

thornado-Hydro (xCFC)

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toolkit for high-order neutrino-radiation hydrodynamics

<https://github.com/endeve/thornado>

- DG
- SSPRK/IMEX^a
- GR (xCFC)
- Hydro^b (Valencia)
- Neutrino transport^c (M1)
- Interfaces to tabulated EoS/Opacities (weaklib: <https://github.com/starkiller-astro/weaklib>)
- Fluid self-gravity via Poseidon: <https://github.com/jrober50/Poseidon>
- GPUs via OpenACC or OpenMP pragmas
- MPI parallelism and AMR via AMReX: <https://github.com/AMReX-Codes/amrex>

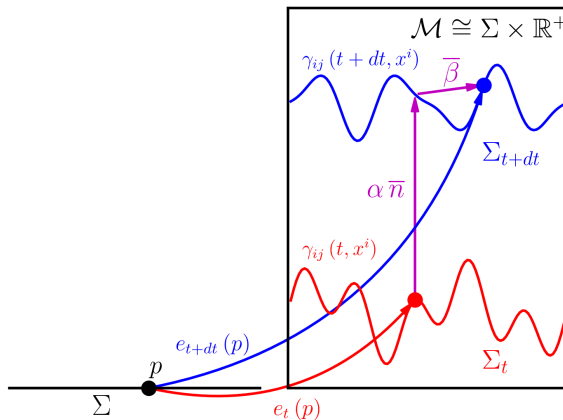
^aChu et al. (2019)

^bDunham et al. (2020); Endeve et al. (2019); Pochik et al. (2021)

^cLaiu et al. (2021); Laiu et al. (2025)

- Capturing diffusion limit for neutrino transport without ad-hoc methods
- High-order accuracy on a compact stencil
- An alternative simulation code than the standard FV/FD methods
- An experiment in applied math: are DG methods effective for CCSN simulations?

3+1 Decomposition



$$ds^2 = g_{\mu\nu} dx^\mu dx^\nu = -\alpha^2 dt^2 + \gamma_{ij} (dx^i + \beta^i dt) (dx^j + \beta^j dt)$$

Conformally-Flat Condition

Developed by Wilson et al. (1996),
extended by Cordero-Carrión et al.
(2009)

Special case: Schwarzschild spacetime
in isotropic coordinates ($G = c = 1$)

$$\gamma_{ij}(x) = \psi^4(x) \bar{\gamma}_{ij}(x^i)$$

$$K = 0, \quad \partial_t K = 0$$

(Always and everywhere)

$$\alpha = \left(1 + \frac{1}{2}\Phi\right) \left(1 - \frac{1}{2}\Phi\right)^{-1}$$

$$\psi = 1 - \frac{1}{2}\Phi$$

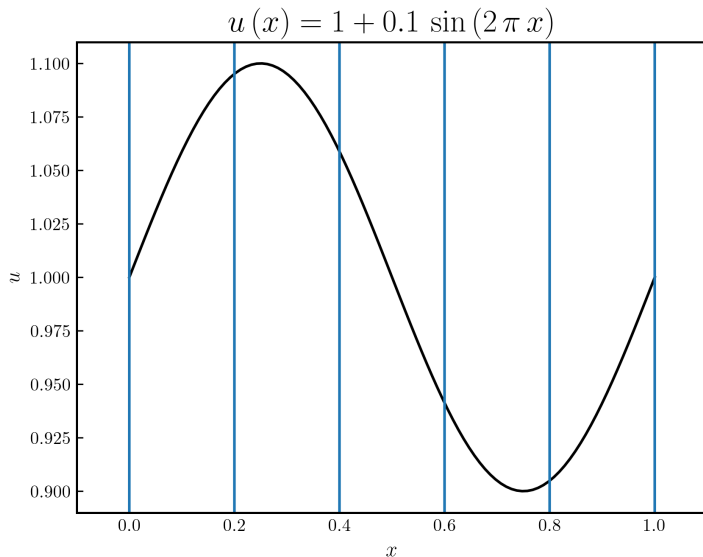
$$\beta^i = 0,$$

- Exact in spherical symmetry!
- Hyperbolic \rightarrow Elliptic equations
- Good for long-time simulations

