Tutorial

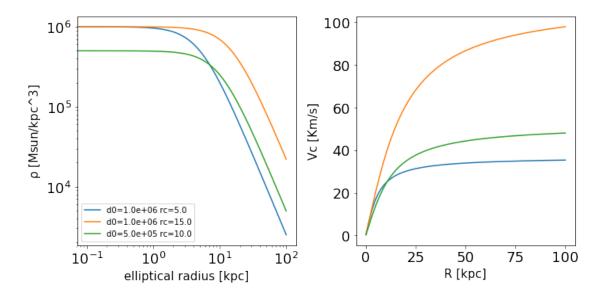
September 27, 2017

```
In [2]: import numpy as np
        import discH
        import discH.dynamic_component as dc
        import matplotlib.pyplot as plt
   HALO MODELS
In [38]: #Isothermal halo
         R=np.linspace(0,100,1000)
         fig=plt.figure(figsize=(10,5))
         axd=fig.add_subplot(121)
         axv=fig.add_subplot(122)
         #d=d0*(1+m*m/rc*rc)^{(-1)}
         mcut=100 #radius where d(m>mcut)=0
         e=0 #ellipticity
         d0=1e6 #Cental density in Msun/kpc3
         rc=5 #Core radius in Kpc
         iso_halo=dc.isothermal_halo(d0=d0, rc=rc, mcut=mcut, e=e)
         dens=iso_halo.dens(R) #3D dens
         vcirc=iso_halo.vcirc(R, nproc=2)
         axd.plot(R,dens[:,1],label='d0=\%.1e rc=\%.1f'\%(d0,rc))
         axv.plot(R,vcirc[:,1],label='d0=\%.1e rc=\%.1f'\%(d0,rc))
         d0=1e6 #Cental density in Msun/kpc3
         rc=15 #Core radius in Kpc
         iso_halo=dc.isothermal_halo(d0=d0, rc=rc, mcut=mcut, e=e)
         dens=iso_halo.dens(R) #3D dens
         vcirc=iso_halo.vcirc(R, nproc=2)
         axd.plot(R,dens[:,1],label='d0=\%.1e rc=\%.1f'\%(d0,rc))
         axv.plot(R,vcirc[:,1],label='d0=%.1e rc=%.1f'%(d0,rc))
         d0=5e5 #Cental density in Msun/kpc3
         rc=10 #Core radius in Kpc
```

```
iso_halo=dc.isothermal_halo(d0=d0, rc=rc, mcut=mcut, e=e)
         dens=iso_halo.dens(R) #3D dens
         vcirc=iso_halo.vcirc(R, nproc=2)
         axd.plot(R,dens[:,1],label='d0=%.1e rc=%.1f'%(d0,rc))
         axv.plot(R,vcirc[:,1],label='d0=%.1e rc=%.1f'%(d0,rc))
         print(iso_halo)
         axd.set_xlabel('elliptical radius [kpc]', fontsize=15)
         axd.set_ylabel('$\\rho$ [Msun/kpc^3]', fontsize=15)
         axv.set_xlabel('R [kpc]', fontsize=15)
         axv.set_ylabel('Vc [Km/s]', fontsize=15)
         axd.set_xscale('log')
         axd.set_yscale('log')
         axd.legend()
         plt.show()
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/discH-3.0.0.dev0-py3.6-macosx-10.
  return np.array(vcirc_iso(R, self.d0, self.rc, self.e, toll=self.toll))
Model: Isothermal halo
d0: 5.00e+05 Msun/kpc3
rc: 10.00
e: 0.000
mcut: 100.000
```

/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/matplotlib/figure.py:1743: UserWa

warnings.warn("This figure includes Axes that are not "



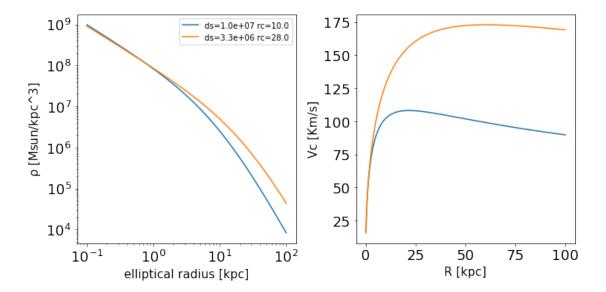
```
In [39]: #NFW halo
        R=np.linspace(0,100,1000)
         fig=plt.figure(figsize=(10,5))
         axd=fig.add_subplot(121)
         axv=fig.add_subplot(122)
         R=np.linspace(0,100,1000)
         #d=d0*((m/rs)^{(-1)})*((1+m/rs)^{(-2)})
         mcut=100 #radius where d(m>mcut)=0
         e=0 #ellipticity
         #Primary use: NFW_halo(d0, rs, mcut=100, e=0)
         \# d=d0/((r/rs)*(1+r/rs)^2)
         #-d0 Scale density in Msun/kpc3
         #-rs Scale length
         #Secondary use: NFW_halo.cosmo(c, V200, H=67 , mcut=100, e=0)
         #-c Concentration parameter
         #-V200 Velocity (km/s) at virial Radius R200 (radius where the density is 200 times th
         #-H Hubble constant (km/s/Mpc)
         d0=1e7 #Scale density in Msun/kpc3
         rs=10 #Scale radius in Kpc
         nfw_halo=dc.NFW_halo(d0=d0, rs=rs, mcut=mcut, e=e) #primary method to call NFW halo
         dens=nfw_halo.dens(R) #3D dens
         vcirc=nfw_halo.vcirc(R, nproc=2)
         axd.plot(R,dens[:,1],label='ds=%.1e rc=%.1f'%(nfw_halo.d0,nfw_halo.rs))
         axv.plot(R,vcirc[:,1],label='ds=%.1e rc=%.1f'%(nfw_halo.d0,nfw_halo.rs))
```

```
V200=150 #Scale radius in Kpc
                       nfw_halo=dc.NFW_halo.cosmo(c=c, V200=V200, mcut=mcut, e=e) #secondary metho do call NFW
                        \#NWF\_halo.cosmo(c, V200, H=67, e=0, mcut=100) \ H \ is the Hubble constant in km/s/Mpc (67) \ Mathematical Research (1998) and the second of the second of
                        dens=nfw_halo.dens(R) #3D dens
                        vcirc=nfw_halo.vcirc(R, nproc=2)
                        axd.plot(R,dens[:,1],label='ds=%.1e rc=%.1f'%(nfw_halo.d0,nfw_halo.rs))
                        axv.plot(R,vcirc[:,1],label='ds=%.1e rc=%.1f'%(nfw_halo.d0,nfw_halo.rs))
                       print(nfw_halo)
                        axd.set_xlabel('elliptical radius [kpc]', fontsize=15)
                        axd.set_ylabel('$\\rho$ [Msun/kpc^3]', fontsize=15)
                        axv.set_xlabel('R [kpc]', fontsize=15)
                        axv.set_ylabel('Vc [Km/s]', fontsize=15)
                        axd.set_xscale('log')
                        axd.set_yscale('log')
                        axd.legend()
                       plt.show()
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/discH-3.0.0.dev0-py3.6-macosx-10.
     return num / den
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/discH-3.0.0.dev0-py3.6-macosx-10.
     return np.array(vcirc_nfw(R, self.d0, self.rc, self.e, toll=self.toll))
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/discH-3.0.0.dev0-py3.6-macosx-10.
     return np.array(vcirc_nfw(R, self.d0, self.rc, self.e, toll=self.toll))
Model: NFW halo
d0: 3.26e+06 Msun/kpc3
rs: 27.99
e: 0.000
mcut: 100.000
```

/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/matplotlib/figure.py:1743: UserWa

c=8 #Scale density in Msun/kpc3

warnings.warn("This figure includes Axes that are not "



```
In [40]: #alfabeta halo
        R=np.linspace(0,100,1000)
        fig=plt.figure(figsize=(10,5))
        axd=fig.add_subplot(121)
        axv=fig.add_subplot(122)
        R=np.linspace(0,100,1000)
        \#d=d0*((m/rs)^{(-alfa)})*((1+m/rs)^{(-(beta-alfa))})
        mcut=100 #radius where d(m>mcut)=0
        e=0 #ellipticity
        d0=1e6 #Scale density in Msun/kpc3
        rs=5 #Scale radius in Kpc
        alfa=1.5 #Inner slope
        beta=2.8 #Outer slope
        ab_halo=dc.alfabeta_halo(d0=d0,alfa=alfa, beta=beta, rs=rs, mcut=mcut, e=e)
        dens=ab_halo.dens(R) #3D dens
        vcirc=ab_halo.vcirc(R, nproc=2)
        axv.plot(R,vcirc[:,1],label='ds=\%.1e rs=\%.1f $$\alpha$=\%.1f $$\beta$=\%.1f'\%(ab_halo.d0,d0,d0)$.
        d0=1e6 #Scale density in Msun/kpc3
        rs=5 #Scale radius in Kpc
        alfa=1.2 #Inner slope
        beta=3.4 #Outer slope
        ab_halo=dc.alfabeta_halo(d0=d0,alfa=alfa, beta=beta, rs=rs, mcut=mcut, e=e)
        dens=ab_halo.dens(R) #3D dens
        vcirc=ab_halo.vcirc(R, nproc=2)
        axd.plot(R,dens[:,1],label='ds=%.1e rs=%.1f $\\alpha$=%.1f $\\beta$=%.1f'%(ab_halo.d0,a
        axv.plot(R,vcirc[:,1],label='ds=\%.1e rs=\%.1f $$\alpha$=\%.1f $$\beta$=\%.1f'\%(ab_halo.d0,d0).
```

