

University of Durham

EXAMINATION PAPER

Examination session:

May/June

Year:

2018

Examination code:

PHYS3661-WE01

Title:

Theoretical Physics 3

Time allowed:	3 hours		
Additional material provided:	None		
Materials permitted:	None		
Calculators permitted:	Yes	Models permitted:	Casio fx-83 GTPLUS or Casio fx-85 GTPLUS
Visiting students may use dictionaries:		No	

Instructions to candidates:

- Answer the compulsory question that heads each of sections A and B. These **two** questions have a total of 15 parts and carry 50% of the total marks for the paper.
- Answer **three** of the other questions with **at least one** from each section. If you attempt more than the required number of questions only those with the lowest question number compatible with the rubric will be marked: **clearly delete** the answers that are not to be marked.
- The marks shown in brackets for the main parts of each question are given as a guide to the weighting the markers expect to apply.
- **ANSWER EACH SECTION IN A SEPARATE ANSWER BOOK.**
- Do **not** attach your answer booklets together with a treasury tag, unless you have used more than one booklet for a single section.
- Slip your booklet for Section B inside your booklet for Section A, before they are collected by the invigilator.

Information

Section A: Relativistic Electrodynamics

Section B: Quantum Theory 3

A list of physical constants is provided on the next page.

Revision:

1

Information

Elementary charge:	$e = 1.60 \times 10^{-19} \text{ C}$
Speed of light:	$c = 3.00 \times 10^8 \text{ m s}^{-1}$
Boltzmann constant:	$k_B = 1.38 \times 10^{-23} \text{ J K}^{-1}$
Bohr magneton:	$\mu_B = 9.27 \times 10^{-24} \text{ J T}^{-1}$
Electron mass:	$m_e = 9.11 \times 10^{-31} \text{ kg}$
Gravitational constant:	$G = 6.67 \times 10^{-11} \text{ N m}^2 \text{ kg}^{-2}$
Proton mass:	$m_p = 1.67 \times 10^{-27} \text{ kg}$
Planck constant:	$h = 6.63 \times 10^{-34} \text{ J s}$
Permittivity of free space:	$\epsilon_0 = 8.85 \times 10^{-12} \text{ F m}^{-1}$
Magnetic constant:	$\mu_0 = 4\pi \times 10^{-7} \text{ H m}^{-1}$
Molar gas constant:	$R = 8.31 \text{ J K}^{-1} \text{ mol}^{-1}$
Avogadro's constant:	$N_A = 6.02 \times 10^{23} \text{ mol}^{-1}$
Gravitational acceleration at Earth's surface:	$g = 9.81 \text{ m s}^{-2}$
Stefan-Boltzmann constant:	$\sigma = 5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$
Astronomical Unit:	$\text{AU} = 1.50 \times 10^{11} \text{ m}$
Parsec:	$\text{pc} = 3.09 \times 10^{16} \text{ m}$
Solar Mass:	$M_\odot = 1.99 \times 10^{30} \text{ kg}$
Solar Luminosity:	$L_\odot = 3.84 \times 10^{26} \text{ W}$

SECTION A: RELATIVISTIC ELECTRODYNAMICS

Answer Question 1 and **at least one** of Questions 2, 3 and 4.

1. (a) Express the electric and magnetic fields $\underline{E}, \underline{B}$ in terms of the 4-potential A^μ . [4 marks]
[Hint: See (e) for $F^{\mu\nu}$ written in terms of the physical fields $\underline{E}, \underline{B}$.]
- (b) A particle of rest mass m moves with velocity $c/2$ along the x -axis. A second particle of rest mass $2m$ has the same spatial part of the 4-momentum. What is its velocity? [4 marks]
- (c) Give the definition and physical consequences of two events with 4-vectors x_1^μ and x_2^μ having a timelike and spacelike separation. [4 marks]
- (d) Compute $a^\mu v_\mu$, where a^μ is the 4-acceleration and v^μ is the 4-velocity of a given particle. [4 marks]
- (e) Express the spatial components of the Maxwell equation $\partial_\mu \tilde{F}^{\mu\nu} = 0$ in terms of the electric and magnetic fields. [4 marks]

$$\left[\begin{array}{l} \text{Hint:} \\ F^{\mu\nu} = \begin{pmatrix} 0 & -E_x & -E_y & -E_z \\ E_x & 0 & -cB_z & cB_y \\ E_y & cB_z & 0 & -cB_x \\ E_z & -cB_y & cB_x & 0 \end{pmatrix} \\ \tilde{F}^{\mu\nu} = \begin{pmatrix} 0 & cB_x & cB_y & cB_z \\ -cB_x & 0 & -E_z & E_y \\ -cB_y & E_z & 0 & -E_x \\ -cB_z & -E_y & E_x & 0 \end{pmatrix} \end{array} \right]$$

- (f) The 4-potential for a parallel-plate capacitor oriented normal to the y -axis is $A^\mu = (Ey, 0, 0, 0)$, where E is the electric field between the plates and y the distance from the negatively charged plate. Show that in a frame moving with velocity v in the x -direction, there is a magnetic field in the z -direction, and calculate its magnitude. [4 marks]
- (g) If $(\underline{E}_1, \underline{B}_1)$ and $(\underline{E}_2, \underline{B}_2)$ are the electromagnetic fields from two different sources, show that $\underline{E}_1 \cdot \underline{E}_2 - c^2 \underline{B}_1 \cdot \underline{B}_2$ and $\underline{E}_1 \cdot \underline{B}_2 + \underline{B}_1 \cdot \underline{E}_2$ are Lorentz invariants. [4 marks]
- (h) Write the gauge-transformation of the 4-potential A^μ in contra-variant form. Use this to show that the elements of the field strength tensor are unchanged under a gauge transformation. [4 marks]

2. (a) A photon with frequency ν moves along the x -axis before it scatters off an electron, with mass m , which is initially at rest. Compute the frequency of the scattered photon, ν' , as a function of the scattering angle of the photon, i.e. the angle the outgoing photon makes with the x -axis. Show that for a non-vanishing scattering angle, the photon always loses energy in the collision. [6 marks]
- (b) Now consider the process where the electron also moves, with velocity $\underline{v} = (-v, 0, 0)$ head-on towards the incoming photon which is travelling in the opposite direction. Compute the frequency ν' of the scattered photon as a function of the scattering angle. [8 marks]
- (c) Find the minimal velocity of the electron such that the process results in a gain of energy for the photon. [6 marks]

3. A particle of charge q and rest mass m travels with a relativistic velocity \underline{v}_0 as it enters a medium, where it is slowed down by a force proportional to the velocity \underline{v} of the particle, $\underline{F} = -\alpha \underline{v}$, where $\alpha > 0$ is a constant. In the following you should ignore the effects of radiation back-reaction.

- (a) Show that for motion along one direction the following holds

$$\frac{d(\gamma m v)}{dt} = m \gamma^3 \frac{dv}{dt} = m \gamma^3 a.$$

Use this to show that after the particle has entered the medium the acceleration $\underline{a} = \frac{dv}{dt}$ is related to the velocity of the particle. [4 marks]

The electromagnetic fields generated by a point charge q in vacuum and in arbitrary motion are given by

$$\underline{E}(\underline{r}, t) = \frac{q}{4\pi\epsilon_0} \frac{R}{(\underline{R} \cdot \underline{u})^3} [(c^2 - v^2)\underline{u} + \underline{R} \times (\underline{u} \times \underline{a})],$$

$$\underline{B}(\underline{r}, t) = \frac{1}{c} \hat{\underline{R}} \times \underline{E}(\underline{r}, t),$$

where \underline{R} is the vector between the point charge and the observer, \underline{v} is the velocity of the point charge, $\underline{u} = c\hat{\underline{R}} - \underline{v}$, and \underline{a} is the acceleration of the point charge. \underline{R} , \underline{u} , \underline{v} , and \underline{a} are all evaluated at the retarded time. $\hat{\underline{R}}$ is a unit vector in the direction of \underline{R} : $\hat{\underline{R}} = \underline{R}/|\underline{R}|$. In the following you can assume that within the medium the radiation propagates as in vacuum.

- (b) Identify the electric radiation field from the equations above, and show that for the trajectory in question it is given by

$$\underline{E}_{\text{rad}}(\underline{r}, t) = \frac{c}{4\pi\epsilon_0} \frac{q}{R} \frac{1}{(\hat{\underline{R}} \cdot \underline{u})^3} \left[\left(\hat{\underline{R}} \cdot \underline{a} \right) \hat{\underline{R}} - \underline{a} \right].$$

[4 marks]

- (c) Show that the Poynting vector for the radiation fields is given by

$$\underline{S}_{\text{rad}} = \frac{\mu_0 q^2 a^2}{16\pi^2 c} \frac{1}{R^2} \frac{\sin^2 \theta}{(1 - \beta \cos \theta)^6} \hat{\underline{R}},$$

where θ is the angle between \underline{R} and \underline{v} , and $\beta = v/c$. [4 marks]

The total power radiated to infinity by the point charge at the time t is given by Liénard's result

$$P = \frac{\mu_0 q^2 \gamma^6}{6\pi c} \left(\underline{a}^2 - \left| \frac{\underline{v} \times \underline{a}}{c} \right|^2 \right).$$

- (d) Use Liénard's result and the result of (a) to calculate the total energy emitted as electromagnetic radiation by the particle as it slows down from the initial velocity v_0 to complete rest. [8 marks]

4. Consider a point charge q of rest mass m in an electromagnetic field. The field strength tensor is given by

$$F^{\mu\nu} = \begin{pmatrix} 0 & -E_x & -E_y & -E_z \\ E_x & 0 & -cB_z & cB_y \\ E_y & cB_z & 0 & -cB_x \\ E_z & -cB_y & cB_x & 0 \end{pmatrix}.$$

The 4-force f^μ acting on the point charge is defined by

$$f^\mu \equiv \frac{dp^\mu}{d\tau},$$

where τ is the proper time and p^μ the 4-momentum of the point charge.

- (a) Show that f^μ is a 4-vector, and demonstrate how it is related to the usual force $\underline{F} = d\underline{p}/dt$. [4 marks]
- (b) The 4-force acting on the point charge due to the electromagnetic field is given by

$$f^\mu = \frac{q}{c} F^{\mu\nu} u_\nu,$$

where u_ν is the 4-velocity of the point charge. Show that the spatial components of this equation correspond to the Lorentz force law. Also, interpret the equation obtained from the 0-component of the above equation. [8 marks]

- (c) The point charge moves under the influence of a uniform magnetic field \underline{B} . Starting with the assumption that the velocity \underline{v} of the point charge is perpendicular to \underline{B} , show that the point charge moves with constant speed in a circle. Ignoring radiation from the charge, compute the radius of this circle in terms of the magnitude of the magnetic field, the mass, m , and speed, v , of the point charge. [6 marks]

What happens if the initial velocity is not perpendicular to \underline{B} ? [2 marks]

SECTION B: QUANTUM THEORY 3

Answer Question 5 and **at least one** of Questions 6, 7 and 8.

5. (a) [This sub-question has been deleted as a result of strike action.]
 (b) The differential cross-section is defined by the well-known formula

$$\frac{d\sigma}{d\Omega} = \frac{1}{F} \frac{dn(\theta, \phi)}{d\Omega}.$$

Define all the symbols in this expression. Verify that the cross-section has the appropriate units. [4 marks]

- (c) What are the common eigenfunctions of the orbital momentum operators \underline{L}^2 and L_z and what are their eigenvalues? Write down the partial wave expansion of a wave-function $\psi(r, \theta)$ in a centrally symmetric potential. [4 marks]
- (d) Imagine that a physical system or a process is described by a wave-function which is a real-valued function of the three-dimensional space. Can such a wave-function describe a scattering process and if so, what would be the values for the total cross-section, the phase shifts, and the scattering amplitudes? [4 marks]
- (e) [This sub-question has been deleted as a result of strike action.]
- (f) [This sub-question has been deleted as a result of strike action.]
- (g) The density matrix operator ρ is defined as follows,

$$\rho = \sum_{\alpha=1}^N |\alpha\rangle W_{\alpha} \langle\alpha|.$$

Explain what we are summing over in the expression above and what is the meaning of W_{α} . What is the value of $\text{Tr}(\rho)$ when pure states $|\alpha\rangle$ are normalised to 1? [4 marks]

