GPUnit

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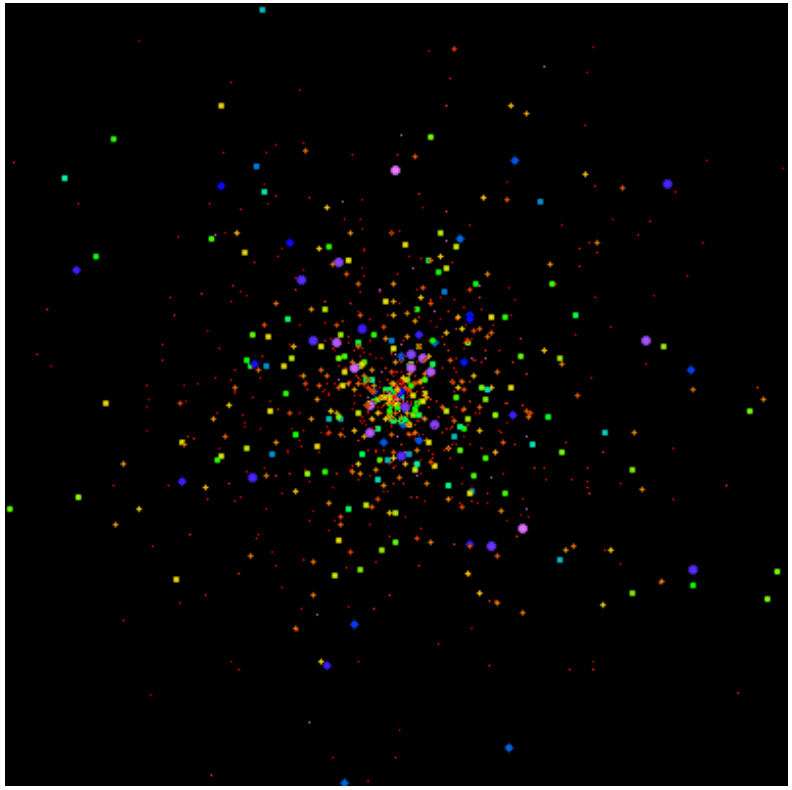
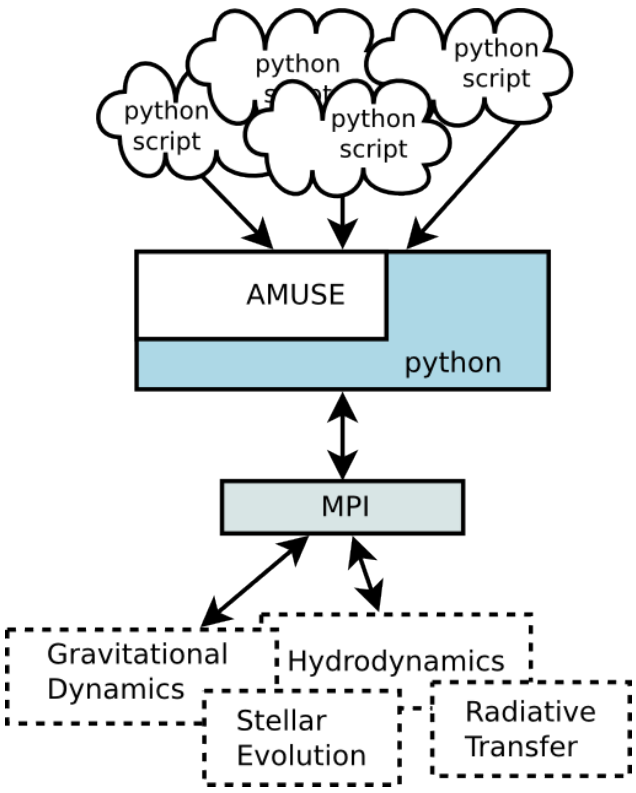
Motivation