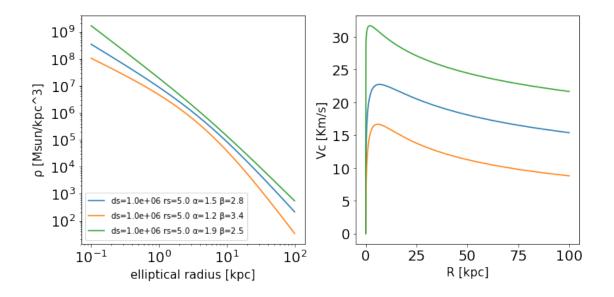
```
d0=1e6 #Scale density in Msun/kpc3
rs=5 #Scale radius in Kpc
alfa=1.9 #Inner slope
beta=2.5 #Outer slope
ab_halo=dc.alfabeta_halo(d0=d0,alfa=alfa, beta=beta, rs=rs, mcut=mcut, e=e)
dens=ab_halo.dens(R) #3D dens
vcirc=ab_halo.vcirc(R, nproc=2)
axd.plot(R,dens[:,1],label='ds=%.1e rs=%.1f $\\alpha$=%.1f $\\beta$=%.1f'%(ab_halo.d0,a
axv.plot(R,vcirc[:,1],label='ds=\%.1e rs=\%.1f $$\alpha$=\%.1f $$\beta$=\%.1f'\%(ab_halo.d0,d0).
axd.set_xlabel('elliptical radius [kpc]', fontsize=15)
axd.set_ylabel('$\\rho$ [Msun/kpc^3]', fontsize=15)
axv.set_xlabel('R [kpc]', fontsize=15)
axv.set_ylabel('Vc [Km/s]', fontsize=15)
axd.set_xscale('log')
axd.set_yscale('log')
axd.legend()
plt.show()
```

/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/discH-3.0.0.dev0-py3.6-macosx-10. return num / den

/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/scipy/integrate/quadpack.py:364: integration interval.

warnings.warn(msg, IntegrationWarning)

/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/matplotlib/figure.py:1743: UserWawarnings.warn("This figure includes Axes that are not "



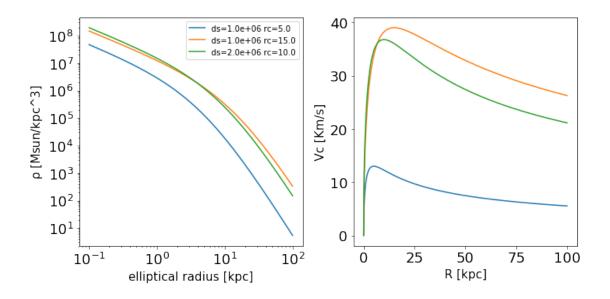
```
In [41]: #hernquist halo
        R=np.linspace(0,100,1000)
         fig=plt.figure(figsize=(10,5))
         axd=fig.add_subplot(121)
         axv=fig.add_subplot(122)
         R=np.linspace(0,100,1000)
         #d=d0*((m/rs)^{(-1)})*((1+m/rs)^{(-2)}
         mcut=100 #radius where d(m>mcut)=0
         e=0 #ellipticity
         d0=1e6 #Scale density in Msun/kpc3
         rs=5 #Scale radius in Kpc
         he_halo=dc.hernquist_halo(d0=d0, rs=rs, mcut=mcut, e=e)
         dens=he_halo.dens(R) #3D dens
         vcirc=he_halo.vcirc(R, nproc=2)
         axd.plot(R,dens[:,1],label='ds=%.1e rc=%.1f'%(he_halo.d0,he_halo.rs))
         axv.plot(R,vcirc[:,1],label='ds=%.1e rc=%.1f'%(he_halo.d0,he_halo.rs))
         d0=1e6 #Scale density in Msun/kpc3
         rs=15 #Scale radius in Kpc
         he_halo=dc.hernquist_halo(d0=d0, rs=rs, mcut=mcut, e=e)
         dens=he_halo.dens(R) #3D dens
         vcirc=he_halo.vcirc(R, nproc=2)
         axd.plot(R,dens[:,1],label='ds=%.1e rc=%.1f'%(he_halo.d0,he_halo.rs))
         axv.plot(R,vcirc[:,1],label='ds=%.1e rc=%.1f'%(he_halo.d0,he_halo.rs))
         d0=2e6 #Scale density in Msun/kpc3
         rs=10 #Scale radius in Kpc
         he_halo=dc.hernquist_halo(d0=d0, rs=rs, mcut=mcut, e=e)
         dens=he_halo.dens(R) #3D dens
         vcirc=he_halo.vcirc(R, nproc=2)
         axd.plot(R,dens[:,1],label='ds=%.1e rc=%.1f'%(he_halo.d0,he_halo.rs))
         axv.plot(R,vcirc[:,1],label='ds=%.1e rc=%.1f'%(he_halo.d0,he_halo.rs))
         print(he_halo)
         axd.set_xlabel('elliptical radius [kpc]', fontsize=15)
         axd.set_ylabel('$\\rho$ [Msun/kpc^3]', fontsize=15)
         axv.set_xlabel('R [kpc]', fontsize=15)
         axv.set_ylabel('Vc [Km/s]', fontsize=15)
         axd.set_xscale('log')
         axd.set_yscale('log')
         axd.legend()
         plt.show()
```

return num / den
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/scipy/integrate/quadpack.py:364:
integration interval.
warnings.warn(msg, IntegrationWarning)

Model: Hernquist halo d0: 2.00e+06 Msun/kpc3

rs: 10.00 e: 0.000 mcut: 100.000

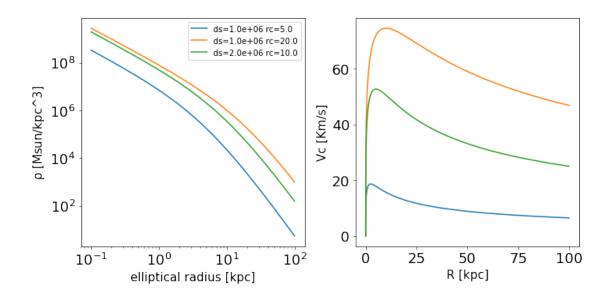
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/matplotlib/figure.py:1743: UserWawarnings.warn("This figure includes Axes that are not "



```
rs=5 #Scale radius in Kpc
         dv_halo=dc.deVacouler_like_halo(d0=d0, rs=rs, mcut=mcut, e=e)
         dens=dv_halo.dens(R) #3D dens
         vcirc=dv_halo.vcirc(R, nproc=2)
         axd.plot(R,dens[:,1],label='ds=%.1e rc=%.1f'%(dv_halo.d0,dv_halo.rs))
         axv.plot(R,vcirc[:,1],label='ds=%.1e rc=%.1f'%(dv_halo.d0,dv_halo.rs))
         d0=1e6 #Scale density in Msun/kpc3
         rs=20 #Scale radius in Kpc
         dv_halo=dc.deVacouler_like_halo(d0=d0, rs=rs, mcut=mcut, e=e)
         dens=dv_halo.dens(R) #3D dens
         vcirc=dv_halo.vcirc(R, nproc=2)
         axd.plot(R,dens[:,1],label='ds=%.1e rc=%.1f'%(dv_halo.d0,dv_halo.rs))
         axv.plot(R,vcirc[:,1],label='ds=%.1e rc=%.1f'%(dv_halo.d0,dv_halo.rs))
         d0=2e6 #Scale density in Msun/kpc3
         rs=10 #Scale radius in Kpc
         dv_halo=dc.deVacouler_like_halo(d0=d0, rs=rs, mcut=mcut, e=e)
         dens=dv_halo.dens(R) #3D dens
         vcirc=dv_halo.vcirc(R, nproc=2)
         axd.plot(R,dens[:,1],label='ds=%.1e rc=%.1f'%(dv_halo.d0,dv_halo.rs))
         axv.plot(R,vcirc[:,1],label='ds=%.1e rc=%.1f'%(dv_halo.d0,dv_halo.rs))
         print(dv_halo)
         axd.set_xlabel('elliptical radius [kpc]', fontsize=15)
         axd.set_ylabel('$\\rho$ [Msun/kpc^3]', fontsize=15)
         axv.set_xlabel('R [kpc]', fontsize=15)
         axv.set_ylabel('Vc [Km/s]', fontsize=15)
         axd.set_xscale('log')
         axd.set_yscale('log')
         axd.legend()
         plt.show()
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/discH-3.0.0.dev0-py3.6-macosx-10.
  return num / den
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/scipy/integrate/quadpack.py:364:
  integration interval.
  warnings.warn(msg, IntegrationWarning)
Model: deVacouler like halo
d0: 2.00e+06 Msun/kpc3
rs: 10.00
e: 0.000
mcut: 100.000
```

d0=1e6 #Scale density in Msun/kpc3

/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/matplotlib/figure.py:1743: UserWawarnings.warn("This figure includes Axes that are not "

