

Thermodynamics – Lecture 13 Recap

- Reviewed Thermodynamics.
- Were introduced to the topic of Statistical Mechanics.
- Consider distinguishable particles and how to count them,

$$\Omega = \frac{N!}{\prod_j n_j!},$$

for a system having N particles, with n_j in a state of energy ϵ_j .

- Defined entropy and temperature statistically,

$$S = k_B \ln \Omega, \quad \beta = \frac{1}{k_B T}.$$

- Saw how Boltzmann distribution, for thermal equilibrium arises.

Thermodynamics – Lecture 14 Aims

- To consider the conditions that result in the Boltzmann distribution.
- To see how the Partition function normalises the Boltzmann distribution and how it is related to classical thermodynamics.
- To look at distribution function for indistinguishable particles.
- To consider Fermion and Bosons, and in particular look at the effect of Bose-Einstein condensation, and the Fermi energy.

Boltzmann – canonical ensemble at thermal equilibrium
 $P(\epsilon_j) \propto \exp(-\epsilon_j / k_B T) = \exp(-\beta \epsilon_j)$

Ensemble – make copies of system to do statistics.

Proof 23.1 Any system has

- Fixed particle number $N = \sum_j n_j$, $dN = \sum_j dn_j = 0$
- Fixed total energy $E = \sum_j \epsilon_j n_j$, $dE = \sum_j \epsilon_j dn_j = 0$
- Most probable macrostate at thermal equilibrium $d\Omega = 0$
 $\Rightarrow \sum_j (\ln n_j) dn_j = 0$

Three conditions are zero together. Lagrange Multiplier

quantities are zero can add arbitrary multiples of each and the sum is still zero

$$0 = \sum_j (\alpha dn_j + \beta \epsilon_j dn_j + \ln n_j) dn_j$$

$$\text{Holds for all } j = \sum_j dn_j (\alpha + \beta \epsilon_j + \ln n_j) = 0$$

$$\alpha + \beta \epsilon_j + \ln n_j = 0$$

$$\ln n_j = -\alpha - \beta \epsilon_j$$

$$n_j = \exp(-\alpha - \beta \epsilon_j) = \underline{A \exp(-\beta \epsilon_j)}$$

$A = \exp(-\alpha)$ is our normalisation

$$\text{Require } \sum_j P(\epsilon_j) = 1 \quad [\text{Sum of probabilities} = 1]$$

$$P(\epsilon_j) = \frac{n_j}{N} \Rightarrow 1 = \frac{1}{N} \sum_j n_j$$

$$1 = \frac{1}{N} \sum_j A \exp(-\beta \epsilon_j)$$

$$A = \frac{N}{\sum_j \exp(-\beta \epsilon_j)}$$

$$Z = \sum_j \exp(-\beta \epsilon_j)$$

Partition function

See DUO notes that $\beta = \frac{1}{k_B T}$ (defines temperature)

24. Partition function

Normalises the probability for particles (distinguishable)

$$Z = \sum_j \exp(-\beta \epsilon_j) = \sum_j \exp\left(-\frac{\epsilon_j}{k_B T}\right)$$

Tells us our N particles are split into i states with state j of energy ϵ_j and the splitting happens in the ratio of the Boltzmann factors

Can relate Z to classical thermodynamics is via the Helmholtz

$$F = -k_B T \ln Z \Rightarrow Z = \exp(-\beta F)$$

If can write partition function for any system (know the state energies) can calculate classical thermodynamic quantities

$$F = U - TS \quad dF = -SdT - pdV \quad p = -\left(\frac{\partial F}{\partial V}\right)_T$$

