

# Computational Études

*A Spectral Approach*

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# Preface

**T**he purpose of this book is twofold: to explain to students why spectral methods work and why they are so remarkably efficient, and to teach them how to implement these methods using modern computational tools.

This text is conceived as a collection of **Computational Études**. In musical education, an étude is a composition designed to perfect a specific technique (be it rapid scales, complex arpeggios, or delicate phrasing) while remaining a pleasing piece of music in its own right. Similarly, each chapter here presents a focused mathematical concept paired with its computational realization: a study that is both instructive and complete.

### *Who Is This Book For?*

We have written primarily for graduate students in applied mathematics, physics, and engineering who seek a hands-on understanding of spectral methods. The reader should be comfortable with calculus, linear algebra, and basic programming. Prior exposure to numerical methods is helpful but not essential; we build the necessary foundations as we proceed.

### *How Is the Book Organized?*

Each chapter is designed to be largely self-contained. We begin with fundamental concepts (interpolation and differentiation) before advancing to time-stepping schemes and applications. The mathematical exposition is deliberately concise, favoring clarity over exhaustive rigor. Proofs are included when they illuminate; otherwise, we direct the reader to authoritative references.

### *Reproducible Science*

This book is also an experiment in **reproducible science**. Every figure, every table, every numerical result you see in these pages was generated by code available in the accompanying repository. We provide implementations in both Python and MATLAB, allowing readers to choose their preferred environment. The Python code emphasizes accessibility and integration with the open-source scientific ecosystem; the MATLAB code leverages its historical significance in numerical computing and, where appropriate, the Advanpix Multiprecision Computing Toolbox for extended precision arithmetic.

### *How to Use This Book*

We invite you to treat this book not as a static reference, but as a workshop. Clone the repository, run the scripts, modify the parameters, break the code, and fix it. That is the only way to truly master the art of spectral methods.

The complete source code and the Typst manuscript are available at:  
<https://github.com/dutykh/computational-etudes/>

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## CHAPTER 1

# Introduction

Differential equations serve as the fundamental language of the physical sciences, describing phenomena ranging from the propagation of sound waves to the flow of heat and the dynamics of fluids. Finding exact analytical solutions to these equations is a luxury rarely afforded in practical applications. Consequently, the scientist and the engineer must turn to numerical approximation.

Broadly speaking, numerical methods for differential equations fall into two categories: local methods and global methods. The former, including Finite Difference and Finite Element Methods, approximate the unknown solution using functions that are non-zero only on small sub-domains (elements or grid stencils). These methods are robust and flexible, handling complex geometries with grace. However, their accuracy is typically algebraic; refining the grid by a factor of two might improve the error by a factor of four or eight, but rarely more. From a computational perspective, local methods are *myopic*: to compute a derivative at a grid point, they look only at immediate neighbors.

Spectral methods represent the global approach. They approximate the solution as a linear combination of continuous, global basis functions, typically trigonometric polynomials (Fourier series) for periodic problems or Chebyshev polynomials for non-periodic ones. In stark contrast to local schemes, spectral methods are *holistic*: the derivative at any single point depends on the function values at *every other point* in the domain. Mathematically, this is equivalent to fitting a single high-degree polynomial through all data points. This global coupling is what allows information to propagate instantly across the grid, granting us the remarkable convergence that we call “spectral accuracy.”

There is, however, a unifying viewpoint that bridges these apparently distinct philosophies. Pseudospectral methods can be understood as the natural limit of finite difference methods as the order of accuracy increases without bound. Imagine a sequence of finite difference stencils: first using two neighbors, then four, then eight, and so on. As the stencil width grows to encompass the entire grid, the local method transforms smoothly into a global one. This perspective, developed by Kreiss and Oliger and elaborated by Fornberg, offers both intuition and practical computational strategies. The theory and practice of spectral methods are comprehensively developed in the classical texts by [1], [2], and [3].

## 1.1 The Spectral Promise

The fundamental argument for spectral methods is one of efficiency. If the solution to a problem is smooth, the coefficients of its expansion in a proper global basis decay exponentially fast. This phenomenon is known as spectral accuracy.

In practical terms, this means that spectral methods can achieve a level of precision with a few dozen degrees of freedom that a finite difference scheme might require thousands of grid points to match. While a fourth-order finite difference method implies that the error  $\varepsilon \sim O(N^{-4})$ , a spectral method boasts  $\varepsilon \sim O(c^{-N})$  for some constant  $c > 1$ . When the solution is analytic, the convergence is explosive; the error drops into the “spectral valley” until it hits the floor of machine precision.

Beyond raw accuracy, spectral methods possess another virtue that is often decisive in physical applications: they are virtually free of both dissipative and dispersive numerical errors. Finite difference schemes, by their very nature, introduce artificial dissipation that can overwhelm the true physical dissipation in problems such as high-Reynolds number fluid flows. They also suffer from dispersive errors that cause different frequency components to propagate

at slightly different speeds, turning sharp gradients into spurious wavetrains. Spectral methods avoid both pathologies. This fidelity to the underlying physics explains their dominance in demanding applications: turbulence modeling, global weather and climate prediction, nonlinear wave dynamics, and seismic analysis all rely heavily on spectral techniques.

This global dependence has a computational consequence: spectral differentiation matrices are *dense*, not sparse. Where a finite difference scheme produces banded matrices that are cheap to store and invert, spectral methods fill in every entry. However, the extraordinary accuracy means we need so few points (often just dozens where finite differences would require thousands) that we can afford this density. The cost per degree of freedom is higher, but the total cost for a given accuracy is dramatically lower.

However, this power is not without its price. Spectral methods are unforgiving regarding grid placement. We cannot simply choose points where we please; for non-periodic problems, the mathematics dictates that points must cluster at boundaries (the celebrated Chebyshev points) to prevent the interpolation from diverging. Attempting high-degree polynomial interpolation on an equispaced grid leads to the notorious Runge phenomenon, where oscillations grow without bound near the boundaries. This sensitivity to geometry is what restricts spectral methods primarily to simple domains, but within those domains, they reign supreme.

This book aims to demystify this “spectral magic.” We will see that it is not magic at all, but a direct consequence of the smoothness of the underlying functions and the careful choice of basis and grid.

## 1.2 A Brief History

Spectral representations have been used for analytic studies of differential equations since the days of Fourier in the early nineteenth century. The idea of employing them for *numerical* computation, however, emerged much later. Lanczos, in the 1930s, pioneered the use of Chebyshev expansions for solving ordinary differential equations numerically. This approach remained somewhat academic until the early 1970s, when Kreiss and Olinger introduced the pseudospectral method for partial differential equations. Their innovation was to work with function values at grid points rather than expansion coefficients, dramatically simplifying the treatment of nonlinear terms. The practical viability of these methods was ensured by the fast Fourier transform algorithm, rediscovered by Cooley and Tukey in 1965, which reduced the cost of transforming between physical and spectral space from  $O(N^2)$  to  $O(N \log N)$  operations. This confluence of theoretical insight and algorithmic efficiency launched spectral methods into the mainstream of computational science.

## 1.3 Limitations and Trade-offs

Honesty compels us to acknowledge that spectral methods are not a panacea. Several factors can limit their applicability or efficiency:

- **Boundary conditions** can be awkward to impose, particularly for problems with complex constraints or time-dependent boundaries.

- **Irregular domains** resist the tensor-product structure that makes spectral methods efficient. While domain decomposition and mapping techniques exist, they sacrifice some of the method’s elegance.
- **Strong shocks and discontinuities** violate the smoothness assumptions that underlie spectral accuracy. The Gibbs phenomenon produces persistent oscillations near discontinuities, requiring filtering or other remediation.
- **Variable resolution requirements** across a large domain are difficult to accommodate. Unlike adaptive mesh refinement in finite element methods, spectral grids are inherently uniform in their polynomial degree.

These limitations explain why finite element and finite difference methods continue to thrive in many applications. The art lies in recognizing when spectral methods are the right tool. When the geometry is simple, the solution is smooth, and high accuracy is paramount, no other approach comes close.

## 1.4 The Philosophy of “Études”

The title of this volume, Computational Études, reflects a specific pedagogical philosophy. In musical education, an étude is a composition designed to practice a particular technical skill (be it rapid scales or complex arpeggios) while remaining a pleasing piece of music in its own right.

In this text, our “technical skills” are not rapid scales or arpeggios, but rather handling stiffness in time-stepping, managing aliasing in nonlinear products, enforcing boundary conditions through tau methods or lifting functions, and filtering spurious oscillations. Just as a Chopin Étude transforms a technical exercise into art, a well-written spectral code transforms a mathematical formula into a robust simulation. The études collected here are designed to cultivate this virtuosity.

We approach spectral methods not through dry, abstract theorems, but through concrete, self-contained studies. Each chapter focuses on a specific mathematical concept (interpolation, differentiation, aliasing, or time-stepping) and explores it through a compact, runnable implementation.

We deliberately restrict our focus primarily to one-dimensional problems. This choice is strategic. The mathematical essence of spectral methods (the treatment of boundaries, the distribution of collocation points, and the structure of differentiation matrices) is fully present in one dimension. Extending these ideas to two or three dimensions usually involves tensor products, which add significant programming overhead without necessarily adding new conceptual depth. By staying in 1D, we keep our code short, readable, and focused on the physics and mathematics.

## 1.5 Collocation: Computing in Physical Space

While the theory of spectral methods relies on orthogonal expansions (Fourier series for periodic problems, Chebyshev series otherwise), the actual computation often proceeds differently. Rather than manipulating expansion coefficients directly (the *modal* or *Galerkin* approach), we

typically work with function values at carefully chosen grid points (the *nodal* or *collocation* approach, also called *pseudospectral*).

The collocation philosophy dominates this book for a practical reason: it handles nonlinear terms with ease. When the governing equation contains products like  $u \cdot u_x$ , the modal approach requires computing convolutions of coefficient sequences, a tedious operation. Collocation simply evaluates the product pointwise on the grid. This directness makes pseudospectral methods the tool of choice for most computational applications, and it is the approach we shall master through these études.

The reader should be aware that both viewpoints (modal and nodal) illuminate the same underlying mathematics. The Fast Fourier Transform provides the bridge, allowing us to move efficiently between coefficient space and physical space as needed.

## 1.6 A Modern Workflow

Finally, this book is an experiment in reproducible science. The days of presenting numerical results as static, unverifiable images are passing. The results you see in these pages were generated by the code available in the accompanying repository. We utilize a dual-language approach:

- **Python:** For accessibility and integration with the vast open-source scientific ecosystem.
- **Matlab:** For its historical significance in this field and its concise matrix syntax, often utilizing the Advanpix Multiprecision Computing Toolbox to explore phenomena that lie beyond standard double precision.

We invite you to treat this book not as a static reference, but as a workshop. Run the scripts, change the parameters, break the code, and fix it. That is the only way to truly learn the art of spectral methods.





## CHAPTER 2

# Classical Second Order PDEs and Separation of Variables

In this opening chapter we derive exact solutions for three classical linear partial differential equations: the *heat equation* (parabolic), the *wave equation* (hyperbolic), and the *Laplace equation* (elliptic). These solutions are found by the *method of separation of variables*, which expresses the solution as an infinite series of eigenfunctions.

Why begin a book on *numerical* methods with *analytical* solutions? Because separation of variables is the theoretical ancestor of spectral methods. When we later truncate these infinite series at some finite  $N$  and compute with only the first  $N$  modes, we are doing exactly what a spectral solver does, but with pen and paper first. This chapter thus serves as the conceptual bridge between classical analysis and modern computation. The methods presented here are classical and thoroughly developed in standard references such as [4].

We treat three model problems:

- heat equation with periodic boundary conditions in one spatial dimension,
- wave equation on a bounded interval,
- Laplace equation in a simple domain.

We begin with a complete analytic solution of the heat equation. The other two examples will follow the same pattern.

## 2.1 Heat Equation with Periodic Boundary Conditions

We consider the one dimensional heat equation on the interval  $[0, 2\pi]$  with periodic boundary conditions. The unknown  $u(x, t)$  represents, for example, the temperature at point  $x$  and time  $t$ .

The problem is

$$\frac{\partial u}{\partial t}(x, t) = \frac{\partial^2 u}{\partial x^2}(x, t), \quad x \in [0, 2\pi], \quad t > 0,$$

with periodic boundary conditions

$$u(x + 2\pi, t) = u(x, t)$$

for all real  $x$  and all  $t > 0$ , and initial condition

$$u(x, 0) = f(x), \quad x \in [0, 2\pi].$$

We assume that  $f$  is smooth and  $2\pi$  periodic:

$$f(x + 2\pi) = f(x).$$

Our goal is to obtain an explicit representation of  $u(x, t)$  as an infinite series and to see how separation of variables leads naturally to a Fourier series in space. The technique we employ here follows the classical approach presented in [4].

### 2.1.1 Step 1: Separation Ansatz

We look for nontrivial solutions of the form

$$u(x, t) = X(x) \cdot T(t),$$

where  $X$  depends only on  $x$  and  $T$  depends only on  $t$ .

Substituting into the PDE gives

$$X(x) \cdot T'(t) = X''(x) \cdot T(t).$$

We assume  $X$  and  $T$  are not identically zero, so we can divide both sides by  $X(x) \cdot T(t)$ :

$$\frac{T'(t)}{T(t)} = \frac{X''(x)}{X(x)}.$$

The left side depends only on  $t$ , the right side only on  $x$ . Therefore both sides must be equal to the same constant, which we denote by  $-\lambda$ :

$$\frac{T'(t)}{T(t)} = \frac{X''(x)}{X(x)} = -\lambda.$$

We obtain two ordinary differential equations:

$$\begin{aligned} T'(t) + \lambda T(t) &= 0, \\ X''(x) + \lambda X(x) &= 0. \end{aligned}$$

The periodic boundary conditions for  $u$  imply periodic conditions for  $X$ :

$$X(0) = X(2\pi), \quad X'(0) = X'(2\pi).$$

We have arrived at a spatial eigenvalue problem for  $X$ . The systematic treatment of such eigenvalue problems is a cornerstone of the theory of partial differential equations; see [4] for a comprehensive exposition.

### 2.1.2 Step 2: Spatial Eigenvalue Problem with Periodic Boundary Conditions

We now solve

$$X''(x) + \lambda X(x) = 0,$$

with

$$X(0) = X(2\pi), \quad X'(0) = X'(2\pi).$$

We consider three cases:  $\lambda < 0$ ,  $\lambda = 0$ , and  $\lambda > 0$ .

#### Case 1: $\lambda < 0$

Write  $\lambda = -\mu^2$  with  $\mu > 0$ . The equation becomes

$$X''(x) - \mu^2 X(x) = 0.$$

The general solution is

$$X(x) = Ae^{\mu x} + Be^{-\mu x}.$$

Imposing periodicity  $X(0) = X(2\pi)$  gives

$$A + B = Ae^{2\mu\pi} + Be^{-2\mu\pi}.$$

Imposing  $X'(0) = X'(2\pi)$  gives

$$\mu(A - B) = \mu(Ae^{2\mu\pi} - Be^{-2\mu\pi}).$$

Since  $\mu > 0$ , we can divide by  $\mu$  and rewrite both conditions as a homogeneous linear system:

$$\begin{aligned} A(1 - e^{2\mu\pi}) + B(1 - e^{-2\mu\pi}) &= 0, \\ A(1 - e^{2\mu\pi}) - B(1 - e^{-2\mu\pi}) &= 0. \end{aligned}$$

Adding these two equations gives

$$2A(1 - e^{2\mu\pi}) = 0.$$

Since  $\mu > 0$ , we have  $e^{2\mu\pi} > 1$ , so  $1 - e^{2\mu\pi} \neq 0$ . Therefore  $A = 0$ .

Subtracting the second equation from the first gives

$$2B(1 - e^{-2\mu\pi}) = 0.$$

Since  $\mu > 0$ , we have  $e^{-2\mu\pi} < 1$ , so  $1 - e^{-2\mu\pi} \neq 0$ . Therefore  $B = 0$ .

Hence the only solution is  $A = B = 0$ , the trivial solution. There are no nontrivial periodic eigenfunctions for  $\lambda < 0$ , and we discard this case.

#### Case 2: $\lambda = 0$

The equation reduces to

$$X''(x) = 0.$$

Its general solution is

$$X(x) = A + Bx.$$

Periodicity  $X(0) = X(2\pi)$  gives

$$A = A + 2\pi B$$

so  $B = 0$ . Then  $X(x) = A$  is constant.

The derivative is  $X'(x) = 0$ , so  $X'(0) = X'(2\pi)$  is automatically satisfied.

Thus  $\lambda = 0$  gives one eigenfunction

$$X_0(x) = 1$$

(up to a multiplicative constant).

**Case 3:  $\lambda > 0$**

Write  $\lambda = k^2$  with  $k > 0$ . The equation becomes

$$X''(x) + k^2 X(x) = 0.$$

The general solution is

$$X(x) = A \cos(kx) + B \sin(kx).$$

We now impose periodicity. First,

$$X(0) = A, \quad X(2\pi) = A \cos(2\pi k) + B \sin(2\pi k).$$

The condition  $X(0) = X(2\pi)$  gives

$$A = A \cos(2\pi k) + B \sin(2\pi k).$$

Next,

$$X'(x) = -Ak \sin(kx) + Bk \cos(kx),$$

so

$$X'(0) = Bk, \quad X'(2\pi) = -Ak \sin(2\pi k) + Bk \cos(2\pi k).$$

The condition  $X'(0) = X'(2\pi)$  gives

$$Bk = -Ak \sin(2\pi k) + Bk \cos(2\pi k).$$

We can divide by  $k$  the last identity (for  $k \neq 0$ ) and write the system as

$$A(1 - \cos(2\pi k)) - B \sin(2\pi k) = 0,$$

$$A \sin(2\pi k) + B(1 - \cos(2\pi k)) = 0.$$

For a nontrivial pair  $(A, B)$  the determinant of the system must vanish:

$$(1 - \cos(2\pi k))^2 + (\sin(2\pi k))^2 = 0.$$

The left side is a sum of squares, so it is zero if and only if

$$1 - \cos(2\pi k) = 0, \quad \sin(2\pi k) = 0.$$

Hence

$$\cos(2\pi k) = 1, \quad \sin(2\pi k) = 0.$$

This happens exactly when  $k$  is an integer:

$$k = n, \quad n \in \mathbb{Z}.$$

The case  $n = 0$  corresponds to  $\lambda = 0$ , which we have already treated. For  $n \geq 1$  we obtain eigenvalues

$$\lambda_n = n^2, \quad n = 1, 2, 3, \dots$$

For each  $n \geq 1$  the corresponding eigenfunctions can be chosen as

$$X_n^{(c)}(x) = \cos(nx), \quad X_n^{(s)}(x) = \sin(nx).$$

These functions are  $2\pi$  periodic, and their derivatives are also  $2\pi$  periodic, so the boundary conditions are satisfied.