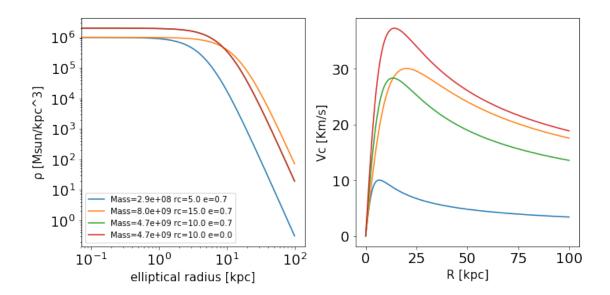


```
In [43]: #Plummer halo
                                    R=np.linspace(0,100,1000)
                                    fig=plt.figure(figsize=(10,5))
                                    axd=fig.add_subplot(121)
                                    axv=fig.add_subplot(122)
                                    R=np.linspace(0,100,1000)
                                    #d=d0*((1+m*m/rs*rs)^{(-5/2)})
                                    mcut=100 #radius where d(m>mcut)=0
                                    e=0.7 #ellipticity
                                    {\tt d0=1e6} \ \textit{\#Central density in Msun/kpc3}
                                    rc=5 #Core radius in Kpc
                                    pl_halo=dc.plummer_halo(d0=d0, rc=rc, mcut=mcut, e=e)
                                    dens=pl_halo.dens(R) #3D dens
                                    vcirc=pl_halo.vcirc(R, nproc=2)
                                    axd.plot(R,dens[:,1],label='Mass=%.1e rc=%.1f e=%.1f'%(pl_halo.mass,pl_halo.rc, pl_halo
                                    axv.plot(R,vcirc[:,1],label='Mass=%.1e rc=%.1f e=%1.f'%(pl_halo.mass,pl_halo.rc, pl_halo.rc, pl_halo.r
                                    d0=1e6 #Central density in Msun/kpc3
                                    rc=15 #Core radius in Kpc
```

pl_halo=dc.plummer_halo(d0=d0, rc=rc, mcut=mcut, e=e)

dens=pl_halo.dens(R) #3D dens

```
vcirc=pl_halo.vcirc(R, nproc=2)
                                               axd.plot(R,dens[:,1],label='Mass=%.1e rc=%.1f e=%.1f'%(pl_halo.mass,pl_halo.rc, pl_halo
                                               axv.plot(R,vcirc[:,1],label='Mass=%.1e rc=%.1f e=%1.f'%(pl_halo.mass,pl_halo.rc, pl_halo.rc, pl_halo.r
                                               d0=2e6 #Central density in Msun/kpc3
                                               rc=10 #Core radius in Kpc
                                              pl_halo=dc.plummer_halo(d0=d0, rc=rc, mcut=mcut, e=e)
                                               dens=pl_halo.dens(R) #3D dens
                                               vcirc=pl_halo.vcirc(R, nproc=2)
                                               axd.plot(R,dens[:,1],label='Mass=%.1e rc=%.1f e=%.1f'%(pl_halo.mass,pl_halo.rc, pl_halo
                                               axv.plot(R,vcirc[:,1],label='Mass=%.1e rc=%.1f e=%1.f'%(pl_halo.mass,pl_halo.rc, pl_halo.rc, pl_halo.r
                                               d0=2e6 #Central density in Msun/kpc3
                                               rc=10 #Core radius in Kpc
                                              pl_halo=dc.plummer_halo(d0=d0, rc=rc, mcut=mcut, e=e)
                                               dens=pl_halo.dens(R) #3D dens
                                               vcirc=pl_halo.vcirc(R, nproc=2)
                                               axd.plot(R,dens[:,1],label='Mass=%.1e rc=%.1f e=%.1f'%(pl_halo.mass,pl_halo.rc, pl_halo
                                               axv.plot(R,vcirc[:,1],label='Mass=%.1e rc=%.1f e=%1.f'%(pl_halo.mass,pl_halo.rc, pl_halo.rc, pl_halo.r
                                              print(pl_halo)
                                               axd.set_xlabel('elliptical radius [kpc]', fontsize=15)
                                               axd.set_ylabel('$\\rho$ [Msun/kpc^3]', fontsize=15)
                                              axv.set_xlabel('R [kpc]', fontsize=15)
                                               axv.set_ylabel('Vc [Km/s]', fontsize=15)
                                               axd.set_xscale('log')
                                               axd.set_yscale('log')
                                               axd.legend()
                                              plt.show()
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/scipy/integrate/quadpack.py:364:
           integration interval.
           warnings.warn(msg, IntegrationWarning)
Model: Plummer halo
Mass: 4.71e+09 Msun
d0: 2.00e+06 Msun/kpc3
rc: 10.00
e: 0.000
mcut: 100.000
```

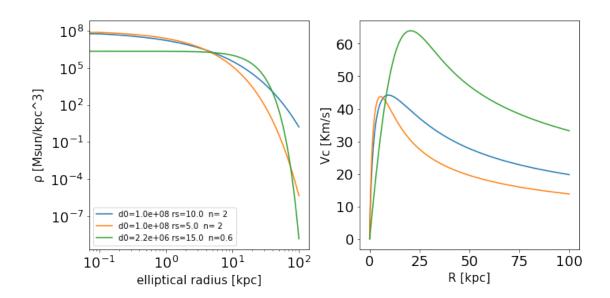


```
In [44]: #Einasto halo
         R=np.linspace(0,100,1000)
         fig=plt.figure(figsize=(10,5))
         axd=fig.add_subplot(121)
         axv=fig.add_subplot(122)
         R=np.linspace(0,100,1000)
         mcut=100 #radius where d(m>mcut)=0
         e=0.0 #ellipticity
         #Primary use: einasto_halo(d0, n, rs, mcut=100, e=0)
         # d=d0*exp(-dn*(r/rs)^{(1/n)})
         #-d0 Central density in Msun/kpc3
         #-n factor n
         #-rs radius containing half the total mass of the halo
         #Secondary use: einasto_halo.de(de, n, rs, mcut=100, e=0)
         # d=de*exp(-2*n*((r/rs)^{(1/n)} - 1))
         #-de Density at rs
         \#-n factor n
         #-rs radius containing half the total mass of the halo
```

```
rs=10 #Radius containing half the total mass of the halo
        ei_halo=dc.einasto_halo(d0=d0, n=n, rs=rs, mcut=mcut, e=e)
        dens=ei_halo.dens(R) #3D dens
        vcirc=ei_halo.vcirc(R, nproc=2)
        axd.plot(R,dens[:,1],label='d0=\%.1e \ rs=\%.1f \ n=\%2.f'\%(ei\_halo.d0,\ ei\_halo.rs,\ ei\_halo.rs)
        axv.plot(R,vcirc[:,1],label='d0='.1e rs=''.1f n=''2.f''(ei_halo.d0, ei_halo.rs, ei_halo.rs)
        d0=1e8 #Central density in Msun/kpc3
        n=1.5 #Core radius in Kpc
        ei_halo=dc.einasto_halo(d0=d0, n=n, rs=rs, mcut=mcut, e=e)
        dens=ei_halo.dens(R) #3D dens
        vcirc=ei_halo.vcirc(R, nproc=2)
        axd.plot(R,dens[:,1],label='d0=%.1e rs=%.1f n=%2.f'%(ei_halo.d0, ei_halo.rs, ei_halo.r
        de=5e5 #Central density in Msun/kpc3
        n=0.6 #Core radius in Kpc
        rs=15
        ei_halo=dc.einasto_halo.de(de=de, n=n, rs=rs, mcut=mcut, e=e)
        dens=ei_halo.dens(R) #3D dens
        vcirc=ei_halo.vcirc(R, nproc=2)
        axd.plot(R,dens[:,1],label='d0=%.1e rs=%.1f n=%.1f'%(ei_halo.d0, ei_halo.rs, ei_halo.r
        axv.plot(R,vcirc[:,1],label='d0=%.1e rs=%.1f n=%.1f'%(ei_halo.d0, ei_halo.rs, ei_halo.
        axd.set_xlabel('elliptical radius [kpc]', fontsize=15)
        axd.set_ylabel('$\\rho$ [Msun/kpc^3]', fontsize=15)
        axv.set_xlabel('R [kpc]', fontsize=15)
        axv.set_ylabel('Vc [Km/s]', fontsize=15)
        axd.set_xscale('log')
        axd.set_yscale('log')
        axd.legend()
        plt.show()
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/scipy/integrate/quadpack.py:364:
  integration interval.
  warnings.warn(msg, IntegrationWarning)
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/matplotlib/figure.py:1743: UserWa
  warnings.warn("This figure includes Axes that are not "
```