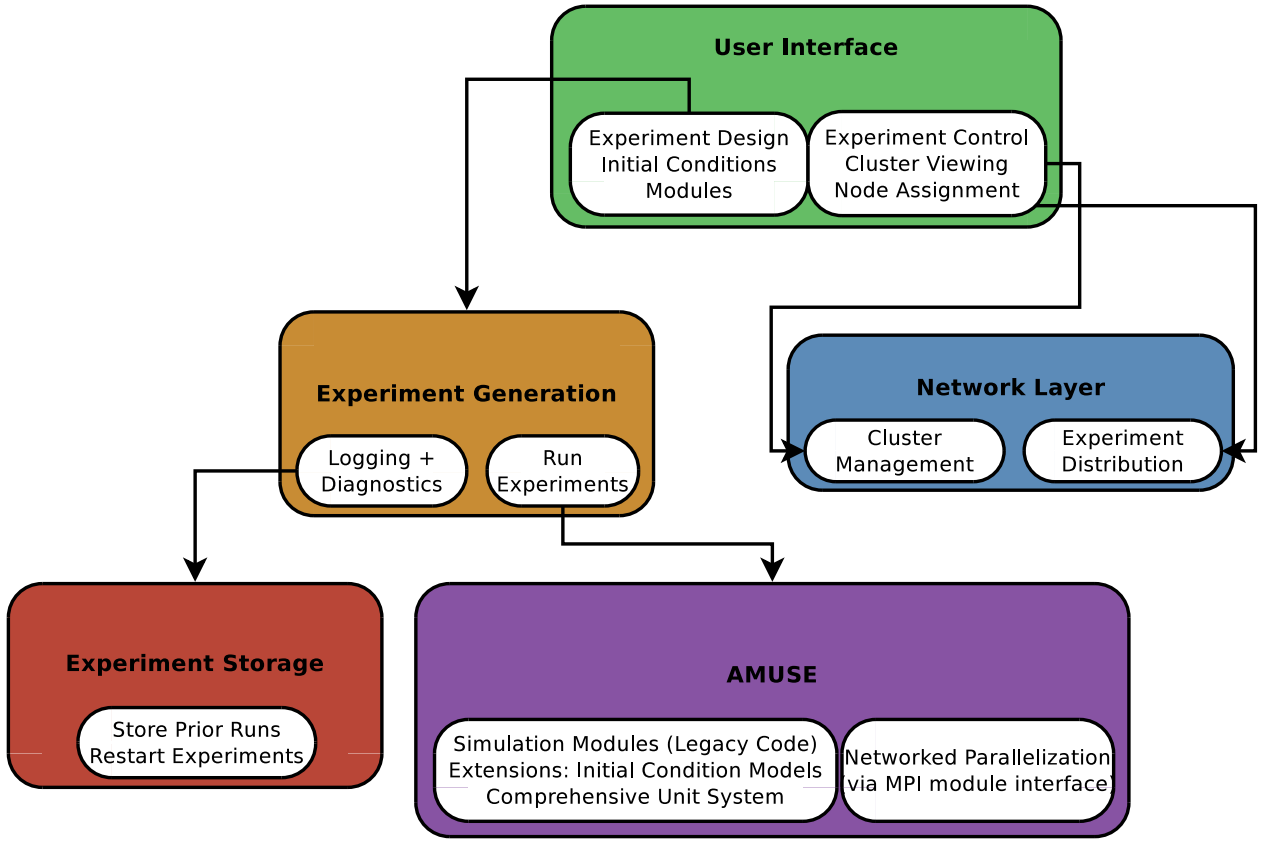
- Astrophysics researchers need to simulate movement and evolution of star clusters and galaxies.
- Every star pulls on all of the others.
  - $O\left(n^2\right)$
- Complex software exists to perform these computations efficiently.
- Our project exists to make running experiments with these tools feasible for a wider audience.



Overview

Astrophysical Multipurpose Software Environment (AMUSE)

