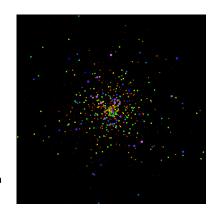
GPUnit

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Motivation

- Astrophysics researchers need to simulate movement and evolution of star clusters and galaxies.
- Every star pulls on all of the others.
 - $ightharpoonup O(n^2)$
- Complex software exists to perform these computations efficiently.
- Our project exists to make running experiments with these tools feasible for a wider audience.



Overview

Introduction

Purpose

Purpose of GPUnit Target Audiences

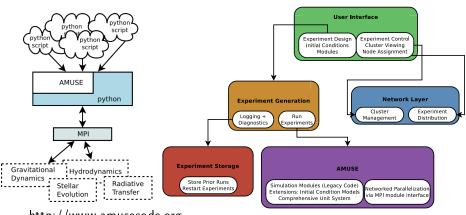
Features and Design

Software Engineering

Impact

Demo

Astrophysical Multipurpose Software Environment (AMUSE)



http://www.amusecode.org

State of AMUSE

- Currently used by researchers to run large-scale simulations.
- Scripts, diagnostics, logging are all written by hand.
- ► AMUSE API/programming knowledge is required to create experiments.
- Still better than separated and opaque FORTRAN codes.

```
del number = get model number(AMUSE id, ierr)
if (evolve failed('get model_number'; ierr, evolve, -3)) return
step loop: do ! may need to repeat this loop for retry or backup
      result = star evolve step(AMUSE id. first try)
     if (result == keep going) result = check model(s, AMUSE id, 0)
if (result == keep going) result = star pick next timestep(AMUSE id)
      if (result == keep going) exit step loop
      model number = get model number(AMUSE id, ierr)
      if (evolve failed('get model number', ierr, evolve, -3)) return
     result reason = get result reason(AMUSE id, ierr)
      if (result == retry) then
              if (evolve failed('get result reason', derry evolves(4)) return
              if (report retries) &
    r Designite(*,'(i6.3x,a,/)') model number, &
                                 retry reason(Va//ntrim(result reason str(result reason))
     else if (result == backup) then
              if (evolve failed(iget result reason), ierr, evolve, 4)) return
              if (report backups) &
                      write(*,'(i6,3x,a,/)') model_number, & number, set to the property of the
      if (result == retry) result = star prepare for retry(AMUSE id)
     if (result == backup) result = star do1 backup(AMUSE id)
     if (result == terminate) then
  evolve = -11 ! Unspecified stop condition reached, or:
              if (s% number of backups in a row > s% max backups in a row ) ther
                      evolve = -14 | max backups reached
```

Purpose of GPUnit

- ► Ease the use of AMUSE
- Create/Design/Modify experiments
- ► Select, configure, swap out modules and initial conditions
- ► Store and restore progress of running experiments.



Target Audiences

- ► Physics Students
- Observational Astrophysicists
- Theoretical Astrophysicists



$$\begin{split} \ddot{X} &= \frac{(K_0^2 + \omega_0)}{K_0^2} \left[-\frac{(K_0^2)(\omega_0^2 + \omega_0^2 \cos 2\theta^2)^2}{(\gamma_m - \frac{R}{K_0^2})} \right] \\ &= \frac{\omega_0^2 K(\theta) \omega_0^2 \cos (\theta - \theta)}{\omega_0^2 \cos (\theta - \theta)} - \sin (\theta - \theta), \\ &= \varepsilon(1 - \theta)^2 \cos (\theta - \theta) - \sin (\theta - \theta), \\ &\gamma_m - \frac{C_0^2}{K_0^2} \gamma_m \frac{(\theta^2)}{(2k_0^2 - \omega_0^2 \cos 2\theta^2)} \gamma_m \\ &= \varepsilon(1 - \theta)^2 \sin (\theta - \theta) - \cos (\phi - \theta), \\ &- \frac{R(\theta^2)}{2M_0^2} \frac{(\omega_0^2 - \omega_0^2 \cos 2\theta^2)}{(2k_0^2 - \omega_0^2 \cos 2\theta^2)} \\ \ddot{\theta} + \frac{K_0^2 + K_0^2 - K_0^2}{(2k_0^2 - \omega_0^2 \cos 2\theta^2)} - \frac{K_0^2}{k_0^2} \\ &= \frac{R(\theta^2)}{2M_0^2} \frac{(\omega_0^2 - \omega_0^2 \cos 2\theta^2)}{k_0^2} \\ &= \frac{R(\theta^2)}{2M_0^2} \frac{(\omega_0^2 - \omega_0^2 \cos 2\theta^2)}{k_0^2} \\ &= \frac{R(\theta^2)}{2M_0^2} \frac{(\omega_0^2 - \omega_0^2 \cos 2\theta^2)}{(2k_0^2 - \omega_0^2 \cos 2\theta^2)} \\ &= \frac{R(\theta^2)}{2M_0^2} \frac{(\omega_0^2 - \omega_0^2 \cos 2\theta^2)}{2M_0^2} \frac{(\omega_0^2 - \omega_0^2 \cos 2\theta^2)}{k_0^2} \\ &= \frac{R(\theta^2)}{2M_0^2} \frac{(\omega_0^2 - \omega_0^2 \cos 2\theta^2)}{(2k_0^2 - \omega_0^2 \cos 2\theta^2)} + \Gamma, \end{split}$$

Features

- Configurable experiments that can be saved and shared.
- Diagnostic tools that compute useful metrics.
- ► Storage of experiment state in case of crashes.
- Custom diagnostics and code generation.
- Provides a display of cluster usage to aid in scheduling.



Design

- ▶ Written in Python using the PyQt4 GUI toolkit.
- ▶ AMUSE is written in Python, streamlines interaction.
- ► C++ was considered as it supports Qt as well.
 - ► Communication w/AMUSE would be cumbersome.
 - ► AMUSE would be in a separate process.
- ▶ Designed APIs for diagnostics, logging and experiment persistence.
 - Users can create new diagnostics easily.
 - Experiments can be stored in a file structure, a remote DB etc...





Tests

► Table of tests that pass.

User Testing

► Tested with customers (Steve/Tim)

Project Plan

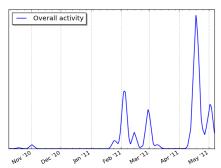
- Mostly waterfall design process.
- Initial phases were spent learning the domain (Physics/AMUSE).
- ► Roles
 - Tim: Physics reference, test subject
 - Andrew/Jason: Experiment and Module design.
 - ► Dan: Diagnostics
 - ► Raj: Logging
 - ► Gabe: Network, GUI.



Team Management

- Used Mercurial as our version control system.
 - Distributed, allows off-line commits.
- ► Team met weekly.
 - ▶ Once to plan work, once to code.
- ▶ Bi-weekly advisor meetings.

GPUnit Commit History



Project Impact

- ► Gives students and physicists easy access to state-of-the-art tools.
- ightharpoonup Simple experiment creation ightharpoonup faster turnaround on experiments.
- ightharpoonup Faster experiments ightarrow more time to study them.
- Current state:
 - ► Software is usable to create simple experiments.
 - Comes with useful diagnostics, from real experimental setups.
 - ► Needs refinement before it will be useful to professional astrophysicists.

Demo

▶ Demonstration of a simulation.

Questions

▶ Questions?