d0=1e8 #Central density in Msun/kpc3

n=2 #Factor n



DISC MODELS

```
In [32]: #Exponential disc
         \#Sigma(R) = Sigma0 * Exp(-R/Rd)
         sigma0=1e6 #Cental surface density in Msun/kpc2
         Rd= 2 #Exponential scale length in kpc
         Rcut= 50 #Cylindrical radius where dens(R>Rcut,z)=0
         zcut= 20 #Cylindrical heigth where dens(R, |z|>zcut)=0
         zlaw='gau' #Vertical density law: it could be gau, sech2, exp
         fig=plt.figure(figsize=(20,5))
         ax_dens=fig.add_subplot(131)
         ax_flare=fig.add_subplot(132)
         ax_vcirc=fig.add_subplot(133)
         R=np.linspace(0,20,100) #Cylidrincal radii where estimate surface density and flare
         #Vertical:
         #razor-thin disc
         ed=dc.Exponential_disc.thin(sigma0=sigma0, Rd=Rd, Rcut=Rcut, zcut=zcut)
         sdens=ed.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
         flare=ed.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
         vcirc=ed.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
         ax_dens.plot(R, sdens[:,1], label=ed.name + ' flare:'+ed.flaw)
         ax_flare.plot(R, flare[:,1])
         ax_vcirc.plot(R, vcirc[:,1])
         \#constant\ scale-heigth
         zd=0.5 #Vertical scale heigth in kpc
         ed=dc.Exponential_disc.thick(sigma0=sigma0, Rd=Rd, Rcut=Rcut, zcut=zcut,zd=zd, zlaw=zla
```

sdens=ed.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]

```
flare=ed.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
vcirc=ed.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
ax_dens.plot(R, sdens[:,1], label=ed.name + ' flare:'+ed.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
#polynomial flare
pcoeff = [0.05, 0.01, 0.005] #Coefficent of the polynomial zd(R) = pcoeff[0] + pcoeff[1] * R + pcoeff[1]
ed=dc.Exponential_disc.polyflare(sigma0=sigma0, Rd=Rd, Rcut=Rcut, zcut=zcut, polycoeff=
sdens=ed.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=ed.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
vcirc=ed.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
ax_dens.plot(R, sdens[:,1], label=ed.name + ' flare:'+ed.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
#Asinh flare
\#zd(R)=h0+c*(Arcsinh(R*R/Rf*Rf))
h0=0.4 #Cental zd in kpc
c=1 #
Rf=15 #Flaring scale length in kpc
ed=dc.Exponential_disc.asinhflare(sigma0=sigma0, Rd=Rd, Rcut=Rcut, zcut=zcut, h0=h0, c=
sdens=ed.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=ed.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
vcirc=ed.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
ax_dens.plot(R, sdens[:,1], label=ed.name + ' flare:'+ed.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
#Tanh flare
\#zd(R)=h0+c*(tanh(R*R/Rf*Rf))
h0=0.4 #Cental zd in kpc
c=1 #
Rf=15 #Flaring scale length in kpc
ed=dc.Exponential_disc.tanhflare(sigma0=sigma0, Rd=Rd, Rcut=Rcut, zcut=zcut, h0=h0, c=c
sdens=ed.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=ed.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
vcirc=ed.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
ax_dens.plot(R, sdens[:,1], label=ed.name + ' flare:'+ed.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
print(ed)
ax_dens.legend()
ax_dens.set_xlabel('R [kpc]',fontsize=15)
ax_vcirc.set_xlabel('R [kpc]',fontsize=15)
ax_flare.set_xlabel('R [kpc]',fontsize=15)
ax_dens.set_ylabel('$\Sigma$ [Msun/kpc^2]',fontsize=15)
```

```
ax_flare.set_ylabel('HWHWM [kpc]',fontsize=15)
ax_vcirc.set_ylabel('Vc [km/s]',fontsize=15)
plt.show()
```

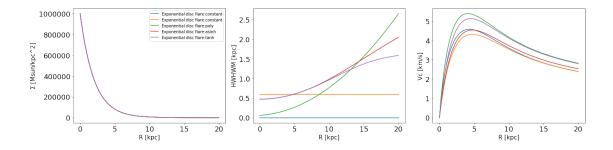
Model: Exponential disc Sigma0: 1.00e+06 Msun/kpc2 Vertical density law: gau Radial density law: epoly

Rd: 2.000 kpc Flaring law: tanh

Fparam: 4.0e-01 1.5e+01 1.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00

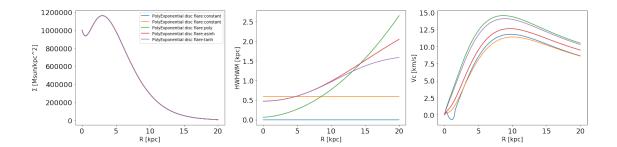
Rcut: 50.000 kpc zcut: 20.000 kpc Rlimit: None

/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/matplotlib/figure.py:1743: UserWawarnings.warn("This figure includes Axes that are not "



```
zlaw='gau' #Vertical density law: it could be gau, sech2, exp
#Vertical:
#razor-thin disc
epd=dc.PolyExponential_disc.thin(sigma0=sigma0, Rd=Rd, coeff=Rcoeff, Rcut=Rcut, zcut=zc
sdens=epd.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=epd.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, co
vcirc=epd.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pl
ax_dens.plot(R, sdens[:,1], label=epd.name + ' flare:'+epd.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
#constant scale-heigth
zd=0.5 #Vertical scale heigth in kpc
epd=dc.PolyExponential_disc.thick(sigma0=sigma0, Rd=Rd, coeff=Rcoeff, Rcut=Rcut, zcut=z
sdens=epd.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=epd.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, co
vcirc=epd.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pl
ax_dens.plot(R, sdens[:,1], label=epd.name + ' flare:'+epd.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
#polynomial flare
pcoeff = [0.05, 0.01, 0.005] #Coefficent of the polynomial zd(R) = pcoeff[0] + pcoeff[1] * R + pcoeff[1]
epd=dc.PolyExponential_disc.polyflare(sigma0=sigma0, Rd=Rd, coeff=Rcoeff, Rcut=Rcut, zc
sdens=epd.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=epd.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, co
vcirc=epd.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pl
ax_dens.plot(R, sdens[:,1], label=epd.name + ' flare:'+epd.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
#Asinh flare
\#zd(R)=h0+c*(Arcsinh(R*R/Rf*Rf))
h0=0.4 #Cental zd in kpc
c=1 #
Rf=15 #Flaring scale length in kpc
epd=dc.PolyExponential_disc.asinhflare(sigma0=sigma0, Rd=Rd, coeff=Rcoeff, Rcut=Rcut, z
sdens=epd.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=epd.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, co
vcirc=epd.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pl
ax_dens.plot(R, sdens[:,1], label=epd.name + ' flare:'+epd.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
#Tanh flare
\#zd(R)=h0+c*(tanh(R*R/Rf*Rf))
h0=0.4 #Cental zd in kpc
c=1 #
```

```
Rf=15 #Flaring scale length in kpc
         epd=dc.PolyExponential_disc.tanhflare(sigma0=sigma0, Rd=Rd, coeff=Rcoeff, Rcut=Rcut, zo
         sdens=epd.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
         flare=epd.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, co
         vcirc=epd.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pl
         ax_dens.plot(R, sdens[:,1], label=epd.name + ' flare:'+epd.flaw)
         ax_flare.plot(R, flare[:,1])
         ax_vcirc.plot(R, vcirc[:,1])
         print(epd)
         ax_dens.legend()
         ax_dens.set_xlabel('R [kpc]',fontsize=15)
         ax_vcirc.set_xlabel('R [kpc]',fontsize=15)
         ax_flare.set_xlabel('R [kpc]',fontsize=15)
         ax_dens.set_ylabel('$\Sigma$ [Msun/kpc^2]',fontsize=15)
         ax_flare.set_ylabel('HWHWM [kpc]',fontsize=15)
         ax_vcirc.set_ylabel('Vc [km/s]',fontsize=15)
         plt.show()
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/scipy/integrate/quadpack.py:364:
  warnings.warn(msg, IntegrationWarning)
Model: PolyExponential disc
Sigma0: 1.00e+06 Msun/kpc2
Vertical density law: gau
Radial density law: epoly
Rd: 2.000 kpc
Polycoeff: 1.0e+00 2.0e-01 4.0e-01 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00
Flaring law: tanh
Fparam: 4.0e-01 1.5e+01 1.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00
Rcut: 50.000 kpc
zcut: 20.000 kpc
Rlimit: None
```



```
In [35]: fig=plt.figure(figsize=(20,5))
         ax_dens=fig.add_subplot(131)
         ax_flare=fig.add_subplot(132)
         ax_vcirc=fig.add_subplot(133)
         R=np.linspace(0,20,100) #Cylidrincal radii where estimate surface density and flare
         #Frat disc
         \#Siqma(R) = Siqma0*Exp(-R/Rd)*(1+R/Rd2)^alfa
         sigma0=1e6 #Cental surface density in Msun/kpc2
         Rd= 3 #Exponential scale length in kpc
         Rd2= 1.5 #Secondary scale length in kpc
         alfa= 1.5 #Exponent
         Rcut= 50 #Cylindrical radius where dens(R>Rcut,z)=0
         zcut= 20 #Cylindrical heigth where dens(R, |z|>zcut)=0
         zlaw='gau' #Vertical density law: it could be gau, sech2, exp
         #Vertical:
         #razor-thin disc
         ed=dc.Frat_disc.thin(sigma0=sigma0, Rd=Rd, Rd2=Rd2,alpha=alfa, Rcut=Rcut, zcut=zcut)
         sdens=ed.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
         flare=ed.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
         vcirc=ed.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
         ax_dens.plot(R, sdens[:,1], label=ed.name + ' flare:'+ed.flaw)
         ax_flare.plot(R, flare[:,1])
         ax_vcirc.plot(R, vcirc[:,1])
         #constant scale-heigth
         zd=0.5 #Vertical scale heigth in kpc
         ed=dc.Frat_disc.thick(sigma0=sigma0, Rd=Rd,Rd2=Rd2,alpha=alfa, Rcut=Rcut, zcut=zcut,zd=
         sdens=ed.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
         flare=ed.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
         vcirc=ed.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
         ax_dens.plot(R, sdens[:,1], label=ed.name + ' flare:'+ed.flaw)
         ax_flare.plot(R, flare[:,1])
```

