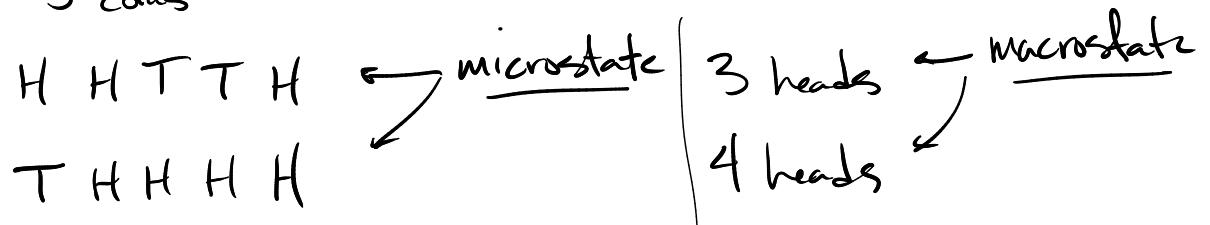


Chapter 2 \rightarrow 2nd Law \rightarrow Entropy

$$P(n) = \frac{n}{N}$$

Combinatorics

Coin flips \rightarrow 5 coins



$$P(3 \text{ heads}) = \frac{\Omega(3)}{\Omega(\text{all})}$$

$$\sum_{n=0}^N \Omega(n)$$

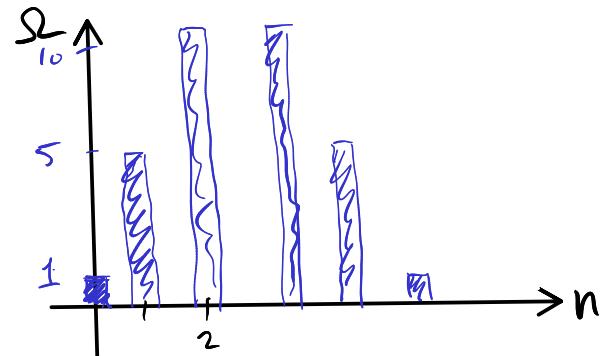
How many microstates are in a macrostate?
 ↳ multiplicity

$$\Omega(n) = \frac{5!}{n!(5-n)!} \leftarrow \text{combinations}$$

$$\begin{array}{c} \# \text{ of} \\ \text{heads} \end{array} \quad \begin{array}{c} 5! \\ \hline 3! \cdot 2! \end{array}$$

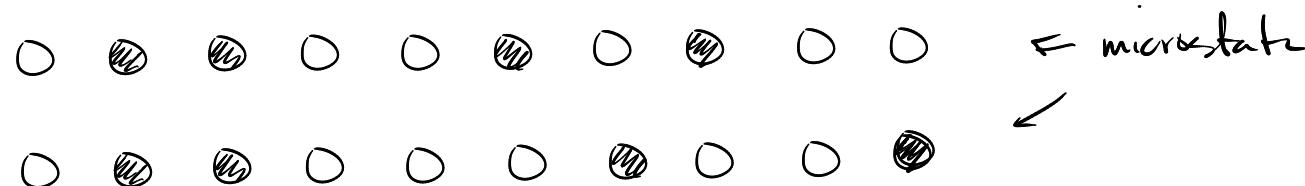
$$\Omega(N, n) = \frac{N!}{n!(N-n)!} \leftarrow \binom{N}{n}$$

$$\begin{array}{c} \# \text{ of} \\ \text{coins} \end{array} \quad \begin{array}{c} \# \text{ of} \\ \text{heads} \end{array}$$



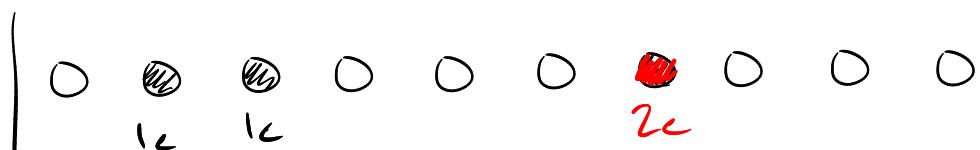
10 atoms \rightarrow each one can have 0 or 1 energy units.

So how many ways are there of arranging 4 units of energy? (10 units of energy)



4 energies vs. 10 energies ← macrostate

What if an atom can have more than 1?

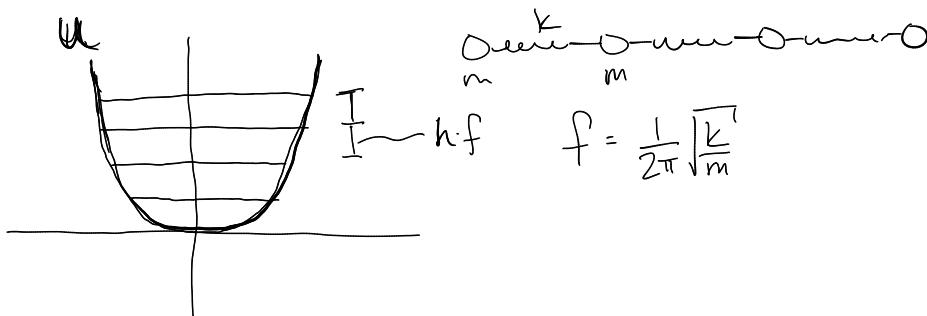


→ need to update the formula

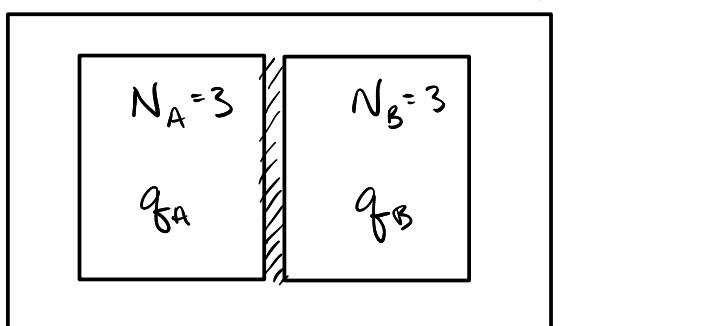
$$\Omega(N, q) = \frac{(q + N - 1)!}{q! (N - 1)!}$$