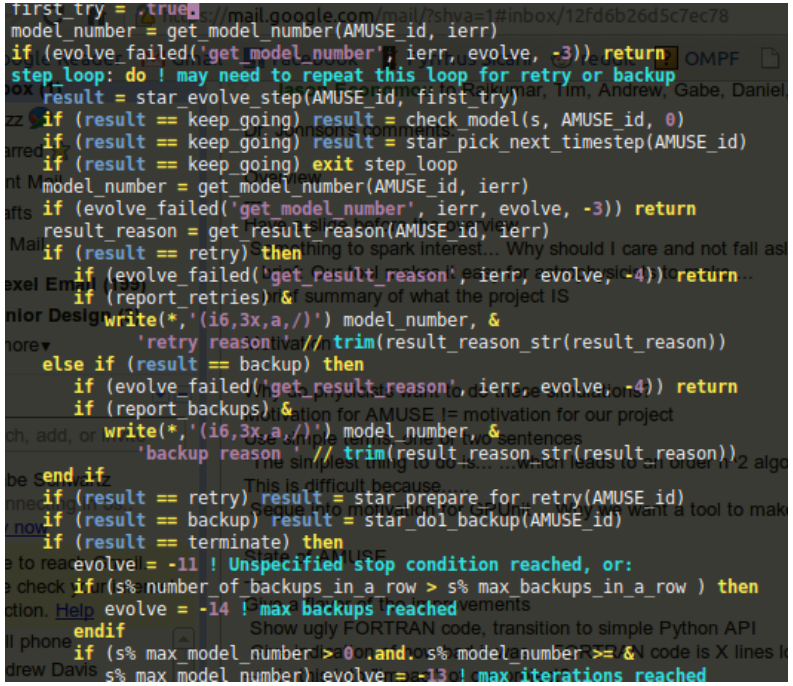




http://www.amusecode.org

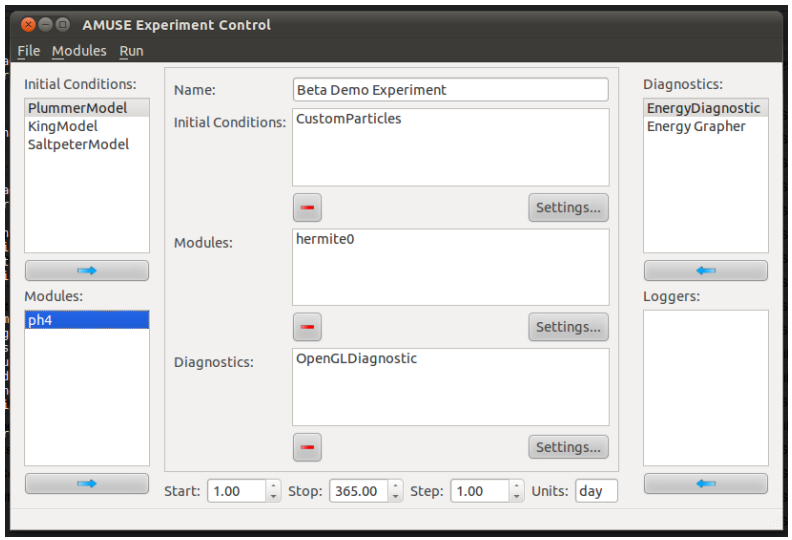
State of AMUSE

- Currently used by researchers to run large-scale simulations.
- Scripts, diagnostics, logging are all written by hand.
- AMUSE API/programming knowledge is required to create experiments.
- Still better than separated and opaque FORTRAN codes.



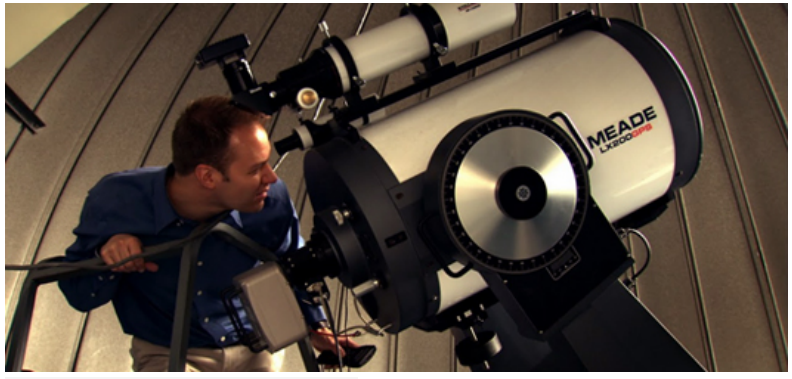
Purpose of GPUnit

- Ease the use of AMUSE
- Create/Design/Modify experiments
- Select, configure, swap out modules and initial conditions
- Store and restore progress of running experiments.