ax_vcirc.plot(R, vcirc[:,1])

```
#polynomial flare
pcoeff = [0.05, 0.01, 0.005] \# Coefficent \ of \ the \ polynomial \ zd(R) = pcoeff[0] + pcoeff[1] * R + pcoe
ed=dc.Frat_disc.polyflare(sigma0=sigma0, Rd=Rd,Rd2=Rd2,alpha=alfa, Rcut=Rcut, zcut=zcu
sdens=ed.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=ed.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
vcirc=ed.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
ax_dens.plot(R, sdens[:,1], label=ed.name + ' flare:'+ed.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
#Asinh flare
\#zd(R)=h0+c*(Arcsinh(R*R/Rf*Rf))
h0=0.4 #Cental zd in kpc
c=1 #
Rf=15 #Flaring scale length in kpc
ed=dc.Frat_disc.asinhflare(sigma0=sigma0, Rd=Rd, Rd2=Rd2,alpha=alfa, Rcut=Rcut, zcut=zc
sdens=ed.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=ed.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
vcirc=ed.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
ax_dens.plot(R, sdens[:,1], label=ed.name + ' flare:'+ed.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
#Tanh flare
\#zd(R)=h0+c*(tanh(R*R/Rf*Rf))
h0=0.4 #Cental zd in kpc
c=1 #
Rf=15 #Flaring scale length in kpc
ed=dc.Frat_disc.tanhflare(sigma0=sigma0, Rd=Rd, Rd2=Rd2,alpha=alfa, Rcut=Rcut, zcut=zcu
sdens=ed.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=ed.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
vcirc=ed.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
ax_dens.plot(R, sdens[:,1], label=ed.name + ' flare:'+ed.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
print(ed)
ax_dens.legend()
ax_dens.set_xlabel('R [kpc]',fontsize=15)
ax_vcirc.set_xlabel('R [kpc]',fontsize=15)
ax_flare.set_xlabel('R [kpc]',fontsize=15)
ax_dens.set_ylabel('$\Sigma$ [Msun/kpc^2]',fontsize=15)
ax_flare.set_ylabel('HWHWM [kpc]',fontsize=15)
ax_vcirc.set_ylabel('Vc [km/s]',fontsize=15)
plt.show()
```

/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/scipy/integrate/quadpack.py:364: warnings.warn(msg, IntegrationWarning)

Model: Frat disc

Sigma0: 1.00e+06 Msun/kpc2 Vertical density law: gau Radial density law: fratlaw

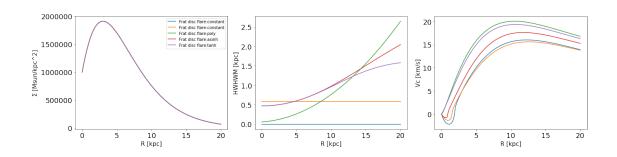
Rd: 3.00 kpc Rd2: 1.50 kpc alpha: 1.50

Flaring law: tanh

Fparam: 4.0e-01 1.5e+01 1.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00

Rcut: 50.000 kpc zcut: 20.000 kpc Rlimit: None

/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/matplotlib/figure.py:1743: UserWawarnings.warn("This figure includes Axes that are not "



In [37]: fig=plt.figure(figsize=(20,5))
 ax_dens=fig.add_subplot(131)
 ax_flare=fig.add_subplot(132)
 ax_vcirc=fig.add_subplot(133)
 R=np.linspace(0,20,100) #Cylidrincal radii where estimate surface density and flare

```
#Gau disc
\#Sigma(R) = Sigma0 * Exp(-0.5 * ((R-R0)/sigmad)^2)
sigma0=1e6 #Cental surface density in Msun/kpc2
RO= 2 #Radius where Sigma reach the peak
sigmad= 2 #Dispersion
Rcut= 50 #Cylindrical radius where dens(R>Rcut,z)=0
zcut= 20 \#Cylindrical height where dens(R, |z|>zcut)=0
zlaw='gau' #Vertical density law: it could be gau, sech2, exp
#Vertical:
#razor-thin disc
gd=dc.Gaussian_disc.thin(sigma0=sigma0, sigmad=sigmad, R0=R0, Rcut=Rcut, zcut=zcut)
sdens=gd.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=gd.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
vcirc=gd.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
ax_dens.plot(R, sdens[:,1], label=gd.name + ' flare:'+gd.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
#constant scale-heigth
zd=0.5 #Vertical scale heigth in kpc
gd=dc.Gaussian_disc.thick(sigma0=sigma0, sigmad=sigmad, R0=R0, Rcut=Rcut, zcut=zcut,zd=
sdens=gd.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=gd.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
vcirc=gd.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
ax_dens.plot(R, sdens[:,1], label=gd.name + ' flare: '+gd.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
#polynomial flare
pcoeff = [0.05, 0.01, 0.005] #Coefficent of the polynomial zd(R) = pcoeff[0] + pcoeff[1] * R + pcoeff[1]
gd=dc.Gaussian_disc.polyflare(sigma0=sigma0, sigmad=sigmad, R0=R0, Rcut=Rcut, zcut=zcu
sdens=gd.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
flare=gd.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
vcirc=gd.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
ax_dens.plot(R, sdens[:,1], label=gd.name + ' flare:'+gd.flaw)
ax_flare.plot(R, flare[:,1])
ax_vcirc.plot(R, vcirc[:,1])
#Asinh flare
\#zd(R)=h0+c*(Arcsinh(R*R/Rf*Rf))
h0=0.4 #Cental zd in kpc
c=1 #
Rf=15 #Flaring scale length in kpc
gd=dc.Gaussian_disc.asinhflare(sigma0=sigma0, sigmad=sigmad, R0=R0, Rcut=Rcut, zcut=zcu
sdens=gd.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
```

```
flare=gd.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
         vcirc=gd.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
         ax_dens.plot(R, sdens[:,1], label=gd.name + ' flare:'+gd.flaw)
         ax_flare.plot(R, flare[:,1])
         ax_vcirc.plot(R, vcirc[:,1])
         #Tanh flare
         \#zd(R)=h0+c*(tanh(R*R/Rf*Rf))
         h0=0.4 #Cental zd in kpc
         c=1 #
         Rf=15 #Flaring scale length in kpc
         gd=dc.Gaussian_disc.tanhflare(sigma0=sigma0, sigmad=sigmad, R0=R0, Rcut=Rcut, zcut=zcut
         sdens=gd.Sdens(R) #sdens 2D array: col-0 R, col-1 Surface density at R [Msun/kpc^2]
         flare=gd.flare(R, HWHM=True) #radial profile of the vertical scale length: col-0 R, col
         vcirc=gd.vcirc(R, nproc=2) #Vcric 2D array: col-0 R, col-1 Circular velocity on the pla
         ax_dens.plot(R, sdens[:,1], label=gd.name + ' flare:'+gd.flaw)
         ax_flare.plot(R, flare[:,1])
         ax_vcirc.plot(R, vcirc[:,1])
         print(gd)
         ax_dens.legend()
         ax_dens.set_xlabel('R [kpc]',fontsize=15)
         ax_vcirc.set_xlabel('R [kpc]',fontsize=15)
         ax_flare.set_xlabel('R [kpc]',fontsize=15)
         ax_dens.set_ylabel('$\Sigma$ [Msun/kpc^2]',fontsize=15)
         ax_flare.set_ylabel('HWHWM [kpc]',fontsize=15)
         ax_vcirc.set_ylabel('Vc [km/s]',fontsize=15)
         plt.show()
/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/scipy/integrate/quadpack.py:364:
  warnings.warn(msg, IntegrationWarning)
Model: Gaussian disc
Sigma0: 1.00e+06 Msun/kpc2
```

Vertical density law: gau

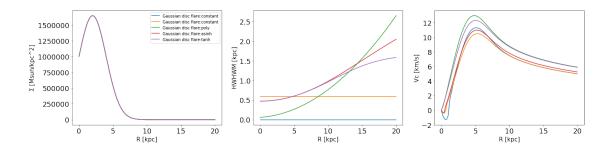
Radial density law: gau

sigmad: 2.000 kpc RO: 2.000 kpc Flaring law: tanh

Fparam: 4.0e-01 1.5e+01 1.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00

Rcut: 50.000 kpc zcut: 20.000 kpc Rlimit: None

/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/matplotlib/figure.py:1743: UserWawarnings.warn("This figure includes Axes that are not "



Notes on disc components class.

