

# University of Durham

## EXAMINATION PAPER

Examination session:

May/June

Year:

2019

Examination code:

PHYS3661-WE01

Title:

Theoretical Physics 3

Time allowed:	3 hours		
Additional material provided:	None		
Materials permitted:	None		
Calculators permitted:	Yes	Models permitted:	Casio fx-83 GTPLUS or Casio fx-85 GTPLUS
Visiting students may use dictionaries:		No	

### Instructions to candidates:

- Attempt **all** questions. The short-answer questions at the start of each section carry 50% of the total marks for the paper. The remaining 50% of the marks are carried by the longer questions, which are equally weighted.
- The marks shown in brackets for the main parts of each question are given as a guide to the weighting the markers expect to apply.
- **ANSWER EACH SECTION IN A SEPARATE ANSWER BOOK.**
- Do **not** attach your answer booklets together with a treasury tag, unless you have used more than one booklet for a single section.
- Slip your booklet for Section B inside your booklet for Section A, before they are collected by the invigilator.

### Information

**Section A:** Relativistic Electrodynamics

**Section B:** Quantum Theory 3

A list of physical constants is provided on the next page.

Revision:

**Information**

Elementary charge:	$e = 1.60 \times 10^{-19} \text{ C}$
Speed of light:	$c = 3.00 \times 10^8 \text{ m s}^{-1}$
Boltzmann constant:	$k_B = 1.38 \times 10^{-23} \text{ J K}^{-1}$
Bohr magneton:	$\mu_B = 9.27 \times 10^{-24} \text{ J T}^{-1}$
Electron mass:	$m_e = 9.11 \times 10^{-31} \text{ kg}$
Gravitational constant:	$G = 6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2}$
Proton mass:	$m_p = 1.67 \times 10^{-27} \text{ kg}$
Planck constant:	$h = 6.63 \times 10^{-34} \text{ J s}$
Permittivity of free space:	$\epsilon_0 = 8.85 \times 10^{-12} \text{ F m}^{-1}$
Magnetic constant:	$\mu_0 = 4\pi \times 10^{-7} \text{ H m}^{-1}$
Molar gas constant:	$R = 8.31 \text{ J K}^{-1} \text{ mol}^{-1}$
Avogadro's constant:	$N_A = 6.02 \times 10^{23} \text{ mol}^{-1}$
Gravitational acceleration at Earth's surface:	$g = 9.81 \text{ m s}^{-2}$
Stefan-Boltzmann constant:	$\sigma = 5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$
Astronomical Unit:	$\text{AU} = 1.50 \times 10^{11} \text{ m}$
Parsec:	$\text{pc} = 3.09 \times 10^{16} \text{ m}$
Solar Mass:	$M_\odot = 1.99 \times 10^{30} \text{ kg}$
Solar Luminosity:	$L_\odot = 3.84 \times 10^{26} \text{ W}$

### SECTION A: RELATIVISTIC ELECTRODYNAMICS

1. (a) Show that the temporal order of two events is the same in all reference frames if and only if they are separated by a time-like interval. [4 marks]
- (b) Calculate the speed of a particle (relative to  $c$ ) if its relativistic kinetic energy is equal to twice its rest-mass  $m$  times  $c^2$ . [4 marks]
- (c) A polarization tensor for a photon is

$$T^{\mu\nu} = \left( g^{\mu\nu} - \frac{p^\mu n^\nu + p^\nu n^\mu}{p \cdot n} \right),$$

where  $p^\mu$  and  $n^\mu$  are lightlike four-vectors. Calculate  $T^\mu{}_\mu$  and  $T^{\mu\nu}T_{\mu\nu}$ . [4 marks]

- (d) Compute  $a^\mu v_\mu$ , where  $a^\mu$  is the 4-acceleration and  $v^\mu$  is the 4-velocity of a given particle. [4 marks]
- (e) Three identical point charges of charge  $q$  are arranged at the corners of an equilateral triangle. Calculate the value measured for  $\underline{E} \cdot \underline{B}$  in the centre of the triangle within the inertial frame that moves with velocity  $v$  along one of the sides of the triangle. [4 marks]
- (f) Use the covariant form of the inhomogeneous Maxwell equation to derive the wave equation in vacuum for the 4-potential  $A^\mu$  in the Lorenz gauge. [4 marks]
- (g) Write the gauge transformation of the 4-potential  $A^\mu$  in contra-variant form. Use this to show that the elements of the field strength tensor are unchanged under a gauge transformation. [4 marks]
- (h) The Lienard-Wiechert potential of a point charge  $q$  with 4-velocity  $u^\mu$  is

$$A^\mu = \frac{q}{4\pi\epsilon_0} \frac{u^\mu}{u^\nu R_\nu},$$

where  $R_\nu$  is the 4-distance between the observer and the point charge. The right-hand side of the expression must be evaluated at the retarded time  $t_{\text{ret}}$ . Evaluate this expression in the instantaneous rest frame of the point charge. Comment on why the result is as expected. [4 marks]

2. Lienard's relativistic generalisation of Larmor's formula for the power radiated from an accelerated, charged, point-like particle can be written as

$$\mathcal{P} = \frac{d\mathcal{W}}{dt} = \frac{\gamma^2 q^2}{6\pi\epsilon_0 m^2 c^3} \left[ \left| \frac{d\mathbf{p}}{dt} \right|^2 - \beta^2 \left( \frac{dp}{dt} \right)^2 \right], \quad (1)$$

where  $q$  and  $m$  are the charge and mass of the particle,  $p = |\mathbf{p}|$  is the magnitude of the relativistic three-momentum  $\mathbf{p} = \gamma m \mathbf{v}$  of the particle,  $\beta = v/c$  where  $v$  is speed of the particle, and  $\gamma$  is the standard relativistic factor  $\gamma = \frac{1}{\sqrt{1-\beta^2}}$ .

Consider a particle of charge  $q > 0$  travelling at relativistic speed along the  $x$ -axis (starting at  $x = -\infty$ ) towards the origin of the coordinate system. There is a point-charge  $q' > 0$  fixed at  $\mathbf{r}' = (x, y) = (0, b)$ . The Coulomb field from the fixed charge is

$$\underline{E}(\mathbf{r}) = \frac{q'}{4\pi\epsilon_0} \frac{\hat{\underline{R}}}{R^2},$$

where  $\underline{R} = \mathbf{r} - \mathbf{r}'$  and  $\hat{\underline{R}}$  as usual denotes a unit vector in the direction of  $\underline{R}$ . During the scattering, the incident particle will emit electromagnetic radiation due to the acceleration it experiences. In the following you can assume that the impact parameter  $b$  and the velocity  $v$  are both so large that the particle travels in a straight line.

- Sketch the scattering including the fixed charge, and a segment of the movement of the scattering charge. The impact parameter  $b$  must be indicated on the sketch, along with the axes of the coordinate system. [2 marks]
- Use the Lorentz force to write  $(d\mathbf{p}/dt)$  in terms of the  $\underline{E}$  and  $q$ . [4 marks]
- Consider a small infinitesimal period of time  $\Delta t$ . Write  $(\Delta \mathbf{p})^2$  in terms of  $\Delta t$ , and by keeping just the first term in  $\Delta t$  show that  $(\Delta \mathbf{p}^2 / \Delta t) = 2\mathbf{p} \cdot \underline{F}$ , where  $\underline{F}$  is the force applied to the particle. Use this to deduce that  $(dp/dt) = \hat{\mathbf{p}} \cdot \underline{F}$ . [4 marks]

The electric field leads to a perturbation of the movement with constant velocity  $v \approx \beta c$ . The lowest order effect in  $\underline{E}$  is obtained by keeping the explicit factors of  $\gamma$  and  $\beta$  constant in Eq. (1), and evaluating the terms of  $d\mathbf{p}/dt$  and  $dp/dt$ .

- Using the approximations above, show that the total energy  $\mathcal{W}$  radiated during the scattering from  $t = -\infty$  to  $t = \infty$  is given by

$$\mathcal{W} = \frac{\gamma^2 q^4 q'^2}{192\pi^2 \epsilon_0^3 m^2 c^4 b^3 \beta} \left( 1 - \frac{\beta^2}{4} \right).$$

[10 marks]

$$\left[ \text{Hint: } \int_{-\infty}^{\infty} \frac{1}{(x^2 + b^2)^2} dx = \frac{\pi}{2b^3} \quad ; \quad \int_{-\infty}^{\infty} \frac{1}{(x^2 + b^2)^2} \frac{x^2}{x^2 + b^2} dx = \frac{\pi}{8b^3} \right]$$

Consider now a separate setup, with point-charges of  $q'/2$  fixed at both  $(x, y) = (0, b)$  and  $(x, y) = (0, -b)$ , and a scattering charge as before.

- e) Sketch the scattering including the fixed charges, and indicate the direction of the total force acting on the scattering particle. [4 marks]
- f) Using the same approximations as above, calculate the total energy  $\mathcal{W}$  radiated during this scattering. [6 marks]

## SECTION B: QUANTUM THEORY 3

3. (a) A beam of particles scatters from a target in a scattering experiment conducted over the time interval  $t_i \leq t \leq t_f$ . The incident flux is  $F(t)$ , and the total cross-section is  $\sigma$ . Derive an expression for the number of scattering events. [4 marks]

- (b) The spherical Bessel functions of the second kind are defined by the Rodrigues formulae,

$$n_l(r) = (-1)^{l+1} r^l \left( \frac{1}{r} \frac{\partial}{\partial r} \right)^l \frac{\cos r}{r}.$$

Derive the expression for  $n_1(r)$  and expand it in powers of  $1/r$  at small  $r$  to the order  $(1/r)^0$ . [4 marks]

- (c) Assume that at large  $r$  the wave-function for a system in a centrally symmetric potential  $V(r)$  is given by,

$$\psi(r, \theta) = N e^{ikr \cos \theta} + \frac{a}{1 + b \sin^2 \theta/2} \frac{e^{ikr}}{r},$$

where  $r$  is the distance from the origin,  $\theta$  is the polar angle and  $N$ ,  $k$ ,  $a$  and  $b$  are constants. Calculate the total cross-section  $\sigma$  for the scattering described by  $\psi(r, \theta)$ . [4 marks]

[Hint:  $\sin \theta = 2 \sin \theta/2 \cos \theta/2$ .]