Target Audiences

- Physics Students
- Observational Astrophysicists
- Theoretical Astrophysicists



$$\ddot{X}+\frac{C}{M}\dot{X}+\omega_n^2\left[1-\frac{f(\Phi)}{2}(\Delta k_1+\Delta k_2\cos 2\Phi)\right]X$$
$$-\frac{\omega_n^2 f(\Phi)\Delta k_2\sin 2\Phi}{2}\left(Y_m-\frac{P}{K}\right)$$
$$=\varepsilon(\Omega+\dot{\phi})^2\cos(\Phi+\delta)+\varepsilon\ddot{\phi}\sin(\Phi+\delta),$$
$$\ddot{Y}_m+\frac{C}{M}\dot{Y}_m-\frac{\omega_n^2 f(\Phi)\Delta k_2\sin 2\Phi}{2}X$$
$$+\omega_n^2\left[1-\frac{f(\Phi)}{2}(\Delta k_1-\Delta k_2\cos 2\Phi)\right]Y_m$$
$$=\varepsilon(\Omega+\dot{\phi})^2\sin(\Phi+\delta)-\varepsilon\ddot{\phi}\cos(\Phi+\delta)$$
$$-\frac{Pf(\Phi)}{2M}(\Delta k_1-\Delta k_2\cos 2\Phi),$$
$$\ddot{\theta}+\frac{K_t+K_c}{I_0}\theta-\frac{K_t}{I_0}\varphi=-\frac{C_t+C_c}{I_0}\dot{\theta}+\frac{C_t}{I_0}\dot{\varphi},$$
$$\ddot{\varphi}+\frac{C_t}{I}\dot{\varphi}-\frac{C_t}{I}\dot{\theta}+\frac{K_t}{I}\varphi-\frac{K_t}{I}\theta$$
$$=\frac{P\varepsilon f(\Phi)}{2I}(\Delta k_1\cos(\Phi+\delta)-\Delta k_2\cos(\Phi-\delta))$$
$$+\frac{P^2}{2KI}\left[\frac{1}{2}\frac{\partial f(\Phi)}{\partial\Phi}(\Delta k_1-\Delta k_2\cos 2\Phi)\right.$$
$$\left.+f(\Phi)\Delta k_2\sin 2\Phi\right]+\Gamma_c,$$

Features

- Configurable experiments that can be saved and shared.
- Diagnostic tools that compute useful metrics.
- Storage of experiment state in case of crashes.
- Custom diagnostics and code generation.
- Provides a display of cluster usage to aid in scheduling.

Design

- Written in Python using the PyQt4 GUI toolkit.
- AMUSE is written in Python, streamlines interaction.
- C++ was considered as it supports Qt as well.
  - Communication w/AMUSE would be cumbersome.
  - AMUSE would be in a separate process.
- Designed APIs for diagnostics, logging and experiment persistence.
  - Users can create new diagnostics easily.
  - Experiments can be stored in a file structure, a remote DB etc...

Tests

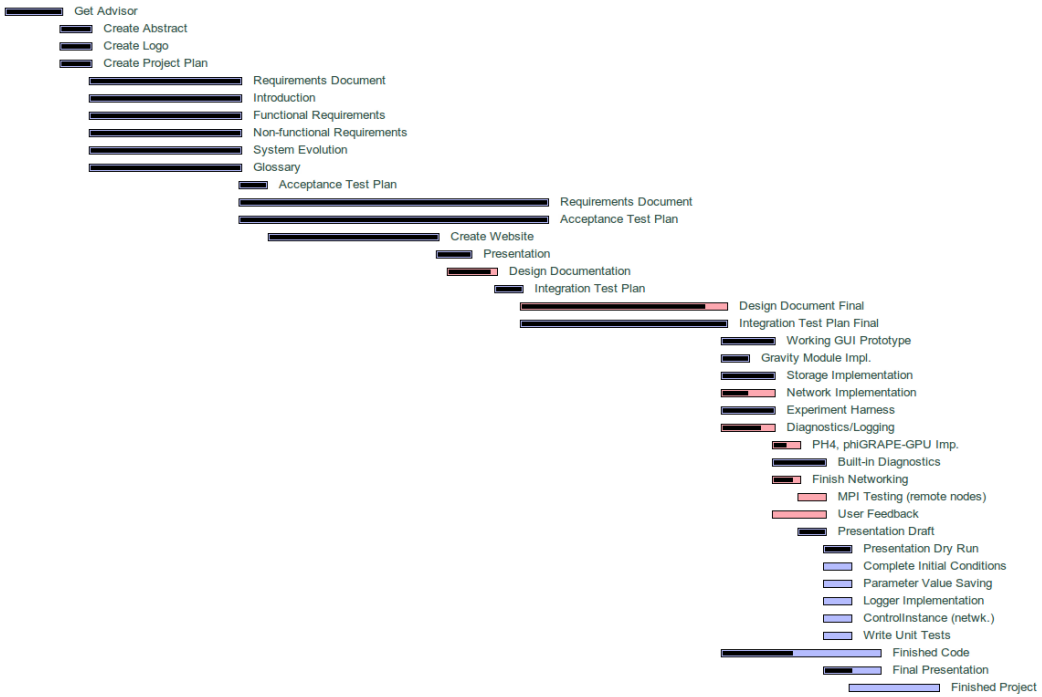
- Table of tests that pass.