We have therefore found a complete set of spatial eigenfunctions for the heat equation with periodic boundary conditions:

- a constant mode  $X_0(x) = 1$  (eigenvalue  $\lambda_0 = 0$ ),
- cosine modes  $X_n^{(c)}(x) = \cos(nx)$ ,
- sine modes  $X_n^{(s)}(x) = \sin(nx)$ ,

with eigenvalues  $\lambda_n = n^2$  for  $n \geq 1$ .

### 2.1.3 Step 3: Time Dependent Factors

For each eigenvalue  $\lambda$  the corresponding time factor satisfies

$$T'(t) + \lambda T(t) = 0.$$

The solution is

$$T(t) = Ce^{-\lambda t}.$$

For the constant mode  $\lambda_0 = 0$  we obtain

$$T_0(t) = C_0$$

(constant in time).

For  $n \geq 1$  we have

$$T_n(t) = C_n e^{-n^2 t}.$$

Combining space and time, we form products  $u(x, t) = X(x) \cdot T(t)$  to obtain separated solutions. For each mode, the arbitrary constant from the time factor can be absorbed into a single new constant. Specifically, we write

$$\begin{aligned} u_0(x, t) &= A_0, \\ u_n^{(c)}(x, t) &= A_n \cos(nx) \cdot e^{-n^2 t}, \\ u_n^{(s)}(x, t) &= B_n \sin(nx) \cdot e^{-n^2 t}, \end{aligned}$$

where we have renamed the constants:  $A_0 = C_0$ , and for  $n \geq 1$ , the constant  $A_n$  absorbs the coefficient from the cosine mode while  $B_n$  absorbs the coefficient from the sine mode. Note that each separated solution carries its own independent arbitrary constant.

Since the heat equation is linear, any linear combination of these separated solutions is again a solution. Therefore the general solution that satisfies the periodic boundary conditions can be written as an infinite series

$$u(x, t) = A_0 + \sum_{n=1}^{\infty} (A_n \cos(nx) + B_n \sin(nx)) e^{-n^2 t}.$$

The coefficients  $A_0, A_n, B_n$  remain to be determined from the initial condition.

### 2.1.4 Step 4: Imposing the Initial Condition and Fourier Series

We impose the initial condition

$$u(x, 0) = f(x).$$

Setting  $t = 0$  in the general solution we obtain

$$u(x, 0) = A_0 + \sum_{n=1}^{\infty} (A_n \cos(nx) + B_n \sin(nx)) = f(x).$$

This is exactly the Fourier series expansion of the  $2\pi$  periodic function  $f$ . Under mild regularity assumptions,  $f$  has a Fourier series

$$f(x) = a_0 + \sum_{n=1}^{\infty} (a_n \cos(nx) + b_n \sin(nx))$$

with Fourier coefficients

$$\begin{aligned} a_0 &= \frac{1}{2\pi} \int_0^{2\pi} f(x) \, dx, \\ a_n &= \frac{1}{\pi} \int_0^{2\pi} f(x) \cos(nx) \, dx, \quad n \geq 1, \end{aligned}$$

$$b_n = \frac{1}{\pi} \int_0^{2\pi} f(x) \sin(nx) dx, \quad n \geq 1.$$

By uniqueness of the Fourier expansion, we must have

$$A_0 = a_0, \quad A_n = a_n, \quad B_n = b_n.$$

Thus the coefficients in the heat equation solution are exactly the Fourier coefficients of the initial data.

### 2.1.5 Step 5: Final Explicit Solution and Interpretation

Substituting these coefficients into the general solution we obtain the explicit formula

$$u(x, t) = a_0 + \sum_{n=1}^{\infty} (a_n \cos(nx) + b_n \sin(nx)) e^{-n^2 t}, \quad t > 0.$$

This infinite sum solves the heat equation with periodic boundary conditions and initial data  $f$ . Each Fourier mode decays exponentially in time at a rate proportional to its eigenvalue  $n^2$ . High frequency modes (large  $n$ ) decay faster, which expresses the smoothing effect of the heat equation.

The constant term  $a_0$  does not decay. It represents the average value of  $f$  on  $[0, 2\pi]$ , which is preserved by the heat flow.

From a spectral viewpoint, the functions

$$1, \cos(x), \sin(x), \cos(2x), \sin(2x), \dots$$

form an eigenbasis of the spatial operator

$$Lu = \frac{\partial^2 u}{\partial x^2}$$

with periodic boundary conditions. The evolution of each eigenmode is independent and given simply by multiplication by  $e^{-n^2 t}$  in time.

Later, in the numerical part of this book, we will approximate  $u(x, t)$  by truncating the sum to finitely many modes:

$$u_{N(x,t)} = a_0 + \sum_{n=1}^N (a_n \cos(nx) + b_n \sin(nx)) e^{-n^2 t}.$$

This truncation is the essence of a Fourier spectral method. The analytic solution derived here is the infinite dimensional limit of that numerical procedure.

## 2.2 Numerical Illustration

To visualize the smoothing effect of heat diffusion, we compute the truncated Fourier series solution for a triangle wave initial condition:


$$f(x) = \pi - |x - \pi|, \quad x \in [0, 2\pi].$$

This function is continuous but has a corner (non-differentiable point) at  $x = \pi$ . Its Fourier series contains only cosine terms with coefficients decaying as  $1/n^2$ :

$$f(x) = \frac{\pi}{2} + \frac{4}{\pi} \sum_{k=1}^{\infty} \frac{\cos((2k-1)x)}{(2k-1)^2}.$$

The key portion of the implementation evaluates the truncated Fourier series at any point in space and time. In Python:

```
1 def heat_solution(x, t, a0, a_n, b_n):
```

 Python

```

2  u = np.full_like(x, a0, dtype=float)
3  n_modes = len(a_n) - 1
4  for n in range(1, n_modes + 1):
5      decay = np.exp(-n**2 * t)
6      u += (a_n[n] * np.cos(n * x) + b_n[n] * np.sin(n * x)) * decay
7  return u

```

The equivalent MATLAB implementation:

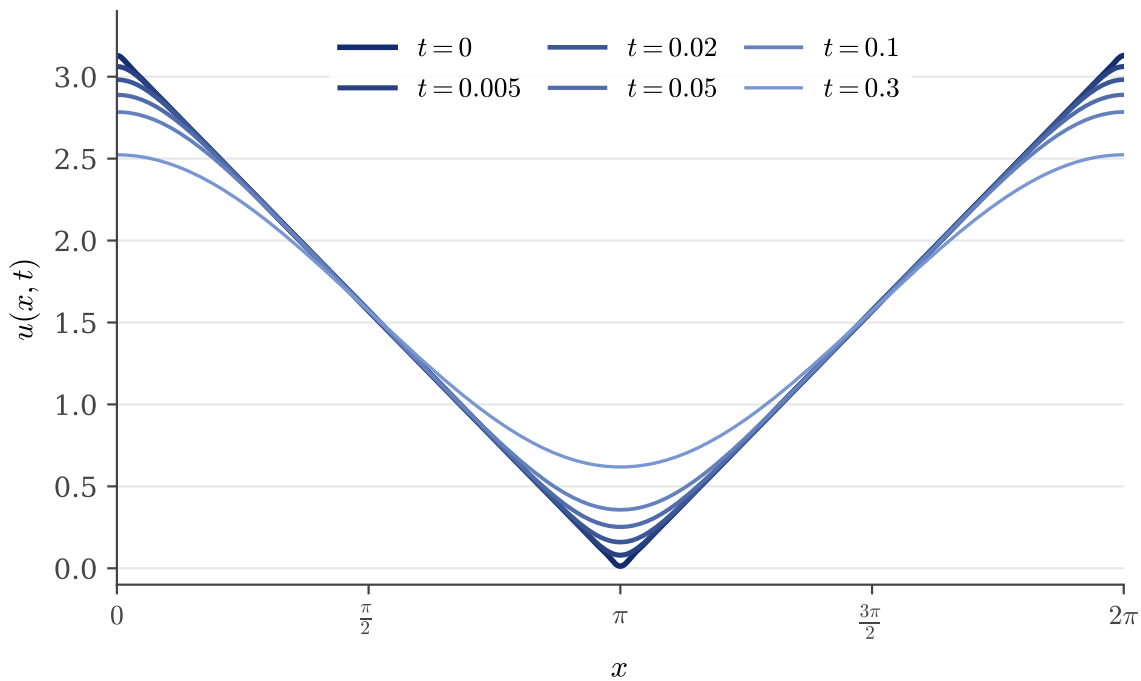
```

1  u = a0 * ones(size(x));
2  for n = 1:N_MODES
3      decay = exp(-n^2 * t);
4      u = u + (a_n(n+1) * cos(n*x) + b_n(n+1) * sin(n*x)) * decay;
5  end

```

Matlab

Figure 1 shows the evolution of  $u_N(x, t)$  with  $N = 50$  modes at several time values. At  $t = 0$  the triangle wave is faithfully reproduced. As time increases, the higher frequency modes decay exponentially faster than the lower ones (the  $n$ -th mode decays as  $e^{-n^2 t}$ ), and the solution rapidly smooths toward the constant equilibrium  $u = \pi/2$ .



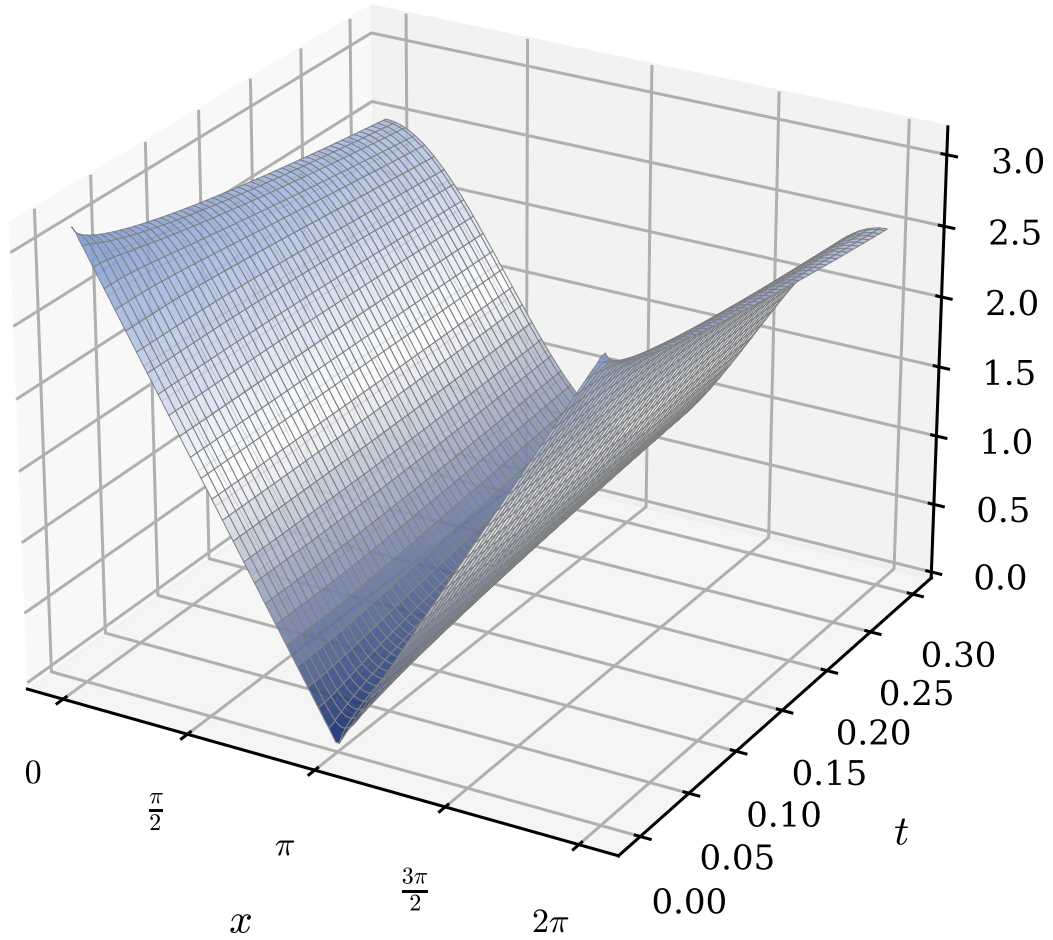
**Figure 1:** Evolution of the heat equation solution with a triangle wave initial condition. The higher frequency modes decay rapidly, smoothing the initial corner at  $x = \pi$ .

The code that generated this figure is available in both Python and MATLAB:

- codes/python/ch02\_classical\_pdes/heat\_equation\_evolution.py
- codes/matlab/ch02\_classical\_pdes/heat\_equation\_evolution.m

A complementary view of the solution is provided by the waterfall plot in Figure 2, which displays the entire space-time evolution as a three-dimensional surface. The smoothing effect of the heat equation is clearly visible: the initial sharp triangle wave rapidly flattens as time progresses, with the solution approaching the constant equilibrium state  $u = \pi/2$ .





**Figure 2:** Waterfall plot showing the complete space-time evolution of the heat equation solution. The initial triangle wave smooths rapidly as higher frequency modes decay exponentially.

## 2.3 Wave Equation with Dirichlet Boundary Conditions

We now consider the one dimensional wave equation on a finite interval with homogeneous Dirichlet boundary conditions. This is the classical model for a vibrating string of length  $L$  with both ends fixed.

Let  $u(x, t)$  denote the vertical displacement of the string at position  $x \in [0, L]$  and time  $t > 0$ . The equation of motion is

$$\frac{\partial^2 u}{\partial t^2}(x, t) = c^2 \frac{\partial^2 u}{\partial x^2}(x, t), \quad 0 < x < L, \quad t > 0,$$

where  $c > 0$  is the wave speed.

The boundary conditions express that the endpoints of the string are clamped:

$$u(0, t) = 0, \quad u(L, t) = 0, \quad t > 0.$$

We prescribe the initial displacement and initial velocity:

$$u(x, 0) = f(x), \quad \frac{\partial u}{\partial t}(x, 0) = g(x), \quad 0 < x < L,$$

with suitable functions  $f$  and  $g$  that vanish at  $x = 0$  and  $x = L$ .

As in the heat equation example, we use separation of variables and obtain a representation of the solution as an infinite series in spatial eigenfunctions. This time the temporal factors are oscillatory instead of decaying. The mathematical theory of the vibrating string is a classical topic covered extensively in [4].

### 2.3.1 Step 1: Separation Ansatz

We look for nontrivial solutions of the form

$$u(x, t) = X(x) \cdot T(t).$$

Substituting into the wave equation gives

$$X(x) \cdot T''(t) = c^2 X''(x) \cdot T(t).$$

Assuming  $X$  and  $T$  are not identically zero, we divide both sides by  $c^2 X(x) \cdot T(t)$ :

$$\frac{T''(t)}{c^2 T(t)} = \frac{X''(x)}{X(x)}.$$

The left side depends only on  $t$ , the right side only on  $x$ . Therefore both sides must be equal to a constant, which we denote by  $-\lambda$ :

$$\frac{T''(t)}{c^2 T(t)} = \frac{X''(x)}{X(x)} = -\lambda.$$

We obtain the pair of ordinary differential equations

$$\begin{aligned} T''(t) + c^2 \lambda T(t) &= 0, \\ X''(x) + \lambda X(x) &= 0, \end{aligned}$$

with boundary conditions

$$X(0) = 0, \quad X(L) = 0.$$

As in the heat equation case, we have a spatial eigenvalue problem for  $X$ .

### 2.3.2 Step 2: Spatial Eigenvalue Problem with Dirichlet Boundary Conditions

We must solve

$$X''(x) + \lambda X(x) = 0, \quad 0 < x < L,$$

with

$$X(0) = 0, \quad X(L) = 0.$$

We again consider three cases:  $\lambda < 0$ ,  $\lambda = 0$ , and  $\lambda > 0$ .

#### Case 1: $\lambda < 0$

Write  $\lambda = -\mu^2$  with  $\mu > 0$ . The equation becomes

$$X''(x) - \mu^2 X(x) = 0.$$

The general solution is

$$X(x) = Ae^{\mu x} + Be^{-\mu x}.$$

The boundary condition at  $x = 0$  gives

$$X(0) = A + B = 0 \quad \Rightarrow \quad B = -A.$$

Then

$$X(L) = Ae^{\mu L} - Ae^{-\mu L} = A(e^{\mu L} - e^{-\mu L}).$$

The condition  $X(L) = 0$  implies

$$A(e^{\mu L} - e^{-\mu L}) = 0.$$

Since  $e^{\mu L} \neq e^{-\mu L}$  for  $\mu > 0$ , we must have  $A = 0$ . Then  $B = 0$  and the solution is trivial. Therefore there are no nontrivial eigenfunctions for  $\lambda < 0$ .

**Case 2:  $\lambda = 0$**

The equation reduces to

$$X''(x) = 0,$$

whose general solution is

$$X(x) = A + Bx.$$

The boundary conditions give

$$\begin{aligned} X(0) &= A = 0, \\ X(L) &= A + BL = BL = 0. \end{aligned}$$

Hence  $B = 0$  and  $X$  is again trivial. There is no nontrivial eigenfunction for  $\lambda = 0$ .

**Case 3:  $\lambda > 0$**

Write  $\lambda = k^2$  with  $k > 0$ . The equation becomes

$$X''(x) + k^2 X(x) = 0.$$

The general solution is

$$X(x) = A \cos(kx) + B \sin(kx).$$

The boundary condition at  $x = 0$  gives

$$X(0) = A = 0.$$

So  $X(x) = B \sin(kx)$ . The boundary condition at  $x = L$  gives

$$X(L) = B \sin(kL) = 0.$$

For a nontrivial solution we need  $B \neq 0$ , so we must have

$$\sin(kL) = 0.$$

Therefore

$$kL = n\pi, \quad n = 1, 2, 3, \dots$$

The corresponding values of  $k$  are

$$k_n = \frac{n\pi}{L}, \quad n = 1, 2, 3, \dots$$

We conclude that the eigenvalues and eigenfunctions are

$$\begin{aligned} \lambda_n &= k_n^2 = \left(\frac{n\pi}{L}\right)^2, \\ X_n(x) &= \sin\left(\frac{n\pi x}{L}\right), \quad n = 1, 2, 3, \dots \end{aligned}$$

Each  $X_n$  vanishes at  $x = 0$  and  $x = L$ , as required by the Dirichlet boundary conditions. The family  $\{X_n\}_{n \geq 1}$  is orthogonal in  $L^2(0, L)$ :

$$\int_0^L \sin\left(\frac{n\pi x}{L}\right) \sin\left(\frac{m\pi x}{L}\right) dx = \begin{cases} 0 & \text{if } n \neq m \\ L/2 & \text{if } n = m. \end{cases}$$

These eigenfunctions will form the spatial basis in our series solution.

