

Computer Graphics Cheatsheet

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January 26, 2022

Computer Graphics - SA. 2021
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Contents

1 Ray Tracing

Ray tracing is a method of graphics rendering that simulates the physical behaviour of light.

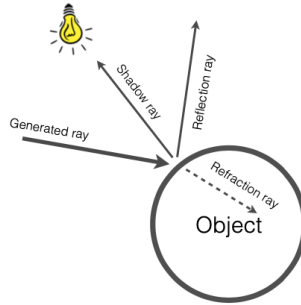
1.1 Whitted Ray Tracing

One ray is traced for every single pixel. These are known as the **primary rays**.

For each of the rays, we need to find their intersection with the scene. After we do so, secondary rays can be generated. Such rays are:

- Shadow Ray
- Reflection Ray
- Refraction Ray

The color of each of the pixels is determined based on the aggregated color of the rays.



The secondary rays are recursively generated. Some termination conditions for this generation are:

- Ray leaves the scene without ever hitting an object
- Maximal recursion depth is reached
- Contribution of the ray to the final color is negligible

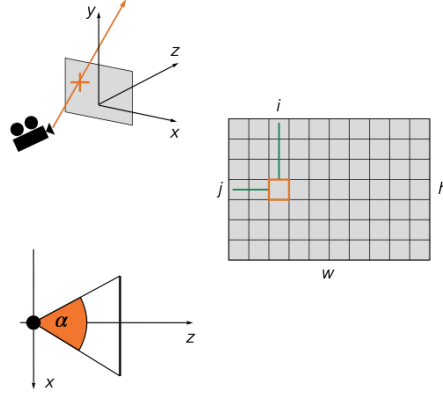
1.2 Camera / Image Definition

We are considering a global coordinate system for the raytracer. Some of the characteristics of the **camera** are:

- It is located at the origin $(0,0,0)$
- It has a horizontal opening α , the FOV (Field Of View)

Some of the characteristics of the **image** are:

- Image plane is located at $z = 1$
- Resolution of the image is of $w \cdot h$ pixels
- Pixels have indices (i, j) and coordinates $p_{ij} = (x_{ij}, y_{ij}, z_{ij})$

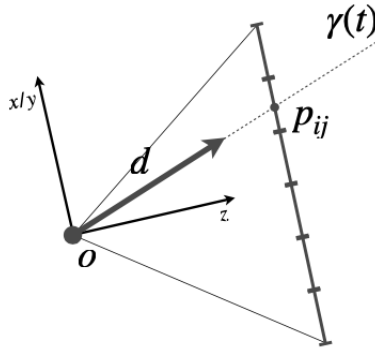


1.3 Ray Computation

The equation of a **ray** is the following:

$$\gamma(t) = o + dt$$

Where t is the distance between the origin and the intersection with an object, o is the point representing the camera position, and d is the vector representing the viewing direction – of magnitude 1.



In order to compute the vector d , we need to do the following:

```

1 s = (2 * tan(alpha / 2)) / 2;
2 X = (-w * s) / 2;
3 Y = (h * s) / 2;
4
5 for (int i = 0; i++; i < w) {
6     for (int j = 0; j++; j < h) {
7         dx = X + (i * s) + (0.5 * s);
8         dy = Y - (j * s) - (0.5 * s);
9         dz = 1;
10
11         d = glm::vec3(dx, dy, dz);
12         d = glm::vec3.normalize(d)
13     }
14 }

```

1.4 Ray-Sphere Intersection

In order to define a **sphere**, we need to set the centre and the radius. Once that is done, we need to check if the sphere intersects with the generated ray. To do so we need to check if there is some t such that:

$$\| \gamma(t) - c \| = r$$

In order to find the t , we use the following procedure.

$$\begin{aligned}
 a &= \langle c, d \rangle \\
 D &= \sqrt{\| c \|^2 - \langle c, d \rangle^2} \\
 t_{1,2} &= \langle c, d \rangle \pm \sqrt{r^2 - D^2}
 \end{aligned}$$

If the ray intersects the sphere and the camera is not inside of the sphere, then the pixel corresponding to the intersection is painted.

2 Lighting Models

In order to compute the color values on an object's surface we can use rules based on the laws of physics – radiometry and photometry. These rules model the effect of:

- **Light Sources**

Position, intensity and colour.

- **Object Surface**

Geometry and reflective properties.

2.1 Illumination Factors

The intensity of a color depends on several different factors:

- The colour and reflective properties of the object
- The position and intensity of the light sources
- The position of the viewer
- The normal of the surface at point p
- The distance of point p to the light source

2.2 Components of the Phong Lighting Model

These components are the **diffuse reflection**, the **ambient illumination**, and the **specular highlight**.

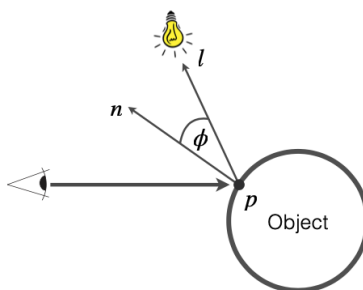
2.2.1 Diffuse Reflection

This type of reflection simulates Lambertian surfaces – i.e. matte surfaces. Here the reflected light is dispersed evenly in all directions.

There is also a material-dependent reflection constant $\rho_d \leq 1$. This is called the **diffuse coefficient**.



The intensity I_d of a surface coming from the diffuse reflection is proportional to the cosine of the angle between the surface normal n and the direction of the light source l . This is known as the **Lambertian Cosine Rule**.



The proper computation of the diffuse reflection is the following – in which the Cosine Rule is applied:

$$\begin{aligned} I_d &= \rho_d \cdot \cos \phi \cdot I \\ &= \rho_d \cdot \langle n, l \rangle \cdot I \end{aligned}$$

Where ρ_d is the diffuse coefficient, n is the normal of the surface at point p , l is the vector pointing towards the light source, and I is the intensity of the light.

An important thing to always check is that $\langle n, l \rangle > 0$.

2.2.2 Ambient Illumination

Ambient illumination is used in order to simulate indirect lighting. This refers to the effect of multiple inter-reflections of light between all objects, and to the isotropy and independence of light sources and viewpoints.

There is a global constant I , and a material-dependent reflection constant $\rho_a \leq 1$. This is called the **ambient coefficient**.

The ambient illumination is given by the following formula:

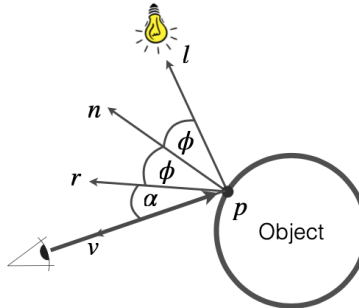
$$I_a = \rho_a \cdot I$$

Where ρ_a is the ambient coefficient, and I is the intensity of the ambient lighting.

2.2.3 Specular Reflection

Specular reflection is used in order to simulate shiny surfaces. The incoming light is reflected in exactly one reflect direction.

The specular reflection makes use of something known as the reflection vector.



The proper formulation of the specular reflection is the following:

$$I_s = \rho_s \cdot \langle r, v \rangle^k \cdot I$$

Where ρ_s is the specular coefficient of the object, r is the reflection ray given by:

$$r = 2n \cdot \langle n, l \rangle - l$$

Furthermore, v is the viewing direction of the camera, $k \geq 1$ is the shininess coefficient, and I is the light intensity.

2.3 Phong Lighting Model

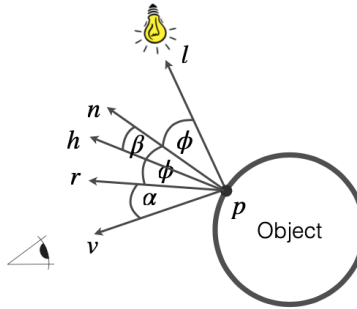
The Phong lighting model is given by the superposition of ambient, diffuse, and specular terms for each light source. The final formula is the following:

$$I = I_e + \rho_a \cdot I_a + \sum_{j=1}^n (\rho_d \cdot \langle n, l_j \rangle + \rho_s \cdot \langle r_j, v \rangle^k) \cdot I_j$$

Where I_e is the self-emitting intensity, ρ_a is the diffuse coefficient, I_a is the ambient intensity, ρ_d is the diffuse coefficient, n is the normal vector, l is the vector pointing to the light source, ρ_s is the specular coefficient, r is the reflection ray, v is the direction vector, k is the shininess, and I_j is the intensity of the j -th light source.

2.4 Blinn-Phong Specular Reflection

Here we use h instead of r in order to compute the specular reflection of an object.



The proper formulation of the Blinn-Phong specular reflection is the following:

$$I_s = \rho_s \cdot \langle n, h \rangle^{4k} \cdot I$$

Where ρ_s is the specular coefficient, n is the normal, k is the shininess coefficient, I is the light intensity, and h is defined as:

$$\frac{1}{2}(l + v)$$

Where l is the light vector, and v is the viewing direction.

2.5 Light Sources

2.5.1 Point Light Sources

Point light sources are a type of source that is isotropic – radiates evenly in all directions, it is specified by a position, and has intensity I .

Directional light sources are a special case of light sources which are only specified by direction. This is because it is an infinite set of rays that all have the same direction.

2.5.2 Spot Light Sources

This type of light source generates a light cone in a specified direction. In order to define the spot light source, we need to specify its position p , its direction d , and its opening angle Θ_L .

The intensity of the light is either maximal in the direction of d , or decreases following the formula:

$$I'(\Theta) = \cos^k \Theta \cdot I$$

And it's 0 when $\Theta > \Theta_L$.

2.6 Distance Attenuation

The more the object is distant from the light source, the less it will be illuminated by the source. Its formal definition is:

$$att(r) = \frac{1}{a_1 + a_2 r + a_3 r^2}$$

Where r is the ray of light, and a_1 , a_2 , and a_3 are constant values.