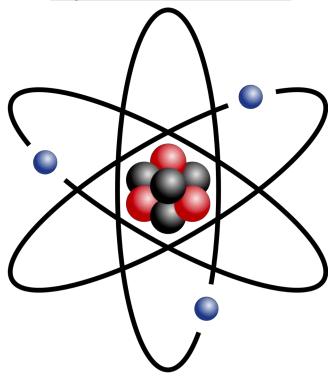
PA1140: Waves and Quanta

# **Unit 4: Atoms and Nuclei**

Tipler, Chapters 36 (36-1 to 36-2) & 40





Dr Eloise Marais (Michael Atiyah Annex, 101)

### Course Rep Nominations for 2019/2020 are Now Open

Course reps are important! They have direct input to the teaching strategy in the department and they influence the way the course is run

If students would like to contribute please nominate yourself before 15 March



### **NOMINATIONS**

Open until Friday 15th March

## **ELECTIONS**

Tuesday 19th- Thursday 21st March

FIND OUT IF YOU'VE WON

Friday 22<sup>nd</sup> March

www.leicesterunion.com/represent/course-reps

## **Lecture 1 Follow Up**

### Why is angular momentum quantized?

Bohr postulated that angular momentum of an electron orbiting a nucleus is quantized for his model to work.

This was later explained by de Broglie

According to de Broglie, an electron moving in a circular orbit behaves like a particle wave

### Reminder of the expression for the de Broglie wavelength:

$$\lambda = \frac{h}{p}$$
 where  $p$  is the momentum of the electron

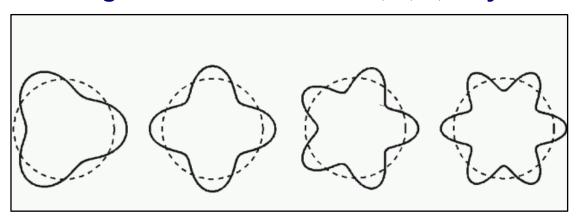


## **Lecture 1 Follow Up (contd)**

For any electron moving in circular orbit n of radius  $r_n$ , the total distance is equal to the circumference of an orbit,  $2\pi r_n$ 

#### This can be illustrated as follows:

Standing wave on a circle for 3, 4, 5, 6 cycles.



**Conceptually**, only a discrete set of rates are possible, since the original wave function has to return to its starting point after each rotation



## **Lecture 1 Follow Up (contd)**

**Mathematically**, this is expressed following these steps:

Relate circumference to wavelength:

$$2\pi r_n = n\lambda$$
 Equation (1)

Define de Broglie wavelength ( $\lambda = \frac{h}{p}$ ) in terms of angular momentum:

$$\lambda = \frac{h}{m\nu_n}$$
 Equation (2)

where  $m\nu_n$  is the momentum of an electron revolving in orbit n

Substitute equation (1) into equation (2) to get an expression for quantized angular momentum

$$m\nu_n r_n = n \frac{h}{2\pi}$$



### Reminder of Bohr's 3<sup>rd</sup> Postulate:

Only orbits are allowed if angular momentum of the electron is an integral multiple of  $h/2\pi$  (quantized angular momentum).

$$h = \frac{h}{2\pi} = 1.055 \times 10^{-34} Js = 6.582 \times 10^{-16} \text{ eV s}$$

where  $\hbar$  is the reduced Planck's constant

The magnitude of a circular orbit is L = mvr, so the 3<sup>rd</sup> postulate is:

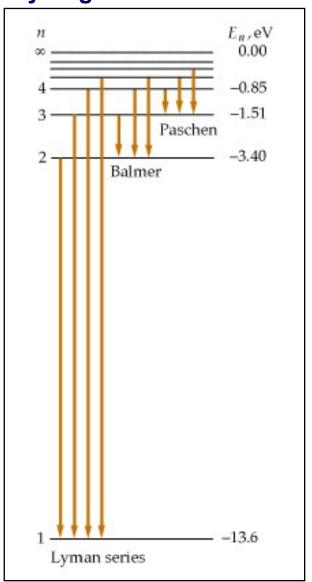
$$m\nu_n r_n = n\hbar$$

where n = 1, 2, 3, ... (the quantum number of the state)