6. A particle of mass m with momentum $\underline{p} = \hbar \underline{k}$ is scattered by the central potential $V(r)$ given by,

$$V(r) = \begin{cases} -V_0 \frac{1}{r} & : \text{ for } 0 \leq r < r_0 \\ -V_1 \frac{e^{-\mu r}}{r} & : \text{ for } r_0 \leq r < \infty \end{cases}$$

where V_0 , r_0 and μ are positive constants, and $V_1 = V_0 e^{\mu r_0}$, so that the potential is continuous at $r = r_0$.

- (a) The expression for the scattering amplitude in the first Born approximation, f_B , is given by

$$f_B = -\frac{1}{\Delta} \int_0^\infty dr r \sin(\Delta r) U(r), \quad \text{where} \quad U(r) = \frac{2m}{\hbar^2} V(r).$$

Define all the quantities in this formula. What is the physical meaning of Δ ? [3 marks]

- (b) Compute the scattering amplitude f_B by evaluating the integral in (a) for the potential $V(r)$ given above. [8 marks]
- (c) Consider the limit $\Delta r_0 \ll 1$ and simplify your expression for the scattering amplitude working to the order $(\Delta r_0)^1$. Hence evaluate the differential cross-section in this approximation. [2 marks]
- (d) Compute the total cross-section for this scattering process in the same limit as in (c). [4 marks]
- (e) Show that at high energies, the cross-section in (d) can be written as

$$\sigma_B(E) = \frac{a}{E} + \frac{b}{E^2} + \mathcal{O}\left(\frac{1}{E^3}\right)$$

where E is the energy of the incoming particle. Give expressions for the constants a and b in terms of V_0 , μ , m and r_0 . [3 marks]

7. Particles of mass m are scattered by the central potential

$$V(r) = \frac{\lambda}{r^2},$$

where λ is a positive constant.

- (a) Write down the time-independent Schrödinger equation for this problem. Explain briefly how the equation

$$\left(\frac{d^2}{dr^2} + \frac{2}{r} \frac{d}{dr} + k^2 - \frac{l(l+1)}{r^2} - \frac{2m\lambda}{\hbar^2} \frac{1}{r^2} \right) R_{kl} = 0$$

for the radial component $R_{kl}(r)$ of the wave function arises and all the symbols used in the expression above. [2 marks]

- (b) Specify the regular solutions for $R_{kl}(r)$ far away from the origin in terms of free spherical waves given by spherical Bessel functions. Show that this leads to the asymptotic expression for $R_{kl}(r)$ at $r \rightarrow \infty$

$$R_{kl}(r) \rightarrow A_l(k) \frac{1}{kr} \sin \left(kr - \frac{\pi l}{2} + \delta_l(k) \right)$$

where $\delta_l(k)$ are the phase-shifts and $A_l(k)$ are some r -independent coefficients. [4 marks]

- (c) By comparing the free case, $\lambda = 0$, with the interacting case, $\lambda \neq 0$, prove that the phase-shifts are given by,

$$\delta_l = \frac{\pi}{2} \left(l + \frac{1}{2} - \sqrt{\left(l + \frac{1}{2} \right)^2 + \frac{2m\lambda}{\hbar^2}} \right)$$

and simplify this expression in the weak-coupling limit $2m\lambda/\hbar^2 \ll 1$. [8 marks]

- (d) Compute the differential cross section using the expression that is valid for $\delta_l \ll 1$,

$$\frac{d\sigma}{d\Omega} = \frac{1}{k^2} \left| \sum_{l=0}^{\infty} (2l+1) \delta_l P_l(\cos \theta) \right|^2.$$

[6 marks]

$$\left[\text{Hint: You can use the summation formula: } \sum_{l=0}^{\infty} P_l(\cos \theta) = \frac{1}{2 \sin(\theta/2)}. \right]$$

8. [This question has been deleted as a result of strike action.]