

Stars Level 2, Durham University Physics

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These equations are given to provide an alternative overview of the concepts that we explored in the stars lectures. Use these to help your understanding rather than as a list to learn by rote. Refer to the lecture notes for the context and usage of these equations.

Equation 1 $L = 4\pi R^2 \sigma T_e^4$

Equation 2 $\frac{N_b}{N_a} = \frac{g_b}{g_a} e^{-(E_b - E_a)/kT}$

Equation 3 $P^2 = \frac{4\pi^2 a^3}{G(m_1 + m_2)}$

Equation 4 $\frac{m_1}{m_2} = \frac{\alpha_2}{\alpha_1} = \frac{a_2}{a_1} = \frac{v_2}{v_1}$

Equation 5 $\frac{L}{L_{sun}} = \left(\frac{M}{M_{sun}} \right)^\alpha$

Equation 6 $\frac{dP}{dr} = -\frac{GM_r}{r^2} \rho$ [SHORT DERIVATION]

Equation 7 $\frac{dM_r}{dr} = 4\pi r^2 \rho$

Equation 8 $P_c = \frac{3}{8\pi} \frac{GM^2}{R^4}$ [DERIVATION]

Equation 9 $P = P_{gas} + P_{rad} = nkT + \frac{1}{3}aT^4$

Equation 10 $K = -\frac{1}{2}U$ [DERIVATION]

Equation 11 $E = \frac{3GM^2}{10R} \left[\frac{1}{R} - \frac{1}{R_{initial}} \right]$ [DERIVATION]

Equation 12 $T_{classical} = \frac{Z_1 Z_2 e_c^2}{6\pi\epsilon_0 k r}$ [SHORT DERIVATION]

Equation 13 $T_{quantum} = \frac{Z_1^2 Z_2^2 e_c^4 \mu_m}{12\pi^2 \epsilon_0^2 k h^2}$ [DERIVATION]

Equation 14 $\varepsilon_{ix} = \varepsilon'_0 X_i X_x \rho^\alpha T^\beta$

Equation 15 $\frac{dL}{dr} = 4\pi r^2 \rho \varepsilon$

Equation 16 $\ell = \frac{1}{n\sigma}$

Equation 17 $N = \left(\frac{d}{\ell}\right)^2$

Equation 18 $\frac{dT}{dr} = -\frac{3}{4ac} \frac{\kappa \rho F_{rad}}{T^3} = -\frac{3}{16\pi ac} \frac{\kappa \rho L_r}{T^3 r^2}$ **[SHORT DERIVATION]**

Equation 19 $I_\lambda = I_{\lambda,0} e^{-\kappa_\lambda \rho s}$ **[DERIVATION]**

Equation 20 $\kappa = \kappa_0 \rho^\alpha T^\beta$

Equation 21 $\left|\frac{dT}{dr}\right|_{sur} > \left(\frac{\gamma_{ad}-1}{\gamma_{ad}}\right) \frac{T}{P} \left|\frac{dP}{dr}\right|_{sur}$ **[DERIVATION]**

Equation 22 $\ell = \alpha H_p$

Equation 23 $\frac{M_{max}}{M_{sun}} = \alpha^{-1} \sqrt{\frac{4\pi c G M_{sun}}{\kappa L_{sun}}}$ **[DERIVATION]**

Equation 24 $\Pi \approx \sqrt{\frac{3\pi}{2\gamma G \rho}}$ **[DERIVATION]**

Equation 25 $M_J \cong \left(\frac{5kT}{G\mu m_H}\right)^{3/2} \left(\frac{3}{4\pi\rho_0}\right)^{1/2}$ **[SHORT DERIVATION]**

Equation 26 $t = \frac{X\xi M c^2}{L}$

Equation 27 $t = 10^{10} \left(\frac{M_{Sun}}{M}\right)^{\alpha-1}$

Equation 28 $P = \frac{\hbar^2}{m_e} \left[\left(\frac{Z}{A}\right) \frac{\rho}{m_H} \right]^{5/3}$ **[DERIVATION]**

Equation 29 $P_{min} = \left(\frac{3\pi}{G\rho}\right)^{1/2}$ **[SHORT DERIVATION]**

Equation 30 $R = \frac{2GM}{c^2} = 2.96 \left(\frac{M}{M_{Sun}}\right) km$

Associated non-numbered equations (which you should know):

Ideal gas law: $P = nKT$ so $P = \frac{\rho kT}{\mu m_H}$

Radiation pressure: $P = \frac{1}{3}aT^4$

Adiabatic pressure: $P = K_a \rho^\gamma$

Adiabatic Sound speed: $v_s = \sqrt{\frac{\gamma P}{\rho}}$

Kinetic energy: $E = \frac{1}{2}mV^2$

Gravitational potential energy: $U = -\frac{GMm}{r}$

Gravitational force: $g = \frac{GM}{r^2}$

Momentum: $p = mv$

Magnitude system: $m = const - 2.5 \times \log(f)$

Parallax: $d(pc) = \frac{1}{\phi(arc\ sec)}$

Wein displacement law: $\lambda_{\max} T = 2.9 \times 10^{-3} mK$