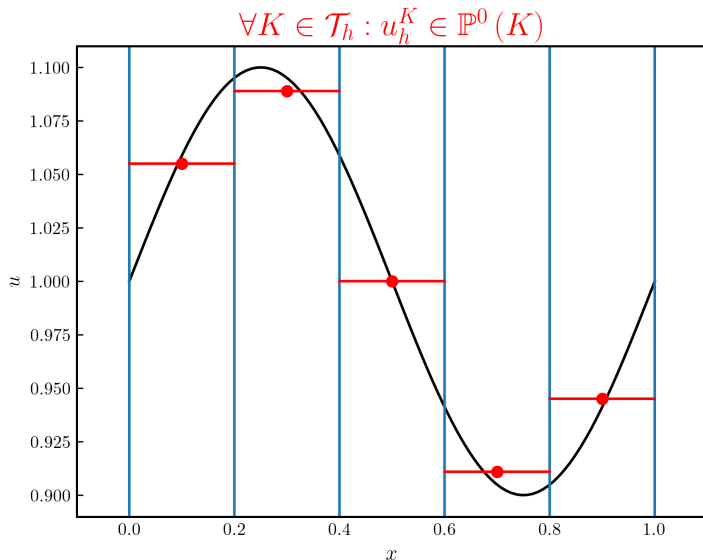
with

$$\Phi(r) := -\frac{M}{r}$$

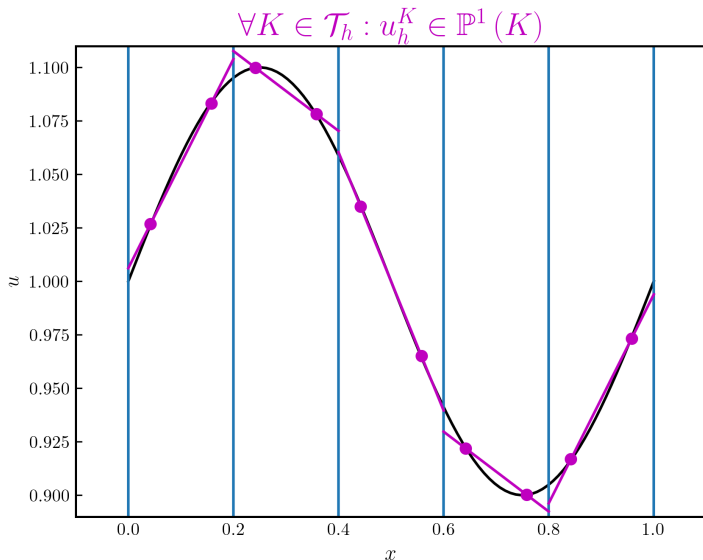
Discontinuous Galerkin (DG) Method



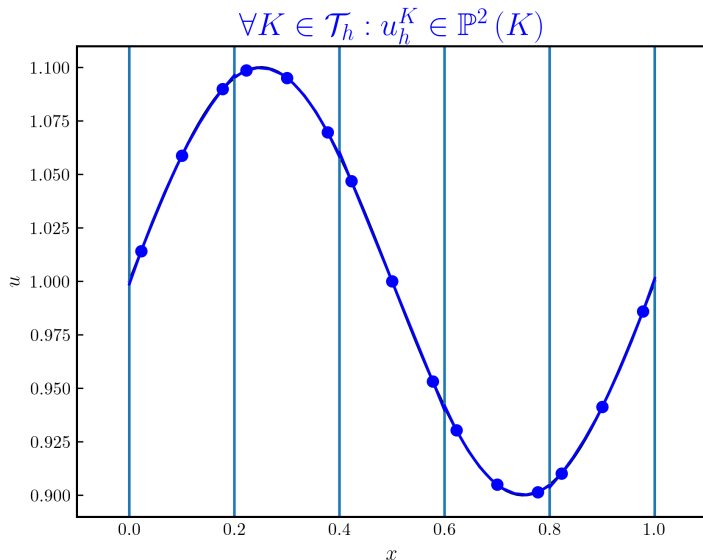
Discontinuous Galerkin (DG) Method



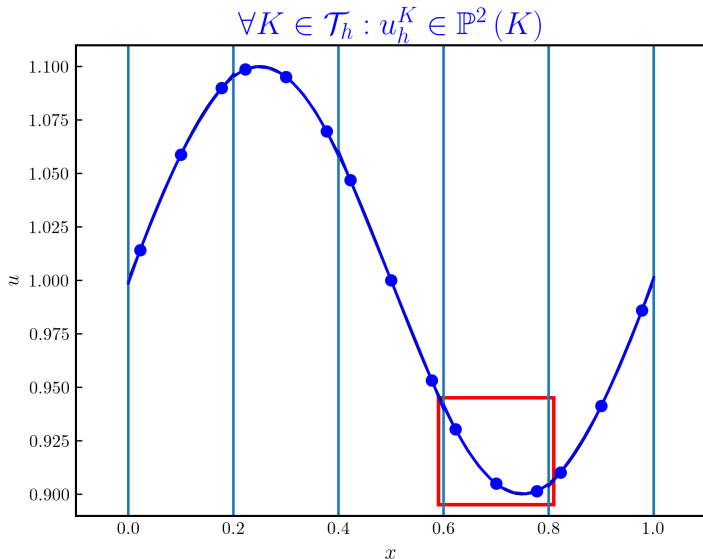
Discontinuous Galerkin (DG) Method



Discontinuous Galerkin (DG) Method

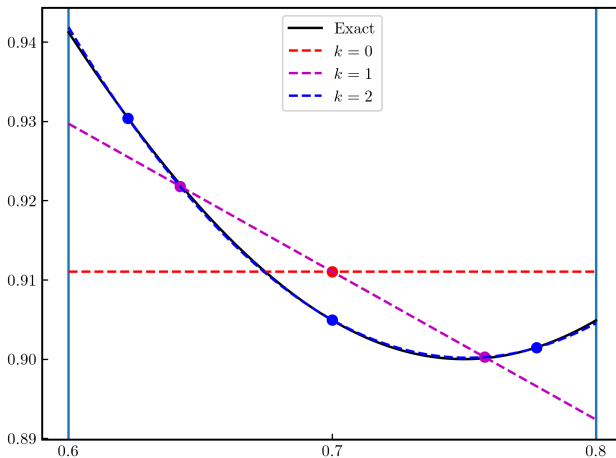


Discontinuous Galerkin (DG) Method



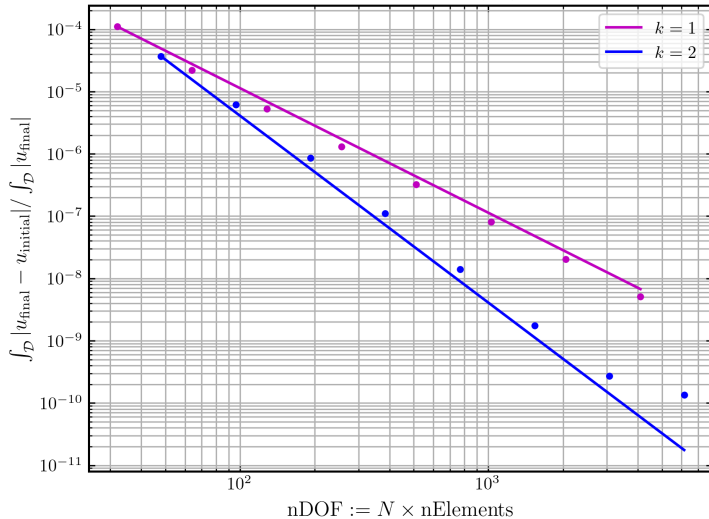
