid-points for any given element.)

#### Conservative Interpolation



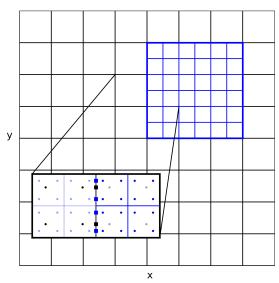
#### • Schaal et al. (2015)

Given a solution on a coarse element K,  $U_h$ , define the solution on the jth  $(j \in \{1, \cdots, 2^d\})$  fine element  $k^{(j)}$ ,  $u_h^{(j)}$  via the minimization of an  $L^2$ -projection:

$$\forall i \in \mathcal{N}: \min_{\tilde{\boldsymbol{u}}_{i}^{(j)}} \int_{k^{(j)}} \left( \tilde{U}_{h} - \tilde{\boldsymbol{u}}_{h}^{(j)} \right)^{2} d^{d}x,$$

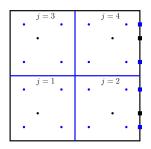
where  $\tilde{U}:=\sqrt{\gamma}\,U.$  For the geometry fields, we use the same procedure but do not pre-multiply with  $\sqrt{\gamma}.$ 

#### Flux Corrections



Samuel J. Dunham

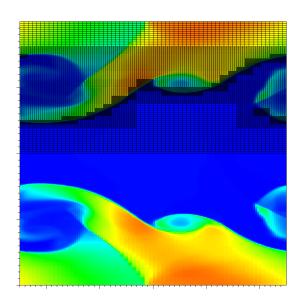
#### Flux Corrections



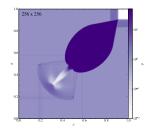
The numerical flux across, say, the  $x^1$  interface on the right side of an element for a 2D problem with no gravity and Cartesian coordinates is given by

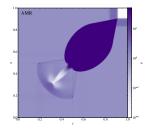
$$\begin{split} & \int_{K^{2}} \widehat{F}^{1} \left( U_{h}^{-}, U_{h}^{+} \right) \, \ell_{k_{1}} \left( x_{\mathrm{H}}^{1} \right) \, \ell_{k_{2}} \left( x^{2} \right) \, dx^{2} \\ & = \sum_{j \in \left\{ 2, 4 \right\}} \int_{k_{(j)}^{2}} \widehat{F}^{1} \left( U_{h}^{-}, U_{h}^{+} \right) \, \ell_{k_{1}} \left( x_{\mathrm{H}}^{1} \right) \, \ell_{k_{2}} \left( x^{2} \right) \, dx^{2} \\ & := \sum_{j \in \left\{ 2, 4 \right\}} \int_{k_{(j)}^{2}} \widehat{F}^{1} \left( u_{h}^{(j), -}, u_{h}^{(j), +} \right) \, \ell_{k_{1}} \left( x_{\mathrm{H}}^{1} \right) \, \ell_{k_{2}} \left( x^{2} \right) \, dx^{2} \end{split}$$

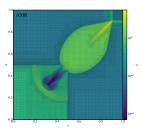
### Kelvin–Helmholtz Instability



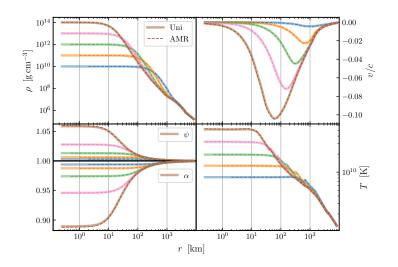
## 2D Blast Wave (Del Zanna & Bucciantini, 2002)



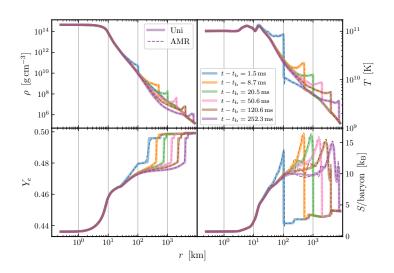




# Adiabatic Collapse (Collapse Phase)



## Adiabatic Collapse (Post-Bounce Phase)



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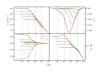
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#### Summary

Can run multi-D pure hydro problems in GR with AMR





Can run 1D hydro+self-gravity problems in GR with AMR

Working on coupling GR transport to existing hydro+gravity modules