User Testing

- Tested with customers (Steve/Tim)

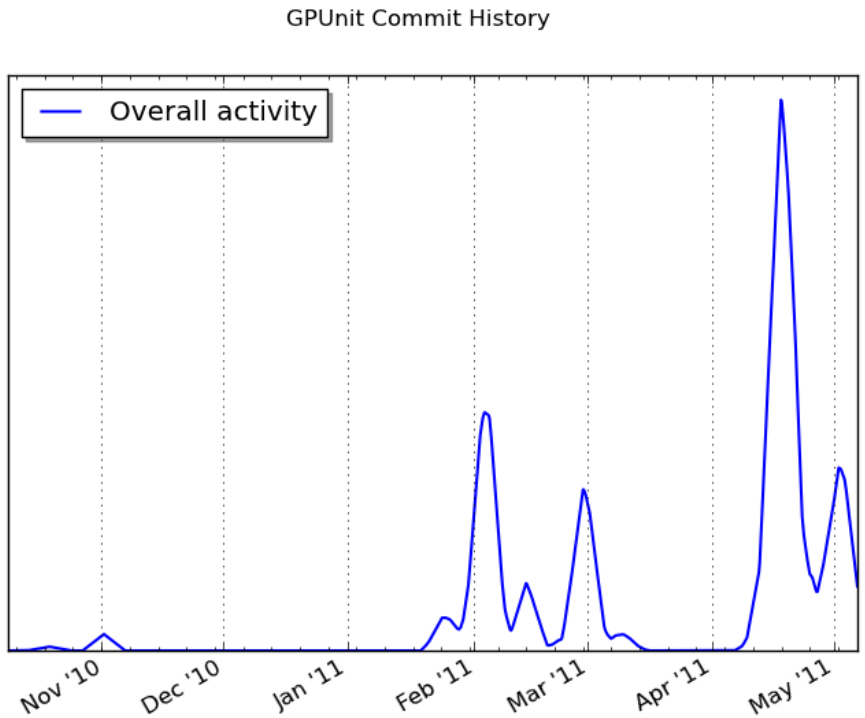
Project Plan

- Mostly waterfall design process.
- Initial phases were spent learning the domain (Physics/AMUSE).
- Roles
  - Tim: Physics reference, test subject
  - Andrew/Jason: Experiment and Module design.
  - Dan: Diagnostics
  - Raj: Logging
  - Gabe: Network, GUI.



Team Management

- Used Mercurial as our version control system.
  - Distributed, allows off-line commits.
- Team met weekly.
  - Once to plan work, once to code.
- Bi-weekly advisor meetings.



Project Impact

- Gives students and physicists easy access to state-of-the-art tools.
- Simple experiment creation $\rightarrow$ faster turnaround on experiments.
- Faster experiments $\rightarrow$ more time to study them.
- Current state:
  - Software is usable to create simple experiments.
  - Comes with useful diagnostics, from real experimental setups.
  - Needs refinement before it will be useful to professional astrophysicists.

Demo

- Demonstration of a simulation.

Questions

- Questions?