

Last Time:

- Quaternions
- Kinetic Energy
- Inertia
- Euler's Equation

Today:

- Stability of spinning bodies
- Numerical simulation of 3D Rotations
- Newton-Euler Dynamics and $SU(3)$

Stability of Spinning Bodies:

* Equilibria

- Assume $\|\boldsymbol{\omega}\| = \text{const}$, rigid body has 6 "equilibrium spins":
 - Equilibrium spin $\Rightarrow \dot{\boldsymbol{\omega}} = \boldsymbol{\tau}_{\boldsymbol{\omega}} = \mathbf{0}$
- $$\underbrace{\boldsymbol{\tau}_{\boldsymbol{\omega}}}_{\mathbf{B}\dot{\boldsymbol{\omega}}} = -\boldsymbol{\omega} \times \mathbf{J}\boldsymbol{\omega} = \mathbf{0} \Rightarrow \boldsymbol{\omega} \times \mathbf{J}\boldsymbol{\omega} = \mathbf{0}$$
- Cross product $= \mathbf{0}$ if $\dot{\boldsymbol{\omega}}$ is parallel to $\mathbf{J}\boldsymbol{\omega}$
 - $\Rightarrow \boldsymbol{\tau}_{\boldsymbol{\omega}} = \lambda \boldsymbol{\omega}, \lambda \in \mathbb{R}$
 - \Rightarrow equilibria are \pm eigenvectors of \mathbf{J}

* Principle Axes

- Since kinetic energy ≥ 0 , J must be positive definite:
 $\frac{1}{2} \omega^T J \omega > 0 \text{ if } \omega \neq 0 \Rightarrow J > 0$
- Positive-definite matrices have positive real eigenvalues
 \Rightarrow we can always diagonalize J
- Eigenbasis of J called "principle axes" of rigid body. Typically, correspond to symmetry axes of body.
- In principle axes, Euler's equation becomes:
$${}^P J_{11} {}^P \dot{\omega}_1 + ({}^P J_{33} - {}^P J_{22}) {}^P \omega_2 {}^P \omega_3 = {}^P \tau_1$$
$$J_{22} \dot{\omega}_2 + (J_{11} - J_{33}) \omega_1 \omega_3 = \tau_2$$
$$J_{33} \dot{\omega}_3 + (J_{22} - J_{11}) \omega_1 \omega_2 = \tau_3$$
- Now equilibria lie along \vec{p}_i basis vectors

* Stability

- Linearize about each axis and check local stability.
- Assume $J_{11} < J_{22} < J_{33}$

$$\star \omega_1 = \omega_0 \Rightarrow \omega_2, \omega_3$$

$$\dot{\omega}_2 = \omega_0 \left(\frac{J_{22} - J_{11}}{J_{22}} \right) \omega_3$$

$\underbrace{\phantom{\dot{\omega}_2 = \omega_0 \left(\frac{J_{22} - J_{11}}{J_{22}} \right) \omega_3}}_{\alpha_1 > 0}$

$$\dot{\omega}_3 = \omega_0 \left(\frac{J_{11} - J_{22}}{J_{22}} \right) \omega_2$$

$\underbrace{\phantom{\dot{\omega}_3 = \omega_0 \left(\frac{J_{11} - J_{22}}{J_{22}} \right) \omega_2}}_{\alpha_2 < 0}$

$$\Rightarrow \frac{d}{dt} \begin{bmatrix} \omega_2 \\ \omega_3 \end{bmatrix} = \begin{bmatrix} 0 & \alpha_1 \\ \alpha_2 & 0 \end{bmatrix} \begin{bmatrix} \omega_2 \\ \omega_3 \end{bmatrix} \quad \left. \begin{array}{l} \text{Looks like SHO} \\ \text{---} \end{array} \right.$$

$\underbrace{\dot{x}}_{x} = \underbrace{A}_{A} \underbrace{x}_{x}$

$$\text{eig}(A) = \pm \sqrt{\alpha_1 \alpha_2} \Rightarrow \text{marginally stable}$$

$\underbrace{\phantom{\text{eig}(A) = \pm \sqrt{\alpha_1 \alpha_2}}_{\text{pure imaginary}}}$

$$\star \omega_2 = \omega_0 \Rightarrow \omega_1, \omega_3:$$

$$\dot{\omega}_1 = \omega_0 \left(\frac{J_{22} - J_{33}}{J_{11}} \right) \omega_3, \quad \dot{\omega}_3 = \omega_0 \left(\frac{J_{11} - J_{22}}{J_{33}} \right) \omega_1$$

$\underbrace{\phantom{\dot{\omega}_1 = \omega_0 \left(\frac{J_{22} - J_{33}}{J_{11}} \right) \omega_3}}_{\alpha_1 < 0}$ $\underbrace{\phantom{\dot{\omega}_3 = \omega_0 \left(\frac{J_{11} - J_{22}}{J_{33}} \right) \omega_1}}_{\alpha_2 < 0}$

$$\Rightarrow \text{eig}(A) = \pm \sqrt{\alpha_1 \alpha_2} \Rightarrow \text{unstable!}$$

$\underbrace{\phantom{\text{eig}(A) = \pm \sqrt{\alpha_1 \alpha_2}}_{\text{pure real}}}$

$$\omega_3 = \omega_0 \Rightarrow \omega_1, \omega_2$$

$$\dot{\omega}_1 = \omega_0 \left(\frac{J_{22} - J_{33}}{J_{11}} \right) \omega_3, \quad \dot{\omega}_2 = \omega_0 \left(\frac{J_{11} - J_{22}}{J_{33}} \right) \omega_3$$

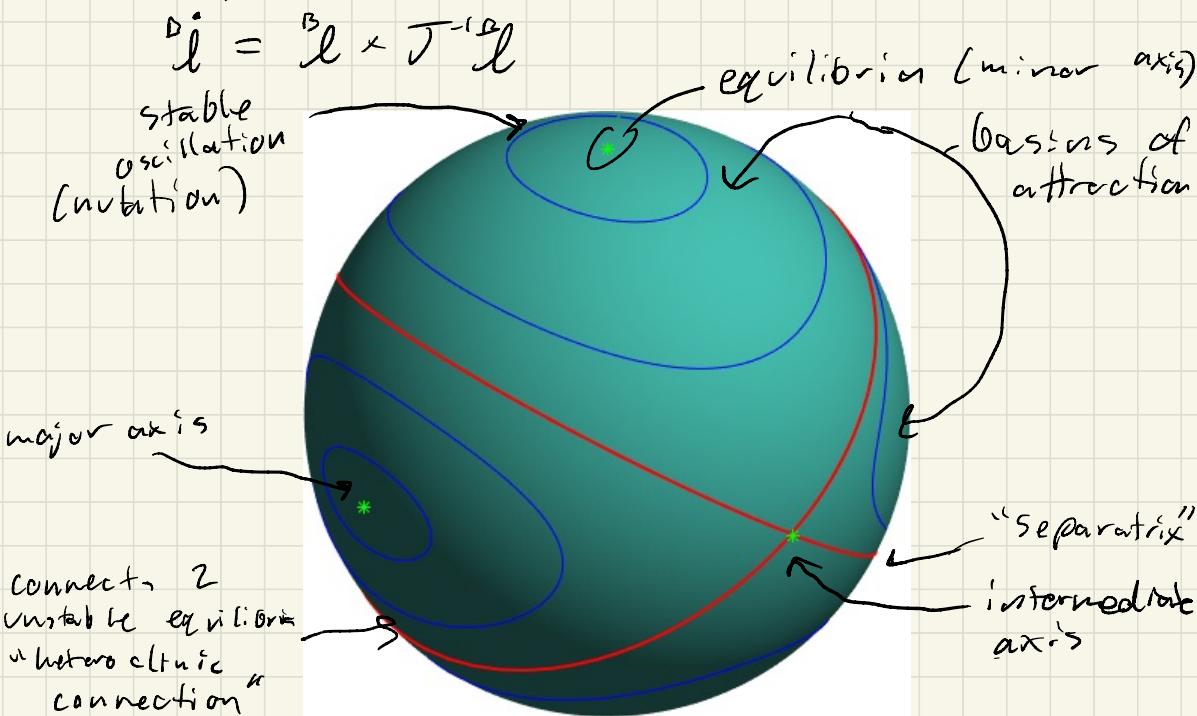
$\underbrace{\phantom{\omega_0 \left(\frac{J_{22} - J_{33}}{J_{11}} \right) \omega_3}}$ $\underbrace{\phantom{\omega_0 \left(\frac{J_{11} - J_{22}}{J_{33}} \right) \omega_3}}$

$\alpha_1 < 0$ $\alpha_2 > 0$

$$\Rightarrow \text{eig}(A) = \pm \sqrt{\alpha_1 \alpha_2} \Rightarrow \text{marginally stable}$$

pure imaginary

- Spin along maximum/minimum axes (a.k.a major/minor axes) of inertia is Lyapunov stable.
- Spin about intermediate axis is always unstable.
- Geometry on Momentum Sphere:



- Oscillatory motion about stable spin axis is called "nutation"
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Numerical Simulation With 3D Rotations:

- Naive use of RK method on Q or q leads to numerical drift off of $SO(3)$ manifold (Q no longer orthogonal, q no longer unit norm)
- Standard Hacks:
 - * Re-Normalize
 - Easy for q , expensive for Q (sud)
 - Only possible if you implement your own RK method
 - * Baumgarte Stabilization
 - Apply proportional or PD control to constraint in the dynamics
 - Easy for q , harder for Q
 - Makes norm Lyapunov stable, but still oscillates about true value
 - Higher gain makes error small but makes ODE stiff.
 - Can apply with standard toolboxes like Differential Equations.jl.

- Better Options:

- * Implicit Method with Explicit Constraint
 - Can combine implicit RK condition with constraint equation and simultaneously solve both with Newton.
 - Can be expensive
- * Runge-Kutta Munthe-Kaas Methods
 - Compute RK step on Lie algebra (axis-angle vectors) and use exp and multiplication to compute Q_{n+1}/q_{n+1}
 - Since group is closed under multiplication
 \Rightarrow guaranteed to stay on group
 - Straight-forward to construct RK-MK method from any standard RK method
 - Can be explicit or implicit

+ Explicit Lie-Euler Method:

- Standard Euler method

$$w_{n+1} = w_n + h \dot{w}_n = h J^{-1} (I_n - \omega_n \times J w_n)$$

$$Q_{n+1} = Q_n + h \dot{Q}_n = Q_n + h Q_n \hat{\omega}_n = Q_n \underbrace{(I + h \hat{\omega}_n)}$$

1st-order Taylor expansion of $\exp(h \hat{\omega}_n)$

- Key idea: replace 1st order Taylor expansion with exact exp function!

$$Q_{n+1} = Q_n \exp(h \hat{w}_n)$$

or

$$q_{n+1} = q_n * \exp\left(h \begin{bmatrix} 0 \\ \hat{w}_n \end{bmatrix}\right) = L(q_n) \exp\left(\frac{h}{2} \begin{bmatrix} 0 \\ \hat{w}_n \end{bmatrix}\right)$$

- Guaranteed to stay on Lie group
- Derivation gets messier for order ≥ 2 , but same idea.