

(1)

Equipartition System of classical (non-interacting) particles at temperature T - the particles have degrees of freedom, e.g.

$$\{x_1, y_1, z_1, p_{x_1}, p_{y_1}, p_{z_1}, x_2, y_2, z_2, p_{x_2}, p_{y_2}, p_{z_2}, \dots\}$$

Let the energy per particle be $\mathcal{E} = ax^2$, $a > 0$ then

the average energy per degree of freedom is

$$\langle \mathcal{E} \rangle = \frac{\int_{-\infty}^{\infty} ax^2 e^{-\beta ax^2} dx}{\int_{-\infty}^{\infty} e^{-\beta ax^2} dx} = \frac{a \sqrt{\pi} / 2 (a/k_B T)^{3/2}}{\sqrt{\pi} / (a/k_B T)^{1/2}} = \frac{1}{2} k_B T$$

If the particle has n degrees of freedom then energy per particle is

$$\langle \mathcal{E} \rangle = \frac{n}{2} k_B T$$

②

Equipartition of energy: the energy is, on average, is divided up equally among the n degrees of freedom, each being $\frac{1}{2} k_B T$.

To count degrees of freedom, count the number of quadratic terms.

$$\text{e.g. } \mathcal{E} = a \underset{\uparrow}{x^2} \text{ (1 degree)} \quad \text{or} \quad \mathcal{E} = a \underset{\uparrow}{x^2} + b \underset{\uparrow}{y^2} + c \underset{\uparrow}{z^2} \text{ (3 degrees)}$$

Note, this is the classical result (MB statistics). You can do a similar analysis on quantum particles, noting in the high T limit both FD and BE tend to MB statistics.

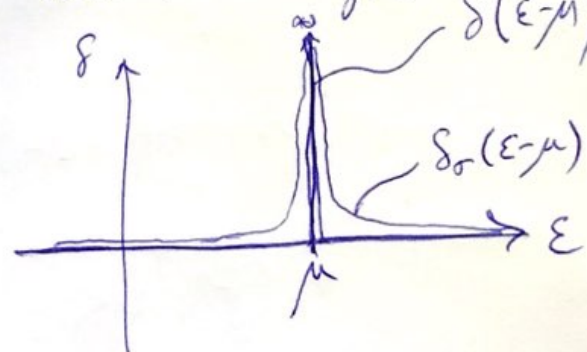
Mathematical Note.

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A δ function centred at μ i.e. $\delta(\epsilon - \mu)$ then the integral

$$\int_{-\infty}^{\infty} \delta(\epsilon - \mu) d\epsilon = 1. \text{ then we also have}$$

$$\int_{-\infty}^{\infty} \delta(\epsilon - \mu) F(\epsilon) d\epsilon = F(\mu).$$



What if instead of δ we had a highly peaked function $\delta_\sigma(\epsilon - \mu)$

$$\text{then } \int_{-\infty}^{\infty} \delta_\sigma(\epsilon - \mu) (\epsilon - \mu)^2 d\epsilon = \sigma^2 \ll 1.$$

In statistical physics the probability ~~p~~ $p(\epsilon)$ is very peaked

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Let's expand some function $F(\epsilon)$ around $\epsilon = \mu$.

$$F(\epsilon) = F(\mu) + (\epsilon - \mu) F'(\mu) + \frac{1}{2} (\epsilon - \mu)^2 F''(\mu) + \dots$$

$$\begin{aligned} \text{Hence } \int p(\epsilon) F(\epsilon) d\epsilon &= F(\mu) \underbrace{\int p(\epsilon) d\epsilon}_1 + F'(\mu) \underbrace{\int p(\epsilon) (\epsilon - \mu) d\epsilon}_0 \\ &\quad + \frac{1}{2} F''(\mu) \underbrace{\int p(\epsilon) (\epsilon - \mu)^2 d\epsilon}_{\sigma^2} + \dots \end{aligned}$$

$$\text{i.e. } \int_{-\infty}^{\infty} p(\epsilon) F(\epsilon) d\epsilon \approx F(\mu) + \frac{\sigma^2}{2} F''(\mu)$$

For the variance $\int p(\epsilon) (\epsilon - \mu)^2 d\epsilon = \sigma^2$ (Book, Appendix C)

$$\left. \int_0^{\infty} \frac{e^y y^2}{(1+e^y)^2} dy = \frac{\pi^2}{6} \right| = F(\mu) + \frac{\pi^2 F''(\mu)}{6} k_B^2 T^2, \quad \sigma^2 = \frac{\pi^2 k_B^2 T^2}{3}$$

Fermi - Dirac Gases.

