

Counting Microstates.

Lecture 8

Recall that for N particles, have a large number of microstates, Ω , correspond to distributions $\{n_1, n_2, n_3, \dots\}$. We're seeking the distribution with the largest number of microstates.

Need to find $\{n_1, n_2, \dots\}$ with constraints

$$\sum_i n_i = N \quad \text{and} \quad \sum_i n_i \epsilon_i = U.$$

$$\text{for which } \Omega(\{n_i\}) = \frac{N!}{\prod_i n_i!} \text{ is maximum.}$$

What changes if energy level ϵ_i has degeneracy g_i .

⑤

- there are more microstates associated with the n_1 particles in ϵ_1 , they can be in any of the g_1 degenerate states.

$$\Omega(\{n_i\}) \rightarrow N! \frac{g_1^{n_1}}{\prod_i n_i!}$$

What about making state ϵ_2 degenerate with degeneracy g_2 ?

By the same reasoning

$$\Omega(\{n_i\}) \rightarrow N! \frac{g_1^{n_1} g_2^{n_2}}{\prod_i n_i!}$$

①

This is generalise with

$$\mathcal{R}(\{n_i\}) \rightarrow$$

$$n! \prod_i \frac{g_i^{n_i}}{n_i!}$$

where n_i is the number of distinguishable particles in state ϵ_i which is g_i -fold degenerate.

Maxwell-Boltzmann distribution - fractional occupancies - most probable distribution of fractional occupancies

$$f_{MB}(\epsilon_i) = n_i/g_i \text{ of levels } \epsilon_i$$

⑤.

To do this we will maximise the entropy.

$$S_0 : \quad \frac{S}{k_B} = \ln \left[\prod_i \frac{g_i^{n_i}}{n_i!} \right] + \overbrace{\ln N!}^{\text{constant, so derivative will be zero. (ignore)}}.$$

$$= \sum_i n_i \ln g_i - \ln n_i! \quad (\text{split with log-rules}).$$

$$\approx \sum_i n_i \ln g_i - (n_i \ln n_i - n_i) \quad (\text{Stirling}).$$

Take constraints of constant particle number and constant internal energy into account with Lagrange multipliers.

⑧.

$$S/k_B - \alpha N - \beta U = \sum_i (n_i \ln g_i - n_i \ln n_i + n_i - \alpha n_i - \beta n_i \epsilon_i)$$

Vary the distribution, i.e. $\partial/\partial n_i \rightarrow 0$ (find maximum by varying the occupation numbers).

$$0 = \frac{\partial}{\partial n_i} \{ \text{above} \} = \ln g_i - \ln n_i - 1 + 1 - \alpha - \beta \epsilon_i$$

$$\Rightarrow \ln \left(\frac{g_i}{n_i} \right) = \alpha + \beta \epsilon_i$$

⑨

$$\Rightarrow \frac{n_i}{g_i} = f_{MB}(\epsilon_i) = e^{-\alpha} e^{-\beta \epsilon_i}.$$

This is the Maxwell Boltzmann distribution function.

As we saw previously, β is inverse temperature, and α is determined by fixed N ,

$$\text{i.e. } \int f_{MB}(\epsilon) d\epsilon = N.$$

This is for distinguishable/classical particles.