### 2.3.3 Step 3: Time Dependent Factors

For each eigenvalue  $\lambda_n$  the time factor  $T_n$  satisfies

$$T_n''(t) + c^2 \lambda_n T_n(t) = 0.$$

Using  $\lambda_n = (n\pi/L)^2$  we can write

$$T_n''(t) + \omega_n^2 T_n(t) = 0,$$

where

$$\omega_n = c \frac{n\pi}{L}, \quad n = 1, 2, 3, \dots$$

The general solution of this second order linear ODE is

$$T_n(t) = A_n \cos(\omega_n t) + B_n \sin(\omega_n t),$$

where  $A_n$  and  $B_n$  are constants.

Combining the space and time factors, we obtain separated solutions

$$u_n(x, t) = (A_n \cos(\omega_n t) + B_n \sin(\omega_n t)) \sin\left(\frac{n\pi x}{L}\right), \quad n = 1, 2, 3, \dots$$

Because the wave equation is linear, any linear combination of these separated solutions is again a solution. Therefore the general solution satisfying the Dirichlet boundary conditions can be written as an infinite series

$$u(x, t) = \sum_{n=1}^{\infty} (a_n \cos(\omega_n t) + b_n \sin(\omega_n t)) \sin\left(\frac{n\pi x}{L}\right),$$

for suitable coefficients  $a_n$  and  $b_n$ .

These coefficients will be determined from the initial conditions.

#### 2.3.4 Step 4: Imposing the Initial Conditions and Sine Series

We now use the initial displacement and velocity.

At  $t = 0$  we have

$$u(x, 0) = \sum_{n=1}^{\infty} (a_n \cos(0) + b_n \sin(0)) \sin\left(\frac{n\pi x}{L}\right) = \sum_{n=1}^{\infty} a_n \sin\left(\frac{n\pi x}{L}\right).$$

The initial condition  $u(x, 0) = f(x)$  becomes

$$f(x) = \sum_{n=1}^{\infty} a_n \sin\left(\frac{n\pi x}{L}\right).$$

This is the Fourier sine series of  $f$  on the interval  $(0, L)$ .

Similarly, we differentiate  $u$  with respect to  $t$ :

$$\frac{\partial u}{\partial t}(x, t) = \sum_{n=1}^{\infty} (-a_n \omega_n \sin(\omega_n t) + b_n \omega_n \cos(\omega_n t)) \sin\left(\frac{n\pi x}{L}\right).$$

Evaluating at  $t = 0$  gives

$$\frac{\partial u}{\partial t}(x, 0) = \sum_{n=1}^{\infty} b_n \omega_n \sin\left(\frac{n\pi x}{L}\right).$$

The initial condition  $\frac{\partial u}{\partial t}(x, 0) = g(x)$  becomes

$$g(x) = \sum_{n=1}^{\infty} b_n \omega_n \sin\left(\frac{n\pi x}{L}\right).$$

So the sequence  $\{a_n\}$  consists of the Fourier sine coefficients of  $f$ , and the sequence  $\{b_n \omega_n\}$  consists of the Fourier sine coefficients of  $g$ .

Using the orthogonality relations, we obtain explicit formulas for the coefficients. For  $n \geq 1$ ,

$$a_n = \frac{2}{L} \int_0^L f(x) \sin\left(\frac{n\pi x}{L}\right) dx,$$

and

$$b_n \omega_n = \frac{2}{L} \int_0^L g(x) \sin\left(\frac{n\pi x}{L}\right) dx.$$

Therefore

$$b_n = \frac{2}{L \omega_n} \int_0^L g(x) \sin\left(\frac{n\pi x}{L}\right) dx = \frac{2}{L} \frac{1}{cn\pi/L} \int_0^L g(x) \sin\left(\frac{n\pi x}{L}\right) dx.$$

In summary,

$$a_n = \frac{2}{L} \int_0^L f(x) \sin\left(\frac{n\pi x}{L}\right) dx,$$

$$b_n = \frac{2}{L\omega_n} \int_0^L g(x) \sin\left(\frac{n\pi x}{L}\right) dx, \quad \omega_n = c \frac{n\pi}{L}.$$

### 2.3.5 Step 5: Final Explicit Solution and Interpretation

Substituting these coefficients into the series, we obtain the explicit solution of the wave equation with Dirichlet boundary conditions:

$$u(x, t) = \sum_{n=1}^{\infty} [a_n \cos(\omega_n t) + b_n \sin(\omega_n t)] \sin\left(\frac{n\pi x}{L}\right),$$

where  $\omega_n = cn\pi/L$  and the Fourier sine coefficients are

$$a_n = \frac{2}{L} \int_0^L f(y) \sin\left(\frac{n\pi y}{L}\right) dy, \quad b_n = \frac{2}{n\pi c} \int_0^L g(y) \sin\left(\frac{n\pi y}{L}\right) dy.$$

Each term in the sum is a normal mode of vibration: a standing wave with spatial shape  $\sin(n\pi x/L)$  and temporal oscillation at frequency  $\omega_n$ . The coefficients of  $\cos(\omega_n t)$  and  $\sin(\omega_n t)$  are determined by the initial displacement  $f$  and initial velocity  $g$  through their Fourier sine coefficients.

Comparing with the heat equation:

- For the heat equation, each mode decayed like  $e^{-n^2 t}$  and the solution became smoother in time.
- For the wave equation, each mode oscillates periodically in time with constant amplitude, reflecting conservation of energy in the undamped string.

From a spectral viewpoint, the functions

$$\sin\left(\frac{\pi x}{L}\right), \sin\left(\frac{2\pi x}{L}\right), \sin\left(\frac{3\pi x}{L}\right), \dots$$

form an eigenbasis of the spatial operator

$$Lu = \frac{\partial^2 u}{\partial x^2}$$

with Dirichlet boundary conditions. In this basis, the evolution is diagonal: each mode evolves independently according to a simple harmonic oscillator in time.

As in the heat equation example, a spectral method will approximate  $u(x, t)$  by truncating the infinite sum. For some integer  $N \geq 1$  we consider the finite approximation

$$u_N(x, t) = \sum_{n=1}^N (a_n \cos(\omega_n t) + b_n \sin(\omega_n t)) \sin\left(\frac{n\pi x}{L}\right).$$

The analytic series above is the infinite dimensional limit of this spectral representation.

## 2.4 Numerical Illustration

To visualize the oscillatory behavior of the vibrating string, we compute the truncated Fourier sine series solution for a plucked string initial condition. The string is plucked at its center, forming a triangular initial displacement:

$$f(x) = \begin{cases} \frac{2h}{L}x & \text{for } 0 \leq x \leq L/2 \\ 2h(1 - x/L) & \text{for } L/2 \leq x \leq L \end{cases}$$

with zero initial velocity  $g(x) = 0$ . Here  $h$  denotes the height of the pluck at the center.

The Fourier sine coefficients of this triangular shape are

$$a_n = \frac{8h}{n^2\pi^2} \sin\left(\frac{n\pi}{2}\right),$$

which gives nonzero values only for odd  $n$ , with alternating signs.

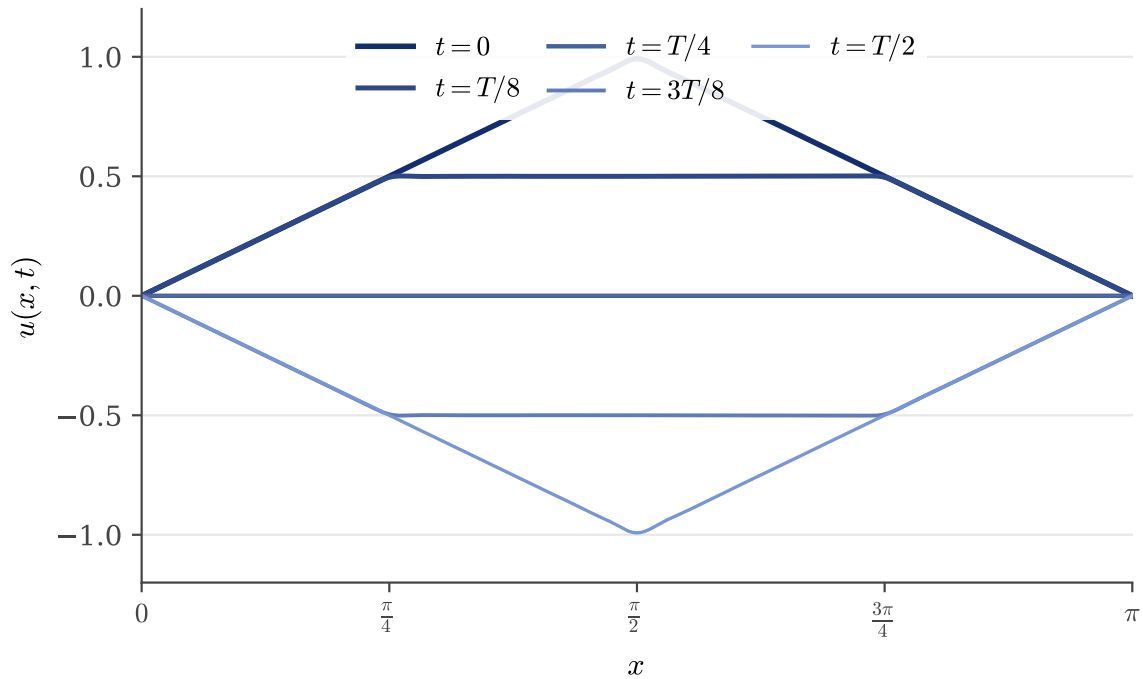
The key portion of the implementation computes the solution at any point in space and time. In Python:

```
1 def wave_solution(x, t, a_n, b_n, L, c):
2     u = np.zeros_like(x, dtype=float)
3     n_modes = len(a_n) - 1
4     for n in range(1, n_modes + 1):
5         omega_n = c * n * np.pi / L
6         spatial = np.sin(n * np.pi * x / L)
7         temporal = a_n[n] * np.cos(omega_n * t) + b_n[n] * np.sin(omega_n * t)
8         u += temporal * spatial
9     return u
```

The equivalent MATLAB implementation:

```
1 u = zeros(size(x));
2 for n = 1:N_MODES
3     omega_n = C * n * pi / L;
4     spatial = sin(n * pi * x / L);
5     temporal = a_n(n+1) * cos(omega_n * t) + b_n(n+1) * sin(omega_n * t);
6     u = u + temporal * spatial;
7 end
```

Figure 3 shows the evolution of  $u_N(x, t)$  with  $N = 50$  modes at several time values within half a period  $T = 2L/c$ . The string oscillates back and forth, with the triangular shape inverting at  $t = T/2$ . Unlike the heat equation, the wave equation preserves energy and the solution does not decay; it continues oscillating indefinitely.

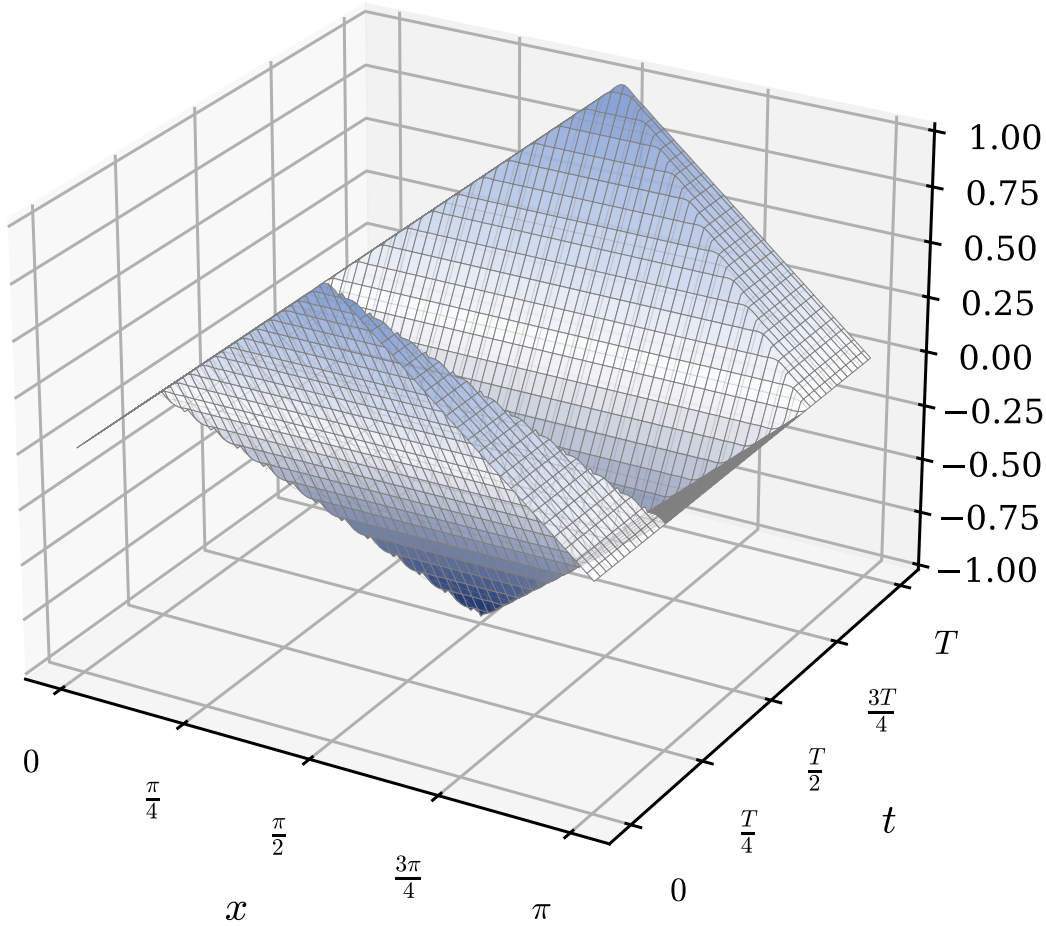


**Figure 3:** Evolution of the wave equation solution with a plucked string initial condition. The string oscillates with period  $T = 2\pi$ , inverting at  $t = T/2$ .

The code that generated this figure is available in both Python and MATLAB:

- `codes/python/ch02_classical_pdes/wave_equation_evolution.py`
- `codes/matlab/ch02_classical_pdes/wave_equation_evolution.m`

The waterfall plot in Figure 4 provides a complete view of the oscillatory dynamics over one full period. Unlike the heat equation, the wave equation conserves energy: the solution oscillates indefinitely without decay, and the periodic nature of the motion is clearly visible in the three-dimensional representation.



**Figure 4:** Waterfall plot showing the complete space-time evolution of the wave equation solution over one period  $T$ . The plucked string oscillates between its initial shape and its mirror image.

## 2.5 Laplace Equation in a Periodic Strip

For the elliptic case we consider the Laplace equation in a simple two dimensional domain that is periodic in one direction and bounded in the other. This setting connects naturally with the periodic heat equation example and again leads to a Fourier series representation in the periodic direction.

Let

$$D = \{(x, y) \in \mathbb{R}^2 : 0 < x < 2\pi, 0 < y < 1\}.$$

We seek a harmonic function  $u(x, y)$  solving

$$u_{xx}(x, y) + u_{yy}(x, y) = 0, \quad (x, y) \in D,$$

with periodic boundary conditions in  $x$

$$u(x + 2\pi, y) = u(x, y), \quad \text{for all real } x, 0 < y < 1,$$



and Dirichlet conditions in  $y$

$$u(x, 0) = f(x), \quad u(x, 1) = 0, \quad 0 \leq x \leq 2\pi.$$

We assume that  $f$  is  $2\pi$  periodic and smooth:

$$f(x + 2\pi) = f(x).$$

As in the parabolic and hyperbolic examples, we apply separation of variables and obtain a representation of  $u$  as an infinite Fourier series in  $x$  with  $y$  dependent coefficients. The theory of harmonic functions and Laplace's equation is presented in detail in [4].

### 2.5.1 Step 1: Separation Ansatz

We look for nontrivial separated solutions of the form

$$u(x, y) = X(x) \cdot Y(y).$$

Substituting into the Laplace equation gives

$$X''(x) \cdot Y(y) + X(x) \cdot Y''(y) = 0.$$

Assuming  $X$  and  $Y$  are not identically zero, we divide by  $X(x) \cdot Y(y)$ :

$$\frac{X''(x)}{X(x)} + \frac{Y''(y)}{Y(y)} = 0.$$

The first term depends only on  $x$ , the second only on  $y$ . Therefore both must be equal to constants whose sum is zero. We introduce a separation constant  $\lambda$  and write

$$\frac{X''(x)}{X(x)} = -\lambda, \quad \frac{Y''(y)}{Y(y)} = \lambda.$$

This leads to the two ordinary differential equations

$$X''(x) + \lambda X(x) = 0,$$

$$Y''(y) - \lambda Y(y) = 0.$$

The periodic boundary conditions for  $u$  in the  $x$  direction imply the periodic conditions

$$X(x + 2\pi) = X(x), \quad \text{for all real } x.$$

The Dirichlet conditions in  $y$  will be enforced later on the full series. For the separated functions  $Y$  we will impose appropriate conditions at  $y = 1$ , while the condition at  $y = 0$  will be handled via the Fourier coefficients of  $f$ .

We have once more an eigenvalue problem in the periodic direction.

### 2.5.2 Step 2: Eigenfunctions in the Periodic Direction

The equation for  $X$  with periodic boundary conditions is exactly the same as in the heat equation example:

$$X''(x) + \lambda X(x) = 0, \quad X(x + 2\pi) = X(x).$$

We recall the result: there is a constant mode with eigenvalue  $\lambda_0 = 0$ ,

$$X_0(x) = 1,$$

and for each integer  $n \geq 1$  there are cosine and sine modes with eigenvalues

$$\lambda_n = n^2,$$

$$X_n^{(c)}(x) = \cos(nx), \quad X_n^{(s)}(x) = \sin(nx).$$

The family

$$1, \cos(x), \sin(x), \cos(2x), \sin(2x), \dots$$

forms an orthogonal basis in  $L^2(0, 2\pi)$  for  $2\pi$  periodic functions.

For each such eigenvalue we now solve the corresponding equation for  $Y$ .

### 2.5.3 Step 3: Equations for the $y$ Dependent Factors

For each eigenvalue  $\lambda$  we have

$$Y''(y) - \lambda Y(y) = 0.$$

We treat separately the constant mode  $\lambda_0 = 0$  and the nonzero modes  $\lambda_n = n^2$  for  $n \geq 1$ .

#### Constant mode $\lambda_0 = 0$

For  $\lambda_0 = 0$  the equation reduces to

$$Y_0''(y) = 0.$$

The general solution is

$$Y_0(y) = A_0 + B_0 y.$$

We want separated solutions that satisfy the homogeneous boundary condition at  $y = 1$ :

$$u(x, 1) = 0 \quad \text{for all } x.$$

For the constant mode this means

$$X_0(x) \cdot Y_0(1) = Y_0(1) = 0,$$

hence

$$Y_0(1) = A_0 + B_0 = 0.$$

We choose a convenient normalization so that  $Y_0(0) = 1$ . Then  $A_0 = 1$  and the relation  $A_0 + B_0 = 0$  gives  $B_0 = -1$ . Thus

$$Y_0(y) = 1 - y.$$

This separated mode

$$u_0(x, y) = X_0(x) \cdot Y_0(y) = 1 - y$$

is harmonic, periodic in  $x$ , and vanishes at  $y = 1$ , with value 1 at  $y = 0$ .

#### Higher modes $\lambda_n = n^2$ for $n \geq 1$

For  $\lambda_n = n^2$  the equation is

$$Y_n''(y) - n^2 Y_n(y) = 0.$$

The general solution can be written in hyperbolic form

$$Y_n(y) = \alpha_n \cosh(ny) + \beta_n \sinh(ny).$$

We require that each separated mode vanish at  $y = 1$ :

$$Y_n(1) = 0.$$

To match later the Fourier coefficients of  $f$  at  $y = 0$ , it is convenient to normalize so that

$$Y_n(0) = 1.$$

Imposing  $Y_n(0) = 1$  gives

$$\alpha_n = 1.$$

Then

$$Y_n(1) = \cosh(n) + \beta_n \sinh(n) = 0$$

so

$$\beta_n = -\frac{\cosh(n)}{\sinh(n)} = -\coth(n).$$

Thus

$$Y_n(y) = \cosh(ny) - \coth(n) \cdot \sinh(ny).$$

An alternative and simpler expression uses the hyperbolic sine function of the distance to the boundary  $y = 1$ . One checks that

$$Y_n(y) = \frac{\sinh(n(1 - y))}{\sinh(n)}$$

satisfies

$$\begin{aligned} Y_n''(y) - n^2 Y_n(y) &= 0, \\ Y_n(1) &= 0, \end{aligned}$$

$$Y_n(0) = 1.$$

Indeed,  $Y_n(1) = \sinh(0)/\sinh(n) = 0$ ,  $Y_n(0) = \sinh(n)/\sinh(n) = 1$ , and

$$Y_n''(y) = n^2 \frac{\sinh(n(1-y))}{\sinh(n)} = n^2 Y_n(y).$$

We will use the form

$$Y_n(y) = \frac{\sinh(n(1-y))}{\sinh(n)}, \quad n \geq 1.$$

#### 2.5.4 Step 4: Building the Series Solution

Each separated solution corresponding to the eigenfunctions in  $x$  and the functions  $Y_n$  in  $y$  has the form

$$u_0(x, y) = C_0 \cdot Y_0(y) = C_0(1-y),$$

$$u_n^{(c)}(x, y) = C_n \cdot \cos(nx) \cdot Y_n(y),$$

$$u_n^{(s)}(x, y) = D_n \cdot \sin(nx) \cdot Y_n(y), \quad n \geq 1,$$

for some constants  $C_0, C_n, D_n$ .

Since the Laplace equation is linear and the boundary condition at  $y = 1$  is homogeneous, any linear combination of these separated solutions is again a solution that vanishes at  $y = 1$ . Therefore a general solution satisfying the periodic condition in  $x$  and the Dirichlet condition  $u(x, 1) = 0$  can be written as

$$u(x, y) = C_0(1-y) + \sum_{n=1}^{\infty} [C_n \cos(nx) + D_n \sin(nx)] \frac{\sinh(n(1-y))}{\sinh(n)}.$$

It remains to impose the boundary condition at  $y = 0$ ,

$$u(x, 0) = f(x).$$

At  $y = 0$  we obtain

$$u(x, 0) = C_0 + \sum_{n=1}^{\infty} [C_n \cos(nx) + D_n \sin(nx)],$$

because  $Y_0(0) = 1$  and  $Y_n(0) = 1$  for  $n \geq 1$ .

Hence the boundary condition  $u(x, 0) = f(x)$  becomes

$$f(x) = C_0 + \sum_{n=1}^{\infty} [C_n \cos(nx) + D_n \sin(nx)].$$

This is exactly the Fourier series expansion of  $f$ . For a  $2\pi$  periodic  $f$  we have

$$f(x) = a_0 + \sum_{n=1}^{\infty} [a_n \cos(nx) + b_n \sin(nx)],$$

with coefficients

$$a_0 = \frac{1}{2\pi} \int_0^{2\pi} f(x) dx,$$

$$a_n = \frac{1}{\pi} \int_0^{2\pi} f(x) \cos(nx) dx, \quad n \geq 1,$$

$$b_n = \frac{1}{\pi} \int_0^{2\pi} f(x) \sin(nx) dx, \quad n \geq 1.$$

By uniqueness of the Fourier expansion, we must have

$$C_0 = a_0, \quad C_n = a_n, \quad D_n = b_n.$$

### 2.5.5 Step 5: Final Explicit Solution and Interpretation

Substituting these coefficients into the series we obtain the explicit representation

$$u(x, y) = a_0(1 - y) + \sum_{n=1}^{\infty} [a_n \cos(nx) + b_n \sin(nx)] \frac{\sinh(n(1 - y))}{\sinh(n)}, \quad 0 < y < 1.$$

Here  $a_0$ ,  $a_n$ , and  $b_n$  are the Fourier coefficients of the boundary data  $f$  as defined above.

This series converges (under mild assumptions on  $f$ ) to the unique harmonic function that is periodic in  $x$ , equal to  $f$  on  $y = 0$ , and zero on  $y = 1$ .

From a spectral viewpoint:

- The functions  $1$ ,  $\cos(nx)$ ,  $\sin(nx)$  are eigenfunctions of the one dimensional Laplacian  $X \mapsto X''$  with periodic boundary conditions in  $x$ , with eigenvalues  $\lambda_0 = 0$  and  $\lambda_n = n^2$ .
- For each spatial frequency  $n$  in the periodic direction, the dependence in the transverse direction  $y$  is determined by the simple ordinary differential equation  $Y'' - \lambda_n Y = 0$  with boundary condition  $Y(1) = 0$  and normalization  $Y(0) = 1$ . This gives the hyperbolic profiles

$$Y_0(y) = 1 - y,$$

$$Y_n(y) = \frac{\sinh(n(1 - y))}{\sinh(n)}, \quad n \geq 1.$$

- The boundary data  $f$  at  $y = 0$  is expanded in the eigenbasis  $\{1, \cos(nx), \sin(nx)\}$  and each Fourier mode is propagated into the interior of the strip with its own  $y$  dependent factor  $Y_n(y)$ .

Analytically, the solution is an infinite sum of separated solutions. In a spectral method we will truncate this sum to finitely many modes in  $x$ ,

$$u_N(x, y) = a_0(1 - y) + \sum_{n=1}^N [a_n \cos(nx) + b_n \sin(nx)] \frac{\sinh(n(1 - y))}{\sinh(n)},$$

and approximate the harmonic function inside the strip by this finite Fourier representation.

## 2.6 Numerical Illustration

To visualize the structure of harmonic functions in the strip, we compute the truncated Fourier series solution for a boundary condition that contains two modes:


$$f(x) = \sin(x) + \frac{1}{2} \sin(3x).$$

For this particular boundary data, the Fourier coefficients are simply  $b_1 = 1$  and  $b_3 = 1/2$ , with all other coefficients zero. The solution can be written explicitly as

$$u(x, y) = \sin(x) \frac{\sinh(1 - y)}{\sinh(1)} + \frac{1}{2} \sin(3x) \frac{\sinh(3(1 - y))}{\sinh(3)}.$$

The key portion of the implementation evaluates the truncated series on a two-dimensional grid. In Python:

```
1 def laplace_solution(x, y, a0, a_n, b_n):
2     u = a0 * (1 - y)
3     n_modes = len(a_n) - 1
4     for n in range(1, n_modes + 1):
```

 Python

```

5     if abs(a_n[n]) < 1e-15 and abs(b_n[n]) < 1e-15:
6         continue
7     y_factor = np.sinh(n * (1 - y)) / np.sinh(n)
8     u += (a_n[n] * np.cos(n * x) + b_n[n] * np.sin(n * x)) * y_factor
9     return u
    
```

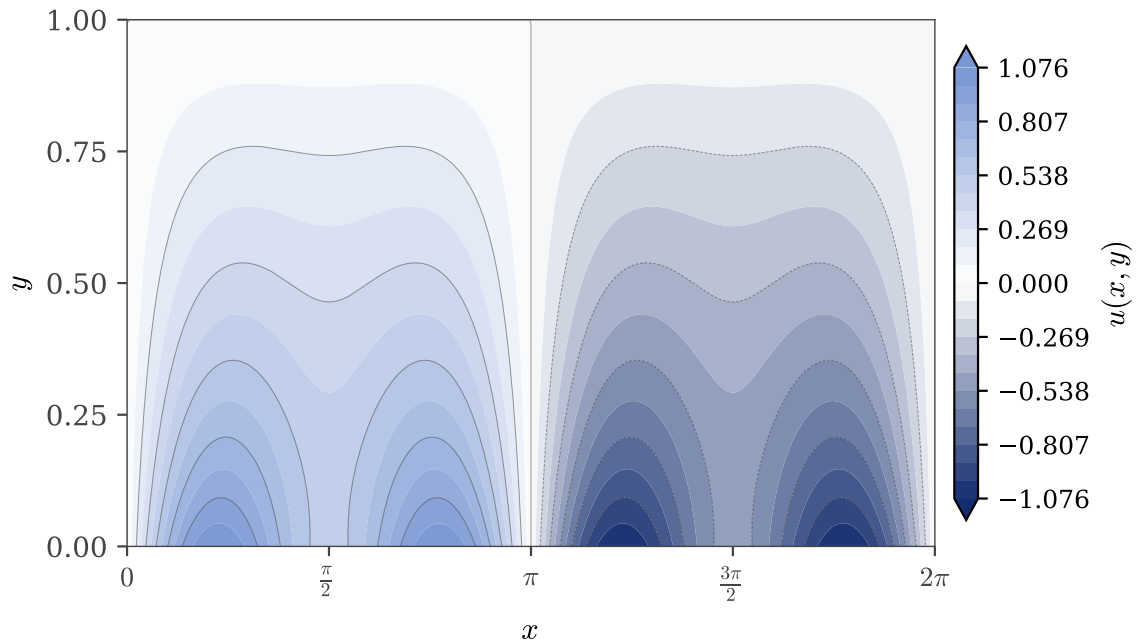
The equivalent MATLAB implementation:

```

1  U = a0 * (1 - Y);
2  for n = 1:N_MODES
3      if abs(a_n(n+1)) < 1e-15 && abs(b_n(n+1)) < 1e-15
4          continue;
5      end
6      y_factor = sinh(n * (1 - Y)) / sinh(n);
7      U = U + (a_n(n+1) * cos(n*X) + b_n(n+1) * sin(n*X)) .* y_factor;
8  end
    
```

Matlab

Figure 5 shows the solution  $u(x, y)$  in the strip  $[0, 2\pi] \times [0, 1]$ . At the bottom boundary  $y = 0$ , the solution matches the prescribed boundary data  $f(x)$ . As  $y$  increases toward the top boundary, the solution decays to zero. Crucially, the higher frequency mode ( $n = 3$ ) decays much faster than the lower frequency mode ( $n = 1$ ), as the hyperbolic factor  $\sinh(n(1 - y))/\sinh(n)$  decreases more rapidly for larger  $n$ . This illustrates the smoothing effect of harmonic extension into the interior.



**Figure 5:** Solution of the Laplace equation in the periodic strip with boundary data  $f(x) = \sin(x) + \frac{1}{2}\sin(3x)$  at  $y = 0$  and  $u = 0$  at  $y = 1$ . Higher frequency modes decay faster toward the interior.

The code that generated this figure is available in both Python and MATLAB:

- codes/python/ch02\_classical\_pdes/laplace\_equation\_2d.py
- codes/matlab/ch02\_classical\_pdes/laplace\_equation\_2d.m

## 2.7 Conclusions

The three examples presented in this chapter (the heat equation, the wave equation, and the Laplace equation) share a common mathematical structure that will guide us throughout the rest of this book. In each case, separation of variables reduces a partial differential equation to a family of ordinary differential equations: an eigenvalue problem in the spatial variable and a simpler equation governing the behavior in the remaining variable (time or the transverse coordinate). The eigenfunctions form an orthogonal basis, and the solution is expressed as an infinite series

$$u(x, t) = \sum_k \hat{u}_k(t) \varphi_k(x)$$

whose coefficients are determined by the initial or boundary data through Fourier projections. This is the DNA of spectral methods.

Let us distill the key ideas:

1. **Separation.** We decompose the solution into modes that evolve independently (or nearly so). Each mode satisfies a simpler equation than the original PDE.
2. **Basis.** The spatial part of each mode is an eigenfunction of a differential operator: Fourier exponentials for periodic problems, trigonometric functions for Dirichlet conditions, Chebyshev or Legendre polynomials for more general settings.
3. **Truncation.** In practice we cannot sum infinitely many terms due to the apparent finitude of our Universe. We retain only the first  $N$  modes, and the accuracy of this approximation depends critically on how fast the coefficients  $\hat{u}_k$  decay.

When the solution is smooth, the coefficients decay *exponentially*, and a modest  $N$  suffices for high accuracy. This is the source of spectral methods' legendary efficiency. For a comprehensive treatment of these classical methods and their mathematical foundations, the reader is referred to [4].

The analytical solutions derived in this chapter are beautiful, but they are also fragile. They apply only to linear equations on simple domains with special boundary conditions. The moment we encounter a nonlinearity, a complicated geometry, or variable coefficients, we must turn to computation. The chapters that follow will develop the computational machinery needed to turn these analytical insights into practical algorithms:

- **FFT and its cousins:** how to move efficiently between physical space and coefficient space.
- **Differentiation matrices:** how to compute derivatives spectrally.
- **Quadrature rules:** how to compute inner products and projections.
- **Time stepping:** how to advance the ODE system for the coefficients.

Armed with these tools, we will be able to solve problems far beyond the reach of pen-and-paper analysis.





## CHAPTER 3

# Mise en Bouche



**I**n French cuisine, a *mise en bouche* is a small appetizer offered by the chef to stimulate the palate before the main courses arrive. In this chapter we offer a similar intellectual appetizer: a compact, self-contained taste of spectral methods that illuminates the essential mechanics before we develop the full machinery of Fourier and Chebyshev approximation. Rather than jumping immediately to high-degree polynomials with  $N = 100$ , we perform hand calculations with just  $N = 2$  or  $N = 3$  unknowns. This low-dimensional setting makes every step transparent. We can verify each formula by direct computation and gain intuition that will guide us through the more sophisticated developments to come.

The techniques presented here follow the classical exposition in [3], adapted to our pedagogical goals. The Method of Weighted Residuals provides the unifying framework that connects the collocation (pseudospectral) approach we favor in this book with the Galerkin methods that dominate finite element analysis.

## 3.1 The Method of Weighted Residuals

### 3.1.1 Series Expansions and the Residual Function

The central idea of spectral methods is to approximate the unknown function  $u(x)$  by a finite sum of basis functions:

$$u(x) \approx u_N(x) = \sum_{n=0}^N a_n \varphi_n(x),$$

where  $\{\varphi_n(x)\}$  are known basis functions and  $\{a_n\}$  are unknown coefficients to be determined.

When we substitute this approximation into a differential equation

$$\mathcal{L}u = f(x),$$

where  $\mathcal{L}$  is a linear differential operator, the result is generally not zero. The *residual function* measures this discrepancy:

$$R(x; a_0, a_1, \dots, a_N) = \mathcal{L}u_N - f.$$

For the exact solution,  $R(x) \equiv 0$ . The challenge is to choose the coefficients  $\{a_n\}$  so that the residual is as small as possible. Different spectral methods correspond to different strategies for minimizing this residual.

### 3.1.2 Two Minimization Strategies

The two most important strategies are:

1. **Collocation (Pseudospectral) Method:** Force the residual to be exactly zero at a set of carefully chosen points  $\{x_j\}$ , called collocation points:

$$R(x_j; a_0, \dots, a_N) = 0, \quad j = 1, 2, \dots, N + 1.$$

This gives  $N + 1$  equations for  $N + 1$  unknowns.

2. **Galerkin Method:** Require the residual to be orthogonal to each basis function in the sense of a weighted inner product:

$$\int_{-1}^1 R(x) \varphi_k(x) w(x) dx = 0, \quad k = 0, 1, \dots, N,$$

where  $w(x)$  is a weight function (often  $w(x) = 1$  for polynomial bases).

Both methods convert the differential equation into a system of algebraic equations for the unknown coefficients. The collocation approach is simpler to implement and handles nonlinear terms easily, while the Galerkin approach often provides better global accuracy in weighted norms.

## 3.2 A First Collocation Example

We illustrate the collocation method with a complete worked example that can be verified by hand calculation.

### 3.2.1 Problem Statement

Consider the boundary value problem on  $[-1, 1]$ :

$$u''(x) - (4x^2 + 2)u(x) = 0, \quad -1 \leq x \leq 1,$$

with boundary conditions

$$u(-1) = 1, \quad u(1) = 1.$$

### 3.2.2 The Exact Solution

The exact solution is

$$u_{\text{exact}}(x) = e^{x^2-1}.$$

Let us verify this claim. The first derivative is

$$u'_{\text{exact}}(x) = 2xe^{x^2-1}.$$

The second derivative is

$$u''_{\text{exact}}(x) = (2 + 4x^2)e^{x^2-1} = (4x^2 + 2)u_{\text{exact}}(x).$$

Substituting into the ODE:

$$u''_{\text{exact}} - (4x^2 + 2)u_{\text{exact}} = (4x^2 + 2)u_{\text{exact}} - (4x^2 + 2)u_{\text{exact}} = 0. \checkmark$$

The boundary conditions are satisfied:

$$u_{\text{exact}}(\pm 1) = e^{1-1} = e^0 = 1. \checkmark$$

### 3.2.3 Trial Function

To satisfy the boundary conditions automatically, we write the approximation in a form that equals 1 at  $x = \pm 1$  regardless of the coefficient values. A convenient choice is:

$$u_2(x) = 1 + (1 - x^2)(a_0 + a_1x + a_2x^2).$$

The factor  $(1 - x^2)$  vanishes at the endpoints, so

$$u_2(\pm 1) = 1 + 0 \cdot (\dots) = 1$$

for any values of  $a_0, a_1, a_2$ . We have three undetermined coefficients.

Expanding the trial function:

$$\begin{aligned} u_2(x) &= 1 + a_0 + a_1x + a_2x^2 - a_0x^2 - a_1x^3 - a_2x^4 \\ &= (1 + a_0) + a_1x + (a_2 - a_0)x^2 - a_1x^3 - a_2x^4. \end{aligned}$$

### 3.2.4 Computing the Residual

The residual is

$$R(x; a_0, a_1, a_2) = u_2''(x) - (4x^2 + 2)u_2(x).$$

Computing the second derivative of  $u_2$ :

$$u_2'(x) = a_1 + 2(a_2 - a_0)x - 3a_1x^2 - 4a_2x^3,$$

$$u_2''(x) = 2(a_2 - a_0) - 6a_1x - 12a_2x^2.$$

Substituting into the residual and simplifying (a calculation best verified with computer algebra), the residual is a polynomial of degree six in  $x$  with coefficients that depend linearly on  $a_0, a_1, a_2$ :

$$\begin{aligned} R(x) = & (-2 - 4a_0 + 2a_2) - 8a_1x + (-4 - 2a_0 - 14a_2)x^2 - 2a_1x^3 \\ & + (4a_0 - 6a_2)x^4 + 4a_1x^5 + 4a_2x^6. \end{aligned}$$

### 3.2.5 Collocation Conditions

We have three unknowns, so we choose three collocation points in the interior of the interval. A simple choice is

$$x_1 = -\frac{1}{2}, \quad x_2 = 0, \quad x_3 = \frac{1}{2}.$$

Setting the residual to zero at these points gives three linear equations:

**At  $x = -1/2$ :**

$$R\left(-\frac{1}{2}\right) = -\frac{17}{4}a_0 + \frac{33}{8}a_1 - \frac{25}{16}a_2 - 3 = 0.$$

**At  $x = 0$ :**

$$R(0) = -4a_0 + 2a_2 - 2 = 0.$$

**At  $x = 1/2$ :**

$$R\left(\frac{1}{2}\right) = -\frac{17}{4}a_0 - \frac{33}{8}a_1 - \frac{25}{16}a_2 - 3 = 0.$$

### 3.2.6 Solving the System

From the second equation:

$$-4a_0 + 2a_2 = 2 \Rightarrow a_2 = 1 + 2a_0.$$

Adding the first and third equations (the  $a_1$  terms cancel):

$$-\frac{17}{2}a_0 - \frac{25}{8}a_2 = 6.$$

Substituting  $a_2 = 1 + 2a_0$ :

$$-\frac{17}{2}a_0 - \frac{25}{8}(1 + 2a_0) = 6$$

$$-\frac{17}{2}a_0 - \frac{25}{8} - \frac{25}{4}a_0 = 6$$

$$-\left(\frac{34}{4} + \frac{25}{4}\right)a_0 = 6 + \frac{25}{8}$$

$$-\frac{59}{4}a_0 = \frac{73}{8}$$

$$a_0 = -\frac{73}{118}.$$

Then

$$a_2 = 1 + 2 \cdot \left( -\frac{73}{118} \right) = 1 - \frac{146}{118} = -\frac{28}{118} = -\frac{14}{59}.$$

Subtracting the first equation from the third:

$$-\frac{33}{4}a_1 = 0 \Rightarrow a_1 = 0.$$

The vanishing of  $a_1$  reflects the symmetry of the problem: both the differential equation and the boundary conditions are symmetric about  $x = 0$ , so the solution must be an even function. An odd coefficient like  $a_1$  would break this symmetry.

### 3.2.7 The Approximate Solution

Substituting the coefficients back:

$$u_2(x) = 1 + (1 - x^2) \left( -\frac{73}{118} - \frac{14}{59}x^2 \right).$$

After simplification, this becomes the even polynomial:


$$u_2(x) = \frac{14}{59}x^4 + \frac{45}{118}x^2 + \frac{45}{118}.$$

The boundary conditions are satisfied:

$$u_2(\pm 1) = \frac{14}{59} + \frac{45}{118} + \frac{45}{118} = \frac{28}{118} + \frac{90}{118} = \frac{118}{118} = 1. \checkmark$$


The implementation of this approximation is straightforward. In Python:

```
1 def u_approx(x):
2     """Evaluate the collocation approximation."""
3     a0, a1, a2 = -73/118, 0, -14/59
4     return 1 + (1 - x**2) * (a0 + a1*x + a2*x**2)
```

 Python

The equivalent MATLAB implementation:

```
1 % Collocation coefficients
2 a0 = -73/118; a1 = 0; a2 = -14/59;
3
4 % Approximate solution (anonymous function)
5 u_approx = @(x) 1 + (1 - x.^2) .* (a0 + a1*x + a2*x.^2);
```

 Matlab

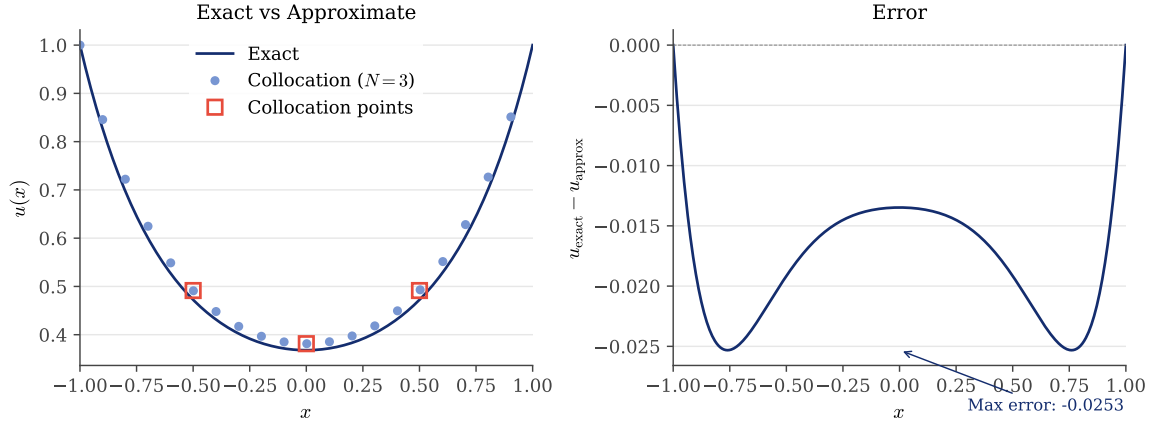
### 3.2.8 Error Analysis

The following table compares the exact and approximate solutions at several points:

$x$	$u_{\text{exact}}(x)$	$u_{\text{approx}}(x)$	Error
-1	1.00000	1.00000	0.00000
-0.5	0.47237	0.49153	-0.01916
0	0.36788	0.38136	-0.01348
0.5	0.47237	0.49153	-0.01916
1	1.00000	1.00000	0.00000

**Table 1:** Comparison of exact and three-coefficient collocation approximation.

The maximum error is approximately  $2 \times 10^{-2}$ , which is remarkably good for such a low-order approximation. Figure 6 shows the solutions graphically.



**Figure 6:** Left: exact solution  $u(x) = e^{x^2-1}$  compared with the three-coefficient collocation approximation. The collocation points  $x = -1/2, 0, 1/2$  are marked with squares. Right: the error  $u_{\text{exact}} - u_{\text{approx}}$ .

The code that generated this figure is available in both Python and MATLAB:

- codes/python/ch03\_mise\_en\_bouche/collocation\_example1.py
- codes/matlab/ch03\_mise\_en\_bouche/collocation\_example1.m

### 3.3 Collocation versus Galerkin

To compare the two main approaches to the Method of Weighted Residuals, we consider a second example where both methods can be applied with explicit hand calculations.

#### 3.3.1 Problem Statement

Consider the reaction-diffusion equation on  $[-1, 1]$ :

$$\frac{d^2 u}{dx^2} - 4u = -1, \quad -1 \leq x \leq 1,$$

with homogeneous Dirichlet boundary conditions:

$$u(-1) = 0, \quad u(1) = 0.$$

#### 3.3.2 The Exact Solution

The homogeneous equation  $u'' - 4u = 0$  has the general solution  $u_h = A \cosh(2x) + B \sinh(2x)$ . A particular solution for the constant forcing  $-1$  is  $u_p = 1/4$ . The general solution is therefore

$$u(x) = A \cosh(2x) + B \sinh(2x) + \frac{1}{4}.$$

The boundary condition  $u(-1) = 0$  gives  $A \cosh(2) - B \sinh(2) + 1/4 = 0$ . The boundary condition  $u(1) = 0$  gives  $A \cosh(2) + B \sinh(2) + 1/4 = 0$ .

Adding these equations:  $2A \cosh(2) + 1/2 = 0$ , so  $A = -1/(4 \cosh(2))$ . Subtracting:  $2B \sinh(2) = 0$ , so  $B = 0$ .

The exact solution is:

$$u_{\text{exact}}(x) = \frac{1}{4} \left( 1 - \frac{\cosh(2x)}{\cosh(2)} \right).$$

The maximum value occurs at  $x = 0$ :

$$u_{\text{exact}}(0) = \frac{1}{4} \left( 1 - \frac{1}{\cosh(2)} \right) \approx 0.1834.$$

### 3.3.3 Trial Function and Basis

Since the boundary conditions are homogeneous, we choose basis functions that automatically vanish at  $x = \pm 1$ . For a symmetric problem like this one, we use even functions:

$$\varphi_0(x) = 1 - x^2, \quad \varphi_1(x) = (1 - x^2)x^2 = x^2 - x^4.$$

Both functions vanish at  $x = \pm 1$  and are even in  $x$ . Our trial function is:

$$u_1(x) = a_0\varphi_0(x) + a_1\varphi_1(x) = a_0(1 - x^2) + a_1(x^2 - x^4).$$

### 3.3.4 The Residual

The operator is  $\mathcal{L}u = u'' - 4u$ . We compute:

$$\varphi_0'' = -2, \quad \varphi_1'' = 2 - 12x^2.$$

Applying the operator to each basis function:

$$\mathcal{L}\varphi_0 = -2 - 4(1 - x^2) = -6 + 4x^2,$$

$$\mathcal{L}\varphi_1 = (2 - 12x^2) - 4(x^2 - x^4) = 2 - 16x^2 + 4x^4.$$

The residual is:

$$\begin{aligned} R(x) &= a_0\mathcal{L}\varphi_0 + a_1\mathcal{L}\varphi_1 - (-1) \\ &= a_0(-6 + 4x^2) + a_1(2 - 16x^2 + 4x^4) + 1. \end{aligned}$$

### 3.3.5 Collocation Method

With two unknowns, we need two collocation points. Due to symmetry, we can use points in  $[0, 1)$ . We choose  $x_1 = 0$  and  $x_2 = 0.5$ .

**At  $x = 0$ :**

$$\mathcal{L}\varphi_0(0) = -6, \quad \mathcal{L}\varphi_1(0) = 2.$$

$$R(0) = -6a_0 + 2a_1 + 1 = 0 \Rightarrow 6a_0 - 2a_1 = 1.$$

**At  $x = 0.5$ :**

$$\mathcal{L}\varphi_0(0.5) = -6 + 4 \cdot 0.25 = -5,$$

$$\mathcal{L}\varphi_1(0.5) = 2 - 16 \cdot 0.25 + 4 \cdot 0.0625 = 2 - 4 + 0.25 = -1.75.$$

$$R(0.5) = -5a_0 - 1.75a_1 + 1 = 0 \Rightarrow 5a_0 + 1.75a_1 = 1.$$

Solving the system:

$$\begin{cases} 6a_0 - 2a_1 = 1 \\ 5a_0 + 1.75a_1 = 1 \end{cases}$$

Multiply the first equation by 5 and the second by 6:

$$30a_0 - 10a_1 = 5, \quad 30a_0 + 10.5a_1 = 6.$$

Subtracting:  $20.5a_1 = 1$ , so  $a_1 \approx 0.04878$ .

Substituting back:  $6a_0 = 1 + 2 \cdot 0.04878 = 1.09756$ , so  $a_0 \approx 0.1829$ .

The collocation solution is:


$$u_{\text{coll}}(x) \approx 0.1829(1 - x^2) + 0.0488(x^2 - x^4).$$

The following Python code assembles and solves the collocation system:

```

1 def solve_collocation():
2     """Solve the collocation system at x = 0 and x = 0.5."""
3     # Operator L[φ] = φ'' - 4φ applied to basis functions
4     L_phi0 = lambda x: -2 - 4*(1 - x**2)
5     L_phi1 = lambda x: (2 - 12*x**2) - 4*(x**2 - x**4)
6
7     # Build system matrix at collocation points
8     A = np.array([[L_phi0(0.0), L_phi1(0.0)],
9                  [L_phi0(0.5), L_phi1(0.5)]])
10    b = np.array([-1.0, -1.0]) # RHS from f = -1
11    return np.linalg.solve(A, b)

```

 Python

The equivalent MATLAB implementation:

```

1 % Operator L = d^2/dx^2 - 4 applied to basis functions
2 L_phi0 = @(x) -2 - 4*(1 - x.^2);
3 L_phi1 = @(x) (2 - 12*x.^2) - 4*(x.^2 - x.^4);
4
5 % Build and solve collocation system
6 A_coll = [L_phi0(0.0), L_phi1(0.0);
7           L_phi0(0.5), L_phi1(0.5)];
8 coeffs = A_coll \ [-1; -1];

```

 Matlab

### 3.3.6 Galerkin Method

The Galerkin conditions require the residual to be orthogonal to each basis function:

$$\int_{-1}^1 R(x) \varphi_k(x) dx = 0, \quad k = 0, 1.$$

This gives a symmetric matrix system  $\mathbf{A}\mathbf{a} = \mathbf{b}$  where:

$$A_{ij} = \int_{-1}^1 \mathcal{L}\varphi_j(x) \cdot \varphi_i(x) dx,$$

$$b_i = \int_{-1}^1 (-1) \cdot \varphi_i(x) dx = - \int_{-1}^1 \varphi_i(x) dx.$$

Computing the integrals (using standard formulas for powers of  $x$ ):

For  $\varphi_0 = 1 - x^2$ :

$$\int_{-1}^1 \varphi_0(x) dx = \int_{-1}^1 (1 - x^2) dx = \left[ x - \frac{x^3}{3} \right]_{-1}^1 = 2 - \frac{2}{3} = \frac{4}{3}.$$

For  $\varphi_1 = x^2 - x^4$ :

$$\int_{-1}^1 \varphi_1(x) dx = \left[ \frac{x^3}{3} - \frac{x^5}{5} \right]_{-1}^1 = \frac{2}{3} - \frac{2}{5} = \frac{4}{15}.$$

The right-hand side is:

$$b_0 = -\frac{4}{3}, \quad b_1 = -\frac{4}{15}.$$

The matrix entries require more computation. Using computer algebra or careful integration:

$$A_{00} = -\frac{104}{15}, \quad A_{01} = A_{10} = -\frac{8}{7}, \quad A_{11} = -\frac{328}{315}.$$

Solving the system yields:

$$a_0 \approx 0.1832, \quad a_1 \approx 0.0550.$$

The Galerkin method requires numerical integration to assemble the system. In Python:

```
1 def solve_galerkin():
2     """Solve using Galerkin:  $\langle R, \phi_k \rangle = 0$  for  $k = 0, 1$ ."""
3     from scipy import integrate
4
5     # Matrix entries:  $A_{ij} = \int L[\phi_j] \phi_i dx$ 
6     A00, _ = integrate.quad(lambda x: L_phi0(x) * phi0(x), -1, 1)
7     A01, _ = integrate.quad(lambda x: L_phi1(x) * phi0(x), -1, 1)
8     A10, _ = integrate.quad(lambda x: L_phi0(x) * phi1(x), -1, 1)
9     A11, _ = integrate.quad(lambda x: L_phi1(x) * phi1(x), -1, 1)
10
11     A = np.array([[A00, A01], [A10, A11]])
12     b0, _ = integrate.quad(lambda x: -phi0(x), -1, 1)
13     b1, _ = integrate.quad(lambda x: -phi1(x), -1, 1)
14     return np.linalg.solve(A, [b0, b1])
```

The equivalent MATLAB implementation uses the built-in integral function:

```
1 % Matrix entries:  $A_{ij} = \int L[\phi_j] \phi_i dx$ 
2 A00 = integral(@(x) L_phi0(x) .* phi0(x), -1, 1);
3 A01 = integral(@(x) L_phi1(x) .* phi0(x), -1, 1);
4 A10 = integral(@(x) L_phi0(x) .* phi1(x), -1, 1);
5 A11 = integral(@(x) L_phi1(x) .* phi1(x), -1, 1);
6
7 A_gal = [A00, A01; A10, A11];
8
9 % RHS:  $b_i = \int f \phi_i dx$  where  $f = -1$ 
10 b0 = integral(@(x) -phi0(x), -1, 1);
11 b1 = integral(@(x) -phi1(x), -1, 1);
12 coeffs = A_gal \ [b0; b1];
```

### 3.3.7 Comparison

The following table compares the two methods at the central point  $x = 0$ :

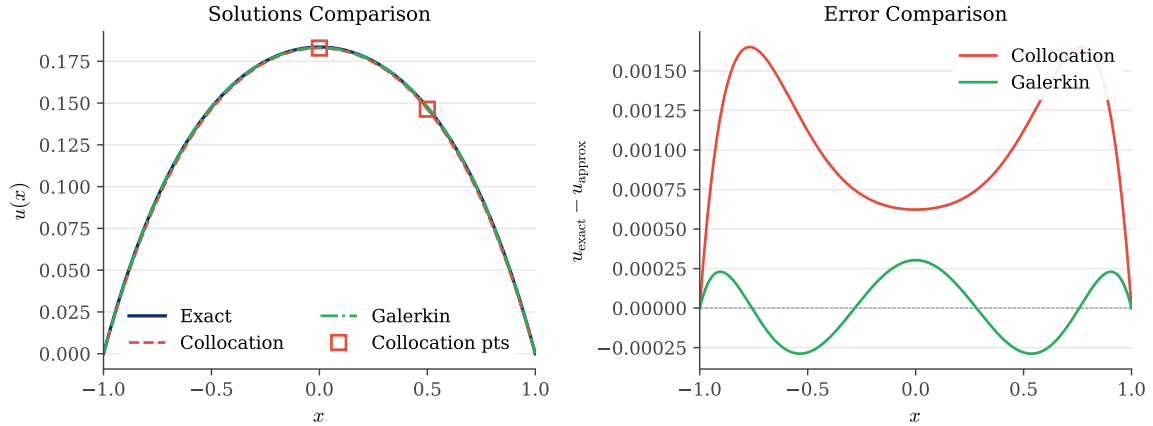
Method	$u(0)$	Absolute Error
Exact	0.1835	0
Collocation ( $N = 2$ )	0.1829	0.0006
Galerkin ( $N = 2$ )	0.1832	0.0003

**Table 2:** Comparison of spectral approximations at the central maximum.



For this problem, the Galerkin method is more accurate both at the central point and in a global sense. This is consistent with the error plot in Figure 7, which shows the Galerkin error (green) remaining closer to zero across the entire interval. The Galerkin method minimizes the error in a root-mean-square sense, which typically leads to better overall accuracy for smooth problems.

Figure 7 shows both approximate solutions compared to the exact solution, along with the error profiles.



**Figure 7:** Left: exact solution compared with collocation and Galerkin approximations. The collocation points  $x = 0$  and  $x = 0.5$  are marked. Right: error profiles for both methods.

The code that generated this figure is available in both Python and MATLAB:

- `codes/python/ch03_mise_en_bouche/collocation_vs_galerkin.py`
- `codes/matlab/ch03_mise_en_bouche/collocation_vs_galerkin.m`

### 3.4 Conclusions and Questions

These simple examples illuminate several important features of spectral methods:

1. **Automatic satisfaction of boundary conditions.** By choosing trial functions that vanish at the boundaries (or equal the prescribed boundary values), we eliminate the boundary conditions from the algebraic system.
2. **Symmetry exploitation.** When the problem has symmetry, the solution inherits that symmetry. In our examples, the vanishing of odd coefficients ( $a_1 = 0$ ) reflects the even symmetry of the exact solution.
3. **Simplicity of implementation.** Even with hand calculations, we can obtain remarkably accurate approximations with just a few coefficients.
4. **Trade-offs between methods.** Collocation is simpler to implement (no integrals to compute), while Galerkin typically provides better accuracy by minimizing a global error measure.

These examples raise important questions that we will address in subsequent chapters:

- **What is the optimal choice of basis functions?** Using simple powers of  $x$  works for small  $N$ , but becomes numerically unstable for large  $N$ . Chebyshev and Fourier bases are far superior.
- **What are the optimal collocation points?** Our ad hoc choices  $x = -1/2, 0, 1/2$  worked well, but there exist optimal point distributions (Gauss and Gauss–Lobatto points) derived from orthogonal polynomial theory.
- **How fast does the error decrease as  $N$  increases?** For smooth solutions, spectral methods achieve exponential convergence (the error decreases like  $c^{-N}$  for some  $c > 1$ ), which is dramatically faster than the algebraic convergence  $O(N^{-p})$  of finite difference and finite element methods.

The following chapters will develop the theory and algorithms needed to answer these questions and to apply spectral methods to a wide range of problems.

## 3.5 A Broader Perspective

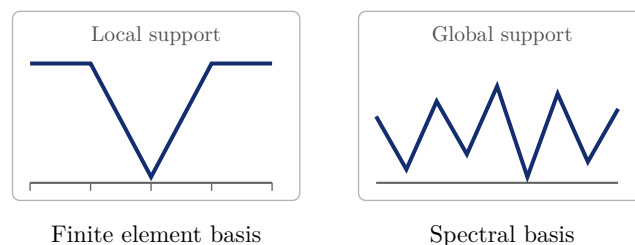
For demanding students who wish to understand how spectral methods fit into the wider landscape of numerical analysis, this section provides a high-level comparison with other approaches, particularly finite element methods. The discussion follows [3, Section 1.3].

### 3.5.1 Local versus Global Basis Functions

The fundamental distinction between spectral methods and finite element methods lies in the *support* of the basis functions. In finite element methods, the computational domain is divided into many small sub-intervals (or triangles, tetrahedra in higher dimensions), and the basis functions  $\varphi_n(x)$  are *local*: they are polynomials of fixed, low degree (typically linear or quadratic) that are non-zero only over one or two adjacent elements.

In contrast, spectral methods use *global* basis functions. Each  $\varphi_n(x)$  is a polynomial (or trigonometric polynomial) of potentially high degree that is non-zero (except at isolated points) over the entire computational domain. This global character is both the source of spectral methods' power and the reason for some of their limitations.

Figure 8 illustrates this distinction schematically. The finite element basis function (left) has compact support and contributes to the solution only locally. The spectral basis function (right) influences the solution everywhere.

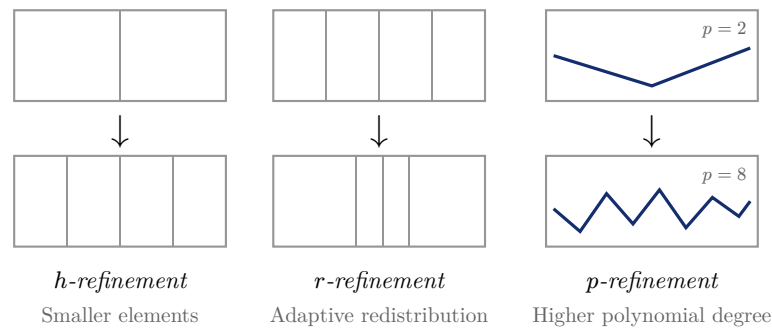


**Figure 8:** Schematic comparison of local (finite element) and global (spectral) basis functions. The finite element “hat function” is non-zero only over two adjacent elements, while the spectral basis function oscillates across the entire domain.

### 3.5.2 Refinement Strategies

When a numerical approximation is insufficiently accurate, there are three fundamentally different strategies to improve it, illustrated schematically in Figure 9:

1.  **$h$ -refinement**: Subdivide each element into smaller pieces, reducing the mesh spacing  $h$  uniformly throughout the domain. This increases the number of elements while keeping the polynomial degree fixed.
2.  **$r$ -refinement** (adaptive): Redistribute the mesh points, clustering them in regions where the solution has steep gradients or other features requiring high resolution. The total number of degrees of freedom remains roughly constant.
3.  **$p$ -refinement**: Keep the mesh fixed while increasing  $p$ , the polynomial degree within each element. For a single-element domain, this is precisely what spectral methods do.



**Figure 9:** Three strategies for improving accuracy in numerical methods. Spectral methods employ  $p$ -refinement: increasing the polynomial degree while using a single element (or few elements) spanning the entire domain.

Spectral methods can be viewed as the extreme form of  $p$ -refinement: a single element spans the entire domain, and accuracy is improved solely by increasing the polynomial degree. This strategy is devastatingly effective when the solution is smooth, but struggles when the solution has discontinuities or sharp gradients.

### 3.5.3 Trade-offs: Sparse versus Full Matrices

The choice between local and global basis functions entails fundamental trade-offs:

**Finite element advantages:**

- *Sparse matrices*: Since each basis function is non-zero over only a few elements, the stiffness matrix has mostly zero entries. Sparse matrix solvers can exploit this structure, dramatically reducing computational cost for large systems.
- *Geometric flexibility*: The small elements (triangles, tetrahedra) can be fitted to irregularly shaped domains like automobile bodies or aircraft wings.

**Finite element disadvantages:**

- *Low accuracy per degree of freedom*: Each basis function is a polynomial of low degree, so many elements are needed for high accuracy.

**Spectral method advantages:**

- *High accuracy for smooth problems*: The high-degree global polynomials capture smooth solutions with far fewer degrees of freedom.

- *Efficiency with iterative solvers:* When fast iterative methods are used, spectral methods can be much more efficient than low-order methods for many problem classes.

### Spectral method disadvantages:

- *Full matrices:* The global basis functions create dense matrices where most entries are non-zero.
- *Geometric limitations:* Spectral methods are most natural on simple domains (intervals, rectangles, disks) and require more sophisticated techniques for irregular geometries.

For problems with smooth solutions on regular domains (many important problems in fluid dynamics, quantum mechanics, and wave propagation fall into this category), the accuracy advantage of spectral methods often outweighs the matrix structure disadvantage.

#### 3.5.4 Spectral Element Methods

A natural question arises: can we combine the geometric flexibility of finite elements with the high accuracy of spectral methods? The answer is yes, through *spectral element methods*.

In spectral element methods, the domain is subdivided into elements (as in finite elements), but within each element, the polynomial degree  $p$  is chosen to be moderately high, typically  $p = 6$  to  $8$ . This hybrid approach inherits several advantages:

- The element subdivision provides geometric flexibility and matrix sparsity.
- The high polynomial degree within each element provides spectral-like accuracy.
- The theoretical framework is essentially the same as for global spectral methods.

Spectral element codes are typically written so that  $p$  is a user-adjustable parameter, allowing practitioners to balance accuracy and cost for their specific application. We will not develop spectral element methods in detail in this book, but the reader should be aware that much of the theory developed for global spectral methods transfers directly to the spectral element context.

#### 3.5.5 The Convergence of Methods at High Order

Perhaps the most profound insight from the comparison between finite element and spectral methods is this: *for sufficiently high polynomial degree, the two approaches become essentially equivalent.*

Low-order finite elements (linear, quadratic) can be derived, justified, and implemented without knowledge of Fourier or Chebyshev convergence theory. However, as the polynomial degree increases, ad hoc approaches become increasingly ill-conditioned and numerically unstable. The only practical way to implement well-behaved high-order finite elements (say, sixth order or higher) is to use the technology of spectral methods: Chebyshev or Legendre basis functions, Gaussian quadrature, and the convergence theory we will develop in subsequent chapters.

Thus, the question “Are finite elements or spectral methods better?” becomes somewhat artificial for high-order approximations. The real question is: *Does the problem at hand require high-order accuracy, or is second or fourth order sufficient?*

When the solution is smooth and high accuracy is needed, the spectral/high-order approach is clearly superior. When the solution has discontinuities, shocks, or boundary layers, or when the geometry is highly irregular, low-order methods with adaptive mesh refinement may be more practical. The wise practitioner chooses the tool appropriate to the problem.





## CHAPTER 4

# The Geometry of Nodes

**B**efore we can differentiate functions numerically using spectral methods, we must first understand how to represent them. Polynomial interpolation—the process of constructing a polynomial that passes through a given set of data points—is the foundation upon which pseudospectral methods are built. In this chapter, we explore a fascinating paradox: while polynomial interpolation seems entirely straightforward, the choice of interpolation nodes determines whether the method succeeds brilliantly or fails catastrophically. The story begins with a surprising discovery by the German mathematician Carl Runge in 1901. Attempting to approximate a simple, smooth function by interpolating polynomials, Runge found that increasing the polynomial degree made the approximation *worse*, not better. This counterintuitive phenomenon, now bearing his name, reveals deep connections between numerical analysis, complex analysis, and potential theory.

Our journey through this “étude in grid geometry” will explain the Runge phenomenon, develop the theoretical framework of potential theory that predicts where interpolation succeeds or fails, and introduce the Chebyshev points that form the cornerstone of practical spectral methods.

## 4.1 The Problem: Polynomial Interpolation

### 4.1.1 Lagrange Interpolation

Given  $N + 1$  distinct points  $\{x_0, x_1, \dots, x_N\}$  on an interval  $[a, b]$  and corresponding function values  $\{f_0, f_1, \dots, f_N\}$  where  $f_j = f(x_j)$ , there exists a unique polynomial  $p_N(x)$  of degree at most  $N$  such that

$$p_N(x_j) = f_j, \quad j = 0, 1, \dots, N.$$

This interpolating polynomial can be written explicitly using the *Lagrange formula*:

$$p_N(x) = \sum_{k=0}^N f_k L_k(x),$$

where the *Lagrange basis polynomials* are

$$L_k(x) = \prod_{j=0, j \neq k}^N \frac{x - x_j}{x_k - x_j}.$$

Each basis polynomial  $L_k(x)$  has the cardinal property: it equals 1 at  $x_k$  and 0 at all other nodes. This property ensures that substituting any node  $x_j$  into the Lagrange formula yields  $f_j$ .

### 4.1.2 Interpolation versus Best Approximation

It is important to distinguish interpolation from best approximation. The Weierstrass Approximation Theorem guarantees that any continuous function on  $[a, b]$  can be uniformly approximated to arbitrary accuracy by polynomials. A constructive proof was given by Sergei Bernstein in 1912 using what are now called *Bernstein polynomials*—explicit polynomial approximations that converge uniformly to any continuous function, though typically more slowly than interpolation-based methods. However, this theorem says nothing about *which* polynomials achieve this approximation or how to construct them efficiently.

Interpolation constructs a specific polynomial by enforcing exact agreement at the nodes. The hope is that as  $N \rightarrow \infty$ , the interpolating polynomials  $p_N$  converge uniformly to  $f$ . As we shall see, this hope is fulfilled for some node distributions but dramatically fails for others.

### 4.1.3 Equispaced Nodes

The most natural choice of interpolation nodes is equally spaced points:

$$x_j = -1 + \frac{2j}{N}, \quad j = 0, 1, \dots, N.$$

These nodes divide the interval  $[-1, 1]$  into  $N$  equal subintervals. For low-degree polynomials and well-behaved functions, equispaced interpolation works reasonably well. However, a fundamental problem emerges as  $N$  increases.

## 4.2 The Runge Phenomenon

### 4.2.1 A Smooth but Troublesome Function

In 1901, Carl Runge studied the interpolation of a deceptively simple function:

$$f(x) = \frac{1}{1 + 25x^2}.$$

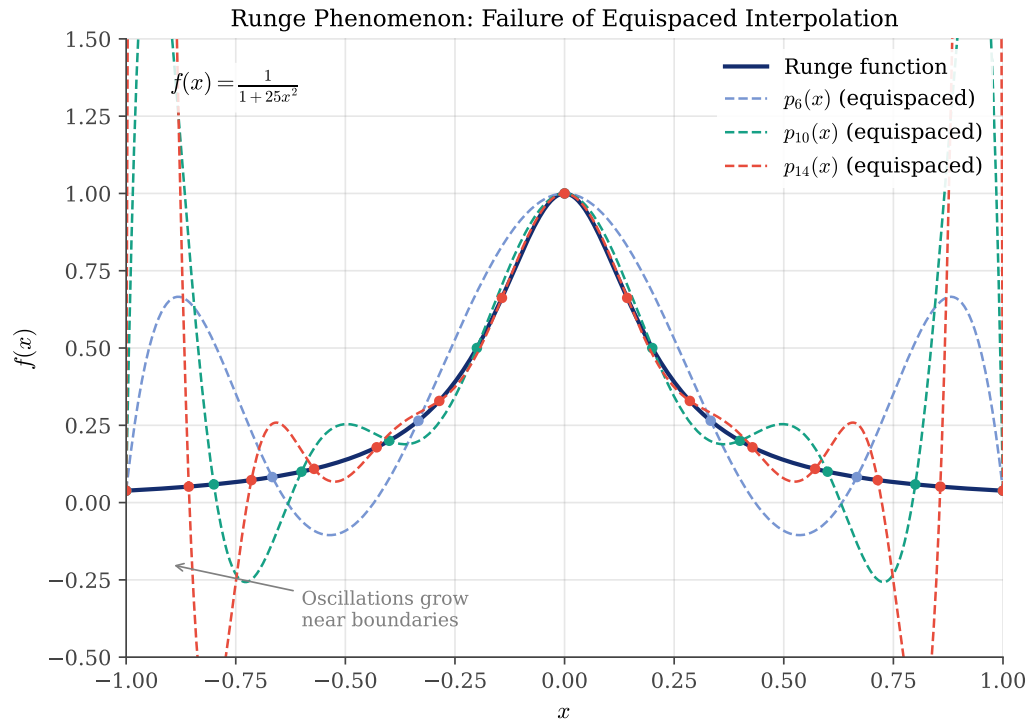
This *Runge function* is infinitely differentiable on the entire real line. Its graph is a smooth bell curve centered at the origin with maximum value  $f(0) = 1$  and asymptotic decay to zero as  $|x| \rightarrow \infty$ .

Runge discovered that polynomial interpolation on equispaced nodes *diverges* for this function. Rather than improving with increasing  $N$ , the interpolating polynomials develop wild oscillations near the endpoints  $x = \pm 1$ .

### 4.2.2 Numerical Demonstration

Figure 10 illustrates this phenomenon. For  $N = 6$ , the interpolant reasonably approximates the function. But for  $N = 10$  and  $N = 14$ , the approximation deteriorates dramatically at the boundaries, with oscillations that grow unboundedly as  $N$  increases.






**Figure 10:** The Runge phenomenon: polynomial interpolation of  $f(x) = 1/(1 + 25x^2)$  on equispaced nodes. As the polynomial degree increases, oscillations near the boundaries  $x = \pm 1$  grow unboundedly. The exact function (solid) and interpolants (dashed) are shown for  $N = 6, 10$ , and  $14$ .

The following Python code computes the Lagrange interpolant for a given set of nodes:

```

1  def lagrange_interpolate(x_nodes, f_nodes, x_eval):
2      """Evaluate the Lagrange interpolating polynomial."""
3      n = len(x_nodes)
4      p_eval = np.zeros_like(x_eval)
5
6      for k in range(n):
7          L_k = np.ones_like(x_eval)
8          for j in range(n):
9              if j != k:
10                 L_k *= (x_eval - x_nodes[j]) / (x_nodes[k] - x_nodes[j])
11             p_eval += f_nodes[k] * L_k
12
13     return p_eval

```


 Python

The equivalent MATLAB implementation:

```

1  function p = lagrange_interp(x_nodes, f_nodes, x_eval)
2      n = length(x_nodes);
3      p = zeros(size(x_eval));
4
5      for k = 1:n
6          L_k = ones(size(x_eval));
7          for j = 1:n

```

 Matlab

```

8         if j ~= k
9             L_k = L_k .* (x_eval - x_nodes(j)) / (x_nodes(k) - x_nodes(j));
10        end
11    end
12    p = p + f_nodes(k) * L_k;
13 end
14 end

```

The code generating Figure 10 is available in:

- codes/python/ch04\_geometry\_of\_nodes/runge\_phenomenon.py
- codes/matlab/ch04\_geometry\_of\_nodes/runge\_phenomenon.m

### 4.2.3 Why Does This Happen?

The Runge phenomenon seems paradoxical: how can adding more information (more interpolation points) make the approximation worse? The answer lies in the *Lebesgue constant*, which we will explore in Section 4.5. But first, let us develop a deeper understanding through potential theory.