⑤

These are quantum objects with $\frac{1}{2}$ -integer spin. They can be fundamental (e.g. electrons) or composite, e.g.

- conducting electrons in metals
- liquid He
- white dwarf stars

Any composite object made up of an odd number of fermions, e.g. proton with 3 quarks and ${}^3\text{He}$ not ${}^4\text{He}$.

We have the Fermi-Dirac distribution function.

$$f_{FD}(\epsilon) = \frac{1}{e^{\alpha} e^{\beta \epsilon} + 1} = \frac{1}{e^{\beta(\epsilon - \mu)}} \quad \text{where} \quad \begin{aligned} \beta &= \frac{1}{k_B T} \\ \alpha &= -\beta \mu. \end{aligned}$$

⑥
The constant particle constraint, α , is now appearing as an energy, μ ,
- this is the Fermi-level, sometimes μ is written as E_F .

The distribution $f_{FD}(\epsilon)$ is the average occupation with energy ϵ ,
but for single particle states ϵ_i

$$f_i = n_i / g_i$$

where g_i is the degeneracy of state with energy ϵ_i , having n_i
particles in it. For Fermions $0 \leq f_i \leq 1$.

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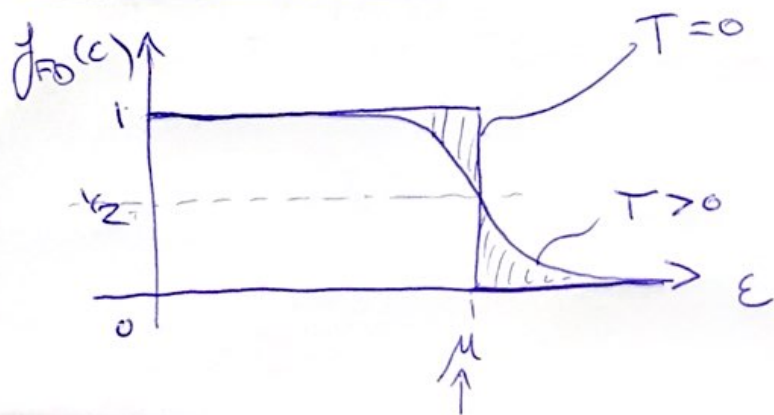
The number of particles in range ϵ to $\epsilon + d\epsilon$ is

$$g(\epsilon) f_{FD}(\epsilon) d\epsilon.$$

The total number of particles N is

$$N = \int_0^{\infty} g(\epsilon) f(\epsilon) d\epsilon \quad (\text{taking } 0 \text{ to be lowest energy}).$$

Properties of $f_{FD}(\epsilon)$



$$f_{FD}(\mu) = 1/2$$

$$\lim_{T \rightarrow 0} f_{FD}(\epsilon < \mu) = 1$$

$$\lim_{T \rightarrow 0} f_{FD}(\epsilon > \mu) = 0$$

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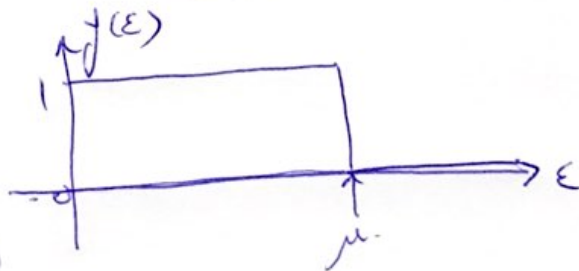
Example. At $(T=0)$ calculate the Fermi level of a system of N particles contained to the 3D ∞ -square well of volume $V=a^3$ (box)

Include spin.

"box" means : $g(\epsilon) d\epsilon = 2V \frac{2\pi}{h^3} (2m)^{3/2} \epsilon^{1/2} d\epsilon$

Note that we need $N = \int_0^{\infty} g(\epsilon) f_{FD}(\epsilon) d\epsilon$. 1 up to μ , 0 otherwise.

$N = \int_0^{\mu} g(\epsilon) d\epsilon$ because



$\Rightarrow N = 2V \cdot \frac{2\pi}{h^3} (2m)^{3/2} \int_0^{\mu} \epsilon^{1/2} d\epsilon$

Do the above integral and rearrange for μ to get

⑨

$$\mu = \frac{\hbar^2}{2m} \left(\frac{3N}{8\pi V} \right)^{2/3}.$$

Recap. if asked for Fermi level we require $g(\epsilon)$ and $f(\epsilon)$ then
integrate $N = \int_0^{\mu} g f d\epsilon$ and rearrange for μ .