$$U = (E) = -\frac{d \ln(Z)}{d\beta} = k_B T^2 \frac{d \ln Z}{dT}$$

↑
Internal energy E
in stat mechanics

Thermodynamics – Handout 14

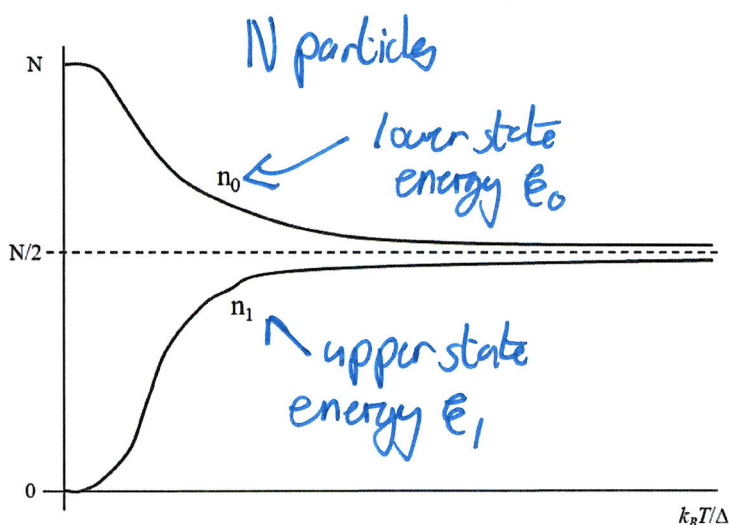


Figure 36: Populations of a system having two energy levels, and how they vary with temperature.

$$\Delta = \epsilon_1 - \epsilon_0$$

$$P(\epsilon_0) = \frac{n_0}{N} \quad N = n_0 + n_1$$

$$= \frac{A \exp\left(\frac{-\epsilon_0}{k_B T}\right)}{A \exp\left(\frac{-\epsilon_0}{k_B T}\right) + A \exp\left(\frac{-\epsilon_1}{k_B T}\right)} = \frac{1}{1 + \exp\left(\frac{-\Delta}{k_B T}\right)}$$

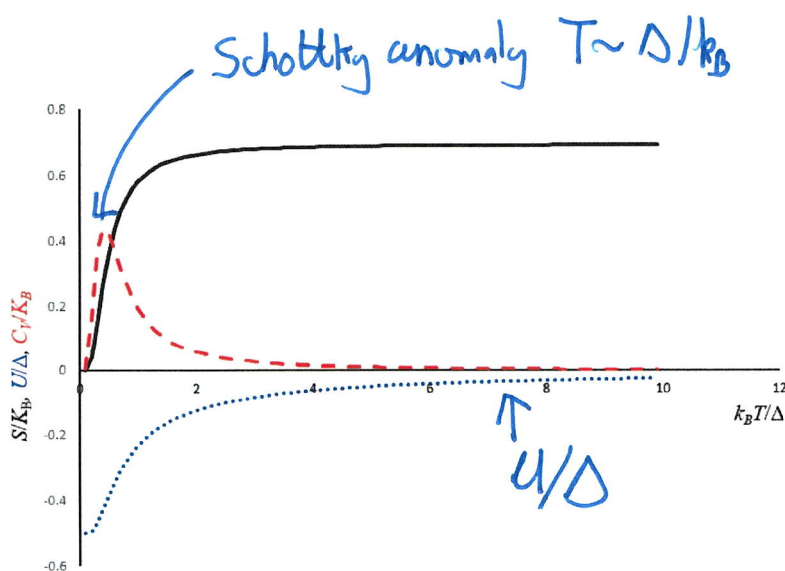
$$P(\epsilon_1) = \frac{n_1}{N} = \frac{\exp\left(\frac{-\epsilon_1}{k_B T}\right)}{1 + \exp\left(\frac{-\Delta}{k_B T}\right)}$$

$$\text{As } T \rightarrow 0, P(\epsilon_0) = \frac{1}{1 + e^{-\infty}} = 1 \quad ; \quad P(\epsilon_1) = \frac{e^{-\infty}}{1 + e^{-\infty}} = 0$$

All at low energy state

$$\text{As } T \rightarrow \infty, P(\epsilon_0) = \frac{1}{1 + e^{-0}} = 1/2 \quad ; \quad P(\epsilon_1) = \frac{e^{-0}}{1 + e^{-0}} = 1/2$$

Equal state occupation



The internal energy (solid), entropy (dotted) and heat capacity (dashed) for a two-state system.

$$Z = \sum_j \exp\left(\frac{-\epsilon_j}{k_B T}\right)$$

$$Z = \exp\left(\frac{-\epsilon_0}{k_B T}\right) + \exp\left(\frac{-\epsilon_1}{k_B T}\right) = 2 \cosh\left(\frac{\Delta}{2k_B T}\right)$$

$$E (= U) = -\frac{d \ln Z}{d\beta} = -\frac{1}{Z} \times \frac{dZ}{d\beta}$$

$$= -\frac{1}{2 \cosh(\beta\Delta/2)} \times 2 \times \frac{\Delta}{2} \sinh(\beta\Delta/2)$$

$$= -\frac{\Delta}{2} \tanh(\beta\Delta/2)$$

$$F = U - TS$$

↑
E

$$F = -k_B T \ln Z$$

$$S = \frac{E - F}{T} = -\frac{\Delta}{2T} \tanh\left(\frac{\beta\Delta}{2}\right) + k_B \ln\left(2 \cosh\left(\frac{\beta\Delta}{2}\right)\right)$$

$$C_V = \left(\frac{\partial E}{\partial T}\right)_V = k_B \left(\frac{\Delta\beta}{2}\right)^2 \text{sech}^2\left(\frac{\beta\Delta}{2}\right)$$

$$C_V = \left(\frac{\partial Q}{\partial T}\right)_V$$

As $T \rightarrow 0$, $E \rightarrow -\Delta/2$ and $S \rightarrow 0$ [System all in lower level]
 $S = k_B \ln \Omega$ with $\Omega = 1$

As $T \rightarrow \infty$, $E \rightarrow 0$ (equal state occupation) $S = k_B \ln 2$

25. Indistinguishable Particles

Cannot label them

- Classical (dilute) gas, particles are spread out in the system.
- Fermions + Bosons, are identical types of quantum particle with differing wave functions.
Quantum effects can change the statistics

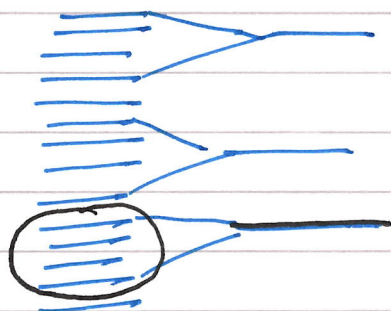
$$\uparrow(1,2) = \frac{1}{\sqrt{2}} [\uparrow_A(1) \uparrow_B(2) \pm \uparrow_A(2) \uparrow_B(1)]$$

Bosons have a symmetric wave function, and allows any number of particles to occupy a given quantum state.
Integer spin.

Fermions have antisymmetric wave function. Satisfy the Pauli exclusion principle, no two particles occupy the same quantum state. Half integer spin.

Density of states

- Gases have many allowed ^{states} ~~ends~~ than particles, group our energy states into levels



True Grouped

ith level energy E_i , has g_i quantum states

Density of states g_i — how many energy states can exist

$$f_i = \frac{n_i}{g_i} \quad \text{distribution function, Number of particles in a given state}$$

Many energies, we can take energy to be continuous, and consider how many particles in the energy $\epsilon \rightarrow \epsilon + d\epsilon$

$$n(\epsilon)d\epsilon = f(\epsilon)g(\epsilon)d\epsilon$$

$$N = \int_0^\infty n(\epsilon)d\epsilon, \quad E = \int \epsilon n(\epsilon)d\epsilon$$

Maxwell Boltzmann

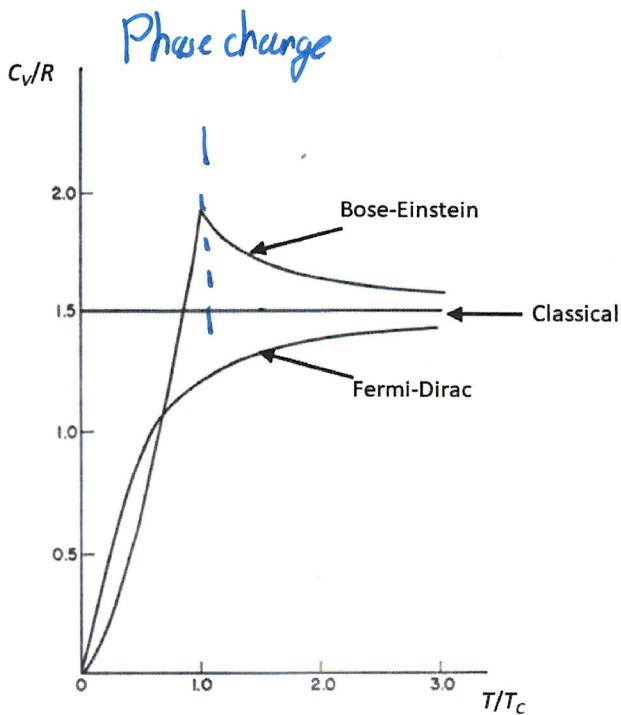
$$f(\epsilon) = A \exp\left(\frac{-\epsilon}{k_B T}\right)$$

Box-Einstein

$$f_{BE}(\epsilon) = \frac{1}{\exp\left(\frac{\epsilon - \mu}{k_B T}\right) - 1}$$

Fermi Dirac

$$f_{FD}(\epsilon) = \frac{1}{\exp\left(\frac{\epsilon - \epsilon_F}{k_B T}\right) + 1}$$



Heat capacity behaviour with temperature for various particles.

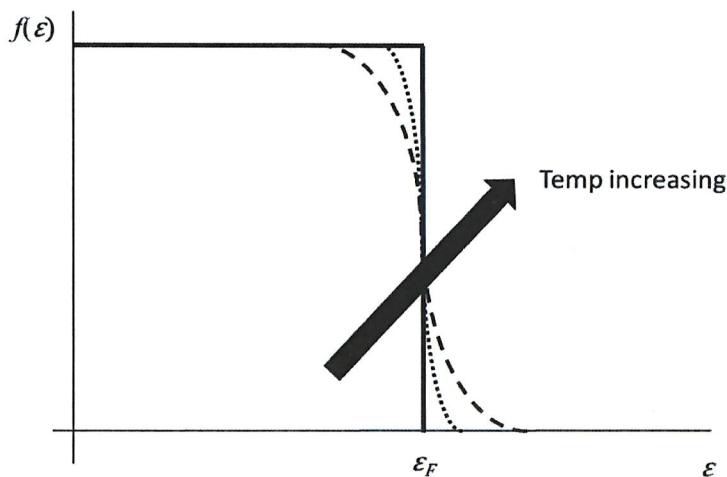


Figure 38: The behaviour of the Fermi-Dirac distribution with temperature.

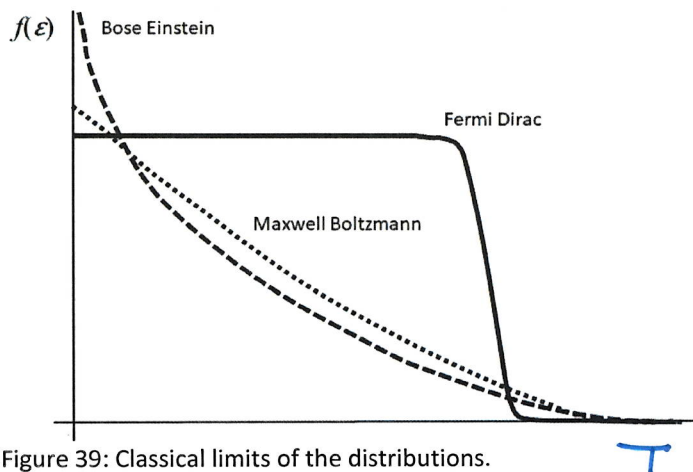


Figure 39: Classical limits of the distributions.

Bose Einstein
 μ is the chemical potential

As $T \rightarrow 0$, $\mu \rightarrow E_0$

$$\lim_{T \rightarrow 0} \exp\left(\frac{E - \mu}{k_B T}\right) = 1$$

$$f_{BE}(E) \rightarrow \frac{1}{1-1} = \frac{1}{0} \rightarrow \infty$$

All Bosons go to the lowest quantum state.
 — Bose-Einstein condensation

E_F - Fermi energy

At $T=0$ $E \rightarrow E_F$

$$f_{FD}(E) = \frac{1}{1+1} = \frac{1}{2}$$

As $T \rightarrow \infty$ $f_{FD} \rightarrow 0$

At $T=0$ all states filled to Fermi energy

At high T all distributions become the same.