## 4.3 Theoretical Explanation: Potential Theory

### 4.3.1 Singularities in the Complex Plane

The key to understanding the Runge phenomenon lies in extending our view from the real line to the complex plane. The Runge function has singularities (poles) where its denominator vanishes:

$$1 + 25z^2 = 0 \quad \Rightarrow \quad z = \pm \frac{i}{5} = \pm 0.2i.$$

These poles lie on the imaginary axis, at a distance of only 0.2 from the real interval  $[-1, 1]$ . This proximity is responsible for the divergence of equispaced interpolation.

### 4.3.2 The Potential Function

The convergence of polynomial interpolation is governed by a *potential function* associated with the node distribution. For a distribution with density  $\mu(x)$  on  $[-1, 1]$ , the potential at a point  $z$  in the complex plane is:

$$\varphi(z) = - \int_{-1}^1 \mu(x) \ln|z - x| \, dx.$$

The *equipotential curves*  $\varphi(z) = \text{const}$  form closed contours around the interval  $[-1, 1]$ . The largest equipotential curve that does not enclose any singularity of  $f$  determines the region of convergence.

### 4.3.3 Uniform versus Chebyshev Density

For equispaced nodes, the density is asymptotically uniform:  $\mu(x) = 1/2$ . The resulting equipotential curves are roughly elliptical, but they extend only a short distance from the interval before reaching the Runge function's poles at  $\pm 0.2i$ .

For Chebyshev nodes (which we introduce in the next section), the density is:

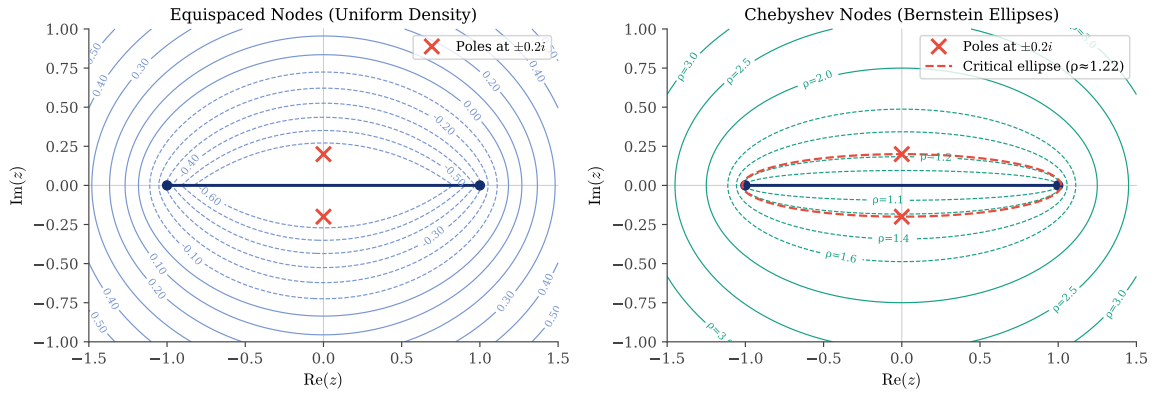
$$\mu(x) = \frac{1}{\pi\sqrt{1-x^2}}.$$

This density diverges at the endpoints, concentrating nodes near  $x = \pm 1$ . The corresponding potential simplifies to:

$$\varphi(z) = \ln|z + \sqrt{z^2 - 1}| - \ln 2 = \ln \rho - \ln 2,$$

where  $\rho = |z + \sqrt{z^2 - 1}|$ . The equipotential curves are *Bernstein ellipses* with foci at  $\pm 1$ .

Figure 11 compares the equipotential curves for both distributions, showing why Chebyshev interpolation succeeds where equispaced interpolation fails.



**Figure 11:** Equipotential curves in the complex plane. Left: uniform node density (equispaced nodes). Right: Chebyshev density, where equipotentials are Bernstein ellipses. The poles of the Runge function at  $z = \pm 0.2i$  are marked with crosses. For Chebyshev interpolation, the critical ellipse passing through the poles determines the convergence rate.

The code generating Figure 11 is available in:

- `codes/python/ch04_geometry_of_nodes/equipotential_curves.py`
- `codes/matlab/ch04_geometry_of_nodes/equipotential_curves.m`

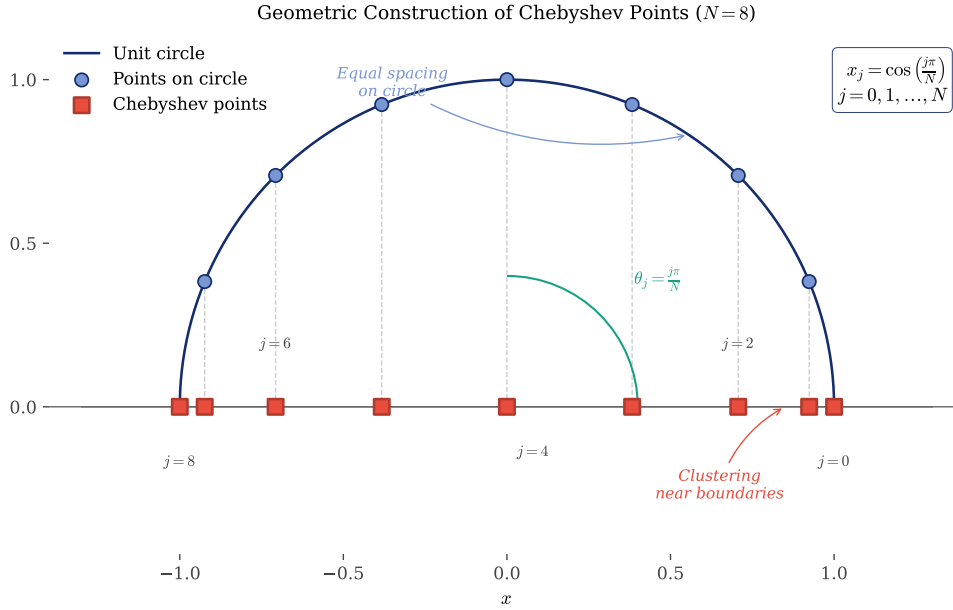
## 4.4 The Solution: Chebyshev Points

### 4.4.1 Definition and Geometric Construction

The *Chebyshev-Gauss-Lobatto points* (often simply called Chebyshev points) are defined by:

$$x_j = \cos\left(\frac{j\pi}{N}\right), \quad j = 0, 1, \dots, N.$$

These points have a beautiful geometric interpretation: they are the projections onto the  $x$ -axis of  $N + 1$  equally spaced points on the upper semicircle of the unit circle. Figure 12 illustrates this construction.



**Figure 12:** Geometric construction of Chebyshev points. Points equally spaced on the upper semicircle (by angle  $\theta_j = j\pi/N$ ) project vertically onto the Chebyshev nodes on the  $x$ -axis. Equal spacing on the circle maps to clustering near the endpoints  $x = \pm 1$ .

#### 4.4.2 Node Density and Clustering

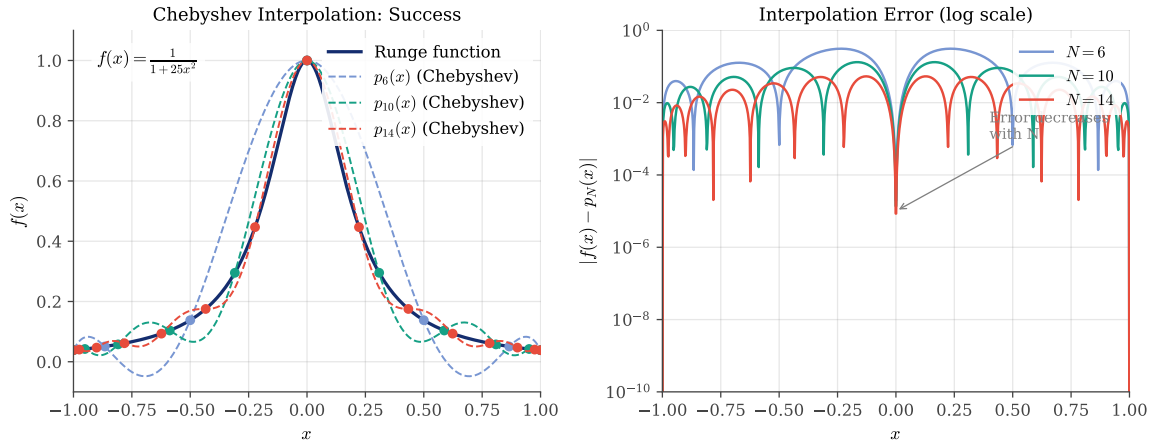
The geometric construction reveals why Chebyshev points cluster near the endpoints. Equal angular spacing on the circle corresponds to denser horizontal spacing near  $x = \pm 1$ . The node density is approximately:

$$\rho(x) \approx \frac{N}{\pi\sqrt{1-x^2}},$$

which diverges as  $x \rightarrow \pm 1$ . This clustering counteracts the growth of interpolation error at the boundaries.

#### 4.4.3 Chebyshev Interpolation of the Runge Function

Figure 13 demonstrates the dramatic improvement when using Chebyshev points instead of equispaced nodes for the Runge function. Where equispaced interpolation diverges, Chebyshev interpolation converges rapidly and uniformly.



**Figure 13:** Chebyshev interpolation of the Runge function. Left: the function and interpolants for  $N = 6, 10$ , and  $14$ . Right: interpolation error on a logarithmic scale. Unlike equispaced interpolation, the error decreases rapidly with increasing  $N$ .

The Chebyshev nodes in Python and MATLAB:

```
1 def chebyshev_nodes(N):
2     """Chebyshev-Gauss-Lobatto points on [-1, 1]."""
3     j = np.arange(N + 1)
4     return np.cos(j * np.pi / N)
```

Python

```
1 % Chebyshev-Gauss-Lobatto points
2 j = 0:N;
3 x_cheb = cos(j * pi / N);
```

Matlab

Note that this formula produces nodes ordered from right to left:  $x_0 = \cos(0) = +1$  down to  $x_N = \cos(\pi) = -1$ . This ordering is natural for the cosine function and is the standard convention in spectral methods.

The code generating Figure 13 is available in:

- codes/python/ch04\_geometry\_of\_nodes/chebyshev\_success.py
- codes/matlab/ch04\_geometry\_of\_nodes/chebyshev\_success.m

## 4.5 Lagrange Basis Functions and Lebesgue Constants

### 4.5.1 The Lebesgue Function

To understand why node choice matters so dramatically, we examine the Lagrange basis functions more closely. The *Lebesgue function* is defined as the sum of absolute values of all basis polynomials:

$$\Lambda_N(x) = \sum_{k=0}^N |L_k(x)|.$$

The *Lebesgue constant* is its maximum over the interval:

$$\Lambda_N = \max_{x \in [-1, 1]} \Lambda_N(x).$$

#### 4.5.2 Error Bounds

The Lebesgue constant bounds the interpolation error. If  $p_N^*$  is the best polynomial approximation of degree  $N$  to  $f$ , and  $p_N$  is the Lagrange interpolant, then:

$$\|f - p_N\|_\infty \leq (1 + \Lambda_N) \|f - p_N^*\|_\infty = (1 + \Lambda_N) E_N(f),$$

where  $E_N(f)$  denotes the best approximation error.

This bound reveals the problem with equispaced nodes: even if  $E_N(f)$  decreases with  $N$  (as guaranteed by the Weierstrass theorem), the factor  $(1 + \Lambda_N)$  may grow so fast that the product increases.

#### 4.5.3 Asymptotic Growth Rates

The growth rate of  $\Lambda_N$  depends critically on the node distribution:

- **Chebyshev nodes:**  $\Lambda_N^{\text{Ch}} = \frac{2}{\pi} \ln N + O(1)$  (logarithmic growth)
- **Legendre nodes:**  $\Lambda_N^{\text{Leg}} = O(\sqrt{N})$  (slow algebraic growth)
- **Equispaced nodes:**  $\Lambda_N^{\text{eq}} \approx \frac{2^{N+1}}{eN \ln N}$  (exponential growth!)

The exponential growth of the Lebesgue constant for equispaced nodes explains the Runge phenomenon: even though the best approximation error decreases geometrically for smooth functions, the exponentially growing factor  $(1 + \Lambda_N)$  eventually dominates.

**Remark.** As shown by Vértési [5], the Lebesgue constant for Chebyshev nodes is remarkably close to the smallest possible Lebesgue constant for *any* node distribution:

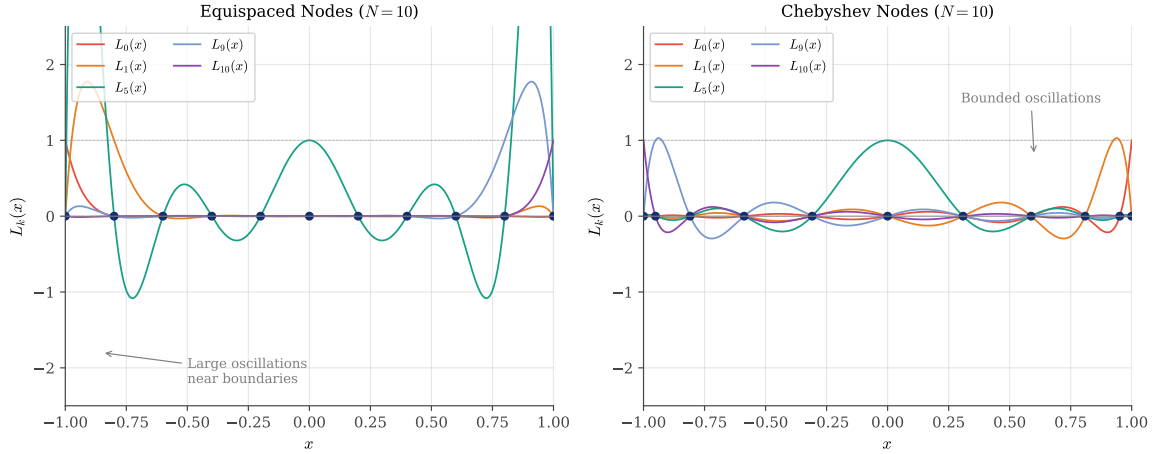
$$\Lambda_N^{\text{Ch}} = \frac{2}{\pi} \left( \ln N + \gamma + \ln \frac{8}{\pi} \right) + o(1),$$

$$\Lambda_N^{\text{min}} = \frac{2}{\pi} \left( \ln N + \gamma + \ln \frac{4}{\pi} \right) + o(1),$$

where  $\gamma \approx 0.5772156649\dots$  is the Euler–Mascheroni constant. The difference between these two expressions is only  $\frac{2}{\pi} \ln 2 \approx 0.44$ , demonstrating that Chebyshev nodes are essentially optimal for polynomial interpolation.

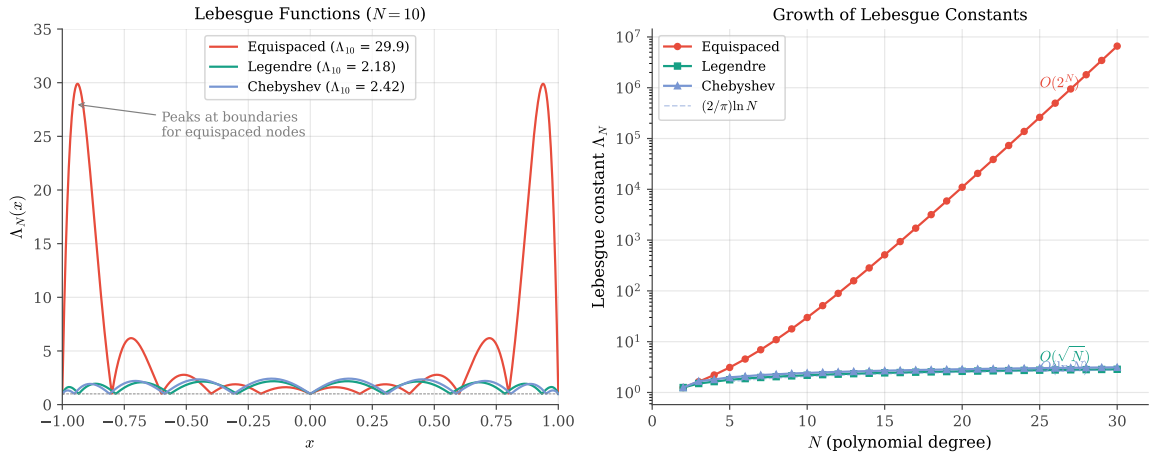
#### 4.5.4 Visualization of Basis Functions

Figure 14 compares the Lagrange basis polynomials for equispaced and Chebyshev nodes. For equispaced nodes, the basis functions near the endpoints develop large oscillations. For Chebyshev nodes, all basis functions remain well-behaved.



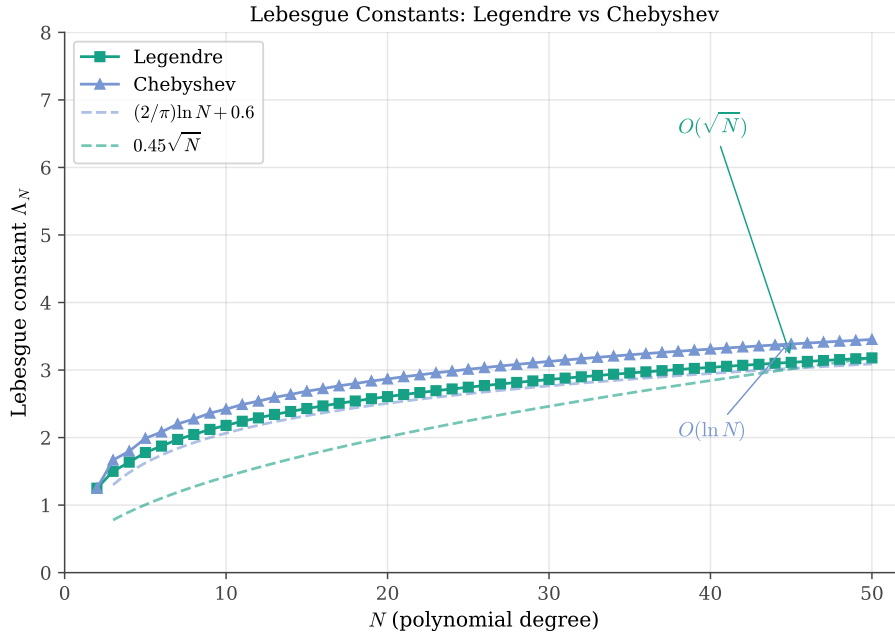
**Figure 14:** Lagrange basis polynomials  $L_{k(x)}$  for  $N = 10$ . Left: equispaced nodes show large oscillations near the boundaries. Right: Chebyshev nodes yield bounded basis functions throughout the interval.

Figure 15 shows the Lebesgue functions and the growth of Lebesgue constants for the three node distributions.



**Figure 15:** Left: Lebesgue functions  $\Lambda_N(x)$  for  $N = 10$  with equispaced, Legendre, and Chebyshev nodes. The equispaced case has large peaks near the boundaries. Right: Growth of Lebesgue constants with  $N$ . The equispaced constant grows exponentially, while Chebyshev grows only logarithmically.

The exponential growth of the equispaced Lebesgue constant dominates Figure 15, making it difficult to distinguish the Legendre and Chebyshev curves. Figure 16 provides a zoomed comparison of these two well-behaved distributions, clearly showing the  $O(\sqrt{N})$  growth of Legendre versus the superior  $O(\ln N)$  growth of Chebyshev.



**Figure 16:** *Lebesgue constants for Legendre and Chebyshev nodes (equispaced omitted for clarity). Chebyshev nodes achieve the slowest possible growth rate  $O(\ln N)$ , while Legendre nodes grow as  $O(\sqrt{N})$ . Both are far superior to equispaced nodes, but Chebyshev remains the optimal choice.*

The code generating these figures is available in:

- codes/python/ch04\_geometry\_of\_nodes/lagrange\_basis.py
- codes/python/ch04\_geometry\_of\_nodes/lebesgue\_functions.py
- codes/python/ch04\_geometry\_of\_nodes/lebesgue\_constants\_zoom.py
- codes/matlab/ch04\_geometry\_of\_nodes/lagrange\_basis.m
- codes/matlab/ch04\_geometry\_of\_nodes/lebesgue\_functions.m

## 4.6 Barycentric Interpolation

### 4.6.1 Numerical Stability Issues

The Lagrange formula, while mathematically elegant, is numerically problematic. Computing each basis polynomial  $L_k(x)$  requires  $N$  multiplications and divisions, and the resulting values can be very large or very small, leading to catastrophic cancellation when summed.

### 4.6.2 The Barycentric Formula

A more stable and efficient approach is the *barycentric interpolation formula*:

$$p_N(x) = \frac{\sum_{j=0}^N \frac{w_j}{x-x_j} f_j}{\sum_{j=0}^N \frac{w_j}{x-x_j}},$$

where the *barycentric weights* are:



$$w_j = \frac{1}{\prod_{k \neq j} (x_j - x_k)}.$$

This formula requires only  $O(N)$  operations to evaluate  $p_{N(x)}$  once the weights are precomputed.

#### 4.6.3 Weights for Chebyshev Points

For Chebyshev-Gauss-Lobatto points, the weights have a particularly simple form:

$$w_j = (-1)^j \delta_j, \quad \text{where } \delta_j = \begin{cases} 1/2 & \text{if } j = 0 \text{ or } j = N \\ 1 & \text{otherwise.} \end{cases}$$

This simplicity, combined with the excellent approximation properties, makes Chebyshev points the standard choice for spectral methods.

## 4.7 Convergence Analysis

### 4.7.1 Geometric Convergence

For functions analytic in a region containing  $[-1, 1]$ , Chebyshev interpolation converges geometrically. If  $f$  is analytic inside the Bernstein ellipse  $\mathcal{E}_\rho$  with parameter  $\rho > 1$ , then:

$$\|f - p_N\|_\infty \leq \frac{C}{\rho^N},$$

where  $C$  depends on  $f$  and  $\rho$  but not on  $N$ .

The parameter  $\rho$  is determined by the location of the nearest singularity: if the closest singularity to  $[-1, 1]$  lies on the ellipse  $\mathcal{E}_\rho$ , then convergence is at rate  $O(\rho^{-N})$ .

### 4.7.2 The Runge Function Revisited

For the Runge function with poles at  $z = \pm 0.2i$ , we can compute:

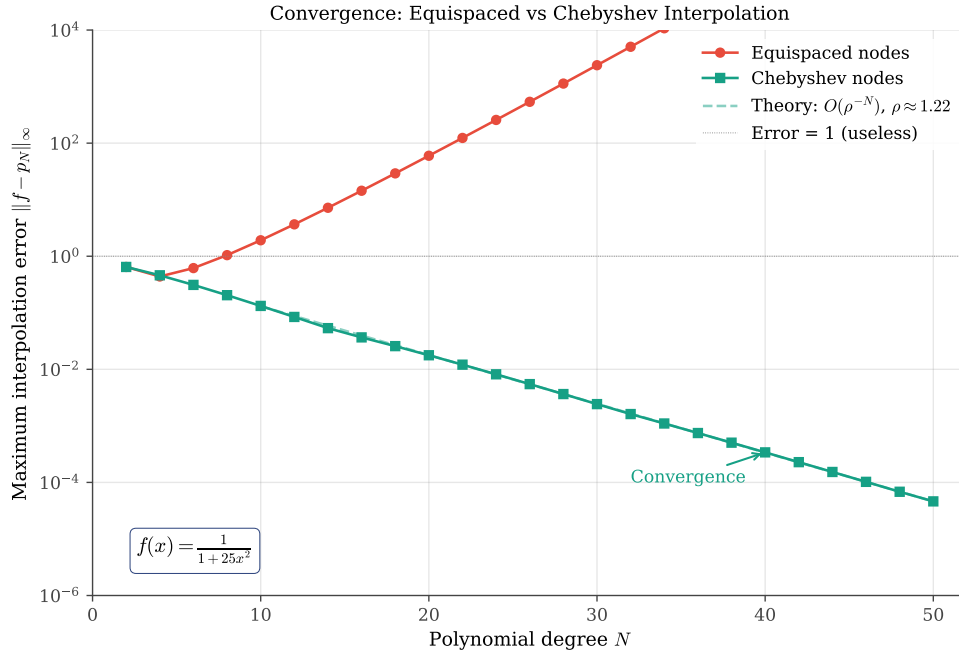
$$\rho = |0.2i + \sqrt{(0.2i)^2 - 1}| = |0.2i + i\sqrt{1.04}| \approx 1.22.$$

Thus Chebyshev interpolation converges at rate  $O(1.22^{-N})$ ; convergence is geometric but modest due to the poles being close to the real axis. In contrast, equispaced interpolation diverges because its effective  $\rho < 1$ .

The divergence of equispaced interpolation can be understood through the error bound  $\|f - p_N\|_\infty \leq (1 + \Lambda_N)E_N(f)$ . For the Runge function, the best approximation error  $E_N(f)$  still decreases geometrically—the function is analytic, after all. However, recall from Section 4.5 that the equispaced Lebesgue constant grows as  $\Lambda_N^{\text{eq}} \approx 2^{N+1}/(eN \ln N)$ . This exponential growth eventually overwhelms the geometric decay of  $E_N(f)$ , causing the interpolation error  $(1 + \Lambda_N)E_N(f)$  to grow without bound.

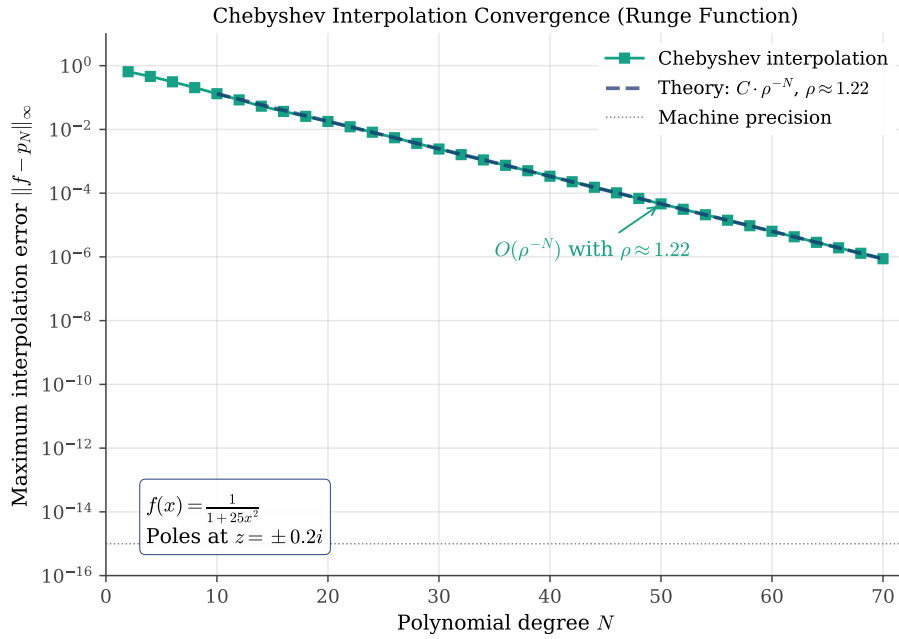
### 4.7.3 Numerical Verification

Figure 17 compares the convergence behavior for both node distributions. The Chebyshev error decreases geometrically, while the equispaced error grows without bound.



**Figure 17:** Convergence comparison for the Runge function. The maximum interpolation error is plotted against polynomial degree  $N$  on a semilogarithmic scale. Chebyshev interpolation (green) converges at rate  $O(\rho^{-N})$  with  $\rho \approx 1.22$ . Equispaced interpolation (red) diverges exponentially.

Figure 18 shows the Chebyshev convergence alone, without the divergent equispaced curve that dominates the vertical scale. This reveals the beautiful geometric convergence: the error decreases by a factor of approximately  $\rho \approx 1.22$  with each increase in polynomial degree. The theoretical rate matches the computed errors almost perfectly.



**Figure 18:** Chebyshev interpolation convergence for the Runge function (equispaced omitted for clarity). The error decreases geometrically at rate  $O(\rho^{-N})$  with  $\rho \approx 1.22$ , matching the theoretical prediction based on the pole locations at  $z = \pm 0.2i$ .

The code generating these figures is available in:

- `codes/python/ch04_geometry_of_nodes/convergence_comparison.py`
- `codes/python/ch04_geometry_of_nodes/convergence_zoom.py`
- `codes/matlab/ch04_geometry_of_nodes/convergence_comparison.m`

## 4.8 Computational Experiment: Random Nodes

Having studied the optimal Chebyshev distribution and the problematic equispaced distribution, a natural question arises: what happens if we choose interpolation nodes *randomly*? This question leads us into the realm of *experimental mathematics*, where we use computation to discover and conjecture mathematical relationships.

### 4.8.1 Motivation

The dramatically different behaviors of equispaced and Chebyshev nodes might suggest that the key factor is *regularity* or *structure* in node placement. Perhaps any “reasonable” arrangement would work? To test this hypothesis, we investigate the simplest unstructured choice: nodes drawn uniformly at random from  $[-1, 1]$ .

This experiment serves two purposes. First, it tests whether the special clustering of Chebyshev points near the endpoints is truly essential, or whether it is merely one of many acceptable distributions. Second, it demonstrates the methodology of computational experimentation—using numerical evidence to formulate conjectures about asymptotic behavior.

### 4.8.2 Experimental Setup


For each polynomial degree  $N$  from 2 to 30, we generate  $N + 1$  random points uniformly distributed on  $[-1, 1]$ , sort them, and compute the Lebesgue constant. We repeat this process  $M = 200$  times to obtain statistical estimates of the mean, standard deviation, and range of  $\Lambda_N$  for random nodes.

The Python implementation uses NumPy’s random number generation:

```

1  def random_nodes(N, rng=None):
2      """Generate N+1 random nodes, sorted, on [-1, 1]."""
3      if rng is None:
4          rng = np.random.default_rng()
5      return np.sort(rng.uniform(-1, 1, N + 1))
6
7  def monte_carlo_lebesgue(N, M=200):
8      """Compute M samples of Lebesgue constant for random nodes."""
9      samples = np.zeros(M)
10     for m in range(M):
11         x_rand = random_nodes(N)
12         samples[m] = lebesgue_constant(x_rand)
13     return samples

```

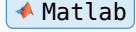
 Python

The equivalent MATLAB code:

```

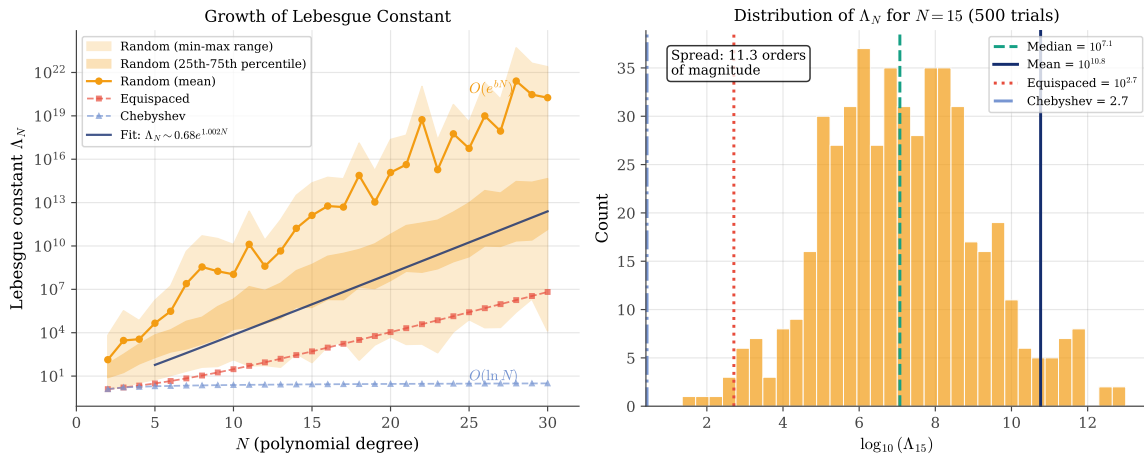
1 % Generate M samples for polynomial degree N
2 samples = zeros(M, 1);
3 for m = 1:M
4     x_rand = sort(2 * rand(N+1, 1) - 1); % Uniform on [-1, 1]
5     samples(m) = max(lebesgue_function(x_rand, x_fine));
6 end

```



### 4.8.3 Results

Figure 19 shows the results of our Monte Carlo experiment. The left panel displays the growth of the Lebesgue constant with polynomial degree, including shaded regions showing the statistical variability. The right panel shows the distribution of  $\Lambda_N$  values for a fixed degree  $N = 15$ .



**Figure 19:** *Lebesgue constant for random nodes. Left: growth with polynomial degree  $N$ , showing mean (solid line) and min-max range (light shading). Right: distribution of  $\log_{10}(\Lambda_{15})$  over 500 random trials on a logarithmic scale, revealing the enormous spread spanning many orders of magnitude. Vertical lines mark the mean, median, equispaced, and Chebyshev values.*

The key observations are striking:

1. **Explosive exponential growth:** Random nodes exhibit exponential growth of the Lebesgue constant, but surprisingly *faster* than equispaced nodes. The mean  $\Lambda_N$  grows roughly as  $e^{bN}$  with  $b \approx 1.0$ , compared to  $b \approx 0.6$  for equispaced.
2. **Extreme variability:** There is enormous variability between different random samples. For  $N = 15$ ,  $\Lambda_N$  can range from around  $10^2$  to  $10^{10}$  or more, spanning many orders of magnitude. This variability increases dramatically with  $N$ .
3. **Worse than equispaced:** Counter to naive intuition, random nodes perform *worse* on average than the structured equispaced distribution. The reason is that random sampling occasionally produces clusters of nearby nodes, creating near-singular interpolation conditions. These worst-case realizations dominate the mean.

### 4.8.4 Empirical Asymptotic Formula

Fitting an exponential model to the mean Lebesgue constants yields an empirical formula:

$$\Lambda_N^{\text{rand}} \approx a \cdot e^{bN},$$

where the fitted constants are approximately  $a \approx 0.7$  and  $b \approx 1.0$ , giving:

$$\Lambda_N^{\text{rand}} \approx 0.7 \cdot 2.7^N.$$

Remarkably, the growth rate for random nodes is *faster* than for equispaced nodes (which grow as  $\approx 1.8^N$ ). This counterintuitive result arises because random sampling occasionally produces clusters of nodes that are extremely close together, causing the Lebesgue constant to explode. The mean is dominated by these worst-case realizations.

The fundamental problem is twofold: (1) the lack of *clustering near the endpoints*, which only carefully designed distributions like Chebyshev points possess, and (2) the risk of *random clustering* in the interior, which creates severely ill-conditioned interpolation problems.

#### 4.8.5 Discussion

This computational experiment provides strong numerical evidence for a fundamental principle: **the clustering of Chebyshev points near the endpoints is not merely convenient but essential**. Random nodes, despite their apparent “fairness” in covering the interval, fail in two ways: they do not provide the boundary resolution needed to control Lagrange basis oscillations, and they risk interior clustering that creates catastrophically ill-conditioned systems.

The enormous variability in  $\Lambda_N$  for random nodes is perhaps the most striking finding. While equispaced nodes are suboptimal, they at least provide *predictable* (if exponentially growing) behavior. Random nodes introduce an additional layer of uncertainty—any given random realization might be acceptable or catastrophic.

The experiment illustrates the power of computational mathematics. By systematic numerical investigation, we have:

- Tested a natural hypothesis (random nodes might work)
- Discovered unexpected behavior (random is *worse* than equispaced)
- Quantified the asymptotic growth through data fitting
- Reinforced our understanding of why structured distributions matter

The code generating Figure 19 is available in:

- `codes/python/ch04_geometry_of_nodes/lebesgue_random_nodes.py`
- `codes/matlab/ch04_geometry_of_nodes/lebesgue_random_nodes.m`

## 4.9 Practical Guidelines and Outlook

### 4.9.1 When to Use Which Grid

Based on the theory developed in this chapter, we can state clear guidelines:

1. **Always prefer Chebyshev points** for polynomial interpolation on a finite interval. The  $O(\ln N)$  growth of the Lebesgue constant ensures stability.
2. **Avoid equispaced nodes** for high-degree polynomial interpolation. The exponential growth of  $\Lambda_N$  makes this a losing proposition.
3. **Consider the singularity structure** of the function being approximated. The location of complex singularities determines the convergence rate.

### 4.9.2 Mapping to General Intervals

Chebyshev points are defined on  $[-1, 1]$ , but can be mapped to any interval  $[a, b]$  via the linear transformation:

$$x_{\text{phys}} = \frac{a+b}{2} + \frac{b-a}{2}x_{\text{ref}},$$

where  $x_{\text{ref}} \in [-1, 1]$  is the reference coordinate.

### 4.9.3 Preview of Differentiation Matrices

The Chebyshev points introduced in this chapter will play a central role in the differentiation matrices we develop in subsequent chapters. The clustering of nodes near the boundaries, far from being a peculiarity, is precisely what enables accurate spectral differentiation. The connection between node distribution, interpolation accuracy, and differentiation stability is one of the beautiful unifying themes of spectral methods.

In the next chapter, we will see how to convert function values at Chebyshev points into accurate approximations of derivatives, building on the geometric insights developed here.





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