# Untitled

### August 22, 2017

```
In [1]: import numpy as np
        import discH
        import discH.dynamic_component as dc
   HALO MODELS
In [2]: #Isothermal halo
        #d=d0*(1+m*m/rc*rc)^{(-1)}
        d0=1e6 #Cental density in Msun/kpc3
        rc=5 #Core radius in Kpc
        mcut=100 #radius where d(m>mcut)=0
        e=0 #ellipticity
        iso_halo=dc.isothermal_halo(d0=d0, rc=rc, mcut=mcut, e=e)
        print(iso_halo)
Model: Isothermal halo
d0: 1.00e+06 Msun/kpc3
rc: 5.00
e: 0.000
mcut: 100.000
In [3]: #NFW halo
        #d=d0*((m/rs)^{(-1)})*((1+m/rs)^{(-2)})
        d0=1e6 #Scale density in Msun/kpc3
        rs=5 #Scale radius in Kpc
        mcut=100 #radius where d(m>mcut)=0
        e=0 #ellipticity
        nfw_halo=dc.NFW_halo(d0=d0, rs=rs, mcut=mcut, e=e)
        print(nfw_halo)
Model: NFW halo
d0: 1.00e+06 Msun/kpc3
rs: 5.00
e: 0.000
mcut: 100.000
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In [4]: #alfabeta halo
        \#d=d0*((m/rs)^{(-alfa)})*((1+m/rs)^{(-(beta-alfa))})
        d0=1e6 #Scale density in Msun/kpc3
       rs=5 #Scale radius in Kpc
       mcut=100 #radius where d(m>mcut)=0
        e=0 #ellipticity
        alfa=1.5 #Inner slope
       beta=2.8 #Outer slope
        ab_halo=dc.alfabeta_halo(d0=d0,alfa=alfa, beta=beta, rs=rs, mcut=mcut, e=e)
        print(ab_halo)
Model: AlfaBeta halo
d0: 1.00e+06 Msun/kpc3
rs: 5.00
alfa: 1.5
alfa: 2.8
e: 0.000
mcut: 100.000
In [5]: #hernquist halo
        \#d=d0*((m/rs)^{(-1)})*((1+m/rs)^{(-2)}
       d0=1e6 #Scale density in Msun/kpc3
       rs=5 #Scale radius in Kpc
       mcut=100 #radius where d(m>mcut)=0
        e=0 #ellipticity
       he_halo=dc.hernquist_halo(d0=d0, rs=rs, mcut=mcut, e=e)
       print(he_halo)
Model: Hernquist halo
d0: 1.00e+06 Msun/kpc3
rs: 5.00
e: 0.000
mcut: 100.000
In [6]: #deVacouler like halo
        \#d=d0*((m/rs)^{(-3/2)})*((1+m/rs)^{(-5/2)})
        #It is an approximation of the R1/4 law
        d0=1e6 #Scale density in Msun/kpc3
       rs=5 #Scale radius in Kpc
       mcut=100 #radius where d(m>mcut)=0
        e=0 #ellipticity
        dv_halo=dc.deVacouler_like_halo(d0=d0, rs=rs, mcut=mcut, e=e)
        print(dv_halo)
Model: deVacouler like halo
d0: 1.00e+06 Msun/kpc3
```

```
In [8]: #Plummer halo
                    #d=d0*((1+m*m/rs*rs)^{(-5/2)})
                    d0=1e6 #Central density in Msun/kpc3
                    rc=5 #Core radius in Kpc
                    mcut=100 #radius where d(m>mcut)=0
                    e=0.7 #ellipticity
                    pl_halo=dc.plummer_halo(d0=d0, rc=rc, mcut=mcut, e=e)
                    print(pl_halo)
Model: Plummer halo
Mass: 2.95e+08 Msun
d0: 1.00e+06 Msun/kpc3
rc: 5.00
e: 0.700
mcut: 100.000
       DISC MODELS
In [9]: #Exponential disc
                    \#Sigma(R) = Sigma0*Exp(-R/Rd)
                    sigma0=1e6 #Cental surface density in Msun/kpc2
                    Rd= 2 #Exponential scale length in kpc
                    Rcut= 50 #Cylindrical radius where dens(R>Rcut,z)=0
                    zcut= 20 #Cylindrical heigth where dens(R, |z|>zcut)=0
                    zlaw='gau' #Vertical density law: it could be gau, sech2, exp
                    #Vertical:
                    #razor-thin disc
                    ed=dc.Exponential_disc.thin(sigma0=sigma0, Rd=Rd, Rcut=Rcut, zcut=zcut)
                    #constant scale-heigth
                    zd=0.5 #Vertical scale heigth in kpc
                    ed=dc.Exponential_disc.thick(sigma0=sigma0, Rd=Rd, Rcut=Rcut, zcut=zcut,zd=zd, zlaw=zlaw
                    #polynomial flare
                    pcoeff=[0.5,1,2] #Coefficent of the polynomial zd(R)=pcoeff[0]+pcoeff[1]*R+pcoeff[2]*R*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]*R+pcoeff[2]
                    ed=dc.Exponential_disc.polyflare(sigma0=sigma0, Rd=Rd, Rcut=Rcut, zcut=zcut, polycoeff=p
                    #Asinh flare
                    \#zd(R)=h0+c*(Arcsinh(R*R/Rf*Rf))
                    h0=0.4 #Cental zd in kpc
                    c=1 #
                    Rf=15 #Flaring scale length in kpc
                    ed=dc.Exponential_disc.asinhflare(sigma0=sigma0, Rd=Rd, Rcut=Rcut, zcut=zcut, h0=h0, c=c
                    #Tanh flare
```