- (d) Consider a four-vector current  $j^\mu(x)$  given by

$$j^\mu = \frac{i}{2m} (\phi^* \partial^\mu \phi - (\partial^\mu \phi^*) \phi)$$

where  $\phi(x)$  is a scalar field of mass  $m$ . Determine whether the current  $j^\mu$  is conserved if  $\phi(x)$  satisfies the Klein-Gordon equation. Briefly explain why  $\phi(x)$  cannot be interpreted as the wave function for a relativistic particle with the probability density given by  $j^0$ . [4 marks]

- (e) Simplify the expression

$$(i\gamma^\mu \partial_\mu + m)(i\gamma^\nu \partial_\nu - m)\psi(x)$$

using the gamma-matrix algebra. Hence prove that any solution of the Dirac equation is automatically a solution to the Klein-Gordon equation. [4 marks]

- (f) Write down the relativistic equation describing the propagation of free electrons in quantum electrodynamics. What is the relation between the energies and momenta of the plane-wave solutions to this equation? Are there negative energy solutions and why is this problematic when interactions between electrons and photons are included? [4 marks]
- (g) Consider an  $n$ -dimensional Hilbert space with the orthonormal basis  $|\psi_1\rangle, |\psi_2\rangle, \dots, |\psi_n\rangle$ . Write down the density matrix  $\rho$  for the system which is in the pure state  $|\psi_2\rangle$  and compute its von Neumann entropy  $S = -\text{tr}(\rho \log \rho)$ . Another system has the density matrix  $\tilde{\rho}$  which is proportional to the unit matrix in this basis,  $\tilde{\rho} = \frac{1}{n} \text{diag}(1, 1, \dots, 1)$ . Compute the von Neumann entropy for  $\tilde{\rho}$  and comment on your result. [4 marks]

4. A particle of mass  $m$  is scattered by the potential  $V(r)$  of the form,

$$V(r) = V_0 \tanh^2(\mu r),$$

where  $V_0$  and  $\mu$  are positive constants.

- (a) Sketch the potential  $V(r)$  as a function of  $r$  and prove that the initial momentum  $p$  of the scattered particle with total energy  $E$  is given by  $p = \sqrt{2m(E - V_0)}$ . Show that the time-independent Schrödinger equation in this potential can be written as

$$(\Delta + k^2 - U(r)) \Psi_k(\underline{r}) = 0,$$

where  $k = p/\hbar$  and  $\lim_{r \rightarrow \infty} U(r) = 0$ . Give an explicit expression for  $U(r)$ . [6 marks]

- (b) Use the separation of variables approach and the expression for the Laplace operator in spherical polar coordinates,

$$\Delta = \frac{1}{r^2} \frac{\partial}{\partial r} \left( r^2 \frac{\partial}{\partial r} \right) - \frac{L^2(\theta, \phi)}{\hbar^2 r^2},$$

to separate the Schrödinger equation into equations for the radial and the angular components of the wave-function. State which functions solve the equation for the angular part and give the corresponding eigenvalues of the operator  $L^2(\theta, \phi)$ . [5 marks]

- (c) Now set  $V_0 = \hbar^2 \mu^2 / m$  and, assuming that  $k^2 > 2\mu^2$ , show that the equation for the  $s$ -wave radial component of the wave-function can be written as,

$$\left( \frac{d^2}{d\rho^2} + \frac{2}{\rho} \frac{d}{d\rho} + \eta^2 + \frac{2}{\cosh^2(\rho)} \right) R(\rho) = 0,$$

where  $\eta$  is a real constant. Determine  $\eta$ . [4 marks]

- (d) Using two particular solutions of the equation above,

$$R_{\pm}(\rho) = \frac{e^{\pm i\eta\rho}}{\rho} (\tanh(\rho) \mp i\eta),$$

find the general solution of the  $s$ -wave radial equation. Impose the appropriate boundary condition at the origin and hence find the radial wave-function up to a total normalisation factor. [5 marks]

- (e) What is the corresponding solution of the  $s$ -wave angular component of the wave-function? [2 marks]
- (f) Show that the wave-function in the  $s$ -wave approximation,  $\Psi_{k,0}(\underline{r})$ , can be written for large  $r$  in the form,

$$\lim_{r \rightarrow \infty} \Psi_{k,0}(r) = \frac{C}{r} \sin(kr + \delta_0(k)),$$

where  $C$  is a normalisation constant, and compute the scattering phase  $\delta_0(k)$ . [4 marks]

- (g) Use your result to express the  $s$ -wave cross-section  $\sigma_0$  as a function of the kinetic energy of the initial particle, given that

$$\sigma_0 = \frac{4\pi}{k^2} \sin^2 \delta_0(k). \quad [4 \text{ marks}]$$