

# IGIMF NOTES

Based on Yan, Jerabkova, & Kroupa (2021) A&A

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## 1 Method

$$\begin{aligned}
 \delta t &= 10^7 \text{ yr} \\
 Z_{\odot} &= 0.0142 \\
 \Delta\alpha &= 63 \\
 m_{\text{max}*} &= 150 M_{\odot} \\
 M_{\text{ecl}, \text{max}*} &= 10^9 M_{\odot} \quad (10^{10} \text{ might be better}) \\
 M_{\text{ecl}, \text{min}} &= 5 M_{\odot}
 \end{aligned}$$

$$\begin{aligned}
 \bar{\psi}_{\delta t} &= \text{const?} \equiv 2 M_{\odot} / \text{yr?} \\
 Z(t) &= \frac{M_Z(t)}{M_{\text{gas}}(t)} \\
 [Z](t) &= \log_{10} (Z(t) / Z_{\odot})
 \end{aligned}$$

Parameters to compute the stellar IMF

$$\begin{aligned}
 \alpha_1 &= 1.3 + \Delta\alpha \cdot (Z - Z_{\odot}), \\
 \alpha_2 &= 2.3 + \Delta\alpha \cdot (Z - Z_{\odot}),
 \end{aligned} \tag{1}$$

$M_{\text{ecl}}$ -dependent exponent

$$\alpha_3 = \begin{cases} 2.3, & x < -0.87, \\ -0.41x + 1.94, & x > -0.87. \end{cases} \tag{2}$$

$$x = -0.14[Z] + 0.99 \log_{10}(\rho_{\text{cl}} / 10^6) \tag{3}$$

$$\log_{10} \rho_{\text{cl}} = 0.61 \log_{10} M_{\text{ecl}} + 2.85, \tag{4}$$

Stellar IMF for individual embedded clusters

$$\xi_{\star}(m, M_{\text{ecl}}, Z) = dN_{\star} / dm = \begin{cases} 2k_{\star} m^{-\alpha_1(Z)}, & 0.08 \leq m / M_{\odot} < 0.50, \\ k_{\star} m^{-\alpha_2(Z)}, & 0.50 \leq m / M_{\odot} < 1.00, \\ k_{\star} m^{-\alpha_3(Z, M_{\text{ecl}})}, & 1.00 \leq m / M_{\odot} < m_{\text{max}}, \end{cases} \tag{5}$$

extract  $k_\star$  and  $m_{\max}$

$$M_{\text{ecl}} = \int_{0.08 \text{ M}_\odot}^{m_{\max}} m \xi_\star(m) dm, \quad (6)$$

$$1 = \int_{m_{\max}}^{m_{\max}^*} \xi_\star(m) dm, \quad (7)$$

Parameter to compute the ECMF

$$\beta = -0.106 \log_{10} \bar{\psi}_{\delta t} + 2. \quad (8)$$

Embedded cluster mass function

$$\xi_{\text{ecl}} = dN_{\text{ecl}}/dM_{\text{ecl}} = k_{\text{ecl}} M_{\text{ecl}}^{-\beta}, \quad 5M_\odot \leq M_{\text{ecl}} < M_{\text{ecl},\max}. \quad (9)$$

extract  $k_{\text{ecl}}$  and  $M_{\text{ecl},\max}$

$$M_{\text{tot}} = \int_{M_{\text{ecl},\min}}^{M_{\text{ecl},\max}} M_{\text{ecl}} \xi_{\text{ecl}}(M_{\text{ecl}}) dM_{\text{ecl}} = \bar{\psi}_{\delta t} \delta t, \quad M_{\text{tot}} - \int_{M_{\text{ecl},\min}}^{M_{\text{ecl},\max}} M_{\text{ecl}} k_{\text{ecl}}(M_{\text{ecl},\max}) M_{\text{ecl}}^{-\beta} dM_{\text{ecl}} = \bar{\psi}_{\delta t} \delta t = 0 \quad (10)$$

$$1 = \int_{M_{\text{ecl},\max}}^{M_{\text{ecl},\max}^*} k_{\text{ecl}} M_{\text{ecl}}^{-\beta} dM_{\text{ecl}} \quad (11)$$

$$k_{\text{ecl}}(M_{\text{ecl},\max}) = \frac{1}{\int_{M_{\text{ecl},\max}}^{M_{\text{ecl},\max}^*} M_{\text{ecl}}^{-\beta} dM_{\text{ecl}}} \quad (12)$$

$$(13)$$

$$k_{\text{ecl}} = \frac{1 - \beta}{M_{\text{ecl}}^{1-\beta} \big|_{M_{\text{ecl},\max}}^{M_{\text{ecl},\max}^*}} \quad (14)$$

$$\frac{M_{\text{ecl}}^{2-\beta} \big|_{M_{\text{ecl},\min}}^{M_{\text{ecl},\max}}}{M_{\text{ecl}}^{1-\beta} \big|_{M_{\text{ecl},\max}}^{M_{\text{ecl},\max}^*}} = \frac{2 - \beta}{1 - \beta} \bar{\psi}_{\delta t} \delta t \quad (15)$$

$$\frac{M_{\text{ecl}}^{2-\beta} \big|_{5M_\odot}^{M_{\text{ecl},\max}}}{M_{\text{ecl}}^{1-\beta} \big|_{M_{\text{ecl},\max}}^{10^9 M_\odot}} = \frac{2 - \beta}{1 - \beta} \bar{\psi}_{\delta t} \delta t \quad (16)$$

$$\frac{M_{\text{ecl},\max}^{2-\beta} - 5M_\odot^{2-\beta}}{M_{\text{ecl},\max}^{1-\beta} - 10^9 M_\odot^{1-\beta}} = \frac{2 - \beta}{1 - \beta} \bar{\psi}_{\delta t} \delta t \quad (17)$$

$$M_{\text{ecl},\max}^{2-\beta} - M_{\text{ecl},\max}^{1-\beta} \frac{2 - \beta}{1 - \beta} \bar{\psi}_{\delta t} \delta t = (5M_\odot)^{2-\beta} - (10^9 M_\odot)^{1-\beta} \frac{2 - \beta}{1 - \beta} \bar{\psi}_{\delta t} \delta t \quad (18)$$

$$(19)$$

IGIMF (or equivalently, gwIMF)

$$\xi_{\text{IGIMF}}(m; t) = dN_\star/dm = \int_{5M_\odot}^{M_{\text{ecl},\max}} \xi_\star(m, M_{\text{ecl}}, [Z/X]) \xi_{\text{ecl}}(M_{\text{ecl}}, \bar{\psi}_{\delta t}) dM_{\text{ecl}}. \quad (20)$$