rs: 5.00 e: 0.000 mcut: 100.000

```
h0=0.4 #Cental zd in kpc
                 c=1 #
                 Rf=15 #Flaring scale length in kpc
                 ed=dc.Exponential_disc.tanhflare(sigma0=sigma0, Rd=Rd, Rcut=Rcut, zcut=zcut, h0=h0, c=c,
                 print(ed)
Model: Exponential disc
Sigma0: 1.00e+06 Msun/kpc2
Vertical density law: gau
Radial density law: epoly
Rd: 2.000 kpc
Flaring law: tanh
Fparam: 4.0e-01 1.5e+01 1.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00
Rcut: 50.000 kpc
zcut: 20.000 kpc
Rlimit: None
In [10]: #Poly Exponential disc
                    \#Sigma(R) = Sigma0*Exp(-R/Rd)*polynomial(R)
                    sigma0=1e6 #Cental surface density in Msun/kpc2
                   Rd= 2 #Exponential scale length in kpc
                   Rcoeff=[1,0.2,0.1] #Coefficent of the polynomial(R)=Rcoeff[0]+Rcoeff[1]*R+Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*R*Rcoeff[2]*Rcoeff[2]*R*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoeff[2]*Rcoef
                                                                   #Rcoeff will be always renormalised to have Rcoeff[0]=1
                   Rcut= 50 #Cylindrical radius where dens(R>Rcut,z)=0
                    zcut= 20 #Cylindrical heigth where dens(R, |z|>zcut)=0
                    zlaw='gau' #Vertical density law: it could be gau, sech2, exp
                    #Vertical:
                    #razor-thin disc
                    epd=dc.PolyExponential_disc.thin(sigma0=sigma0, Rd=Rd, coeff=Rcoeff, Rcut=Rcut, zcut=zc
                    #constant scale-heigth
                   zd=0.5 #Vertical scale heigth in kpc
                    epd=dc.PolyExponential_disc.thick(sigma0=sigma0, Rd=Rd, coeff=Rcoeff, Rcut=Rcut, zcut=z
                    #polynomial flare
                   pcoeff=[0.5,1,2] #Coefficent of the polynomial zd(R)=pcoeff[0]+pcoeff[1]*R+pcoeff[2]*R*
                    epd=dc.PolyExponential_disc.polyflare(sigma0=sigma0, Rd=Rd, coeff=Rcoeff, Rcut=Rcut, zc
                    #Asinh flare
                    \#zd(R)=h0+c*(Arcsinh(R*R/Rf*Rf))
                   h0=0.4 #Cental zd in kpc
                    c=1 #
                   Rf=15 #Flaring scale length in kpc
                    epd=dc.PolyExponential_disc.asinhflare(sigma0=sigma0, Rd=Rd, coeff=Rcoeff, Rcut=Rcut, z
                    #Tanh flare
                    \#zd(R)=h0+c*(tanh(R*R/Rf*Rf))
                   h0=0.4 #Cental zd in kpc
                    c=1 #
```

#zd(R)=h0+c\*(tanh(R\*R/Rf\*Rf))

```
epd=dc.PolyExponential_disc.tanhflare(sigma0=sigma0, Rd=Rd, coeff=Rcoeff, Rcut=Rcut, zc
         print(epd)
Model: PolyExponential disc
Sigma0: 1.00e+06 Msun/kpc2
Vertical density law: gau
Radial density law: epoly
Rd: 2.000 kpc
Polycoeff: 1.0e+00 2.0e-01 1.0e-01 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00
Flaring law: tanh
Fparam: 4.0e-01 1.5e+01 1.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00
Rcut: 50.000 kpc
zcut: 20.000 kpc
Rlimit: None
In [11]: #Frat disc
         \#Sigma(R) = Sigma0*Exp(-R/Rd)*(1+R/Rd2)^alfa
         sigma0=1e6 #Cental surface density in Msun/kpc2
         {\tt Rd=\ 5}\ \textit{\#Exponential scale length in kpc}
         Rd2= 2 #Secondary scale length in kpc
         alfa= 2 #Exponent
         Rcut= 50 #Cylindrical radius where dens(R>Rcut,z)=0
         zcut= 20 #Cylindrical heigth where dens(R, |z|>zcut)=0
         zlaw='gau' #Vertical density law: it could be gau, sech2, exp
         #Vertical:
         #razor-thin disc
         ed=dc.Frat_disc.thin(sigma0=sigma0, Rd=Rd, Rd2=Rd2,alpha=alfa, Rcut=Rcut, zcut=zcut)
         #constant scale-heigth
         zd=0.5 #Vertical scale heigth in kpc
         ed=dc.Frat_disc.thick(sigma0=sigma0, Rd=Rd,Rd2=Rd2,alpha=alfa, Rcut=Rcut, zcut=zcut,zd=
         #polynomial flare
         pcoeff=[0.5,1,2] #Coefficent of the polynomial zd(R)=pcoeff[0]+pcoeff[1]*R+pcoeff[2]*R*
         ed=dc.Frat_disc.polyflare(sigma0=sigma0, Rd=Rd,Rd2=Rd2,alpha=alfa, Rcut=Rcut, zcut=zcu
         #Asinh flare
         \#zd(R)=h0+c*(Arcsinh(R*R/Rf*Rf))
         h0=0.4 #Cental zd in kpc
         c=1 #
         Rf=15 #Flaring scale length in kpc
         ed=dc.Frat_disc.asinhflare(sigma0=sigma0, Rd=Rd, Rd2=Rd2,alpha=alfa, Rcut=Rcut, zcut=zc
         #Tanh flare
         \#zd(R)=h0+c*(tanh(R*R/Rf*Rf))
         h0=0.4 #Cental zd in kpc
         c=1 #
         Rf=15 #Flaring scale length in kpc
         ed=dc.Frat_disc.tanhflare(sigma0=sigma0, Rd=Rd, Rd2=Rd2,alpha=alfa, Rcut=Rcut, zcut=zcu
         print(ed)
```