-Initialize a class with data:

It is possible to define a disc component fitting some data. If we want to fit the surface density we must define a disc model using the parameter rfit_array, while if we want to fit the flaring we must use the ffit_array. In both cases the array should be an array containing the R in the first column the data in the second and if present the data error on the third column. If the chosen flaring law is polynomial we must provide also the degree of the polynomial with the keyword fitdegree. Examples below

In [13]: #We want a razor-thin disc with a exponential surface density law obtained fittig some
 #oberserved data
R=np.linspace(0.1,30,20)
sigma_o=1e6*np.exp(-R/4)
observed_data=np.zeros(shape=(20,2))
observed_data[:,0]=R
observed_data[:,1]=sigma_o
 #define the model
ed=dc.Exponential_disc.thin(rfit_array=observed_data)
print(ed)

#We want an exponential disc with a polynomial flare
#flaring data
zd=lambda R,a1,a2,a3: a1+a2*R+a3*R*R

zd_o=zd(R,0.4,0.01,0.2)
observed_dataf=np.zeros(shape=(20,2))
observed_dataf[:,0]=R
observed_dataf[:,1]=zd_o
ed=dc.Exponential_disc.polyflare(rfit_array=observed_data,ffit_array=observed_dataf,fit
print(ed)

Model: Exponential disc Sigma0: 1.00e+06 Msun/kpc2 Vertical density law: dirac Radial density law: epoly

Rd: 4.000 kpc

Flaring law: constant

Fparam: 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00

Rcut: 50.000 kpc zcut: 30.000 kpc Rlimit: None

Model: Exponential disc Sigma0: 1.00e+06 Msun/kpc2 Vertical density law: gau Radial density law: epoly

Rd: 4.000 kpc Flaring law: poly

Fparam: 4.0e-01 1.0e-02 2.0e-01 -9.4e-18 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00

Rcut: 50.000 kpc zcut: 30.000 kpc Rlimit: None

POTENTIAL ESTIMATE

In [14]: #Estimate the potential of a single component

#Define model
d0=1e6 #Cental density in Msun/kpc3
rc=5 #Core radius in Kpc
mcut=100 #radius where d(m>mcut)=0
e=0 #ellipticity
iso halo=dc isothermal halo(d0=d0 r

iso_halo=dc.isothermal_halo(d0=d0, rc=rc, mcut=mcut, e=e)

#Estimate potential

R=[0.1,2,10] #List with the cylindrical radial coordinates in Kpc Z=[0,0.1,1] #List with the cylindrical vertical coordinates in Kpc grid=True #If True create a grid from R and Z, otherwise estimate the potential in the nproc=2 #Number of processors to use for parallel computation toll=1e-4 #Relative and absolute Tollerance for the potential integration

```
potential_grid=iso_halo.potential(R=R,Z=Z,grid=grid,nproc=2)
        print(potential_grid)
         #First Column -R
         \#Second\ Column\ -Z
         #Third Column Potenzial in Kpc^2/Myr^2
[[ 1.00000000e-01 0.0000000e+00 -4.23552484e-03]
 [ 1.00000000e-01 1.00000000e-01 -4.23543065e-03]
 [ 1.00000000e-01 1.00000000e+00 -4.22621602e-03]
 [ 2.00000000e+00 0.00000000e+00 -4.19961401e-03]
 [ 2.00000000e+00 1.00000000e-01 -4.19952792e-03]
 [ 2.00000000e+00 1.00000000e+00 -4.19109439e-03]
 [ 1.00000000e+01 0.0000000e+00 -3.72924475e-03]
 [ 1.00000000e+01 1.00000000e-01 -3.72921321e-03]
 [ 1.00000000e+01 1.00000000e+00 -3.72609958e-03]]
In [15]: #Estimate the potential of a ensemble of dynamic components
         from discH.dynamics import galpotential
         #Step1: Define the components
         #Halo
         d0=1e6
        rs=5
        mcut=100
        e=0
        halo=dc.NFW_halo(d0=d0, rs=rc, mcut=mcut, e=e)
         #Bulge
        d0=3e6
        rs=1
        mcut=10
        e = 0.6
        bulge=dc.hernquist_halo(d0=d0, rs=rc, mcut=mcut, e=e)
         #Stellar disc
        sigma0=1e6
        Rd=3
         z_0d=0.4
        zlaw='sech2'
        Rcut=50
        zcut=30
        disc=dc.Exponential_disc.thick(sigma0=sigma0, Rd=Rd, zd=zd, zlaw=zlaw, Rcut=Rcut, zcut=
         #Step2: Initialize galpotential class
         ga=galpotential(dynamic_components=(halo,disc,bulge))
         #If you want to check the properties of the component:
```

```
print('#########STEP2########")
print('Components info')
ga.dynamic_components_info()
print('###############")
#Step3
#Calculate potential at R-Z
R=np.linspace(0.1,30,10) #List with the cylindrical radial coordinates in Kpc
Z=np.linspace(0,5,10) #List with the cylindrical vertical coordinates in Kpc
grid=True #If True create a grid from R and Z, otherwise estimate the potential in the
nproc=2 #Number of proccesor to use for parallel computation
toll=1e-4 #Relative and absolute Tollerance for the potential integration
Rcut=None #If not None, set the Rcut of all the disc components to this value
zcut=None #If not None, set the zcut of all the disc components to this value
mcut=None #If not None, set the mcut of all the halo components to this value
external_potential=None #If not None, this should be an array matching the dimension of
print('#########STEP2########")
print('Estimate Potential')
hp=ga.potential(R,Z,grid=grid, nproc=nproc, toll=toll, Rcut=Rcut, zcut=zcut, mcut=mcut,
#Return a grid with O-R 1-Z 2-Total Potential in kpc^2/Myr^2
print('\nReturn a grid 0-R 1-Z 2-Total Potential in kpc^2/Myr^2, e.g.:')
print(hp[:5])
print('#################")
#Step4 Use the results or save them in files:
#The potential information can be accessed with
pot_grid=ga.potential_grid
#Array with col-0: R in kpc, col-1: Z in kpc, col-2: Total potential in kpc^2/Myr^2
pot_grid_complete=ga.potential_grid_complete
#Array with col-0: R in kpc, col-1: Z in kpc, col-i+1: Potential of the single (i+1)th
#col-ncomponent+2: External potential col-ncomponent+3: Total potential
#e.q:
pot_disc=pot_grid_complete[:,3]
#To save in file
complete=True #If True save the pot_grid_complete array (see above), if False the pot_g
filename='potential.dat' #File where to store the data
ga.save(filename=filename, complete=complete)
```

##########STEP2##########

Components info Components: 0 Model: NFW halo

d0: 1.00e+06 Msun/kpc3

rs: 5.00 e: 0.000

```
mcut: 100.000
Components: 1
Model: Exponential disc
Sigma0: 1.00e+06 Msun/kpc2
Vertical density law: sech2
Radial density law: epoly
Rd: 3.000 kpc
Flaring law: constant
Fparam: 4.0e-01 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00
Rcut: 50.000 kpc
zcut: 30.000 kpc
Rlimit: None
Components: 2
Model: Hernquist halo
d0: 3.00e+06 Msun/kpc3
rs: 5.00
e: 0.600
mcut: 10.000
##############################
##########STEP2###########
Estimate Potential
External potential: No
Calculating Potential of the 1th component (NFW halo)...Done (0.01 s)
Calculating Potential of the 2th component (Exponential disc)...Done (1.78 s)
Calculating Potential of the 3th component (Hernquist halo)...Done (0.04 s)
Return a grid O-R 1-Z 2-Total Potential in kpc^2/Myr^2, e.g.:
[[ 1.00000000e-01 0.0000000e+00 -2.99033196e-03]
 [ 1.00000000e-01 5.5555556e-01 -2.75253289e-03]
 [ 1.00000000e-01 1.11111111e+00 -2.51901029e-03]
 [ 1.00000000e-01 1.66666667e+00 -2.32376449e-03]
 [ 1.00000000e-01 2.2222222e+00 -2.15828749e-03]]
##############################
  ESTIMATE SCALE HEIGHT The scale height of a disc can be obtained using the class dis-
cHeight
In [19]: from discH.dynamics import discHeight
         ##STEP: 1
         #Define all the fixed components
         #Halo
```

d0=1e6 rs=5

```
mcut=100
e=0
halo=dc.NFW_halo(d0=d0, rs=rc, mcut=mcut, e=e)
#Bulge
d0=3e6
rs=1
mcut=10
e = 0.6
bulge=dc.hernquist_halo(d0=d0, rs=rc, mcut=mcut, e=e)
#Stellar disc
sigma0=5e6
Rd=3
zd=0.4
zlaw='sech2'
Rcut=50
zcut=30
disc=dc.Exponential_disc.thick(sigma0=sigma0, Rd=Rd, zd=zd, zlaw=zlaw, Rcut=Rcut, zcut=
galaxy=(bulge,disc,halo)
#STEP 2: Define the disc model
#Gas disc
g_sigma0=1e6
g_Rd=5
g_Rd2=5
g_alpha=1
Rcut=60
zcut=30
gas_disc=dc.Frat_disc.thin(sigma0=g_sigma0, Rd=g_Rd, Rd2=g_Rd2, alpha=g_alpha, Rcut=Rcu
#NB, Here the definition of the flaring model is not important, because then it will be
#scale height calculation, so the use of thin is useful to avoid to insert useless info
#STEP 3: Initialize the discHeight class
h=discHeight(dynamic_components=galaxy, disc_component=gas_disc)
#Step 4: Estimat height
zlaw='gau' #Vertical zlaw, it could be 'gau', 'sech2' or 'exp' default=gau
flaw='poly' #Flaring law, it could be 'poly', 'asinh', 'tanh', default=poly
polyflare_degree=5 #If flaw='poly' this is the degree of the polynomial, otherwise it is
#Vel dispersion
#Velocity dispersion, we assume that the disc component as an isotropic velocity disper
#isothermal in the vertical direction, so vdisp=vdisp(R).
#There are different option:
```

```
#1-Constant velocity dispersion
vdisp=10
#2-Function of R, e.g.
vdisp=lambda R: 10 + 5/(1+R)
#3-Array of values with col-0 R col-1 v(R)
vdisp_array=np.array([[0,1,4,5,10],[15,12,10,9,8]])
vdisp=vdisp_array
#In this internally, vidsp=vdisp_func(R), where vdisp_func is the interpolating function
#R array
#These three quantities define the cylindrical R coordinates that will be used to estim
Rpoints=30 #Number of R points, or list of Rpoints, default=30
Rinterval='linear' #interval type, default=linear
                   \#Min-max\ R, default=(0.01,30)
Rrange=(0.01,30)
#If Rpoints is a number, the R grid is defined as np.linspace(Rrange[0], Rrange[30], Rpoi
#If Rpoints is a list a tuple or np.ndarray use the points inside the list
#Z array
#These three quantities define the cylindrical z coordinates that will be used to estim
             #Number of z points, or list of zpoints, default=30
                 #nterval type, default=log
Zinterval='log'
Zrange=(0,10) #Min-max z, default=(0,10)
#If Zpoints is a number, the z grid is defined as np.linspace(Zrange[0], Zrange[30], Zpoi
#If Zpoints is a list a tuple or np.ndarray use the points inside the list
#NB, Zrange[0] must be always 0 to have a good estimate of the vertical profile of the
#The estimate of zd is iterative. The iteration stop when one of the following is True
#Number of iteration < Niter
#Maximum Absolute residual between two sequential estiamates of zd lower than flaretoll
#Maximum Relative residual between two sequential estiamates of zd lower than flaretoll
Niter=10 #Max number of iteration, default=10
flaretollabs=1e-4 # default=1e-4
flaretollrel=1e-4 # default=1e-4
nproc=2 #Number of processors to use for parallel computation, default=2
Rcut=None #If not None, set the Rcut of all the disc components to this value, default=
zcut=None #If not None, set the zcut of all the disc components to this value, default=
mcut=None #If not None, set the mcut of all the halo components to this value, default=
```

inttoll=1e-4 #Relative and absolute Tollerance for the potential integration, default=1 external_potential=None #External potential, default=None outdir='gasHeight_res' #Folder where to save the outputs, default='gasHeight' diagnostic=True #If True, save figures and tables to see all results of the iterations

Rlimit='max' #If not None, set a limit Radius for the flaring, i.e. the radius where ze #this could be useful when the flare is fitted with an high degree polynomial that can

#if 'max', Rlimit=max(R), where R is defined using Rpoints (see above)

```
Calculating fixed potential
External potential: No
Calculating Potential of the 1th component (Hernquist halo)...Done (0.08 s)
Calculating Potential of the 2th component (Exponential disc)...Done (16.69 s)
Calculating Potential of the 3th component (NFW halo)...Done (0.01 s)
Fixed potential Done
Iter-0: Massless disc
**************
            START FITZPROFILE
*************
Number of Radii: 30
Number of Vertical points: 30
Number of the used distributions: 1
                             ['gau']
nplot 2
---Fitting---
Working on radius: 0.01
Working on radius: 1.04
Plotting
Working on radius: 2.08
Working on radius: 3.11
Plotting
Working on radius: 4.15
Working on radius: 5.18
Plotting
Working on radius: 6.21
Working on radius: 7.25
Plotting
Working on radius: 8.28
Working on radius: 9.32
Plotting
Working on radius: 10.35
Working on radius: 11.39
Plotting
Working on radius: 12.42
Working on radius: 13.45
Plotting
Working on radius: 14.49
Working on radius: 15.52
Plotting
Working on radius: 16.56
Working on radius: 17.59
```

```
Plotting
Working on radius: 18.62
Working on radius: 19.66
Plotting
Working on radius: 20.69
Working on radius: 21.73
Plotting
Working on radius: 22.76
Working on radius: 23.80
Plotting
Working on radius: 24.83
Working on radius: 25.86
Working on radius: 26.90
Working on radius: 27.93
Working on radius: 28.97
Working on radius: 30.00
Save figures
Writing table
DONE in 0.104 minutes
Output data files in gasHeight_res/diagnostic/run0/dat
Output images in gasHeight_res/diagnostic/run0/image
***************
           END FITZPROFILE
**************
***************
            START FITFLARE
*************
Start fitting
Writing table
Save table
Make plot
Save plot
data in gasHeight_res/diagnostic/run0/flare/fitflare_par.dat
image in gasHeight_res/diagnostic/run0/flare/flare.pdf
*************
             END FITFLARE
**************
Tter-0: Done
Iter-1:
External potential: Yes
Calculating Potential of the 1th component (Frat disc)...Done (25.28 s)
**************
           START FITZPROFILE
**************
```

Number of Radii: 30

```
Number of Vertical points: 30
Number of the used distributions: 1 ['gau']
1
nplot 2
---Fitting---
Working on radius: 0.01
Working on radius: 1.04
Plotting
Working on radius: 2.08
Working on radius: 3.11
Plotting
Working on radius: 4.15
Working on radius: 5.18
Plotting
Working on radius: 6.21
Working on radius: 7.25
Plotting
Working on radius: 8.28
Working on radius: 9.32
Plotting
Working on radius: 10.35
Working on radius: 11.39
Plotting
Working on radius: 12.42
Working on radius: 13.45
Plotting
Working on radius: 14.49
Working on radius: 15.52
Plotting
Working on radius: 16.56
Working on radius: 17.59
Plotting
Working on radius: 18.62
Working on radius: 19.66
Plotting
Working on radius: 20.69
Working on radius: 21.73
Plotting
Working on radius: 22.76
Working on radius: 23.80
Plotting
Working on radius: 24.83
Working on radius: 25.86
Working on radius: 26.90
Working on radius: 27.93
Working on radius: 28.97
Working on radius: 30.00
Save figures
```

```
Writing table
DONE in 0.093 minutes
Output data files in gasHeight_res/diagnostic/run0/dat
Output images in gasHeight_res/diagnostic/run0/image
**************
           END FITZPROFILE
*************
*************
           START FITFLARE
***************
Start fitting
Writing table
Save table
Make plot
Save plot
data in gasHeight_res/diagnostic/run0/flare/fitflare_par.dat
image in gasHeight_res/diagnostic/run0/flare/flare.pdf
***************
            END FITFLARE
*************
Iter-1: Done
Max Absolute residual=1.28e+00
Max Relative residual=1.25e-01
Iter-2:
External potential: Yes
Calculating Potential of the 1th component (Frat disc)...Done (24.32 s)
*************
          START FITZPROFILE
**************
Number of Radii: 30
Number of Vertical points: 30
Number of the used distributions: 1
                         ['gau']
nplot 2
---Fitting---
Working on radius: 0.01
Working on radius: 1.04
Plotting
```

/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/matplotlib/pyplot.py:524: Runtime

max_open_warning, RuntimeWarning)