# **Practice Problem: Atomic Spectra**

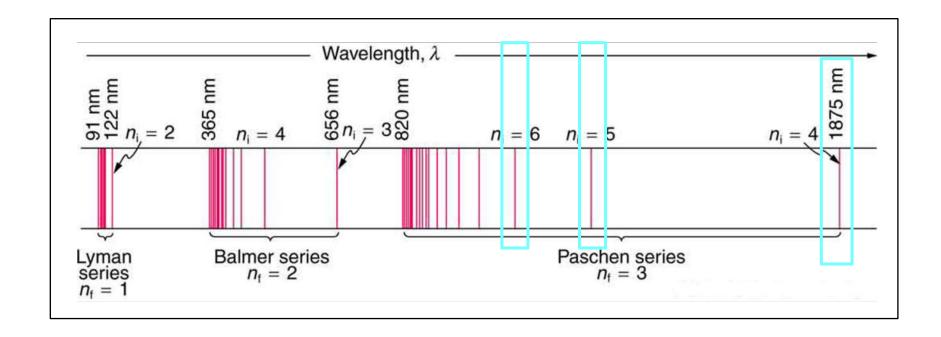
### **Hydrogen emission lines**



- a) Find the photon energy and wavelength for the series limit (shortest wavelength) in the Paschen series ( $n_2 = 3$ ).
- b) Calculate the wavelengths for the three longest wavelengths in this series and indicate their positions on a horizontal linear scale.



## Let's check our answer:





# **Practice Problem: Ionization Energy**

Calculate the ionization energy for:

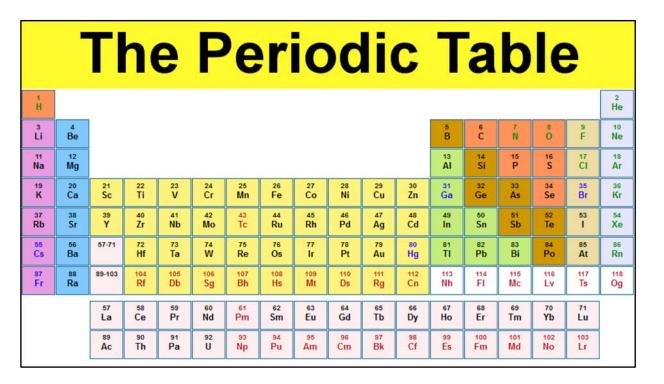
- a) the hydrogen atom
- b) He+
- c) Li<sup>2+</sup>



# **Practice Problem: ionization energy**

Calculate the ionization energy for:

- a) the hydrogen atom
- b) He+
- c) Li<sup>2+</sup>





# Lecture 2: Properties of Nuclei & Radioactivity

## Relevant Terminology

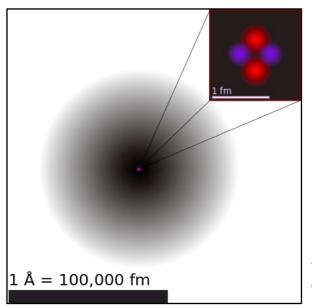
- Nucleus: incudes protons and neutrons (masses are about equal)
- Number of protons: Z or atomic number
- Number of neutrons: N
- N = Z for lighter nuclei
  N > Z for heavier nuclei (difference increases as mass increases)
- Number of nucleons: A or nucleon or mass number (A = N + Z)
- Nuclide: name given to a nuclear species
- Atomic symbol: unique name for an atom (H for hydrogen)
- Isotope: 2 or more nuclides with the same Z, but different N (and A)



# **Strong Nuclear Force**

- One of four fundamental forces
- Nucleons exert strong attractive forces on other nucleons
- Stronger than electrostatic force of repulsion between protons
- Much stronger than gravitational forces (negligible)

#### **Nucleus of He**



- About the same between 2 neutrons, 2 protons, a neutron and proton
- Decrease with distance to zero at a distance of a few fentometer (1 fm = 10<sup>-15</sup> m)

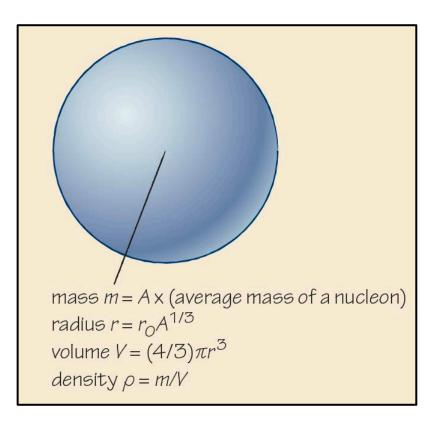
2 protons stay together due to nuclear force



# **Nuclear Size, Shape, and Density**

Assuming spherical shape, nuclear radii can be calculated as follows:

$$R = r_0 A^{1/3}$$



where A = N + Z is the atomic mass and  $r_0 \sim 1.2 \text{ fm}$  (empirical constant)

$$R \propto A^{1/3}$$

So:  $V \propto A$ 

Nuclear mass:

$$m = A \times m_{nucleus}$$

Because  $m_{nucleus} \propto A$ , densities of all nuclei are roughly the same



## **Practice Problem: Density of nuclear matter**

Let's prove that densities of nuclear matter is roughly the same.



## N and Z Numbers

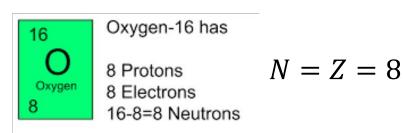
### Light nuclei:

• stable for  $N \approx Z$ 

#### Heavier nuclei:

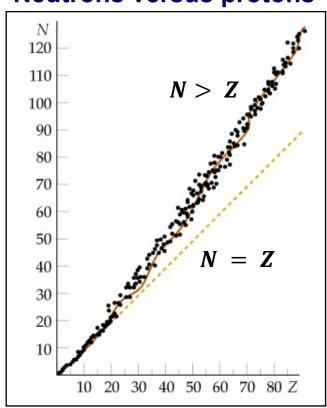
- instability caused by electrostatic repulsion between protons
- instability minimized by having more neutrons than protons

## **Example:**





### **Neutrons versus protons**





# Mass and Binding Energy

Mass of nucleus = sum of parts  $-E_b/c^2$ 

E<sub>b</sub>: binding energy (energy needed to break all bonds)

When 2 or more nucleons fuse together to form a nucleus: energy is released and mass decreases (opposite when break apart)

atomic mass units (a.m.u): mass convention for atoms and nuclei

1 a.m.u = one-twelfth the mass of carbon-12

Rest energy of 1 a.m.u (using  $E=mc^2$ ), and  $m_{a.m.u}$  = 1.66054 x 10<sup>-27</sup> kg:

$$(1u)c^2 = 931.5 \text{ MeV}$$



# **Worked Example**

<sup>4</sup>He includes 2 protons and 2 neutrons. The mass of <sup>4</sup>He measured with a mass spectrometer is **4.002603 u**. The mass of a proton (<sup>1</sup>H) is 1.007825 u and of a neutron is 1.008665 u.

The mass of 2 protons and 2 neutrons = 4.03298 u

(which is greater than the measured mass of <sup>4</sup>He)

We can use this to estimate the binding energy,  $E_b$ :

$$E_b = (4.03298u \ c^2 - 4.002603u \ c^2) \frac{931.5 \text{ MeV}}{1 \ u \ c^2}$$

$$E_b = 28.3 \text{ MeV}$$



# **Binding Energy (contd)**

### General expression for $E_b$ :

$$E_b = (ZM_H + NM_N - M_A)c^2$$

where:  $M_H$  is the mass of a proton

 $M_N$  is the mass of a neutron  $M_A$  is the final atomic mass

Electrons not included. 1000 times lighter than proton:

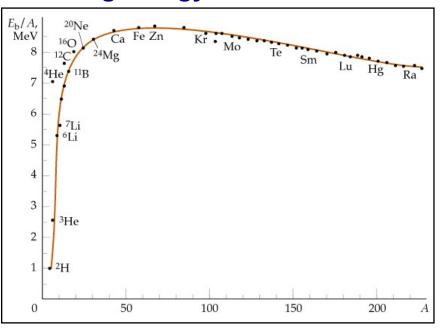
Mass of proton: of the order 10<sup>-27</sup> kg

Mass of electron: of the order 10<sup>-30</sup> kg



# **Binding Energy (contd)**

### Binding energy vs mass number



A < 20: Increase in nearest neighbours and in number of bonds per nucleon

**50 < A < 200:** Saturation of nuclear forces

**A > 200:** Coulomb repulsion so great that nucleus becomes unstable.

A > 300: Nucleus unstable and undergoes spontaneous fission.



# Radioactivity

# Relevant Terminology

- Radioactive nuclei: decay into other nuclei by emitting particles (photons, electrons, neutrons, particles)
- $\alpha$  decay: release of <sup>4</sup>He atom (i.e.  $\alpha$  particle)
- $\beta$  decay: release of electrons ( $\beta^-$ ) or positrons ( $\beta^+$ )
- positron: positively charged particle with the same mass as an electron
- γ decay: release of photons