Rf=15 #Flaring scale length in kpc

Sigma0: 1.00e+06 Msun/kpc2 Vertical density law: gau Radial density law: fratlaw Rd: 5.00 kpc Rd2: 2.00 kpc alpha: 2.00 Flaring law: tanh Fparam: 4.0e-01 1.5e+01 1.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 Rcut: 50.000 kpc zcut: 20.000 kpc Rlimit: None In [12]: #Gau disc  $\#Sigma(R) = Sigma0*Exp(-0.5*((R-R0)/sigmad)^2)$ sigma0=1e6 #Cental surface density in Msun/kpc2 RO= 5 #Radius where Sigma reach the peak sigmad= 2 #Dispersion Rcut= 50 #Cylindrical radius where dens(R>Rcut,z)=0 zcut= 20 #Cylindrical heigth where dens(R, |z|>zcut)=0 zlaw='gau' #Vertical density law: it could be gau, sech2, exp #Vertical: #razor-thin disc gd=dc.Gaussian\_disc.thin(sigma0=sigma0, sigmad=sigmad, R0=R0, Rcut=Rcut, zcut=zcut) #constant scale-heigth zd=0.5 #Vertical scale heigth in kpc gd=dc.Gaussian\_disc.thick(sigma0=sigma0, sigmad=sigmad, R0=R0, Rcut=Rcut, zcut=zcut,zd= #polynomial flare pcoeff=[0.5,1,2] #Coefficent of the polynomial zd(R)=pcoeff[0]+pcoeff[1]\*R+pcoeff[2]\*R\*gd=dc.Gaussian\_disc.polyflare(sigma0=sigma0, sigmad=sigmad, R0=R0, Rcut=Rcut, zcut=zcu #Asinh flare #zd(R)=h0+c\*(Arcsinh(R\*R/Rf\*Rf))h0=0.4 #Cental zd in kpc c=1 # Rf=15 #Flaring scale length in kpc gd=dc.Gaussian\_disc.asinhflare(sigma0=sigma0, sigmad=sigmad, R0=R0, Rcut=Rcut, zcut=zcu #Tanh flare #zd(R)=h0+c\*(tanh(R\*R/Rf\*Rf))h0=0.4 #Cental zd in kpc c=1 # Rf=15 #Flaring scale length in kpc gd=dc.Gaussian\_disc.tanhflare(sigma0=sigma0, sigmad=sigmad, R0=R0, Rcut=Rcut, zcut=zcut print(gd) Model: Gaussian disc Sigma0: 1.00e+06 Msun/kpc2

Model: Frat disc

Vertical density law: gau Radial density law: gau

sigmad: 2.000 kpc
R0: 5.000 kpc
Flaring law: tanh

Fparam: 4.0e-01 1.5e+01 1.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00

Rcut: 50.000 kpc zcut: 20.000 kpc Rlimit: None

Notes on disc components class.

-Initialize a class with data:

It is possible to define a disc component fitting some data. If we want to fit the surface density we must define a disc model using the parameter rfit\_array, while if we want to fit the flaring we must use the ffit\_array. In both cases the array should be an array containing the R in the first column the data in the second and if present the data error on the third column. If the chosen flaring law is polynomial we must provide also the degree of the polynomial with the keyword fitdegree. Examples below

```
In [13]: #We want a razor-thin disc with a exponential surface density law obtained fittig some
         #oberserved data
         R=np.linspace(0.1,30,20)
         sigma_o=1e6*np.exp(-R/4)
         observed_data=np.zeros(shape=(20,2))
         observed_data[:,0]=R
         observed_data[:,1]=sigma_o
         #define the model
         ed=dc.Exponential_disc.thin(rfit_array=observed_data)
         print(ed)
         #We want an exponential disc with a polynomial flare
         #flaring data
         zd=lambda R,a1,a2,a3: a1+a2*R+a3*R*R
         zd_o=zd(R,0.4,0.01,0.2)
         observed_dataf=np.zeros(shape=(20,2))
         observed_dataf[:,0]=R
         observed_dataf[:,1]=zd_o
         ed=dc.Exponential_disc.polyflare(rfit_array=observed_data,ffit_array=observed_dataf,fit
         print(ed)
```