Working on radius: 2.08

```
Working on radius: 3.11
Plotting
Working on radius: 4.15
Working on radius: 5.18
Plotting
Working on radius: 6.21
Working on radius: 7.25
Plotting
Working on radius: 8.28
Working on radius: 9.32
Plotting
Working on radius: 10.35
Working on radius: 11.39
Plotting
Working on radius: 12.42
Working on radius: 13.45
Plotting
Working on radius: 14.49
Working on radius: 15.52
Plotting
Working on radius: 16.56
Working on radius: 17.59
Plotting
Working on radius: 18.62
Working on radius: 19.66
Plotting
Working on radius: 20.69
Working on radius: 21.73
Plotting
Working on radius: 22.76
Working on radius: 23.80
Plotting
Working on radius: 24.83
Working on radius: 25.86
Working on radius: 26.90
Working on radius: 27.93
Working on radius: 28.97
Working on radius: 30.00
Save figures
Writing table
DONE in 0.088 minutes
Output data files in gasHeight_res/diagnostic/run1/dat
Output images in gasHeight_res/diagnostic/run1/image
*************
               END FITZPROFILE
***************
***************
```

START FITFLARE

```
*****************
Start fitting
Writing table
Save table
Make plot
Save plot
data in gasHeight_res/diagnostic/run1/flare/fitflare_par.dat
image in gasHeight_res/diagnostic/run1/flare/flare.pdf
*************
              END FITFLARE
***************
Iter-2: Done
Max Absolute residual=5.39e-02
Max Relative residual=9.96e-03
Iter-3:
External potential: Yes
Calculating Potential of the 1th component (Frat disc)...Done (24.57 s)
*************
            START FITZPROFILE
******************
Number of Radii: 30
Number of Vertical points: 30
Number of the used distributions: 1
                            ['gau']
nplot 2
---Fitting---
Working on radius: 0.01
Working on radius: 1.04
Plotting
Working on radius: 2.08
Working on radius: 3.11
Plotting
Working on radius: 4.15
Working on radius: 5.18
Plotting
Working on radius: 6.21
Working on radius: 7.25
Plotting
Working on radius: 8.28
Working on radius: 9.32
Plotting
Working on radius: 10.35
Working on radius: 11.39
Plotting
```

Working on radius: 12.42

```
Working on radius: 13.45
Plotting
Working on radius: 14.49
Working on radius: 15.52
Plotting
Working on radius: 16.56
Working on radius: 17.59
Plotting
Working on radius: 18.62
Working on radius: 19.66
Plotting
Working on radius: 20.69
Working on radius: 21.73
Plotting
Working on radius: 22.76
Working on radius: 23.80
Plotting
Working on radius: 24.83
Working on radius: 25.86
Working on radius: 26.90
Working on radius: 27.93
Working on radius: 28.97
Working on radius: 30.00
Save figures
Writing table
DONE in 0.094 minutes
Output data files in gasHeight_res/diagnostic/run2/dat
Output images in gasHeight_res/diagnostic/run2/image
*************
              END FITZPROFILE
**************
***************
               START FITFLARE
*************
Start fitting
Writing table
Save table
Make plot
Save plot
data in gasHeight_res/diagnostic/run2/flare/fitflare_par.dat
image in gasHeight_res/diagnostic/run2/flare/flare.pdf
*************
               END FITFLARE
*************
Iter-3: Done
Max Absolute residual=9.68e-02
Max Relative residual=1.33e-02
```

```
Iter-4:
External potential: Yes
Calculating Potential of the 1th component (Frat disc)...Done (24.32 s)
*************
             START FITZPROFILE
***************
Number of Radii: 30
Number of Vertical points: 30
Number of the used distributions: 1
                                 ['gau']
1
nplot 2
---Fitting---
Working on radius: 0.01
Working on radius: 1.04
Plotting
Working on radius: 2.08
Working on radius: 3.11
Plotting
Working on radius: 4.15
Working on radius: 5.18
Plotting
Working on radius: 6.21
Working on radius: 7.25
Plotting
Working on radius: 8.28
Working on radius: 9.32
Plotting
Working on radius: 10.35
Working on radius: 11.39
Plotting
Working on radius: 12.42
Working on radius: 13.45
Plotting
Working on radius: 14.49
Working on radius: 15.52
Plotting
Working on radius: 16.56
Working on radius: 17.59
Plotting
Working on radius: 18.62
Working on radius: 19.66
Plotting
Working on radius: 20.69
Working on radius: 21.73
Plotting
```

Working on radius: 22.76

```
Working on radius: 23.80
Plotting
Working on radius: 24.83
Working on radius: 25.86
Working on radius: 26.90
Working on radius: 27.93
Working on radius: 28.97
Working on radius: 30.00
Save figures
Writing table
DONE in 0.084 minutes
Output data files in gasHeight_res/diagnostic/run3/dat
Output images in gasHeight_res/diagnostic/run3/image
**************
           END FITZPROFILE
***************
**************
            START FITFLARE
*************
Start fitting
Writing table
Save table
Make plot
Save plot
data in gasHeight_res/diagnostic/run3/flare/fitflare_par.dat
image in gasHeight_res/diagnostic/run3/flare/flare.pdf
*************
             END FITFLARE
***************
Iter-4: Done
Max Absolute residual=1.33e-03
Max Relative residual=1.82e-04
Iter-5:
External potential: Yes
Calculating Potential of the 1th component (Frat disc)...Done (24.23 s)
**************
           START FITZPROFILE
*************
Number of Radii: 30
Number of Vertical points: 30
Number of the used distributions: 1
                          ['gau']
nplot 2
---Fitting---
Working on radius: 0.01
```

```
Working on radius: 1.04
Plotting
Working on radius: 2.08
Working on radius: 3.11
Plotting
Working on radius: 4.15
Working on radius: 5.18
Plotting
Working on radius: 6.21
Working on radius: 7.25
Plotting
Working on radius: 8.28
Working on radius: 9.32
Plotting
Working on radius: 10.35
Working on radius: 11.39
Plotting
Working on radius: 12.42
Working on radius: 13.45
Plotting
Working on radius: 14.49
Working on radius: 15.52
Plotting
Working on radius: 16.56
Working on radius: 17.59
Plotting
Working on radius: 18.62
Working on radius: 19.66
Plotting
Working on radius: 20.69
Working on radius: 21.73
Plotting
Working on radius: 22.76
Working on radius: 23.80
Plotting
Working on radius: 24.83
Working on radius: 25.86
Working on radius: 26.90
Working on radius: 27.93
Working on radius: 28.97
Working on radius: 30.00
Save figures
Writing table
DONE in 0.092 minutes
Output data files in gasHeight_res/diagnostic/run4/dat
Output images in gasHeight_res/diagnostic/run4/image
```

```
*****************
***************
          START FITFLARE
*************
Start fitting
Writing table
Save table
Make plot
Save plot
data in gasHeight_res/diagnostic/run4/flare/fitflare_par.dat
image in gasHeight_res/diagnostic/run4/flare/flare.pdf
**************
           END FITFLARE
**************
Iter-5: Done
Max Absolute residual=6.51e-05
Max Relative residual=8.95e-06
```

1 Results of the functions:

0-final_gas_model: The final disc model, with the Radial surface density law given in inputr and the vertical profiles obtained in the iterative process

1-tab_zd: A tabel with 0-R [kpc] 1-Zd [kpc]

2-flare_func: The interpolating function of tab_zd, zd(R)=flare_func(R)

3-fit_func: The best-fit function (as defined with flaw) to the last zd estimate.

In the output folder you can find:

- -finalflare_zd.pdf: a figure with the zd estimate at each iterative step (gray lines), the last estimate is shown by blue points and the red curve is the last best-fit function
 - -finalflare_hwhm.pdf: The final zd estimate, but the value in y is the HWHM
 - -tabflare.dat: 0-Col R[kpc], 1-Col zd[kpc], 2-Col HWHM[kpc]
 - -tab_fixedpotential.dat: Tab with the potentials of the fixed dynamic components
 - -tab_totpotential.dat: Tab with the potential of the final disc component
 - -My suggestion is to use:

```
Rlimit='max'
flaw='poly'
polyflare_degree_degree=5

In [22]: ##An example of use: estimate of the scale height for the HI disc and H2 disc

##Fixed component

##halo

#halo=dc.isothermal_halo(....)

##bulge

#bulge=dc.hernquist_halo(....)
```

```
##stellar disc
#disc=dc.Exponential_disc.thick(...)

##0bserved intrinsic HI surface density
#HI_tab=[RHI,Sigma_HI]
#HI_disc=dc.Frat_disc.thin(rfit_array=HI_tab,....)

##0bserved intrinsic H2 surface density
#HII_tab=[RHII,Sigma_HII]
#HII_disc=dc.Frat_disc.thin(rfit_array=HII_tab,....)

#galaxy=(halo,bulge,disc)

#h=discHeight(dynamic_components=galaxy, disc_component=HI_disc)
#HI_disc=h.height(....)[0]

##galaxy_new=(halo,bulge,disc,HI_disc)

#h=discHeight(dynamic_components=galaxy_new, disc_component=HII_disc)
#HII_disc=h.height(....)[0]
```

In []: