

Should geophysicists use gravity anomaly or gravity disturbance?

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SUMMARY

In applied geophysics, the gravity anomaly has long been used for representing the gravitational effect produced by gravity sources, which are defined as the density contrasts between the internal density distribution of the Earth and the normal Earth. The gravity anomaly is defined by the difference between the gravity (produced by the Earth) on the Geoid and the normal gravity (produced by the normal Earth) on reference ellipsoid, both at the same geodetic latitude and longitude. Because these quantities are not defined at the same point, the gravity anomaly does not represent the gravitational effect due to the gravity sources and, consequently, cannot be considered a harmonic function. On the contrary, the gravity disturbance is defined by the difference between the gravity and normal gravity at the same point. Consequently, the centrifugal effects can be neglected, the gravity disturbance can be considered a harmonic function representing the gravitational effect due to the gravity sources and then it satisfies the premise behind most of the processing techniques of potential-field data (e.g., upward/downward continuations). Then, rigorously, the gravity disturbance is more appropriated to be used for estimating density variations in subsurface. The use the gravity anomaly for most of the geophysicists carries the implicit meaning that it can be considered a good approximation of the gravity disturbance. It is important to bear in mind, however, that in Brazil, for example, the maximum absolute difference between the gravity disturbance (on or close to the Earth's surface) and the Free-air anomaly reaches 9.3 mGal.

Key words: potential fields – gravity disturbance – gravity anomaly – gravity modelling.

1 INTRODUCTION

The resultant of gravitational force and centrifugal force acting on a body at rest on the Earth's surface is called *gravity vector* and its intensity is called *gravity* (Heiskanen & Moritz 1967; Hofmann-Wellenhof & Moritz 2005). In the case of gravimetry on moving platforms (e.g., airplanes, helicopters, marine vessels), there are additional non-gravitational accelerations due to the vehicle motion, such as Coriolis acceleration and high-frequency vibrations (Glenie et al. 2000; Nabighian et al. 2005; Baumann et al. 2012). Geophysicists use gravity for estimating the Earth's internal density distribution whereas geodesists use gravity to estimate the Geoid (Li & Götze 2001). Hence, geophysicists are usually interested in the gravitational component of the observed gravity, which is produced by the Earth's internal density distribution.

For isolating this gravitational component consists in removing the non-gravitational effects due to the vehicle motion and also the time variations such as Earth tides, instrumental drift and barometric pressure changes, for example. If these effects are properly removed, the resultant gravity data can be considered as the sum of a centrifugal component due to the Earth's rotation and a gravi-

tational component produced by the whole Earth's internal density distribution. The isolation of this particular gravitational component and its subsequent use for estimating density distributions related to geological structures in subsurface are the main goals in applied geophysics (Blakely 1996). The computation of the gravitational effect produced by the geological structures is commonly known in geophysics as *gravity modelling*. This term is different from *gravity field modelling*, which is commonly used in geodesy to denote the characterization of the gravity field in a local, regional or global scales.

Based on well-established concepts of the literature, we present a discussion aiming at critically examining the following question: in geophysical applications, should we use the gravity disturbance or gravity anomaly? It seems that this theoretical issue has been debated within the scientific community from a more geodetic than geophysical point of view (LaFehr 1991; Chapin 1996; Li & Götze 2001; Fairhead et al. 2003; Hackney & Featherstone 2003; Hinze et al. 2005). Our reasoning suggests that the gravity disturbance is more appropriated than the gravity anomaly for geophysical purposes.

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2 NORMAL EARTH AND NORMAL GRAVITY

Traditionally, the Earth's gravity field is approximated by the gravity field produced by a reference ellipsoid (or level ellipsoid), which is a rigid and geocentric model. This reference ellipsoid has the minor axis b coincident with the mean rotating axis of the Earth Z , the same total mass (including the atmosphere) and also the same angular velocity of the Earth (Heiskanen & Moritz 1967; Vaníček & Krakiwsky 1987; Hofmann-Wellenhof & Moritz 2005; Torge & Müller 2012). Another characteristic of this model is that its limiting surface coincides with a particular equipotential of its own gravity field. Here, we follow (Torge & Müller 2012) and call this model as *normal Earth*.