Model: Exponential disc Sigma0: 1.00e+06 Msun/kpc2 Vertical density law: dirac Radial density law: epoly

Rd: 4.000 kpc

Flaring law: constant

Fparam: 0.0e+00 Rcut: 50.000 kpc zcut: 30.000 kpc Rlimit: None Model: Exponential disc Sigma0: 1.00e+06 Msun/kpc2 Vertical density law: gau Radial density law: epoly Rd: 4.000 kpc Flaring law: poly Fparam: 4.0e-01 1.0e-02 2.0e-01 -9.4e-18 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 Rcut: 50.000 kpc zcut: 30.000 kpc Rlimit: None POTENTIAL ESTIMATE In [14]: #Estimate the potential of a single component #Define model d0=1e6 #Cental density in Msun/kpc3 rc=5 #Core radius in Kpc mcut=100 #radius where d(m>mcut)=0 e=0 #ellipticity iso\_halo=dc.isothermal\_halo(d0=d0, rc=rc, mcut=mcut, e=e) #Estimate potential R=[0.1,2,10] #List with the cylindrical radial coordinates in Kpc Z=[0,0.1,1] #List with the cylindrical vertical coordinates in Kpc grid=True #If True create a grid from R and Z, otherwise estimate the potentail in the nproc=2 #Number of proccesors to use for parallel computation toll=1e-4 #Relative and absolute Tollerance for the potential integration potential\_grid=iso\_halo.potential(R=R,Z=Z,grid=grid,nproc=2) print(potential\_grid) #First Column -R #Second Column -Z #Third Column Potenzial in Kpc^2/Myr^2 [ 1.00000000e-01 0.0000000e+00 -4.23552484e-03] [ 1.00000000e-01 1.00000000e-01 -4.23543065e-03] [ 1.00000000e-01 1.00000000e+00 -4.22621602e-03] [ 2.00000000e+00 0.00000000e+00 -4.19961401e-03] [ 2.00000000e+00 1.00000000e-01 -4.19952792e-03]

[ 2.00000000e+00 1.00000000e+00 -4.19109439e-03] [ 1.00000000e+01 0.0000000e+00 -3.72924475e-03]

```
[ 1.0000000e+01
                   1.00000000e-01 -3.72921321e-03]
 [ 1.00000000e+01 1.00000000e+00 -3.72609958e-03]]
In [15]: #Estimate the potential of a ensemble of dynamic components
        from discH.dynamics import galpotential
         #Step1: Define the components
         #Halo
         d0=1e6
        rs=5
        mcut=100
         e=0
        halo=dc.NFW_halo(d0=d0, rs=rc, mcut=mcut, e=e)
         #Bulge
        d0=3e6
        rs=1
        mcut=10
        bulge=dc.hernquist_halo(d0=d0, rs=rc, mcut=mcut, e=e)
         #Stellar disc
        sigma0=1e6
        Rd=3
        zd = 0.4
        zlaw='sech2'
        Rcut=50
        zcut=30
         disc=dc.Exponential_disc.thick(sigma0=sigma0, Rd=Rd, zd=zd, zlaw=zlaw, Rcut=Rcut, zcut=
         #Step2: Initialize galpotential class
        ga=galpotential(dynamic_components=(halo,disc,bulge))
         #If you want to check the properties of the component:
        print('##########STEP2########")
        print('Components info')
        ga.dynamic_components_info()
        print('############")
         #Step3
         #Calculate potential at R-Z
        R=np.linspace(0.1,30,10) #List with the cylindrical radial coordinates in Kpc
         Z=np.linspace(0,5,10) #List with the cylindrical vertical coordinates in Kpc
         grid=True #If True create a grid from R and Z, otherwise estimate the potential in the
        nproc=2 #Number of proccesor to use for parallel computation
         toll=1e-4 #Relative and absolute Tollerance for the potential integration
        Rcut=None #If not None, set the Rcut of all the disc components to this value
```

```
zcut=None #If not None, set the zcut of all the disc components to this value
         mcut=None #If not None, set the mcut of all the halo components to this value
         external_potential=None #If not None, this should be an array matching the dimension of
         print('#########STEP2########")
         print('Estimate Potential')
        hp=ga.potential(R,Z,grid=grid, nproc=nproc, toll=toll, Rcut=Rcut, zcut=zcut, mcut=mcut,
         #Return a grid with O-R 1-Z 2-Total Potential in kpc^2/Myr^2
         print('\nReturn a grid 0-R 1-Z 2-Total Potential in kpc^2/Myr^2, e.g.:')
        print(hp[:5])
        print('##############")
         #Step4 Use the results or save them in files:
         #The potential information can be accessed with
        pot_grid=ga.potential_grid
         #Array with col-0: R in kpc, col-1: Z in kpc, col-2: Total potential in kpc^2/Myr^2
        pot_grid_complete=ga.potential_grid_complete
         #Array with col-0: R in kpc, col-1: Z in kpc, col-i+1: Potential of the single (i+1)th
         #col-ncomponent+2: External potential col-ncomponent+3: Total potential
         #e.q:
        pot_disc=pot_grid_complete[:,3]
         #To save in file
         complete=True #If True save the pot_grid_complete array (see above), if False the pot_g
         filename='potential.dat' #File where to store the data
         ga.save(filename=filename, complete=complete)
##########STEP2###########
Components info
Components: 0
Model: NFW halo
d0: 1.00e+06 Msun/kpc3
rs: 5.00
e: 0.000
mcut: 100.000
Components: 1
Model: Exponential disc
Sigma0: 1.00e+06 Msun/kpc2
Vertical density law: sech2
Radial density law: epoly
Rd: 3.000 kpc
Flaring law: constant
Fparam: 4.0e-01 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00 0.0e+00
Rcut: 50.000 kpc
zcut: 30.000 kpc
```

