

Stars and Galaxies  
**Observational Techniques Workshop 3**

- 1) What are the four main components of a spectrograph?
- 2) State the resolving power,  $R$ , in terms of wavelength. An astronomical slit spectrograph has a resolving power of  $R = 7000$ . What is the smallest wavelength difference that can be discerned at a wavelength of 700 nm?
- 3) The spectrum of a nearby star peaks at a wavelength of 3000 Å. It has a parallax of 379 milli-arcseconds and a  $V$ -band magnitude of  $V = -1.5$ . Sketch a HR diagram and place the star on it. What type of star is it? [The Sun has a  $V$ -band absolute magnitude of 4.8]
- 4) Write down the reflection grating equation of a spectrograph, and define the symbols. The sodium D line is a doublet of wavelengths 589.0 nm and 589.6 nm. A grating produces a first order spectrum where these two lines are diffracted at angles of  $42.69^\circ$  and  $42.78^\circ$  (with respect to the grating normal). Calculate the ruling density of the grating (in units of lines  $\text{mm}^{-1}$ ).
- 5) The two Keck telescopes are each 10 m across and separated by 85 meters. What is the maximum distance from Earth that an extra-solar planet will be resolved if observed with (i) a single telescope; (ii) both telescopes when used as an interferometer? Assume the extra-solar planet is located 1 AU from the star ( $1 \text{ AU} = 1.50 \times 10^{11} \text{ m}$ ).
- 6) (i) Write down two properties of a star or galaxy that can be measured with a spectrograph.  
(ii) The resolving power,  $R$ , of a slit spectrograph that employs a diffraction grating is given by  $R = m\rho\lambda W / (\chi D_T)$ . Define all the terms in this expression for  $R$ .  
(iii) For a spectrograph equipped with a reflection grating of length 10.4 cm with input angle  $\alpha = 20^\circ$  and output angle  $\beta = 15^\circ$ , find the spectrograph's resolving power if used on an 8-m telescope with a slit width of 0.5 arcsec.  
(iv) Calculate the spectral resolution at a wavelength of 500 nm if the telescope is put in space.
- 7) [from 2013 June Exam]

An astronomical slit spectrograph has a resolving power of  $R = 7000$ . What is the smallest wavelength difference that can be discerned at a wavelength of 700 nm? [3 marks]

The same spectrograph is used to observe an active galaxy nuclear emission line with flux  $1 \times 10^{-20} \text{ W m}^{-2}$ . What is the total power that reaches the spectrograph detector when it is used on a telescope with a 6-m diameter mirror? You may assume that the telescope and spectrograph combination have 100% quantum efficiency. [4 marks]

If the emission line is just contained within a single pixel on the spectrograph CCD detector, what is the signal-noise ratio ( $S/N$ ) of the measurement of the emission line strength in a  $t = 1000 \text{ s}$  exposure if the sky flux is 10 times higher than the total emission line flux? You may assume that the dark noise is negligible and that the read-out noise is  $\pm 5$  electrons per pixel. [6 marks]

Is the  $S/N$  of the observation approximately sky-noise or read-out noise limited? [2 marks]

If the slit width doubles what is the new resolving power,  $R$ , of the spectrograph? [2 marks]

What is the new  $S/N$  of the measured emission line strength? [3 marks]