The resultant of the virtual gravitational and centrifugal forces exerted by the normal Earth on a body at rest at a point P is called *normal gravity vector* and its intensity is called *normal gravity*. In geodesy, any model used to represent the normal gravity field can be arbitrarily defined for the only purpose of keeping the difference from the actual gravity field as small as possible (Vaníček & Krakiwsky 1987). It is worth noting that, although the normal Earth has the same total mass (including the atmosphere) of the Earth, its internal density distribution is unknown. The search for physically meaningful mass distributions that generate a required normal gravity field has geophysical rather than geodetic motives (Marussi et al. 1974). The only condition imposed on its internal density distribution is that it produces a gravity field having a particular equipotential which coincides with its limiting surface. For convenience, we denote any density distribution satisfying this condition as a *normal density distribution*.

3 TERRESTRIAL REFERENCE SYSTEMS USED IN GRAVITY MODELLING

For geophysical purposes, there are three important Terrestrial Reference Systems used in gravity modelling. They rotate with the Earth and are used for describing positions and movements of objects on and close to the Earth's surface (Torge & Müller 2012).

The first one is a geocentric system of Cartesian coordinates having the Z -axis coincident with the mean Earth's rotational axis, the X -axis pointing to the Greenwich meridian and the Y -axis directed so as to obtain a right-handed system (Fig. A1). This reference system can be found in the literature with different names: Mean Terrestrial System (e.g., Soler 1976), Earth-fixed geocentric Cartesian system (e.g., Torge & Müller 2012) or Earth-centered Earth-fixed system (e.g., Bouman et al. 2013), for example. Here, we opted for simply using the term Geocentric Cartesian System (GCS).

Another important reference system is a geocentric system of geodetic coordinates, which is defined by the reference ellipsoid used in the Normal Earth model (Heiskanen & Moritz 1967; Soler 1976; Torge & Müller 2012; Bouman et al. 2013). In this coordinate system, the position of a point is defined by the *geometric height* h , *geodetic latitude* φ and *longitude* λ (Fig. A1). For convenience, we call this system Geocentric Geodetic System (GGS). At a given point (h, φ, λ) , there are three mutually-orthogonal unit vectors (Fig. A1) given by (Soler 1976):

$$\hat{\mathbf{u}} = \begin{bmatrix} \cos \varphi \cos \lambda \\ \cos \varphi \sin \lambda \\ \sin \varphi \end{bmatrix} \quad \hat{\mathbf{v}} = \begin{bmatrix} -\sin \varphi \cos \lambda \\ -\sin \varphi \sin \lambda \\ \cos \varphi \end{bmatrix} \quad \hat{\mathbf{w}} = \begin{bmatrix} -\sin \lambda \\ \cos \lambda \\ 0 \end{bmatrix}. \quad (1)$$

The required equations to convert coordinates (h, φ, λ) referred to

the GGS into coordinates (X, Y, Z) referred to the GCS (Fig. A1) and vice versa can be easily found in the literature (e.g., Heiskanen & Moritz 1967; Torge & Müller 2012; Bouman et al. 2013).

In local or regional studies, geophysicists commonly use a topocentric Cartesian coordinate system (TCS) with origin at a point P on or close to the Earth's surface and axes x, y and z (Fig. A2). In the TCS, the axes x and y are parallel to the unit vectors $\hat{\mathbf{v}}_P$ and $\hat{\mathbf{u}}_P$, respectively, whereas the z -axis is opposite to the unit vector $\hat{\mathbf{u}}_P$ (eq. 1) and points downward. Consider, for example, a TCS (Fig. A2) with origin at a point P with coordinates (X_P, Y_P, Z_P) referred to the GCS (Fig. A1). In this case, the relationship between the coordinates (X, Y, Z) of a point in the GCS and the coordinates (x, y, z) of the same point in the TCS is given by:

$$\begin{bmatrix} x \\ y \\ z \end{bmatrix} = \mathbf{R}^T \begin{bmatrix} X - X_P \\ Y - Y_P \\ Z - Z_P \end{bmatrix}, \quad (2)$$

where \mathbf{R} is a 3×3 orthogonal matrix whose first, second and third columns are defined, respectively, by the unit vectors $\hat{\mathbf{v}}$, $\hat{\mathbf{w}}$ and $-\hat{\mathbf{u}}$ (eq. 1), evaluated at the point P .

4 RELATIONSHIP BETWEEN GRAVITY DISTURBANCE AND GRAVITY ANOMALY

Let \mathbf{g}_P and γ_P be, respectively, the gravity vector (corrected from non-gravitational effects due to vehicle motion and time variations such as Earth tides, instrumental drift and barometric pressure changes, for instance) and the normal gravity vector at a point P . In this case, the gravity vector \mathbf{g}_P represents the gradient of a scalar potential called *gravity potential*, which is the sum of a scalar gravitational potential and a scalar centrifugal potential. Similarly, the normal gravity vector γ_P represents the gradient of a scalar potential called *normal potential*, which is also the sum of a scalar gravitational potential and a scalar centrifugal potential. By definition, the centrifugal part of the normal potential is equal to that of the gravity potential.

The difference between \mathbf{g}_P and γ_P , at the same point P , defines a quantity called *gravity disturbance vector*, which is given by:

$$\delta \mathbf{g}_P = \mathbf{g}_P - \gamma_P. \quad (3)$$

Because the centrifugal part of the normal gravity vector is equal to the centrifugal component of the gravity vector, the gravity disturbance vector $\delta \mathbf{g}_P$ represents a purely gravitational effect caused by contrasts between the actual internal density distribution of the Earth and the internal density distribution of the normal Earth. ergo, the gravity disturbance vector $\delta \mathbf{g}_P$ (eq. 3) is harmonic. It is important to stress that both density distributions are unknown. In applied geophysics, these density differences are generally called *anomalous masses* (e.g., Hammer, 1945; LaFehr, 1965), *density anomalies* (e.g., Forsberg, 1984) or *gravity sources* (e.g., Blakely, 1996). Here, we opted for using the last term. Fig. A3 illustrates the gravity vector \mathbf{g}_P , normal gravity vector γ_P and the gravity disturbance vector $\delta \mathbf{g}_P$ at a point P located on the Earth's surface.

The difference between the magnitudes of gravity vector $g_P = \|\mathbf{g}_P\|$ and the normal gravity vector $\gamma_P = \|\gamma_P\|$, at the same point P , is called *gravity disturbance* (Heiskanen & Moritz 1967; Hofmann-Wellenhof & Moritz 2005) and can be represented as follows:

$$\delta g_P = g_P - \gamma_P. \quad (4)$$

Notice that the gravity disturbance δg_P is not equivalent to the magnitude of the gravity disturbance vector $\delta \mathbf{g}_P$ (Barthelmes 2013; Sansò & Sideris 2013).

On the other hand, the *gravity anomaly vector* at P is the difference between the gravity vector at a point Q' on the Geoid (a particular equipotential surface of the gravity potential) and the normal gravity vector at a point Q on the ellipsoid surface, both at the same geodetic latitude and longitude (Fig. A3) as in

$$\Delta \mathbf{g}_P = \mathbf{g}_{Q'} - \gamma_Q. \quad (5)$$

Similarly to the gravity disturbance expression (eq. 4), the gravity anomaly is given by

$$\Delta g_P = g_{Q'} - \gamma_Q, \quad (6)$$

where $g_{Q'} = \|\mathbf{g}_{Q'}\|$ is the magnitude of the gravity vector $\mathbf{g}_{Q'}$ on the Geoid, at Q' , and $\gamma_Q = \|\gamma_Q\|$ is the magnitude of the normal gravity vector γ_Q on the reference ellipsoid, at the point Q (Fig. A3). As properly pointed out by Barthelmes (2013) the gravity anomaly, by definition, depends on longitude and latitude only and is not a function of height. Because the gravity is on the Geoid and the normal gravity is on reference ellipsoid, the gravity anomaly Δg_P (eq. 6) is not a purely gravitational effect due to the gravity sources, but rather a combination of gravitational and centrifugal effects. Consequently, its upward continuation, for example, could not be rigorously done.

However, many authors in the literature have computed the upward continuation of gravity anomalies. All these authors have implicitly considered that the gravity anomaly is an approximation of the gravity disturbance, which in turn can be represented by a harmonic function and can then being upward continued.

Different gravity anomalies can be calculated, depending on the corrections applied to them. These corrections are usually called *gravity reductions*. The Free-air anomaly, for example, may be defined as follows (Blakely 1996; Hofmann-Wellenhof & Moritz 2005):

$$\Delta g_P^F = g_P - \left(\gamma_Q + \frac{\partial \gamma}{\partial h} H_P \right), \quad (7)$$

where $\frac{\partial \gamma}{\partial h} \approx -0.3086$ mGal/m is the derivative of the normal gravity with respect to the geometric height h and H_P is the orthometric height H (Fig. A3) at the point P . Notice that the term between parenthesis represents an approximation of the normal gravity at the point Q' on the Geoid. By using a similar approach, we can approximate the normal gravity γ_P , at a point P on the Earth's surface, so that the gravity disturbance δg_P (eq. 4) can be rewritten as follows:

$$\delta g_P \approx g_P - \left(\gamma_Q + \frac{\partial \gamma}{\partial h} h \right). \quad (8)$$

By inspection, one easily concludes that the absolute difference between the approximated gravity disturbance δg_P (eq. 8), at a point P on or close to the Earth's surface, and the Free-air anomaly Δg_P^F (eq. 7), is given by:

$$|\delta g_P - \Delta g_P^F| \approx \left| \frac{\partial \gamma}{\partial h} N \right|, \quad (9)$$

where $N \approx h - H$ is the geoidal undulation (Fig. A3). This approximation assumes that the Geoid and the surface of the reference ellipsoid are approximately parallel at P and the surrounding area. We know empirically that the geoidal undulation N in the world is $\approx \pm 1$ m on the oceans and reaches a maximum absolute value of ≈ 120 m (e.g., Torge & Müller 2012; Sansò & Sideris 2013).

In Brazil, for example, the geoidal undulation reaches $\approx \pm 30$ m (IBGE 2015). Consequently, the maximum absolute differences between gravity disturbance and Free-air anomaly, according to eq. 9, is ≈ 9.258 mGal. The approximations defined by eqs 8 and 9 are commonly used in geodesy to define the *fundamental equation of physical geodesy* (Hofmann-Wellenhof & Moritz 2005).

The simple Bouguer anomaly is commonly used by geophysicists as the gravitational effect produced by the gravity sources. It is defined, over the continents, as follows (Blakely 1996; Hofmann-Wellenhof & Moritz 2005):

$$\Delta g_P^B = \Delta g_P^F - 2\pi G \rho_t H_P, \quad (10)$$

where Δg_P^F is the Free-air anomaly (eq. 7) and the last term on the right side is the simple Bouguer correction. This term approximates the gravitational attraction that all topographic masses above the Geoid exerts at the point P by the gravitational attraction of a homogeneous, infinitely extended slab of constant density ρ_t and thickness equal to H_P , which is the orthometric height H (Fig. A3) at the point P . By computing the simple Bouguer correction at a set of points of a survey around P , we approximately remove, from the gravity measured at the points of the survey, the gravitational effect of a homogeneous model representing the topographic masses above the Geoid. It is evident from eqs 7 to 9 that the simple Bouguer anomaly Δg_P^B (eq. 10) approximates the gravity disturbance δg_P (eq. 4) minus the gravitational effect produced by this homogeneous topographic model. Although this approximation is valid for most practical applications, it is important to bear in mind not only the terminology changes, but also the conceptual assumptions.

There was a certain lack of comprehension regarding the geophysical meaning of gravity anomalies until the mid 90's. As properly pointed out by Chapin (1996) at that time, "although the corrections which bring about a Bouguer gravity anomaly are well established, the reasons for doing them are not well understood. One cause of this common misunderstanding is that the subject has been poorly presented in many of the basic texts". In his seminal book, Blakely (1996) brought some light on the geophysical meaning of gravity anomalies from the perspective of applied geophysics. Blakely (1996) correctly defined gravity sources as density contrasts between the actual internal density distribution of the Earth and the internal density distribution of the normal Earth. However, he did not stress that, by removing the normal gravity evaluated on the ellipsoid from the gravity measured on the Earth's surface, the remaining anomalous field will reflect not only the effect produced by the gravity sources, but also a small combination of gravitational and centrifugal effects. This additional, non-harmonic and undesired effect is simply due to the calculation of the normal gravity at a point other than that where the gravity is measured.

5 MATHEMATICAL DESCRIPTION OF THE GRAVITY DISTURBANCE AT THE TCS

From eq. 3, the gravity vector \mathbf{g}_i , at a point (x_i, y_i, z_i) in the TCS (Fig. A2a), can be represented by:

$$\mathbf{g}_i = \gamma_i + \delta \mathbf{g}_i, \quad (11)$$

where γ_i and $\delta \mathbf{g}_i$ are, respectively, the normal gravity vector and the gravity disturbance vector produced by the anomalous masses at the point (x_i, y_i, z_i) .

The gravity \mathbf{g}_i can be approximated by a first order Taylor's

expansion as follows (Sansò & Sideris 2013):

$$g_i \approx \gamma_i + \hat{\gamma}_i^\top \delta \mathbf{g}_i, \quad (12)$$

where $^\top$ denotes transposition, $\hat{\gamma}_i = -\hat{\mathbf{u}}_i$ is a unit vector with the same direction as the normal gravity vector γ_i at the point (x_i, y_i, z_i) , in the topocentric Cartesian coordinate system (Fig. A2a), and $\hat{\mathbf{u}}_i$ is the unit vector $\hat{\mathbf{u}}$ (eq. 1) evaluated at the corresponding point $(h_i, \varphi_i, \lambda_i)$ in the GGS (Fig. A1). This approximation can be made because, fortunately, the condition $\gamma_i \gg \|\delta \mathbf{g}_i\|$ is met at all points located above or on the Earth's surface. This approximation is known in geodesy (e.g., Sansò & Sideris 2013) and a similar approximation is largely used in applied geophysics for representing total-field anomalies (e.g., Blakely 1996). Notice that, in local- or regional-gravity studies, the unit vector $\hat{\gamma}_i$ (eq. 12) may be considered constant throughout the study area and parallel to the z axis of the TCS (Fig. A2a). By using the approximation defined in eq. 12, the gravity disturbance (eq. 4) can be rewritten as follows

$$\delta g_i \approx \hat{\gamma}_P^\top \delta \mathbf{g}_i, \quad (13)$$

where $\hat{\gamma}_P$ represents the unit vector with the same direction as the normal gravity vector γ_P at the origin P of the TCS (Fig. A2a). This equation shows that the gravity disturbance δg_i (eq. 4) is different from the magnitude of the gravity disturbance vector $\delta \mathbf{g}_i$ produced by the gravity sources. Rather, it represents the component of $\delta \mathbf{g}_i$ on the direction of the normal gravity vector (Hofmann-Wellenhof & Moritz 2005; Sansò & Sideris 2013).

In the TCS (Fig. A2a), the gravity disturbance δg_i (eq. 13) can be defined as the z -component of the gravitational attraction exerted by the gravity sources at the point (x_i, y_i, z_i) . As a consequence, the gravity disturbance produced by a gravity source can be represented by the following harmonic function:

$$d_i = \int \int \int_v \frac{c_g G \Delta \rho(x', y', z') (z_i - z') dv'}{\sqrt{(x_i - x')^2 + (y_i - y')^2 + (z_i - z')^2}}, \quad (14)$$

where G is the Newtonian constant of gravitation (in $m^3 kg^{-1} s^{-2}$), $c_g = 10^5$ is a constant factor transforming from $m s^{-2}$ to milligal (mGal), $\Delta \rho(x', y', z')$ is the density contrast (in $kg m^{-3}$) at a point (x', y', z') within the volume v of the gravity source and the integration is conducted over x' , y' and z' . This equation can be easily generalized for the case of multiple gravity sources. Practically, all articles accomplishing gravity modelling use the quantity d_i (eq. 14) to represent the gravity anomaly produced by gravity sources (e.g., Blakely 1996). Consequently, almost all geophysicists use the gravity anomaly as an approximation of the gravity disturbance produced by the gravity sources. Notice that d_i (eq. 14) does not depend on the height with respect to the Geoid (orthometric height). Rather, it depends on the relative position of the observation points (x_i, y_i, z_i) with respect to the gravity sources in the TCS (Fig. A2a). Finally, it is important to notice that, contrary to what is written in some basic texts, gravimeters do not measure the quantity d_i (eq. 14).

6 CONCLUSIONS

Although the gravity disturbance is a well-known quantity in geodesy, its use in the applied geophysics is uncommon. For years, geophysicists have been using gravity anomaly for modelling the gravitational effects due to geological bodies. The question "Should geophysicists use gravity anomaly or gravity disturbance for modelling the gravitational effect of gravity sources?" is of more than academic interest.

Theoretically, the gravity anomaly is the difference between the gravity on the Geoid and the normal gravity on the reference ellipsoid, both at the same geodetic latitude and longitude. On the other hand, the gravity disturbance is defined by the difference between gravity and normal gravity at the same point. Because these quantities are at the same point, the gravity disturbance is a purely gravitational effect due to the gravity sources whereas the gravity anomaly is a combination of gravitational and centrifugal effects. Consequently, the gravity disturbance is a harmonic function whereas the gravity anomaly is not. Because the gravity disturbance is a purely gravitational effect, it can be represented by a harmonic function and satisfies the premise required by the most techniques for processing potential-field data (e.g., upward/downward continuation, data processing with equivalent layer, conversions between gravity and magnetic data, computation of vertical derivatives via Fourier and Hilbert transforms). Hence, the gravity disturbance is the more appropriated quantity for representing the gravity effect produced by the gravity sources.

In practical problems of applied geophysics, geophysicists perform the gravity-forward modelling by computing the vertical component of the gravitational attraction exerted by the sources at the observation points, regardless of the Geoid or reference ellipsoid positions, for example. This means that, actually, they compute gravity disturbance to fit a gravity anomaly (e.g., Bouguer anomaly) and this is a clear inconsistency. Moreover, upward or downward continuations, for example, are commonly applied to gravity anomalies. In this case, the gravity anomaly is wrongly considered a harmonic function and this is another inconsistency.

By adopting these practical procedures, geophysicists are implicitly taking the gravity anomaly as a good approximation of the gravity disturbance. But why should geophysicists care about this usual procedure? Because it may result in wrong interpretations of quantitative nature. In Brazil, for example, the maximum difference between the gravity disturbance and the free-air anomaly reaches 9.3 mGal. Such difference on the Earth's surface may lead, for example, to an error of about xxx m in the depth-to-center estimate of a sphere-like source with true depth-to-center at xxx m, radius xxxx m and density contrast of $xxx kg m^{-3}$.

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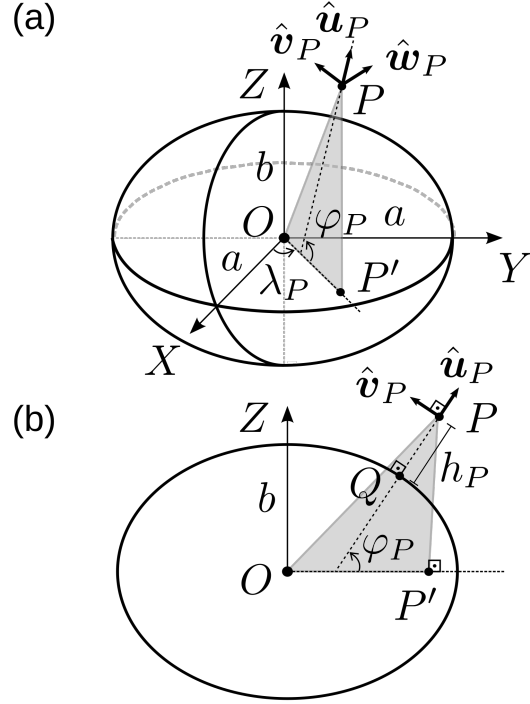


Figure A1. Schematic representation of the Geocentric Cartesian System (GCS) and the Geocentric Geodetic System (GGG). The GCS has the Z -axis coincident with the mean Earth's rotational axis, the X -axis pointing to the Greenwich meridian and the Y -axis directed so as to obtain a right-handed system. The GGG is defined by an oblate ellipsoid with semi-minor axis b , coincident with the Z -axis of GCS, and a semi-major axis a . In this coordinate system, the position of a point is determined by the geometric height h , geodetic latitude φ and longitude λ . The Earth's center of mass is represented by O , P represents a point $(h_P, \varphi_P, \lambda_P)$ and P' its projection onto the plane XY (Equatorial plane). The plane containing O , P and P' is represented in gray in (a) and (b). The unit vectors \hat{u}_P , \hat{v}_P and \hat{w}_P define three mutually orthogonal directions at P (eq. 1). In (b), Q represents a point $(h_Q, \varphi_Q, \lambda_Q)$, which is the projection of P onto the reference ellipsoid, at the same latitude and longitude ($\varphi_Q = \varphi_P$ and $\lambda_Q = \lambda_P$).

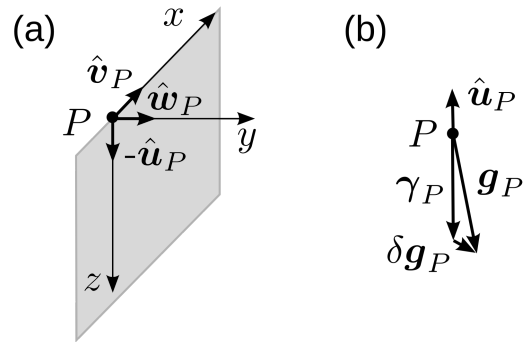


Figure A2. (a) Schematic representation of a topocentric Cartesian coordinate system (TCS) with origin at a point P . The axes x and y are parallel to the unit vectors \hat{v}_P and \hat{w}_P (eq. 1 and Fig. A1), respectively. On the other hand, the z axis is opposite to the unit vector \hat{u}_P (eq. 1 and Fig. A1) and points downward. The gray plane is the same shown in Fig. A1. (b) Schematic representation of the gravity vector \mathbf{g}_P , normal gravity vector γ_P , gravity disturbance vector $\delta \mathbf{g}_P$ (eq. 3) and unit vector \hat{u}_P (eq. 1) at a point P .

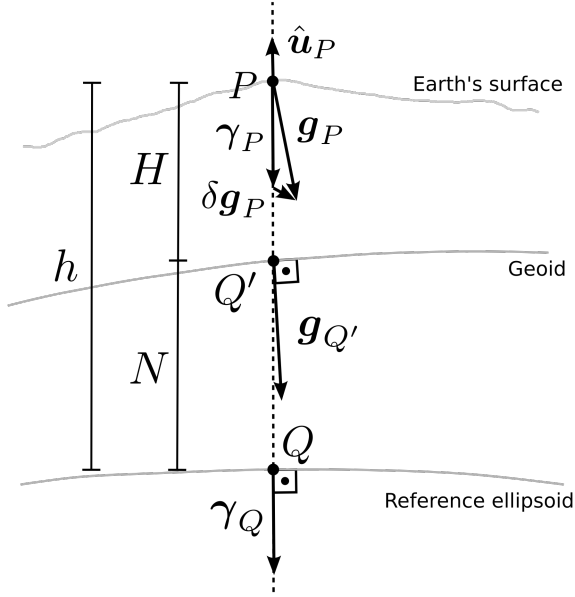


Figure A3. Schematic representation of the gravity vector \mathbf{g}_P , normal gravity vector γ_P , gravity disturbance vector $\delta\mathbf{g}_P$ (eq. 3), unit vector $\hat{\mathbf{u}}_P$ (eq. 1) at a point P on the surface of the Earth, gravity vector $\mathbf{g}_{Q'}$ at a point Q' on the Geoid, normal gravity vector γ_Q at a point Q on the reference ellipsoid, geometric height h , orthometric height H and geoidal undulation N (Heiskanen & Moritz 1967). The dashed line passing through Q , Q' and P is normal to the surface of the reference ellipsoid at Q . This figure shows a commonly used approximation in which the ellipsoid surface and the Geoid are represented as parallel surfaces, so that $h \approx H + N$.