Rlimit: None

```
Components: 2
Model: Hernquist halo
d0: 3.00e+06 Msun/kpc3
rs: 5.00
e: 0.600
mcut: 10.000
##############################
##########STEP2###########
Estimate Potential
External potential: No
Calculating Potential of the 1th component (NFW halo)...Done (0.01 s)
Calculating Potential of the 2th component (Exponential disc)...Done (1.78 s)
Calculating Potential of the 3th component (Hernquist halo)...Done (0.04 s)
Return a grid O-R 1-Z 2-Total Potential in kpc^2/Myr^2, e.g.:
[[ 1.00000000e-01 0.00000000e+00 -2.99033196e-03]
 [ 1.00000000e-01 5.5555556e-01 -2.75253289e-03]
 [ 1.00000000e-01 1.11111111e+00 -2.51901029e-03]
 [ 1.00000000e-01 1.66666667e+00 -2.32376449e-03]
 [ 1.00000000e-01 2.2222222e+00 -2.15828749e-03]]
#####################################
```

ESTIMATE SCALE HEIGHT The scale heigth of a disc can be obtained using the class discHeight

```
In [19]: from discH.dynamics import discHeight
```

```
##STEP: 1
#Define all the fixed components
#Halo
d0=1e6
rs=5
mcut=100
e=0
halo=dc.NFW_halo(d0=d0, rs=rc, mcut=mcut, e=e)
#Bulge
d0=3e6
rs=1
mcut=10
e = 0.6
bulge=dc.hernquist_halo(d0=d0, rs=rc, mcut=mcut, e=e)
#Stellar disc
sigma0=5e6
```

```
Rd=3
zd = 0.4
zlaw='sech2'
Rcut=50
zcut=30
disc=dc.Exponential_disc.thick(sigma0=sigma0, Rd=Rd, zd=zd, zlaw=zlaw, Rcut=Rcut, zcut=
galaxy=(bulge,disc,halo)
#STEP 2: Define the disc model
#Gas disc
g_sigma0=1e6
g_Rd=5
g_Rd2=5
g_alpha=1
Rcut=60
zcut=30
gas_disc=dc.Frat_disc.thin(sigma0=g_sigma0, Rd=g_Rd, Rd2=g_Rd2, alpha=g_alpha, Rcut=Rcu
#NB, Here the definition of the flaring model is not important, because then it will be
#scale height calculation, so the use of thin is useful to avoid to insert useless info
#STEP 3: Initialize the discHeight class
h=discHeight(dynamic_components=galaxy, disc_component=gas_disc)
#Step 4: Estimat height
zlaw='gau' #Vertical zlaw, it could be 'qau', 'sech2' or 'exp' default=qau
flaw='poly' #Flaring law, it could be 'poly', 'asinh', 'tanh', default=poly
polyflare_degree=5 #If flaw='poly' this is the degree of the polynomial, otherwise it is
#Vel dispersion
#Velocity dispersion, we assume that the disc component as an isotropic velocity disper
#isothermal in the vertical direction, so vdisp=vdisp(R).
#There are different option:
#1-Constant velocity dispersion
vdisp=10
\#2\text{-Function of }R, e.g.
vdisp=lambda R: 10 + 5/(1+R)
#3-Array of values with col-0 R col-1 v(R)
vdisp_array=np.array([[0,1,4,5,10],[15,12,10,9,8]])
vdisp=vdisp_array
#In this internally, vidsp=vdisp_func(R), where vdisp_func is the interpolating function
#R array
#These three quantities define the cylindrical R coordinates that will be used to estim
Rpoints=30 #Number of R points, or list of Rpoints, default=30
Rinterval='linear' #interval type, default=linear
```

```
Rrange=(0.01,30)
                          \#Min\text{-}max R, default=(0.01,30)
        #If Rpoints is a number, the R grid is defined as np.linspace(Rrange[0], Rrange[30], Rpoi
        #If Rpoints is a list a tuple or np.ndarray use the points inside the list
        #Z array
        #These three quantities define the cylindrical z coordinates that will be used to estin
                     #Number of z points, or list of zpoints, default=30
        Zinterval='log'
                        #nterval type, default=log
        Zrange=(0,10) #Min-max z, default=(0,10)
        #If Zpoints is a number, the z grid is defined as np.linspace(Zrange[0], Zrange[30], Zpoi
        #If Zpoints is a list a tuple or np.ndarray use the points inside the list
        #NB, Zrange[0] must be always 0 to have a good estimate of the vertical profile of the
        #The estimate of zd is iterative. The iteration stop when one of the following is True
        #Number of iteration < Niter
        #Maximum Absolute residual between two sequential estiamates of zd lower than flaretoll
        #Maximum Relative residual between two sequential estiamates of zd lower than flaretoll
        Niter=10 #Max number of iteration, default=10
        flaretollabs=1e-4 # default=1e-4
        flaretollrel=1e-4 # default=1e-4
        nproc=2 #Number of processors to use for parallel computation, default=2
        Rcut=None #If not None, set the Rcut of all the disc components to this value, default=
        zcut=None #If not None, set the zcut of all the disc components to this value, default=
        mcut=None #If not None, set the mcut of all the halo components to this value, default=
        Rlimit='max' #If not None, set a limit Radius for the flaring, i.e. the radius where zo
        #this could be useful when the flare is fitted with an high degree polynomial that can
        #if 'max', Rlimit=max(R), where R is defined using Rpoints (see above)
        inttoll=1e-4 #Relative and absolute Tollerance for the potential integration, default=1
        external_potential=None #External potential, default=None
        outdir='gasHeight_res' #Folder where to save the outputs, default='qasHeight'
        diagnostic=True #If True, save figures and tables to see all results of the iterations
        final_gas_model, tab_zd,flare_func,fit_func=h.height(flaw=flaw, zlaw=zlaw, polyflare_de
Calculating fixed potential
External potential: No
Calculating Potential of the 1th component (Hernquist halo)...Done (0.08 s)
Calculating Potential of the 2th component (Exponential disc)...Done (16.69 s)
Calculating Potential of the 3th component (NFW halo)...Done (0.01 s)
Fixed potential Done
```

