

Thermodynamics – Lecture 12 Recap

- Looked at adiabatic cooling, both via expansion and demagnetization.
- Were introduced to the Third Law of Thermodynamics.
- Looked at example of Thermodynamics in action, including elastic rods and real gases:

$$pV \rightarrow xY \quad ; \quad \delta W = -pdV \rightarrow \delta W = x dY$$

$$\delta W = +f dx, m dB, \dots$$

$$pV = RT \rightarrow f(p, V, T) = 0$$

$$\left(p + \frac{a}{V^2}\right)(V - b) = RT \dots \quad \left(\frac{\partial p}{\partial V}\right)_T = \dots$$

Thermodynamics – Lecture 13 Aims

- To review Thermodynamics.
- To introduce Statistical Mechanics.
- To consider distinguishable particles.
- To consider how we can define entropy and temperature statistically.
- To see how the Boltzmann distribution, for thermal equilibrium arises.

Three Laws + Zeroth

- 0 – Temperature + thermal equilibrium
- 1 – Energy conservation
- 2 – Entropy of the Universe increases [Engines aren't 100%]
- 3 – Can't get to absolute zero

$$dU = \delta Q + \delta W$$

2). Statistical Mechanics

Considered bulk properties – impossible to look at the individual behaviour of all atoms
– statistical methods

Macrostate – how a system behaves on average

Microstate – most detailed description of a system we can imagine

Example 21.1 Flip 4 coins, get Head or Tail

Macrostate = number of heads + number of tails
 Microstate = individual arrangement

Macro	Micro	
4 H	HHHH	1
3 H 1 T	HHHT, HHTH, HTHH, THHH	4
2 H 2 T		6
3 T 1 H		4
4 T		1

Particles are distinguishable, the number of microstates in the macrostate

$$\Omega = \frac{N!}{n! (N-n)!}$$

N = Total number of particles

n = Number of particles in state 1

$N-n$ = Number of particles in state 2

$$\Omega(2H, 2T) = \frac{4!}{2! 2!} = 6$$

$$\text{Total number of arrangements } \sum \Omega = 2^N$$

Two possible outcomes

More generally, N particles in total with n_j particles in a state energy ϵ_j

$$N = \sum_j n_j$$

$$\Omega = \frac{N!}{\prod_j n_j!}$$

$$\prod_j n_j! = n_1! \times n_2! \times \dots \times n_m! \quad \left[\begin{array}{l} \text{Extended} \\ \text{product} \end{array} \right]$$

$$N! = N \times (N-1) \times (N-2) \times \dots \times 1$$

N is large, and often taking natural logs
 $S = k_B \ln \Omega$

Stirling approximation $\ln X! \approx X \ln X - X$

$$\ln X! = \ln X + \ln(X-1) + \ln(X-2) + \dots + \ln 1$$

$$= \sum_{x=1}^X \ln x \approx \int_1^X \ln x dx$$

$$= [x \ln x - x]_1^X \sim X \ln X - X$$

Thermodynamics – Handout 13

21.2 100 coins in a box, can be arranged in one of $2^{100} \approx 10^{30}$ ways, each equally likely,

$p \sim \frac{1}{10^{30}}$

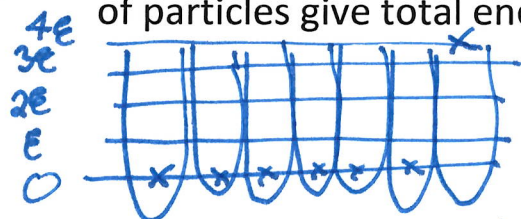
Macro $x H, y T (x+y=100)$,

Micro $HTHH TTTH \dots$

Particular microstate requires each coin to be individually labelled – simpler to measure macrostate!

Macrostate	Number of Microstates
50 H and 50 T	$\frac{100!}{50! 50!} \sim 4 \times 10^{27}$
53 H and 47 T	$\frac{100!}{53! 47!} \sim 3 \times 10^{27}$
90 H and 10 T	$\frac{100!}{90! 10!} \sim 1 \times 10^{13}$
99 H and 1 T	$\frac{100!}{99! 1!} \sim 100$

21.3 7 potential wells with energy levels of $0, \epsilon, 2\epsilon, 3\epsilon, 4\epsilon$. What arrangements of particles give total energy of 4ϵ ?



Any one well has energy 4ϵ , 7 other wells have energy 0

$\Omega = \frac{N!}{n_1! n_2! \dots n_r!}, N=7$

$\Omega = \frac{7!}{(6! 0! 0! 0! 1!)} = 7$

$\Omega = \frac{7!}{4! 2! 1! 0! 0!} = 105$

n_j depend on the internal arrangement.

Energy	0	ϵ	2ϵ	3ϵ	4ϵ
n_j	6	0	0	0	1
	4	2	1	0	0
	5	1	0	1	0

21.4 $\ln 5! = \ln 120 = 4.79$. Stirling, $5 \ln 5 - 5 = 3.04$, $\sim 40\%$ under

$\ln 100! = 364$, Stirling = 361 1% underestimate

$\ln 1000! = 5912$, Stirling = 5908 $\sim 0.1\%$ underestimate

$N \sim 10^{23}$ for one mole, error in Stirling approx is negligible

22. Temperature + Entropy Statistically

Thermal equilibrium defined by some temperature

- Macrostate you see will most likely have the most associated microstates

[Ergodic hypothesis, microstates are continually changing]

Proof 22.1

Two systems thermally isolated from World
 $E = E_1 + E_2$ [Fixed total energy]

First system has $\Omega_1(E_1)$ microstates
Second " $\Omega_2(E_2)$ "

Total Microstates $\Omega_1(E_1) \Omega_2(E_2)$

Slightly change energy E_1 , maximise # of microstates

$$\frac{d}{dE_1} (\Omega_1(E_1) \Omega_2(E_2)) = \Omega_2(E_2) \frac{d\Omega_1(E_1)}{dE_1} + \Omega_1(E_1) \frac{d\Omega_2(E_2)}{dE_2} \frac{dE_2}{dE_1}$$

$$dE = dE_1 + dE_2 = 0 \text{ as total energy fixed}$$
$$\therefore \frac{dE_1}{dE_2} = -1$$

$$\frac{1}{\Omega_1(E_1)} \frac{d\Omega_1(E_1)}{dE_1} = \frac{1}{\Omega_2(E_2)} \frac{d\Omega_2(E_2)}{dE_2}$$

$$\frac{d \ln \Omega_1(E_1)}{dE_1} = \frac{d \ln \Omega_2(E_2)}{dE_2} = \frac{1}{k_B T}$$

Must be a temperature - thermal equilibrium

Proof 22.2

$$S = k_B \ln \Omega$$

1st Law $du = Tds - pdv$
 $\left(\frac{\partial u}{\partial s}\right)_v = T$

$$\frac{1}{T} = \left(\frac{\partial s}{\partial u}\right)_v ; \quad \frac{1}{k_B T} = \frac{d \ln \Omega}{dE}$$

Statistically energy E corresponds to internal energy, U

$$\Rightarrow dS = k_B d \ln \Omega \Rightarrow S = k_B \ln \Omega$$

23. Boltzmann distribution

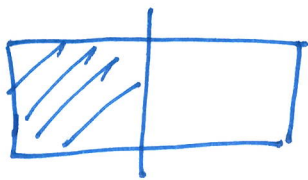
Canonical distribution - an ensemble of systems, which exchange energy with a heat reservoir

Thermal equilibrium for distinguishable particles
- system at same temperature as the reservoir

$$P(E_j) = \frac{\exp(-E_j/k_B T)}{\sum_j \exp(-E_j/k_B T)}$$

$$\sum_j \exp(-E_j/k_B T) = Z, \quad \text{Normalisation}$$

↑
Partition function.



Remove partition
 $V \rightarrow 2V$

Joule expansion, with volume doubling for one mole ideal gas isothermally.

$U = U(T)$, so $dU = 0$, and the first law ($dU = \delta Q + \delta W$) is $TdS = pdV$

$$\Delta S = \int_V^{2V} \frac{p}{T} dV = \int_V^{2V} \frac{R}{V} dV = R \ln 2.$$

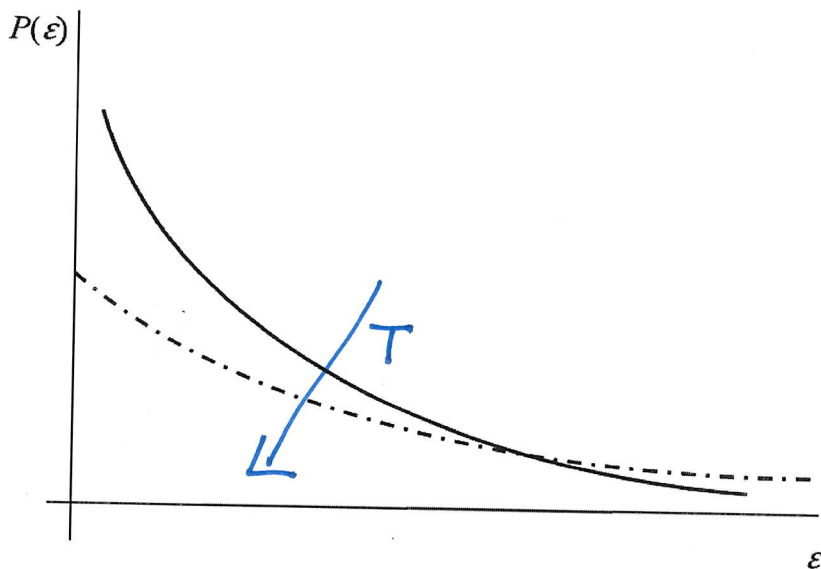
Any single molecule now in two places, either left or right

Statistically the number of possible microstates becomes larger by a multiplicative factor of 2^{N_A}

$$\Omega_{After} = 2^{N_A} \Omega_{Before}$$

$$\begin{aligned} S_{After} &= k_B \ln \Omega_{After} = k_B \ln(2^{N_A} \Omega_{Before}) = k_B \ln(2^{N_A}) + \ln \Omega_{Before} \\ &= N_A k_B \ln 2 + k_B \ln \Omega_{Before}. \end{aligned}$$

$$\Delta S = N_A k_B \ln 2 = R \ln 2.$$



$$P(\epsilon_j) \propto \exp\left(-\frac{\epsilon_j}{k_B T}\right).$$

Increase temperature,
 distribution has more chance
 of finding a particle with
 higher energy

For a given temperature, energy takes a probability distribution
 'System has' reasonable chance of having $\epsilon < k_B T$
 The exponential quickly reduces the chance of $\epsilon > k_B T$