Discontinuous Galerkin (DG) Method

$$\mathbb{P}^k \ni u_h(x, t) := \sum_{i=1}^{k+1} u_i(t) \ell_i(x)$$

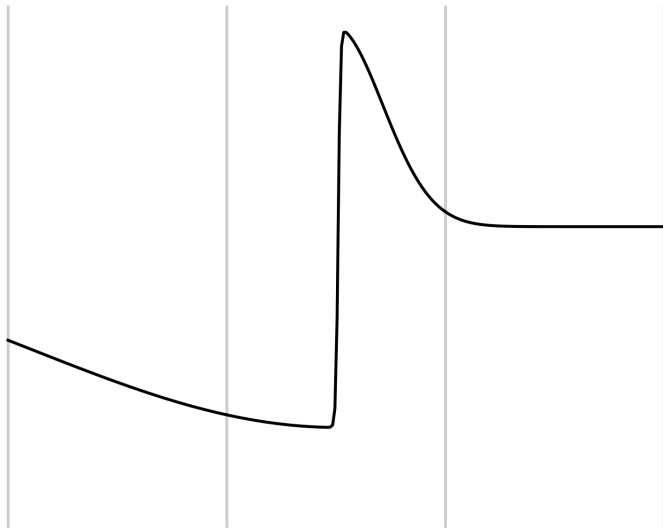


- For hydro-only, strong-stability-preserving Runge–Kutta (SSPRK) methods
 - Convex combinations of forward-Euler timesteps
- For hydro+neutrinos, realizability-preserving ImEx methods (Chu et al. (2019))

Convergence Rates for Sine Wave Advection (1D)

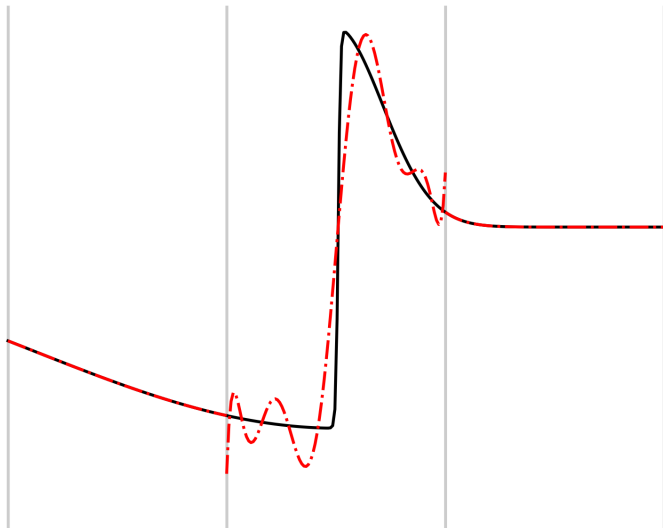


Slope Limiter (MinMod)

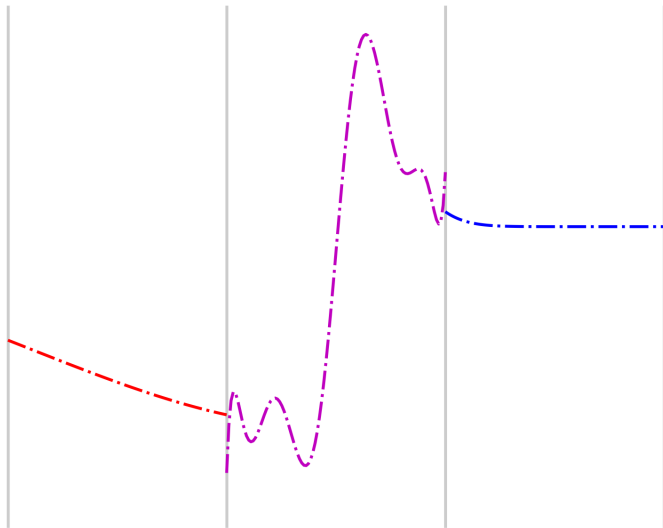


Slope Limiter (MinMod)

$$k = 9$$

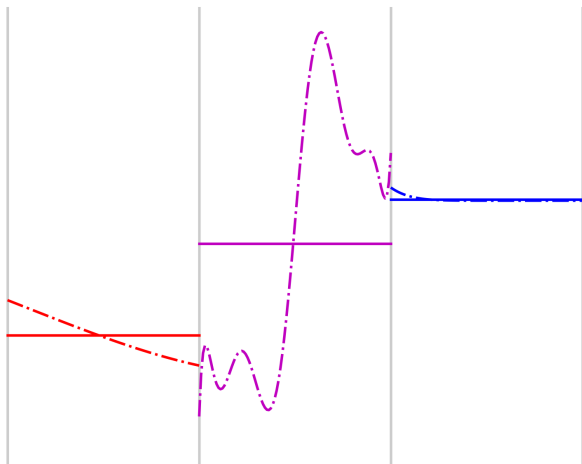


Slope Limiter (MinMod)



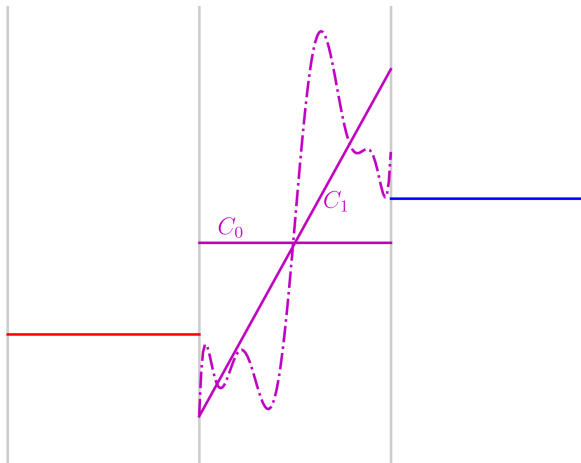
Slope Limiter (MinMod)

$$\bar{u}(t) := \frac{1}{V_h(t)} \int_K u_h(x, t) \sqrt{\gamma_h}(x, t) dx, \quad V_h(t) := \int_K \sqrt{\gamma_h}(x, t) dx$$



Slope Limiter (MinMod)

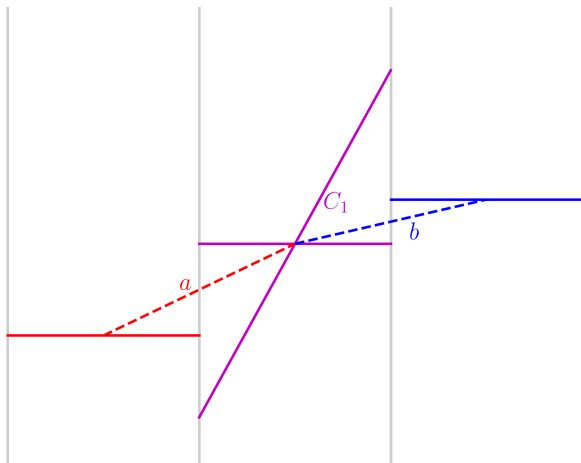
$$u_h(x, t) = \sum_{i=1}^{k+1} u_i(t) \ell_i(x) = \sum_{n=0}^k C_n(t) P_n(x)$$



Slope Limiter (MinMod)

$$a := \bar{u}_K - \bar{u}_{K-1}$$

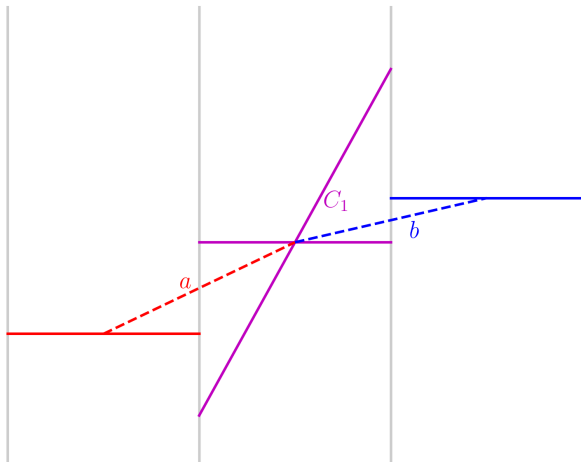
$$b := \bar{u}_{K+1} - \bar{u}_K$$



Slope Limiter (MinMod)