Iter-0: Massless disc

# \*\*\*\*\*\*\*\*\*\*\*\*\*\* START FITZPROFILE \*\*\*\*\*\*\*\*\*\*\*\*\* Number of Radii: 30 Number of Vertical points: 30 Number of the used distributions: 1 ['gau'] nplot 2 ---Fitting---Working on radius: 0.01 Working on radius: 1.04 Plotting Working on radius: 2.08 Working on radius: 3.11 Plotting Working on radius: 4.15 Working on radius: 5.18 Plotting Working on radius: 6.21 Working on radius: 7.25 Plotting Working on radius: 8.28 Working on radius: 9.32 Plotting Working on radius: 10.35 Working on radius: 11.39 Plotting Working on radius: 12.42 Working on radius: 13.45 Plotting Working on radius: 14.49 Working on radius: 15.52 Plotting Working on radius: 16.56 Working on radius: 17.59 Plotting Working on radius: 18.62 Working on radius: 19.66 Plotting Working on radius: 20.69 Working on radius: 21.73 Plotting Working on radius: 22.76 Working on radius: 23.80 Plotting Working on radius: 24.83

Working on radius: 25.86 Working on radius: 26.90

```
Working on radius: 27.93
Working on radius: 28.97
Working on radius: 30.00
Save figures
Writing table
DONE in 0.104 minutes
Output data files in gasHeight_res/diagnostic/run0/dat
Output images in gasHeight_res/diagnostic/run0/image
**************
           END FITZPROFILE
**************
***************
            START FITFLARE
**************
Start fitting
Writing table
Save table
Make plot
Save plot
data in gasHeight_res/diagnostic/run0/flare/fitflare_par.dat
image in gasHeight_res/diagnostic/run0/flare/flare.pdf
**************
             END FITFLARE
*************
Tter-0: Done
Iter-1:
External potential: Yes
Calculating Potential of the 1th component (Frat disc)...Done (25.28 s)
***************
           START FITZPROFILE
***************
Number of Radii: 30
Number of Vertical points: 30
Number of the used distributions: 1
nplot 2
---Fitting---
Working on radius: 0.01
Working on radius: 1.04
Plotting
Working on radius: 2.08
Working on radius: 3.11
Plotting
Working on radius: 4.15
Working on radius: 5.18
```

```
Plotting
Working on radius: 6.21
Working on radius: 7.25
Plotting
Working on radius: 8.28
Working on radius: 9.32
Plotting
Working on radius: 10.35
Working on radius: 11.39
Plotting
Working on radius: 12.42
Working on radius: 13.45
Plotting
Working on radius: 14.49
Working on radius: 15.52
Plotting
Working on radius: 16.56
Working on radius: 17.59
Plotting
Working on radius: 18.62
Working on radius: 19.66
Plotting
Working on radius: 20.69
Working on radius: 21.73
Plotting
Working on radius: 22.76
Working on radius: 23.80
Plotting
Working on radius: 24.83
Working on radius: 25.86
Working on radius: 26.90
Working on radius: 27.93
Working on radius: 28.97
Working on radius: 30.00
Save figures
```

/Users/Giuliano/anaconda/envs/py36/lib/python3.6/site-packages/matplotlib/pyplot.py:524: Runtime max\_open\_warning, RuntimeWarning)

\*\*\*\*\*\*\*\*\*\*\*\*\*\*

```
***************
             START FITFLARE
*************
Start fitting
Writing table
Save table
Make plot
Save plot
data in gasHeight_res/diagnostic/run0/flare/fitflare_par.dat
image in gasHeight_res/diagnostic/run0/flare/flare.pdf
**************
             END FITFLARE
*************
Iter-1: Done
Max Absolute residual=1.28e+00
Max Relative residual=1.25e-01
Iter-2:
External potential: Yes
Calculating Potential of the 1th component (Frat disc)...Done (24.32 s)
***************
           START FITZPROFILE
*************
Number of Radii: 30
Number of Vertical points: 30
Number of the used distributions: 1
                           ['gau']
nplot 2
---Fitting---
Working on radius: 0.01
Working on radius: 1.04
Plotting
Working on radius: 2.08
Working on radius: 3.11
Plotting
Working on radius: 4.15
Working on radius: 5.18
Plotting
Working on radius: 6.21
Working on radius: 7.25
Plotting
Working on radius: 8.28
Working on radius: 9.32
Plotting
Working on radius: 10.35
```

Working on radius: 11.39

Plotting Working on radius: 12.42 Working on radius: 13.45 Plotting Working on radius: 14.49 Working on radius: 15.52 Plotting Working on radius: 16.56 Working on radius: 17.59 Plotting Working on radius: 18.62 Working on radius: 19.66 Plotting Working on radius: 20.69 Working on radius: 21.73 Plotting Working on radius: 22.76 Working on radius: 23.80 Plotting Working on radius: 24.83 Working on radius: 25.86 Working on radius: 26.90 Working on radius: 27.93 Working on radius: 28.97 Working on radius: 30.00 Save figures Writing table DONE in 0.088 minutes Output data files in gasHeight\_res/diagnostic/run1/dat Output images in gasHeight\_res/diagnostic/run1/image \*\*\*\*\*\*\*\*\*\*\*\*\*\* END FITZPROFILE \*\*\*\*\*\*\*\*\*\*\*\*\*\* \*\*\*\*\*\*\*\*\*\*\*\*\*\*\* START FITFLARE \*\*\*\*\*\*\*\*\*\*\*\*\* Start fitting Writing table Save table Make plot Save plot data in gasHeight\_res/diagnostic/run1/flare/fitflare\_par.dat image in gasHeight\_res/diagnostic/run1/flare/flare.pdf \*\*\*\*\*\*\*\*\*\*\*\*\*\*\* END FITFLARE \*\*\*\*\*\*\*\*\*\*\*\*\*\*\* Iter-2: Done

Max Absolute residual=5.39e-02

```
Max Relative residual=9.96e-03
Iter-3:
External potential: Yes
Calculating Potential of the 1th component (Frat disc)...Done (24.57 s)
***************
            START FITZPROFILE
*************
Number of Radii: 30
Number of Vertical points: 30
Number of the used distributions: 1
nplot 2
---Fitting---
Working on radius: 0.01
Working on radius: 1.04
Plotting
Working on radius: 2.08
Working on radius: 3.11
Plotting
Working on radius: 4.15
Working on radius: 5.18
Plotting
Working on radius: 6.21
Working on radius: 7.25
Plotting
Working on radius: 8.28
Working on radius: 9.32
Plotting
Working on radius: 10.35
Working on radius: 11.39
Plotting
Working on radius: 12.42
Working on radius: 13.45
Plotting
Working on radius: 14.49
Working on radius: 15.52
Plotting
Working on radius: 16.56
Working on radius: 17.59
Plotting
Working on radius: 18.62
Working on radius: 19.66
Plotting
Working on radius: 20.69
```

Working on radius: 21.73

```
Plotting
Working on radius: 22.76
Working on radius: 23.80
Plotting
Working on radius: 24.83
Working on radius: 25.86
Working on radius: 26.90
Working on radius: 27.93
Working on radius: 28.97
Working on radius: 30.00
Save figures
Writing table
DONE in 0.094 minutes
Output data files in gasHeight_res/diagnostic/run2/dat
Output images in gasHeight_res/diagnostic/run2/image
***************
            END FITZPROFILE
**************
***************
            START FITFLARE
*************
Start fitting
Writing table
Save table
Make plot
Save plot
data in gasHeight_res/diagnostic/run2/flare/fitflare_par.dat
image in gasHeight_res/diagnostic/run2/flare/flare.pdf
**************
             END FITFLARE
**************
Iter-3: Done
Max Absolute residual=9.68e-02
Max Relative residual=1.33e-02
Iter-4:
External potential: Yes
Calculating Potential of the 1th component (Frat disc)...Done (24.32 s)
*************
           START FITZPROFILE
*************
Number of Radii: 30
Number of Vertical points: 30
Number of the used distributions: 1
nplot 2
```

---Fitting---

Working on radius: 0.01 Working on radius: 1.04

Plotting

Working on radius: 2.08 Working on radius: 3.11

Plotting

Working on radius: 4.15 Working on radius: 5.18

Plotting

Working on radius: 6.21 Working on radius: 7.25

Plotting

Working on radius: 8.28 Working on radius: 9.32

Plotting

Working on radius: 10.35 Working on radius: 11.39

Plotting

Working on radius: 12.42 Working on radius: 13.45

Plotting

Working on radius: 14.49 Working on radius: 15.52

Plotting

Working on radius: 16.56 Working on radius: 17.59

Plotting

Working on radius: 18.62 Working on radius: 19.66

Plotting

Working on radius: 20.69 Working on radius: 21.73

Plotting

Working on radius: 22.76 Working on radius: 23.80

Plotting

Working on radius: 24.83
Working on radius: 25.86
Working on radius: 26.90
Working on radius: 27.93
Working on radius: 28.97
Working on radius: 30.00

Save figures Writing table

DONE in 0.084 minutes

Output data files in gasHeight\_res/diagnostic/run3/dat Output images in gasHeight\_res/diagnostic/run3/image

```
***************
           END FITZPROFILE
**************
*************
            START FITFLARE
*************
Start fitting
Writing table
Save table
Make plot
Save plot
data in gasHeight_res/diagnostic/run3/flare/fitflare_par.dat
image in gasHeight_res/diagnostic/run3/flare/flare.pdf
*************
            END FITFLARE
**************
Iter-4: Done
Max Absolute residual=1.33e-03
Max Relative residual=1.82e-04
Iter-5:
External potential: Yes
Calculating Potential of the 1th component (Frat disc)...Done (24.23 s)
*************
          START FITZPROFILE
**************
Number of Radii: 30
Number of Vertical points: 30
Number of the used distributions: 1
nplot 2
---Fitting---
Working on radius: 0.01
Working on radius: 1.04
Plotting
Working on radius: 2.08
Working on radius: 3.11
Plotting
Working on radius: 4.15
Working on radius: 5.18
Plotting
Working on radius: 6.21
Working on radius: 7.25
Plotting
Working on radius: 8.28
Working on radius: 9.32
```

```
Plotting
Working on radius: 10.35
Working on radius: 11.39
Plotting
Working on radius: 12.42
Working on radius: 13.45
Plotting
Working on radius: 14.49
Working on radius: 15.52
Plotting
Working on radius: 16.56
Working on radius: 17.59
Plotting
Working on radius: 18.62
Working on radius: 19.66
Plotting
Working on radius: 20.69
Working on radius: 21.73
Plotting
Working on radius: 22.76
Working on radius: 23.80
Plotting
Working on radius: 24.83
Working on radius: 25.86
Working on radius: 26.90
Working on radius: 27.93
Working on radius: 28.97
Working on radius: 30.00
Save figures
Writing table
DONE in 0.092 minutes
Output data files in gasHeight_res/diagnostic/run4/dat
Output images in gasHeight_res/diagnostic/run4/image
*************
              END FITZPROFILE
***************
***************
               START FITFLARE
*************
Start fitting
Writing table
Save table
Make plot
Save plot
data in gasHeight_res/diagnostic/run4/flare/fitflare_par.dat
image in gasHeight_res/diagnostic/run4/flare/flare.pdf
***************
```

END FITFLARE

```
**********

Iter-5: Done

Max Absolute residual=6.51e-05

Max Relative residual=8.95e-06
```

#### 1 Results of the functions:

0-final\_gas\_model: The final disc model, with the Radial surface density law given in inputr and the vertical profiles obtained in the iterative process

```
1-tab_zd: A tabel with 0-R [kpc] 1-Zd [kpc]
```

2-flare\_func: The interpolating function of tab\_zd, zd(R)=flare\_func(R)

3-fit\_func: The best-fit function (as defined with flaw) to the last zd estimate.

In the output folder you can find:

- -finalflare\_zd.pdf: a figure with the zd estimate at each iterative step (gray lines), the last estimate is shown by blue points and the red curve is the last best-fit function
  - -finalflare\_hwhm.pdf: The final zd estimate, but the value in y is the HWHM
  - -tabflare.dat: 0-Col R[kpc], 1-Col zd[kpc], 2-Col HWHM[kpc]
  - -tab\_fixedpotential.dat: Tab with the potentials of the fixed dynamic components
  - -tab\_totpotential.dat: Tab with the potential of the final disc component
  - -My suggestion is to use:

```
Rlimit='max'
flaw='poly'
polyflare_degree_degree=5
In [22]: ##An example of use: estimate of the scale height for the HI disc and H2 disc
         ##Fixed component
         ##halo
         #halo=dc.isothermal_halo(....)
         ##bulge
         #bulge=dc.hernquist_halo(....)
         ##stellar disc
         \#disc=dc.Exponential\_disc.thick(...)
         ##Observed intrinsic HI surface density
         \#HI\_tab = [RHI, Sigma\_HI]
         #HI_disc=dc.Frat_disc.thin(rfit_array=HI_tab,....)
         ##Observed intrinsic H2 surface density
         #HII_tab=[RHII,Siqma_HII]
         #HII_disc=dc.Frat_disc.thin(rfit_array=HII_tab,....)
         #galaxy=(halo,bulge,disc)
```

```
#h=discHeight(dynamic_components=galaxy, disc_component=HI_disc)
#HI_disc=h.height(....)[0]

##galaxy_new=(halo,bulge,disc,HI_disc)

#h=discHeight(dynamic_components=galaxy_new, disc_component=HII_disc)
#HII_disc=h.height(....)[0]
```

#### In []: