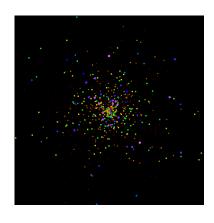


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Advisor: Prof. Jeremy Johnson Stakeholders: Prof. Steve McMillan Alfred Whitehead The Leiden Observatory

Motivation

- ► Why simulate astrophysics?
- Why make tools to help run simulations?
- Simulations are large and complex.
 - ► 1K 1M particles
 - Many steps.
- Complex software (AMUSE) exists to perform these computations efficiently.
 - ► Hardware-specific algorithms
 - ► More/less accurate algorithms



Target Audiences

- ► Physics Students
- Observational Astrophysicists
- ▶ Theoretical Astrophysicists



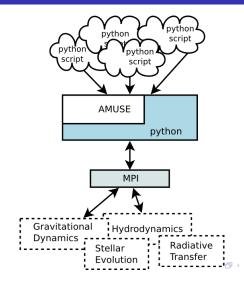
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\begin{split} \ddot{X} + \frac{c}{2N} X &= \frac{c}{2R} \left[ 1 - \frac{d(S)}{2} \left( \Delta b_1 + \Delta b_2 \cos 2\Phi \right) \right] V \\ &= \frac{c}{2R} \left( \theta \left( \Delta b_2 \cos 2\Phi \right) - \frac{c}{2R} \left( \theta \left( -\Phi \right) - \frac{c}{2R} \right) \right) \\ &= \left( 2 \left( 1 - \Phi \right)^2 \cos \left( \Phi - \Phi \right) - \cos \left( \Phi - \Phi \right) - \frac{c}{2R} \left( \Phi - \Phi \right) \right) \\ &= \left( 2 \left( 1 - \Phi \right)^2 \cos \left( \Phi - \Phi \right) - \frac{c}{2R} \left( 1 - \Phi \right) \right) \\ &= \left( 2 \left( 1 - \Phi \right)^2 \sin \left( \Phi - \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \right) \left[ \frac{c}{2R} \right] \\ &= \left( 2 \left( 1 - \Phi \right)^2 \sin \left( \Phi - \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 - \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \left( 1 + \Phi \right) - \frac{c}{2R} \cos 2\Phi \right) \\ &= \frac{c}{2R} \left( \frac{c}{2R} \cos
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<u>Overvie</u>w

Demo

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Introduction
   Motivation
   Target Audiences
Overview
   AMUSE
   Purpose of GPUnit
Features and Design
   Features
   Architecture
   Design
Software Engineering
   Testing
   Planning
Impact
```

Astrophysical Multipurpose Software Environment (AMUSE)



State of AMUSE

- Currently used by researchers to run large-scale simulations.
- Scripts, diagnostics, logging are all written by hand.
- ► AMUSE API/programming knowledge is required to create experiments.
- Still better than separated and opaque FORTRAN codes.

```
del number = get_model_number(AMUSE_id, ierr)
if (evolve failed('get model_number'; ierr, evolve, -3)) return
step loop: do ! may need to repeat this loop for retry or backup
  result = star evolve step(AMUSE id. first try)
  if (result == keep going) result = check model(s, AMUSE id, 0)
if (result == keep going) result = star pick next timestep(AMUSE id)
  if (result == keep going) exit step loop
  model number = get model number(AMUSE id, ierr)
  if (evolve failed('get model number', ierr, evolve, -3)) return
  result_reason = get_result_reason(AMUSE_id, ierr)
  if (result == retry) then
     if (evolve failed('get result reason', derry evolves(4)) return
     if (report retries) &
  r Designite(*,'(i6.3x,a,/)') model number, &
            retry reason(Va//ntrim(result reason str(result reason))
  else if (result == backup) then
     if (evolve failed(iget result reason), ierr, evolve, 4)) return
     if (report backups) &
        if (result == retry) result = star prepare for retry(AMUSE id)
  if (result == backup) result = star_dol_backup(AMUSE_id)
  if (result == terminate) then
  evolve = -11 ! Unspecified stop condition reached, or:
     if (s% number of backups in a row > s% max backups in a row ) ther
        evolve = -14 | max backups reached
```

Purpose of GPUnit

- ► Ease the use of AMUSE
- Create/Design/Modify experiments
- ► Select, configure, swap out modules and initial conditions
- Store and restore progress of running experiments.

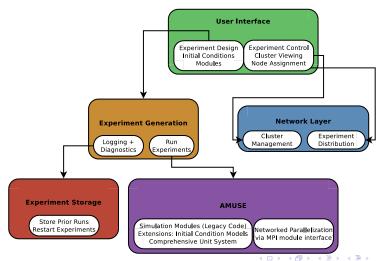


Features

- Configurable experiments that can be saved and shared.
- Diagnostic tools that compute and display useful measurements.
- Storage of experiment state in case of crashes.
- Custom diagnostics and code generation.
- Provides a display of cluster usage to aid in scheduling.



Architecture



Design

- AMUSE is the only integrated simulation environment available.
- ▶ AMUSE is written in Python, streamlines interaction.
- ▶ C++ was considered as it supports Qt as well.
 - Communication w/AMUSE would be cumbersome.
 - AMUSE would be in a separate process.
- ▶ Designed APIs for diagnostics, logging and experiment persistence.
 - Users can create new diagnostics easily.
 - Experiments can be stored in a file structure, a remote DB etc...





Tests

- ▶ GUI / Integration tests were performed manually.
 - Created and ran a simple experiment from scratch to ensure functionality.
- Unit testing performed using Python's unittest module (PyUnit).
- ► Tests:
 - networking
 - built-in diagnostics
 - object serialization
 - experiment storage
 - experiment running

Project Plan

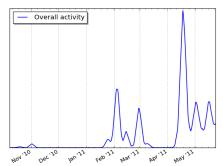
- Mostly waterfall design process.
- Initial phases were spent learning the domain (Physics/AMUSE).
- ► Roles
 - Tim: Physics reference, test subject
 - Andrew/Jason: Experiment and Module design.
 - ► Dan: Diagnostics
 - ► Raj: Logging
 - ► Gabe: Network, GUI.



Team Management

- Used Mercurial as our version control system.
 - Distributed, allows off-line commits.
- ► Team met weekly.
 - ▶ Once to plan work, once to code.
- ▶ Bi-weekly advisor meetings.

Commits to GPUnit Repository vs. Time



Project Impact

- Gives students and physicists easy access to state-of-the-art tools.
- ightharpoonup Simple experiment creation ightharpoonup faster turnaround on experiments.
- ightharpoonup Faster experiments ightarrow more time to study them.
- Current state:
 - Software is usable to create simple experiments.
 - Comes with useful diagnostics, from real experimental setups.
 - Ready to get feedback from more advanced users.
 - Capability/APIs already exist to provide more advanced features.

Demo

▶ Demonstration of a simulation.

Questions

▶ Questions?