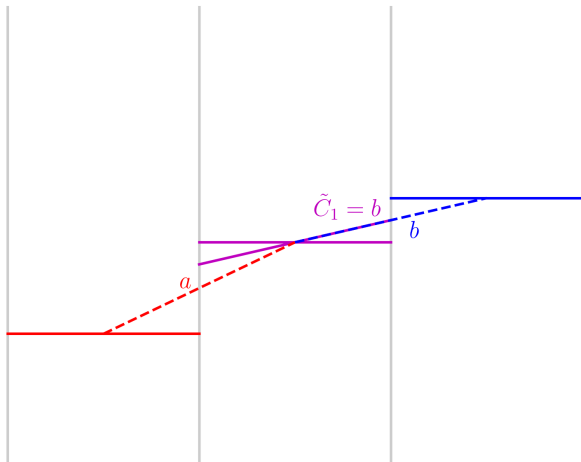
$$\tilde{C}_1 := \text{minmod}(a, b, C_1),$$

$$\text{minmod}(a, b, C_1) := \begin{cases} \text{sgn}(a) \times \min(|a|, |b|, |C_1|), & \text{sgn}(a) = \text{sgn}(b) = \text{sgn}(C_1) \\ 0, & \text{else} \end{cases}$$



Slope Limiter (MinMod)

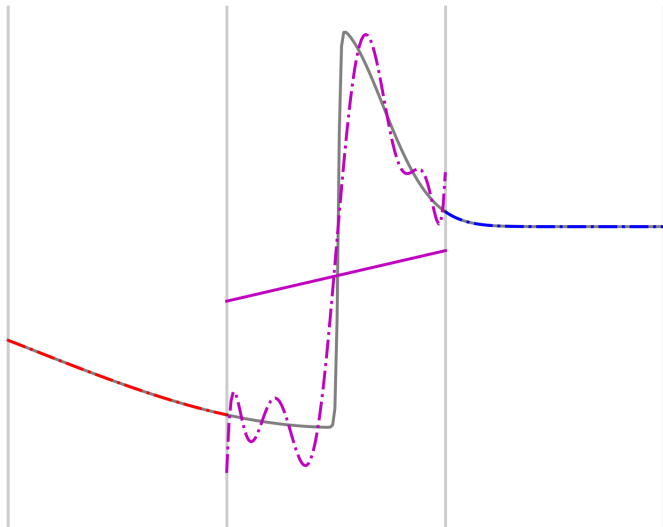
$$\sum_{n=1}^k C_n P_n = u_h \rightarrow \tilde{u}_h := \Lambda_{\text{SL}}(u_h) := C_0 P_0 + \tilde{C}_1 P_1$$



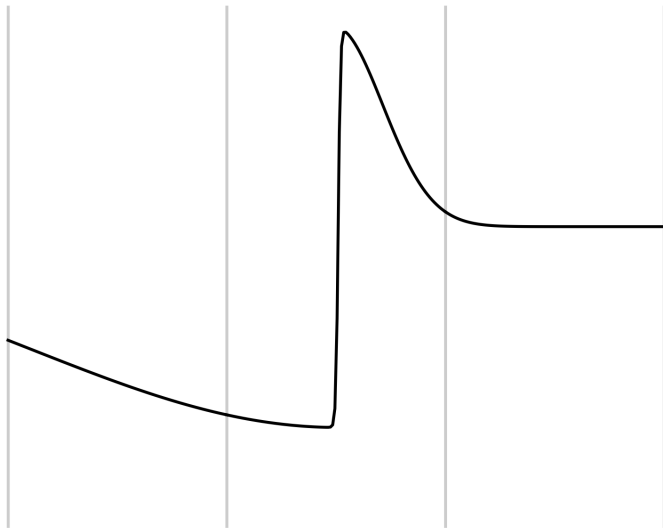
Slope Limiter (MinMod)

$$k = 9$$

$$\tilde{k} = 1$$

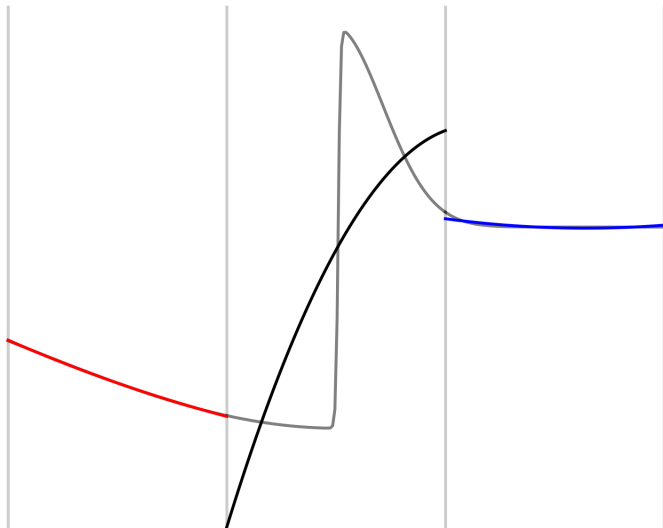


Troubled Cell Indicator (Fu & Shu, 2017)

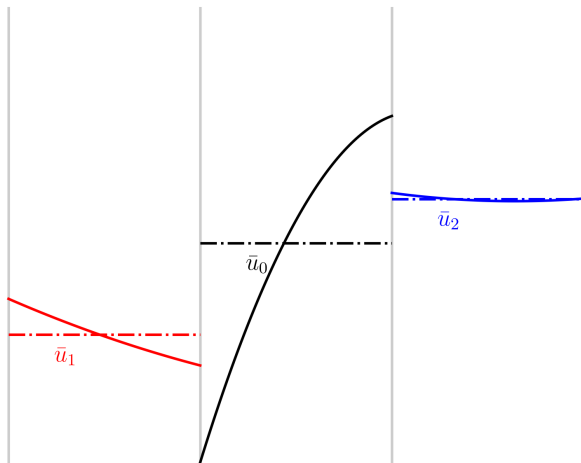


Troubled Cell Indicator (Fu & Shu, 2017)

$$k = 2$$

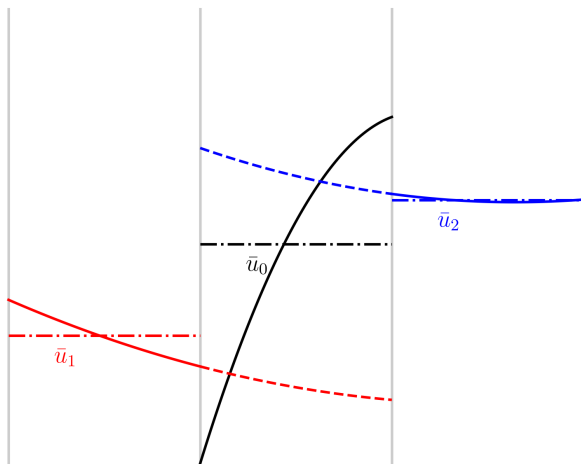


Troubled Cell Indicator (Fu & Shu, 2017)

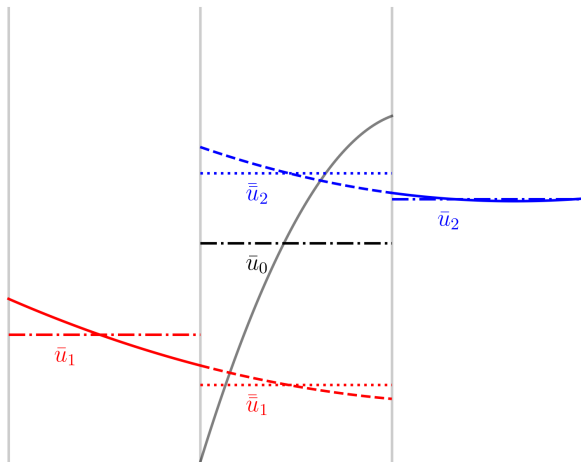


Troubled Cell Indicator (Fu & Shu, 2017)

$$u_h(x, t) = \sum_{i=1}^{k+1} u_i \ell_i(x), \quad \bar{u}_h(x, t) = \sum_{i=1}^{k+1} \bar{u}_i \ell_i(x)$$

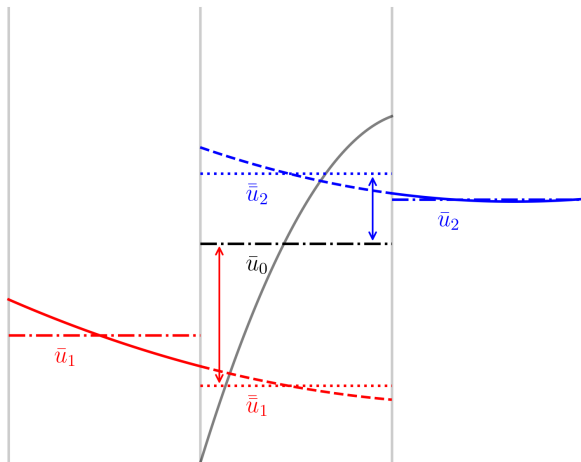


Troubled Cell Indicator (Fu & Shu, 2017)



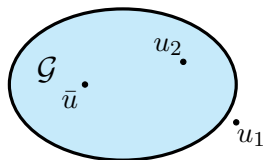
Troubled Cell Indicator (Fu & Shu, 2017)

$$I_K := \left(\max_{j \in \{0, \dots, 2^d\}} (|\bar{u}_j|) \right)^{-1} \sum_{j=1}^{2^d} |\bar{u}_0 - \bar{\bar{u}}_j|$$



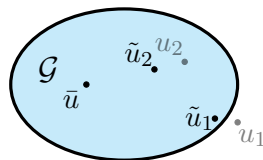
Bound-Enforcing Limiter (Λ_{BEL})

- Proposed in Qin et al. (2016)
- Define \mathcal{G} as the set of physically-admissible states; i.e., positive density and energy density, subluminal fluid velocity (for an ideal EoS, \mathcal{G} is convex)



$$u_h = u_1 \ell_1 + u_2 \ell_2$$

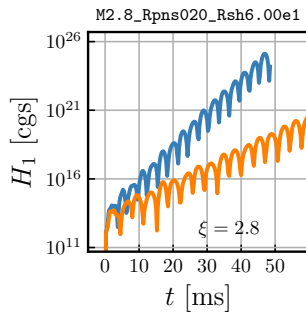
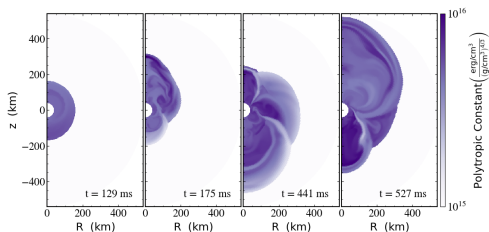
$$\xrightarrow{\Lambda_{\text{BEL}}}$$



$$\tilde{u}_h := \Lambda_{\text{BEL}}(u_h) = \tilde{u}_1 \ell_1 + \tilde{u}_2 \ell_2$$

Standing Accretion Shock Instability

Used thornado to investigate the role of GR on the SASI¹



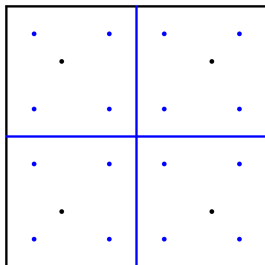
¹Dunham et al. (2020, 2024)

Mesh Refinement with AMReX

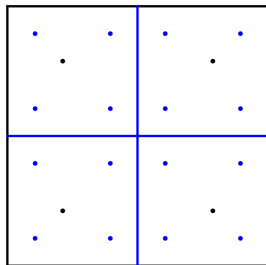


<https://github.com/AMReX-Codes/amrex>

Interpolation



Given $U_h(x, t) = \sum_{i \in \mathcal{N}} U_i(t) \ell_i(x)$, how to compute $u_h^{(j)}(x, t) = \sum_{i \in \mathcal{N}} u_i^{(j)}(t) \ell_i^{(j)}(x)$, or vice-versa? (\mathcal{N} is the set of all grid-points for any given element.)



- Schaal et al. (2015)

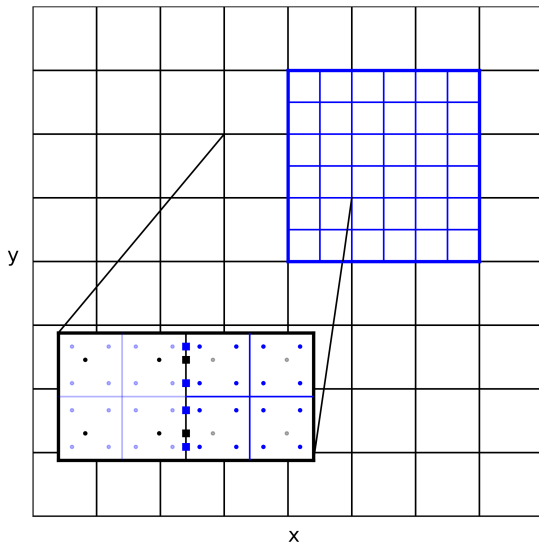
Given a solution on a coarse element K , U_h , define the solution on the j th

($j \in \{1, \dots, 2^d\}$) fine element $k^{(j)}$, $u_h^{(j)}$ via the minimization of an L^2 -projection:

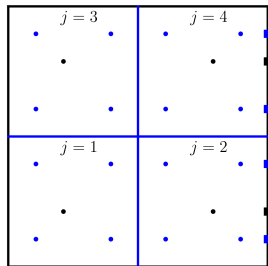
$$\forall i \in \mathcal{N} : \min_{\tilde{u}_i^{(j)}} \int_{k^{(j)}} \left(\tilde{U}_h - \tilde{u}_h^{(j)} \right)^2 d^d x,$$

where $\tilde{U} := \sqrt{\gamma} U$. For the geometry fields, we use the same procedure but do not pre-multiply with $\sqrt{\gamma}$.

Flux Corrections



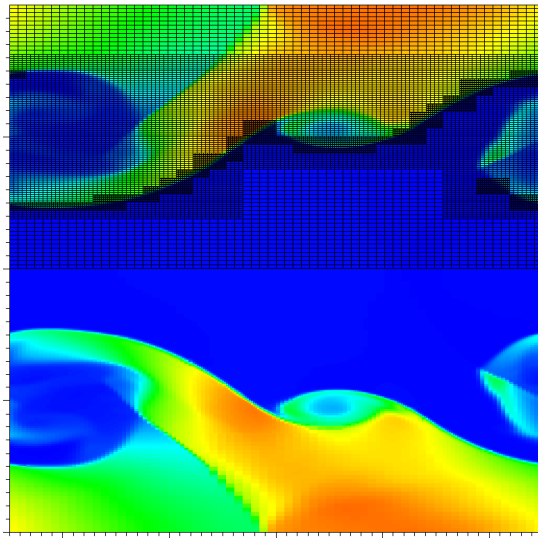
Flux Corrections



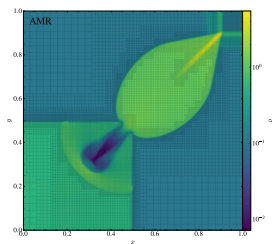
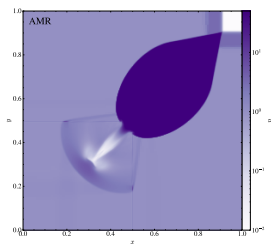
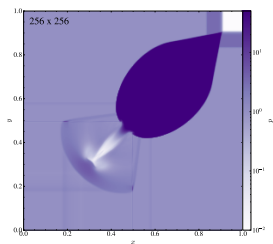
The numerical flux across, say, the x^1 interface on the right side of an element for a 2D problem with no gravity and Cartesian coordinates is given by

$$\begin{aligned}
 & \int_{K^2} \hat{F}^1 \left(U_h^-, U_h^+ \right) \ell_{k_1} \left(x_H^1 \right) \ell_{k_2} \left(x^2 \right) dx^2 \\
 &= \sum_{j \in \{2,4\}} \int_{k_{(j)}^2} \hat{F}^1 \left(U_h^-, U_h^+ \right) \ell_{k_1} \left(x_H^1 \right) \ell_{k_2} \left(x^2 \right) dx^2 \\
 &:= \sum_{j \in \{2,4\}} \int_{k_{(j)}^2} \hat{F}^1 \left(u_h^{(j),-}, u_h^{(j),+} \right) \ell_{k_1} \left(x_H^1 \right) \ell_{k_2} \left(x^2 \right) dx^2
 \end{aligned}$$

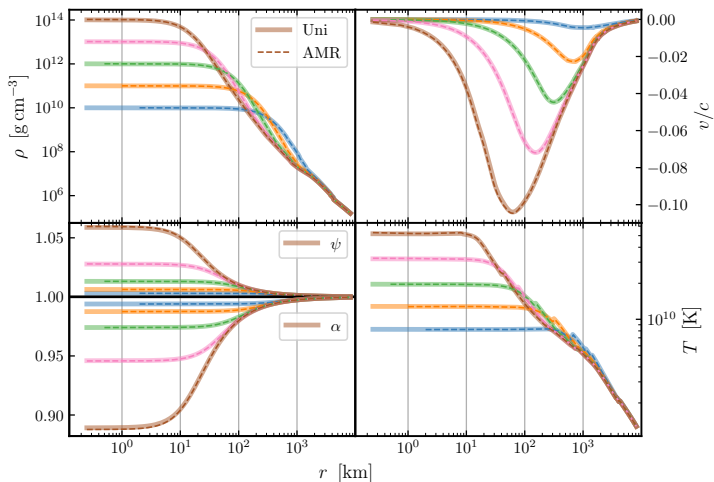
Kelvin–Helmholtz Instability



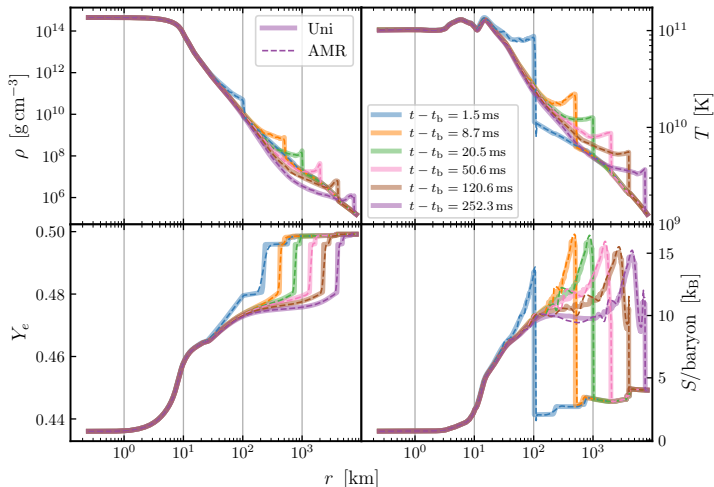
2D Blast Wave (Del Zanna & Bucciantini, 2002)



Adiabatic Collapse (Collapse Phase)



Adiabatic Collapse (Post-Bounce Phase)

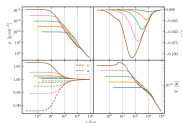
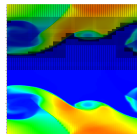


Bibliography

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Summary

Can run multi-D pure hydro problems in GR with AMR



Can run 1D hydro+self-gravity problems in GR with AMR

Working on coupling GR transport to existing hydro+gravity modules