This idea of treating a solid like

this is known as an Einstein solid ← Debye model



Two systems



$g = 6 \leftarrow$ six quanta of energies that can be shared

- Large Number \rightarrow addition of small #s does not matter

$$10^{23} + 23 = 10^{23}$$

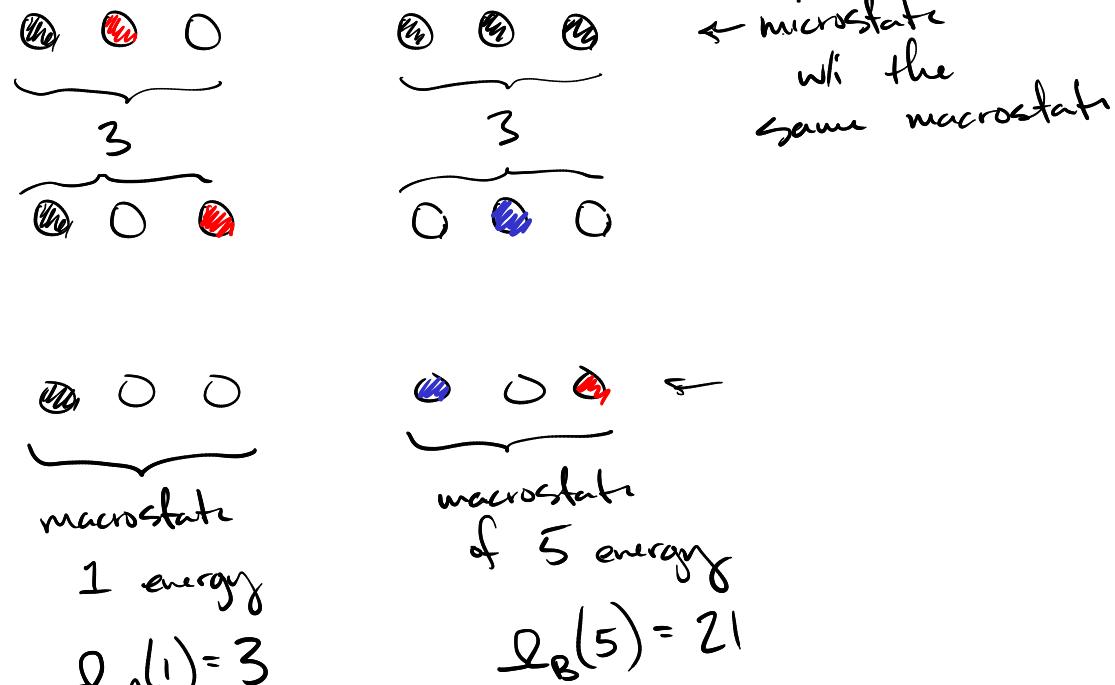
- Very Large Number

$$10^{10^{23}} \times 10^{23} = 10^{10^{23} + 23} = 10^{10^{23}}$$

Stirling's Approximation

$$N! \approx N^N e^{-N} \sqrt{2\pi N} \approx N^N e^{-N} = \frac{N^N}{e^N}$$

$$\ln N! \approx N \ln N - N$$



$$\Omega_{\text{total}} = \Omega_A \cdot \Omega_B = 63$$

Anaconda \rightarrow jupyter lab

Fundamental Assumption of Statistical Mechanics

↳ all microstates are possible and equally probable.

But, that does not mean that every microstate will occur.

Not all macrostates are equally probable.

```
[1]: %matplotlib widget
```

```
[2]: import numpy as np
import pandas as pd
import matplotlib.pyplot as plt
import scipy.special as sp
```

```
[3]: pd.set_option('display.max_colwidth', None)
pd.set_option('display.max_columns', None)
pd.set_option('display.max_rows', None)
```

```
[12]: def factorial(x):
    return sp.factorial(x, exact=True)

def multiplicity(N, q):
    return factorial(q+N-1)//factorial(q)//factorial(N-1)

def logArray(array):
    import math
    return [math.log(x) for x in array]
```

two einstein solids

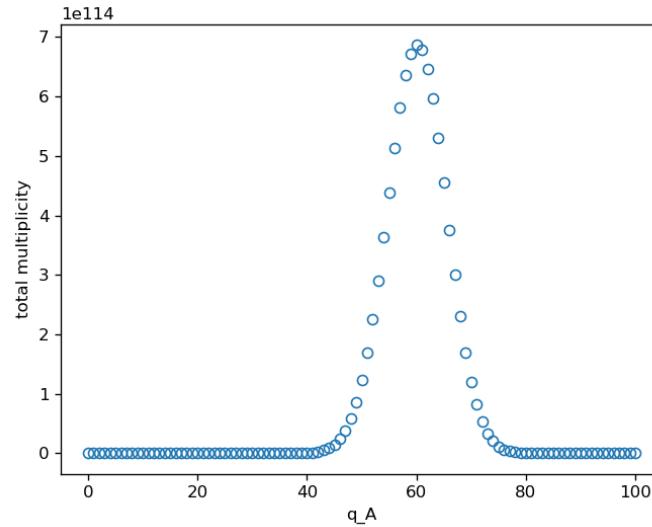
3 particles each and 6 energy units

```
[24]: table0 = pd.DataFrame()
table0['q_A'] = range(0,6+1, 1)
table0['q_B'] = range(6, 0-1, -1)
table0['mult_A'] = [multiplicity(3, i) for i in table0['q_A']]
table0['mult_B'] = [multiplicity(3, i) for i in table0['q_B']]
table0['mult_total'] = table0['mult_A']*table0['mult_B']
```

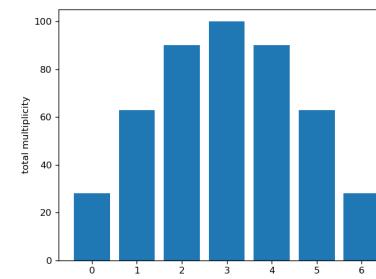
```
[25]: table0
```

	q_A	q_B	mult_A	mult_B	mult_total
0	0	6	1	28	28
1	1	5	3	21	63
2	2	4	6	15	90
3	3	3	10	10	100
4	4	2	15	6	90
5	5	1	21	3	63
6	6	0	28	1	28

Figure 1



- increase to
 $n_A = 300, n_B = 200$
 $q_f = 100$



So, now lets apply Stirlings Approximation to the Multiplicity

$$\Omega(N, q) = \frac{(q+N-1)!}{q!(N-1)!} \approx \frac{(q+N)!}{q! N!} \Rightarrow \ln \Omega = N \ln N - N$$

$$\begin{aligned}\ln \Omega &= \ln(q+N)! - \ln q! - \ln N! \rightarrow \ln N! = N \ln N - N \\ &\rightarrow \ln q! = q \ln q - q \\ &\rightarrow \ln(q+N)! = (q+N) \ln(q+N) - (q+N)\end{aligned}$$

$$\ln \Omega = (q+N) \ln(q+N) - q \cancel{- N} - q \ln q + \cancel{q} - N \ln N + \cancel{N}$$

$$\boxed{\ln \Omega = (q+N) \ln(q+N) - q \ln q - N \ln N}$$

high temperature limit $\rightarrow q \gg N$

$$\ln \Omega = (q+N) \ln(q+N) - q \ln q - N \ln N$$

$$= q \underbrace{\ln(q+N)}_{\downarrow} + N \underbrace{\ln(q+N)}_{\downarrow} - q \ln q - N \ln N$$

$$= \ln \left[q \cdot \left(1 + \frac{N}{q} \right) \right]$$

$$= \ln q + \underbrace{\ln \left(1 + \frac{N}{q} \right)}_{\substack{\hookrightarrow \ln(1+x) \approx x \text{ for small } x \\ \Rightarrow \frac{N}{q}}}$$

$$\Rightarrow \ln q + \frac{N}{q} \approx \ln(q+N)$$

$$\ln \Omega = \cancel{q \ln q} + N + \cancel{N \ln q} + \cancel{\frac{N^2}{q}} - \cancel{q \ln q} - \cancel{N \ln N}$$

$$= N \ln \left(\frac{q}{N} \right) + N + \underbrace{\frac{N^2}{q}}_{\text{small}}$$

$$\ln \Omega (q \gg N) = N \ln \left(\frac{q}{N} \right) + N$$

$$\Omega(q \gg N) = e^{N \ln \left(\frac{q}{N} \right) + N} = e^{N \ln \left(\frac{q}{N} \right)} \cdot e^N = \underbrace{\left(e^{\ln \left(\frac{q}{N} \right)} \right)^N}_{\frac{q}{N}} \cdot e^N$$

$$\Omega(g \gg N) = \left(\frac{g}{N}\right)^N \cdot e^N$$

$$\Omega(g \gg N) = \left(\frac{e g}{N}\right)^N \quad \leftarrow \text{Einstein solid at high temps}$$

- problem 2.18 |

High temperature 2 solids in thermal contact

$$\Omega = \left(\frac{e g_A}{N_A}\right)^{N_A} \cdot \left(\frac{e g_B}{N_B}\right)^{N_B} \quad N_A = N_B = N$$

$$\Omega = \left(\frac{e}{N}\right)^{2N} \cdot (g_A \cdot g_B)^N$$

$$g_A = g_B = \frac{g}{2}$$

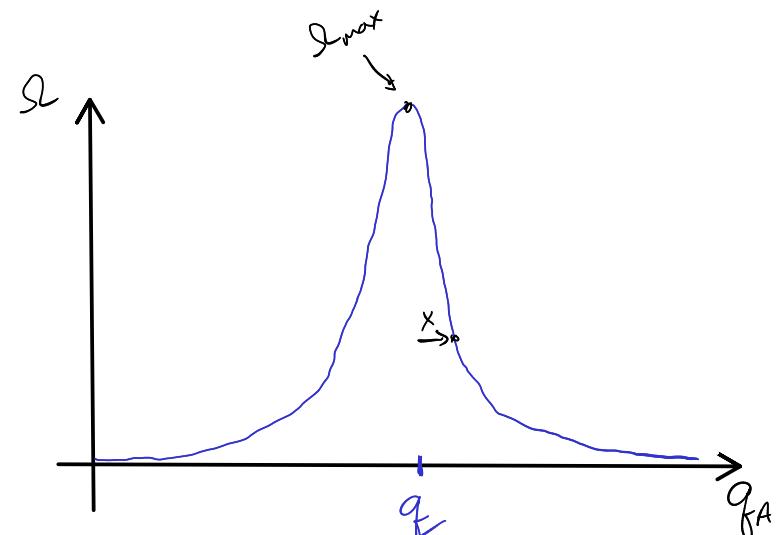
$$\Omega_{\max} = \left(\frac{e}{N}\right)^{2N} \left(\frac{g}{2}\right)^{2N}$$

$$g_A = \frac{g}{2} + x$$

$$g_B = \frac{g}{2} - x$$

$$\Omega = \left(\frac{e}{N}\right)^{2N} \left[\left(\frac{g}{2}\right)^2 - x^2\right]^N$$

$$\ln \left[\left(\frac{g}{2}\right)^2 - x^2\right]^N = N \cdot \ln \left[\left(\frac{g}{2}\right)^2 - x^2\right]$$



back to the graph:

$$g_A = 10^{28}$$

$$\text{width} = 10^{16}$$

$$\frac{0.6 \text{ m}}{10^{28}} \cdot 10^{16} = 0.6 \cdot 10^{-12}$$

$$= \underbrace{6 \cdot 10^{-13} \text{ m}}$$

sharp peak!

$$= N \ln \left[\left(\frac{q}{2} \right)^2 \cdot \left(1 - \frac{x^2}{\left(\frac{q}{2} \right)^2} \right) \right]$$

$$= N \left(\ln \left(\frac{q}{2} \right)^2 + \underbrace{\ln \left(1 - \frac{x^2}{\left(\frac{q}{2} \right)^2} \right)}_{\text{Takaway! There is one macrostate where we will ever measure this system to be in.}} \right)$$

$\ln(1+x) \approx x$, for small x

$$\ln \left[\left(\frac{q}{2} \right)^2 - x^2 \right]^N = N \left(\ln \left(\frac{q}{2} \right)^2 - \left(\frac{2x}{q} \right)^2 \right)$$

$$\Omega = \left(\frac{e}{N} \right)^{2N} e^{N \cdot \ln \left(\frac{q}{2} \right)^2 - N \left(\frac{2x}{q} \right)^2}$$

$$\Omega = \left(\frac{e}{N} \right)^{2N} e^{\underbrace{N \cdot \ln \left(\frac{q}{2} \right)^2}_{\left(\frac{q}{2} \right)^{2N}}} e^{-N \left(\frac{2x}{q} \right)^2}$$

$$\Omega = \underbrace{\left(\frac{e}{N} \right)^{2N} \cdot \left(\frac{q}{2} \right)^{2N}}_{\Omega_{\max}} \cdot e^{-N \left(\frac{2x}{q} \right)^2}$$

$$I = I_{\max} e^{-N \left(\frac{2x}{q}\right)^2}$$

Gaussian function

$$\frac{I}{I_{\max}} = e^{-N \left(\frac{2x}{q}\right)^2}$$

$$\frac{1}{e} = e^{-1}$$

$$e^{-1} = e^{-N \left(\frac{2x}{q}\right)^2}$$

$$-1 = -N \left(\frac{2x}{q}\right)^2$$

$$\frac{1}{N} = \left(\frac{2x}{q}\right)^2$$

how far
from the
central peak → $x = \frac{q}{2\sqrt{N}}$

suppose $N = 10^{22}$
 $q = 10^{28}$

} width = $2x = \frac{q}{\sqrt{N}} = \frac{10^{28}}{\sqrt{10^{22}}} = \frac{10^8}{10^4} = 10^4$

2.18 + 2.22

$$\underline{2.18} \quad \text{show: } \Omega = \frac{\left(\frac{q+N}{q}\right)^q \left(\frac{q+N}{N}\right)^N}{\sqrt{2\pi q(q+N)/N}}$$

$$\Omega = \frac{(q+N)!}{q! N!}$$

$$\Omega = \frac{(q+N)^{q+N} \cdot e^{-q-N} \sqrt{2\pi(q+N)}}{q^q e^{-q} \sqrt{2\pi q} \cdot N^N e^{-N} \sqrt{2\pi N}}$$

$$\Omega = \frac{(q+N)^q \cdot (q+N)^N \cdot \sqrt{(q+N)}}{q^q \cdot N^N \cdot \sqrt{2\pi q N}}$$

$$\Omega = \frac{\left(\frac{q+N}{q}\right)^q \cdot \left(\frac{q+N}{N}\right)^N}{\sqrt{\frac{2\pi q N}{q+N}}} \quad \leftarrow \text{not working}$$

Stirling's Approximation

$$N! \approx N^N e^{-N} \sqrt{2\pi N} \approx N^N e^{-N} = \frac{N^N}{e^N}$$

$$\ln N! \approx N \ln N - N$$

$$\Omega = \frac{(q+N-1)!}{q! (N-1)!}$$

$$N! = N \underbrace{(N-1)(N-2)\dots}_{(N-1)!}$$

$$(N-1)! = \frac{N!}{N}$$

$$\Omega = \frac{N(q+N-1)!}{q! N!}$$

Entropy and the Second Law of Thermodynamics

Any large system in equilibrium will be found in the macrostate with the largest multiplicity

2nd Law of Thermodynamics
↓

Multiplicity tends to increase

Multiplicities are large numbers! Take the natural log of them

$$\text{entropy} \rightarrow S = k_B \ln \Omega$$

$$\frac{S}{k_B} = \ln \Omega$$

Ex: entropy of an Einstein solid.

$$\Omega = \left(\frac{eq}{N}\right)^N \quad q \gg N$$

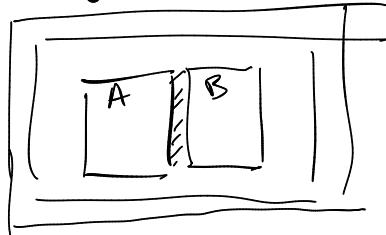
$$N = 10^{23} \quad q = 10^{25}$$

$$S = k_B \ln \left(\left(\frac{eq}{N} \right)^N \right) = N k_B \ln \left(\frac{eq}{N} \right) = N k_B \left(\ln(e) + \ln \left(\frac{q}{N} \right) \right)$$

$$S = N k_B \left(1 + \ln \left(\frac{q}{N} \right) \right)$$

$$S = N k_B \left(1 + \ln(10^2) \right) = 1.38 (1 + 4.6) = 7.7 \text{ J/K}$$

Entropy of a composite system



$$\Omega_{\text{total}} = \Omega_A \cdot \Omega_B$$

$$S_{\text{total}} = k_B \ln \Omega_{\text{total}} = k_B \ln(\Omega_A \cdot \Omega_B) = k_B \ln \Omega_A + k_B \ln \Omega_B$$

$$S_{\text{total}} = S_A + S_B$$

Chapter 3 - temperature \rightarrow thermal equilibrium

For an Einstein solid:

$$\frac{\partial S_{\text{total}}}{\partial q_A} = 0$$

(generalize)

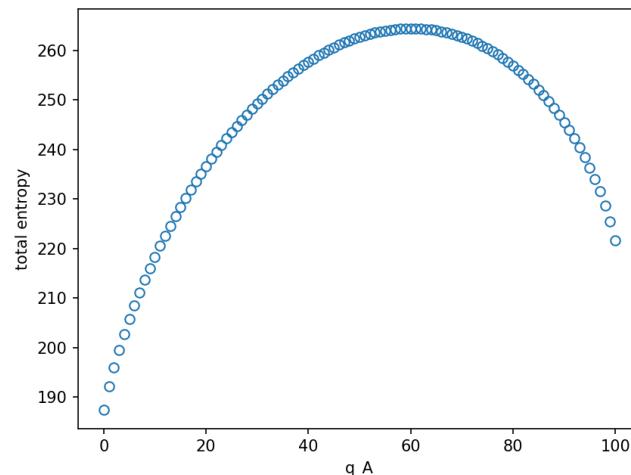
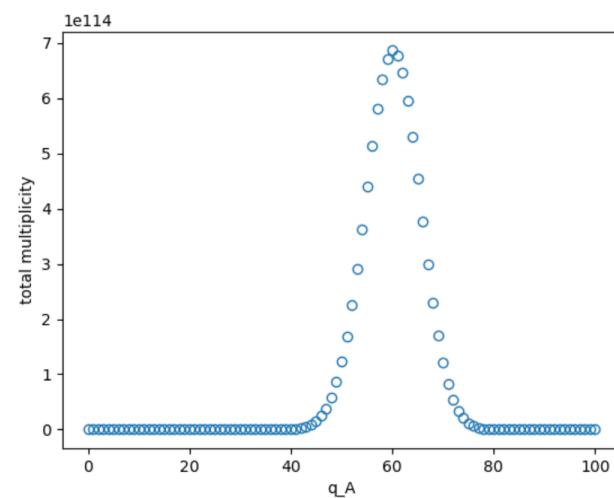
$$\frac{\partial S_{\text{total}}}{\partial U_A} = 0$$

$$\frac{\partial (S_A + S_B)}{\partial U_A} = 0$$

$$\frac{\partial S_A}{\partial U_A} + \frac{\partial S_B}{\partial U_A} = 0$$

$$U_B = U - U_A$$

$$dU_B = -dU_A$$

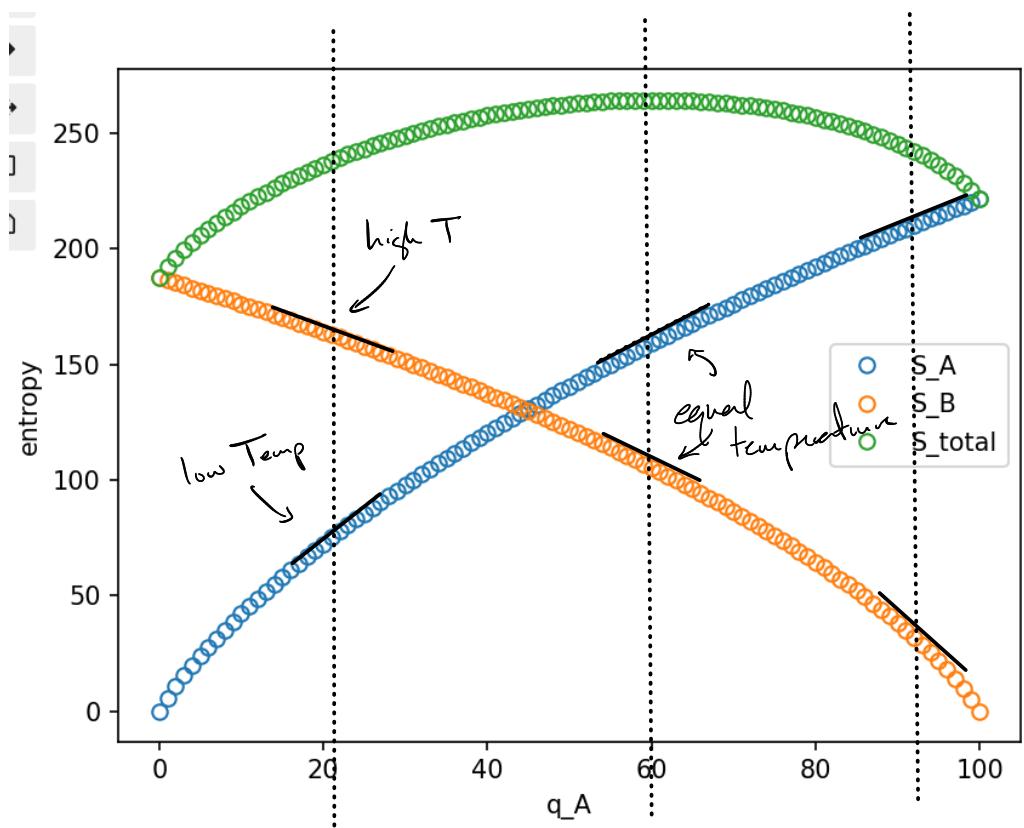


$$\frac{\partial S_A}{\partial U_A} - \frac{\partial S_B}{\partial U_B} = 0$$

$$\frac{\partial S_A}{\partial U_A} = \frac{\partial S_B}{\partial U_B}$$

$$\frac{1}{T} = \frac{\partial S}{\partial U}$$

$$T = \left(\frac{\partial S}{\partial U} \right)_{N,V}$$



HW: 2.29 + 2.30 \rightarrow look at 3.1

Total HW: 2.18, 22, 29, 30, 3.1

Let's apply this to an Einstein solid at high temp: $\mathcal{Q} = \left(\frac{e\epsilon}{N}\right)^N$

$$T = \left(\frac{\partial S}{\partial U}\right)_{N,V}$$

$$U = q \cdot \epsilon \longrightarrow q = \frac{U}{\epsilon}$$

↓
 # of energy packets
 ↓
 energy of one packet

$$S = k_B \ln \mathcal{Q} = k_B \ln \left(\frac{e\epsilon}{N} \right)^N = N k_B \ln \left(\frac{e\epsilon}{N} \right) = N k_B \left(1 + \ln \left(\frac{\epsilon}{N} \right) \right)$$

$$S = N k_B \left(1 + \ln \left(\frac{\epsilon}{N} \right) \right)$$

$$S = N k_B \left(1 + \ln \left(\frac{U}{N\epsilon} \right) \right)$$

$$S = N k_B + \underline{N k_B \ln U} - N k_B \ln N \epsilon$$

$$\frac{1}{T} = \left(\frac{\partial S}{\partial U} \right) = \frac{\partial (N k_B \ln U)}{\partial U} = \frac{N k_B}{U}$$

$$\underline{U = N k_B T} \longleftrightarrow U = \frac{f}{2} N k_B T$$

↗ 1D kinetic energy

$f=2 \quad \left. \begin{array}{l} \{ \text{Einstein} \rightarrow 1D \text{ spring potential energy} \end{array} \right.$

$$C_v = \left(\frac{\partial U}{\partial T} \right)_{N,V} = Nk_B \quad \leftarrow \begin{array}{l} \text{experimentally} \\ \text{testable} \\ \text{quantity} \end{array}$$

Review our process so far:

1. Used combinatorics + QM to find Ω in terms of N, V, U etc. } could be impossible!

2. $S = k_B \ln \Omega$ to get entropy

3. $T = \left(\frac{\partial S}{\partial U} \right)_{N,V,\text{etc}}$

4. solve step 3 for U as a function of T

5. $C_v = \frac{\partial U}{\partial T}$

statistical mechanics
gives an alternative
approach to arriving
at step #4.

But, going backwards to measure entropy is easy

$$dS = \underbrace{\frac{\partial S}{\partial U} dU}$$

$$\frac{1}{T}$$

$$dS = \frac{dU}{T}$$

→ use constant volume process (isochoric)

$$\begin{aligned} dV &= 0 \\ dW &= 0 \end{aligned} \quad \left. \begin{array}{l} \{ \\ \} \end{array} \right. \quad \begin{aligned} dU &= dQ + dW \\ dU &= dQ \end{aligned}$$

$$\boxed{dS = \frac{dQ}{T}} \rightarrow \text{important! need to be the definition of entropy}$$

→ T does not change very much w/ a $\pm Q$ amount of heat!

$$dS = \frac{C_v dT}{T}$$

$$\Delta S = \int_{T_i}^{T_f} \frac{C_v}{T} dT \quad \begin{array}{l} \xrightarrow{\text{C}_v \text{ can be constant, or}} \\ \text{can be a function of temp} \end{array}$$

$$S(T) - S(0) = \int_0^T \frac{C_v}{T} dT \quad \begin{array}{l} \xrightarrow{\text{but, you need to know this function all the way to zero Kelvin!}} \\ C_v \rightarrow 0 \text{ as } T \rightarrow 0 \end{array}$$

much experimental data
has been gathered for
many substances

could be 0 or residual entropy
but it is a constant

3rd Law of
Thermodynamics

2.22] a) $g \rightarrow 2N$

$$N_A = N_B = N$$

$$\underline{\text{macrostates}} = g + 1 = \underline{\underline{2N+1}}$$

b) $\mathcal{Q}(N, g) = \frac{\left(\frac{g+N}{g}\right)^g \left(\frac{g+N}{N}\right)^N}{\sqrt{2\pi g(g+N)/N}}$

$$\mathcal{Q}_{\text{total}} = \mathcal{Q}_A \cdot \mathcal{Q}_B = \frac{\left(\frac{g+N}{g}\right)^g \left(\frac{g+N}{N}\right)^N}{\sqrt{2\pi g(g+N)/N}} \cdot \frac{\left(\frac{g+N}{g}\right)^g \left(\frac{g+N}{N}\right)^N}{\sqrt{2\pi g(g+N)/N}}$$

$$g = 2N$$

$$\frac{\left(\frac{2N+N}{2N}\right)^{2N} \left(\frac{2N+N}{N}\right)^N}{\sqrt{2\pi 2N(2N+N)/N}} \cdot \frac{\left(\frac{2N+N}{2N}\right)^{2N} \left(\frac{2N+N}{N}\right)^N}{\sqrt{2\pi 2N(2N+N)/N}}$$

Let's go back to section 2.5

2.5] Ideal Gas Multiplicity

↳ the number of microstates in a macrostate.

↳ total energy, volume

which defines a macrostate
for an ideal gas

Ω = distribution of particles in space . distribution of energy among particles

For a single particle

$\Omega_1 \propto V \cdot$ combinations of momentums that has some energy
volume of momentum space

$$\Omega_1 \propto V \cdot V_p$$

$$\Omega_1 = \frac{L_x \cdot L_y \cdot L_z}{\Delta x \cdot \Delta y \cdot \Delta z} \cdot \frac{L_{p_x} \cdot L_{p_y} \cdot L_{p_z}}{\Delta p_x \cdot \Delta p_y \cdot \Delta p_z}$$

Heisenberg Uncertainty Principle

$$\Delta x \cdot \Delta p_x \approx h \leftarrow \text{Plank constant}$$

↑
uncertainty
in position uncertainty
in momentum

Inversely proportional

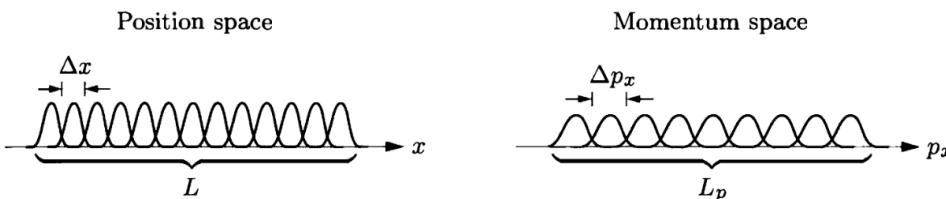


Figure 2.9. A number of “independent” position states and momentum states for a quantum-mechanical particle moving in one dimension. If we make the wavefunctions narrower in position space, they become wider in momentum space, and vice versa.

Let's go back to one particle - as all of the energy of the system, U .

$$U = \frac{1}{2}mv^2 = \frac{1}{2}m(V_x^2 + V_y^2 + V_z^2) \left(\frac{m}{\cancel{m}}\right)^2$$

$$U = \frac{1}{2m}(P_x^2 + P_y^2 + P_z^2)$$

$$\mathcal{D}_1 = \frac{\sqrt{V} \cdot \sqrt{V_p}}{h^3}$$

So, what about two particles?

$$\mathcal{D}_2 = \mathcal{D}_A \cdot \mathcal{D}_B$$

$$\mathcal{D}_2 = \frac{\sqrt{V} \cdot \sqrt{V_p}}{(h^3)^2} \cdot \frac{1}{2}$$

So, what about N particles?

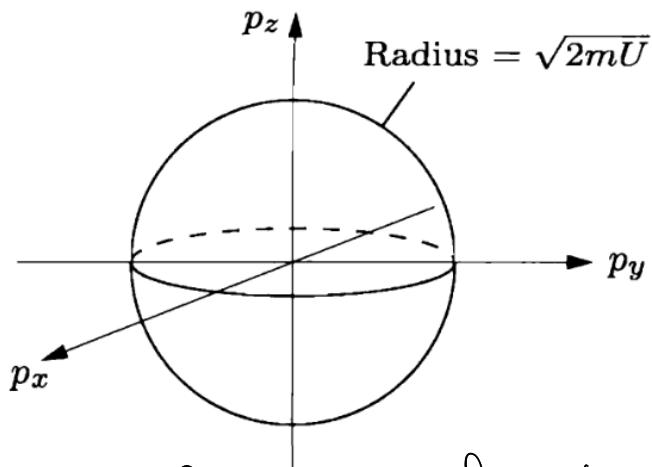
$$\mathcal{D}_N = \frac{1}{N!} \frac{\sqrt{V}^N \sqrt{V_p}^N}{(h^3)^N}$$

$$P_x^2 + P_y^2 + P_z^2 = (\underbrace{\sqrt{2mU}}_{{\text{if this is}}\text{ constant}})^2$$

equation of a sphere
in momentum space

that has a radius $r_p = \sqrt{2mU}$

Figure 2.8. A sphere in momentum space with radius $\sqrt{2mU}$. If a molecule has energy U , its momentum vector must lie somewhere on the surface of this sphere.



So the energy is really the surface area of this sphere
 $\rightarrow 4\pi r^2 \rightsquigarrow 4\pi(\sqrt{2mU})^2$

So, for two particles constrained to having U total energy.

$$p_{1x}^2 + p_{1y}^2 + p_{1z}^2 + p_{2x}^2 + p_{2y}^2 + p_{2z}^2 = (\sqrt{2mU})^2$$

6 dimensional sphere? \rightarrow hypersphere!

$$\text{"surface area"} = \frac{2\pi^{\frac{d}{2}} \cdot r^{d-1}}{\left(\frac{d}{2} - 1\right)!}$$

TEST: $d=2$

$$2\pi r ? = \frac{2 \cdot \pi^{\frac{2}{2}} \cdot r^{2-1}}{\left(\frac{2}{2} - 1\right)!} \\ = 2 \cdot \pi r \checkmark$$

Now go to N particles

$$\Omega_N = \frac{1}{N!} \frac{V^N}{h^{3N}} \cdot \frac{2\pi^{\frac{3N}{2}} \cdot (2m\mu)^{\frac{3N-1}{2}}}{\left(\frac{3N}{2} - 1\right)!}$$

Let's throw away what does not matter!

$$\boxed{\Omega_N = \frac{1}{N!} \frac{V^N}{h^{3N}} \cdot \frac{\pi^{\frac{3N}{2}} \cdot (2m\mu)^{\frac{3N}{2}}}{\left(\frac{3N}{2}\right)!}}$$

$$\underline{\Omega_N} = f(N) \cdot V^N \cdot U^{\frac{3N}{2}}$$

Put two ideal gases in thermal contact
(barrier exchange energy + volume)

$$\begin{aligned} \underline{\Omega_{\text{total}}} &= f(N) V_A^N U_A^{\frac{3N}{2}} \cdot f(N) V_B^N U_B^{\frac{3N}{2}} \\ &= (f(N))^2 \cdot (V_A V_B)^N \cdot (U_A \cdot U_B)^{\frac{3N}{2}} \end{aligned}$$

TEST: $d = 3$

$$4\pi r^2 = \frac{2\pi^{\frac{3}{2}} r^{3-1}}{\left(\frac{3}{2} - 1\right)!} = 2 \left[\frac{\pi^{\frac{3}{2}}}{\left(\frac{1}{2}\right)!} \right] r^2$$

need to dig in!

$$\frac{1}{2}! = \Gamma\left(\frac{1}{2} + 1\right)$$

$$\frac{1}{\left(\frac{1}{2}\right)!} = \frac{2}{\sqrt{\pi}} \quad \checkmark \checkmark$$

$$= 4\pi r^2$$

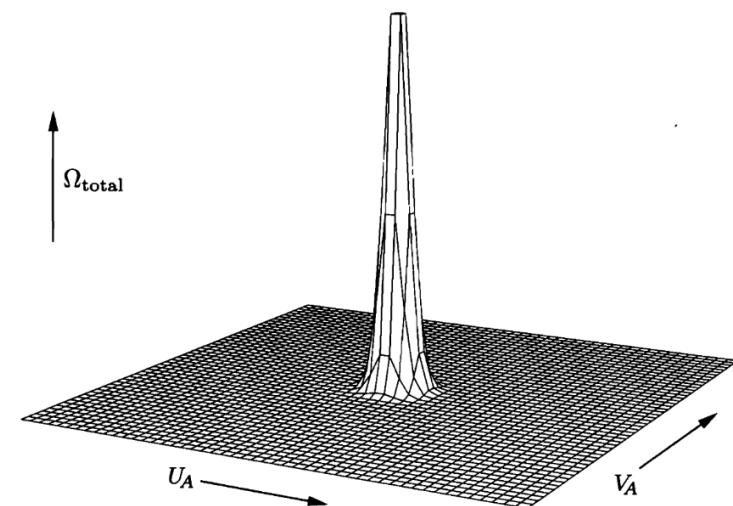


Figure 2.12. Multiplicity of a system of two ideal gases, as a function of the energy and volume of gas A (with the total energy and total volume held fixed). If the number of molecules in each gas is large, the full horizontal scale would stretch far beyond the edge of the page.

Entropy of an ideal gas:

$$\Omega_N = \frac{1}{N!} \frac{\nabla^N}{h^{3N}} \cdot \frac{\pi^{\frac{3N}{2}} \cdot (2mU)^{\frac{3N}{2}}}{\left(\frac{3N}{2}\right)!}$$

$$S = k_B \ln \Omega = k_B \ln \left(\frac{1}{N!} \frac{\nabla^N}{h^{3N}} \cdot \frac{\pi^{\frac{3N}{2}} \cdot (2mU)^{\frac{3N}{2}}}{\left(\frac{3N}{2}\right)!} \right)$$

$$S = k_B \left[-\ln N! - \ln \left(\frac{3N}{2}\right)! + N \ln \nabla + \frac{3N}{2} \ln \left(2\pi mU\right) - \frac{3N}{2} \ln (h^2) \right]$$

$\xrightarrow{-\ln N! = -N \ln N + N}$ $\xrightarrow{-\frac{3N}{2} \ln \left(\frac{3N}{2}\right) + \frac{3N}{2}}$
 $= N \ln \frac{1}{N} + N$ $\frac{3N}{2} \ln \left(\frac{2}{3N}\right) + \frac{3N}{2}$

$$S = k_B N \left(\ln \left[\frac{\nabla}{N} \cdot \left(\frac{2\pi mU}{h^2} \cdot \frac{2}{3N} \right)^{3/2} \right] + \frac{3}{2} + 1 \right)$$

$$S = k_B N \left(\ln \left[\frac{\nabla}{N} \left(\frac{4\pi mU}{3Nh^2} \right)^{3/2} \right] + \frac{5}{2} \right) \quad \xrightarrow{\text{Sackur-Tetrode Equation}}$$

$\frac{\nabla}{N} \rightarrow$ volume per particle $\rightarrow V$ $\frac{U}{N} \rightarrow$ energy per particle $\rightarrow u$
 little red little blue (new)

apply to one mole of N_2 gas at room temp and atmospheric pressure.

$$n_m = 1 \quad \text{or} \quad N = N_A \quad N_2 \rightarrow \text{diatomic} \rightarrow f = 5$$

$$\bullet U = \frac{5}{2} N k_B T = \frac{5}{2} n_m R T = \frac{5}{2} P V$$

$$\bullet \langle V \rangle = \frac{N k_B T}{P} = \frac{n_m R T}{P}$$

$$m_{N_2} = 28 \frac{\text{g}}{\text{mol}} = 0.028 \frac{\text{kg}}{\text{mol}} \cdot \frac{1 \text{ mol}}{6.022 \cdot 10^{23} \text{ part.}} = 4.65 \cdot 10^{26} \frac{\text{kg}}{\text{part}}$$

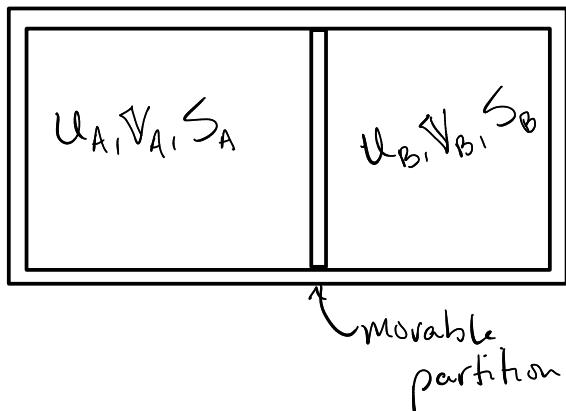
$$S = k_B N \left(\ln \left[\frac{\langle V \rangle}{N} \left(\frac{4\pi m U}{3N h^2} \right)^{3/2} \right] + \frac{5}{2} \right)$$

$$S = \underbrace{k_B N_A}_{R} \left(\ln \left[\frac{N_A k_B T}{N A P} \left(\frac{4\pi (4.65 \cdot 10^{26} \text{ kg}) \cdot \frac{5}{2} \cdot 8.31 \cdot 293}{3 \cdot 1 \text{ mol} \cdot (6.626 \cdot 10^{34})^2} \right)^{3/2} \right] + \frac{5}{2} \right) = ?$$

3.4] Mechanical Equilibrium + Pressure

- Exchange of energy is governed by temperature.
- Exchange of volume is governed by pressure difference.

know: $\frac{1}{T} = \left(\frac{\partial S}{\partial U}\right)_{N,V}$ ← is pressure related to entropy also?



equilibrium $\left\{ \begin{array}{l} \frac{\partial S_{\text{total}}}{\partial U} = 0 \\ \text{this is the definition} \\ \text{of the same temp} \\ \text{for } A + B \end{array} \right.$

propose: $\frac{\partial S_{\text{total}}}{\partial V_A} = 0$ $S_{\text{total}} = S_A + S_B$

$$\frac{\partial(S_A + S_B)}{\partial V_A} = 0$$

$$\frac{\partial S_A}{\partial V_A} + \frac{\partial S_B}{\partial V_A} = 0$$

$$dV_A = -dV_B$$

$$\frac{\partial S_A}{\partial V_A} - \frac{\partial S_B}{\partial V_B} = 0$$

↖

$$\frac{\partial S_A}{\partial V_A} = \frac{\partial S_B}{\partial V_B}$$

At equilibrium, pressure is the same, b/c otherwise the partition would move

$$P = T \left(\frac{\partial S}{\partial V} \right)_{U,N}$$

units?

$$\frac{J}{K m^3} = \frac{Nm}{K m^3 2} = \frac{Pa}{K} \quad \text{units work!}$$

is this right?

$$\Omega = f(N) \cdot V^N \cdot U^{\frac{3N}{2}} \quad \leftarrow \text{multiplicity of an ideal gas}$$

$$S = k_B \left(\ln f(N) + N \ln V + \frac{3N}{2} \ln U \right)$$

$$P = T \left(\frac{\partial S}{\partial V} \right)_{U,N} = k_B T \cdot N \frac{\partial \ln V}{\partial V}$$

$$\boxed{P = \frac{N k_B T}{V}} \quad \leftarrow \text{ideal gas law!}$$

new problems: 3.10, 3.14 ←
+ Total HW: 2.18, 22, 29, 30, 3.1

New total HW:

2.26, 32

New New total HW (totally new)

3.24 + 3.25

