## **GA** 4

(1) Consider a spherical density distribution, and denote by M(< R) the mass enclosed in a sphere of radius R. The speed,  $V_c(R)$ , of a particle on a circular orbit with radius R in such a density distribution, is given by

$$V_c^2(R) = \frac{GM(\langle R)}{R} \,. \tag{1}$$

Demonstrate that a constant circular velocity,  $V_c = V_{c,0}$ , implies that the density distribution is  $\rho(R) = V_{c,0}^2/(4\pi GR^2)$ . [2 marks]

(2) The surface density,  $\sigma(R)$ , of stars in a spiral galaxy falls with distance R to the centre as

$$\sigma(R) = \sigma_d \exp(-R/R_d)$$
.

Here,  $\sigma_d$  is the central surface density, and  $R_d$  is the scale-length. Both are constants. Demonstrate that the disc mass enclosed in a sphere of radius R is

$$M_d(< R) = 2\pi \sigma_d R_d^2 (1 - (1 + x) \exp(-x)),$$

where 
$$x \equiv R/R_d$$
. [2 marks]

The relation between enclosed mass and circular velocity given in Eq.(1) is only approximately valid for a disc. However you may assume it does hold in what follows.

- (3) Take  $R_d=3$  kpc,  $M_d(R\to\infty)=2\times 10^{10}~M_\odot$ , and assume that the circular velocity of disc and dark halo combined is  $V_c$ =220 km s<sup>-1</sup> at distance R=100 kpc. Evaluate  $V_c$  at R=5 kpc and R=10 kpc. [3 marks]
- (4) This galaxy is at a distance of d=10 Mpc, and its disc is tilted by 30 degrees (with 90 degrees corresponding to the case where the galaxy is face on). An observer uses a radio telescope to detect gas in the disc moving with the circular velocity using the HI 21-cm line. Sketch the detected wavelength of the line as function of angle from the centre of the galaxy. [3 marks]

$$[1 \text{ pc} = 3.09 \times 10^{16} \text{ m}, M_{\odot} = 2.0 \times 10^{30} \text{ kg}, G = 6.67 \times 10^{-11} \text{ m}^3 \text{ kg}^{-1} \text{ s}^{-2}]$$