

# Chapter 2 - 2<sup>nd</sup> Law of Thermodynamics

↳ heat spontaneously flows from high temp to low temp

Einstein Solid

Ideal Gas

↳ Entropy

→ Anacanda python  
install language  
Monty Python

## Combinatorics

one coin  
 $P(\text{heads}) = \frac{1}{2}$

multiple coins

$$P(n) = \frac{n}{N}$$

$$P(3 \text{ heads}) = \frac{\Omega(3)}{\Omega(\text{all})}$$
$$= \frac{\Omega(3)}{\sum_{n=0}^n \Omega(n)}$$

← microstate in the macrostate of 3  
← all of the microstates

5 coins

microstates { H H T T H — 3 H } macrostates  
                  T H H H H — 4 H }

how many microstates are in a macrostate?  
↳ multiplicity

$$\Omega(n) = \frac{5!}{n!(5-n)!}$$

↑  
number of heads

$$\begin{aligned} \Omega(0) &= 1 \\ \Omega(1) &= 5 \\ \Omega(2) &= 10 \\ \Omega(3) &= 10 \\ \Omega(4) &= 5 \\ \Omega(5) &= 1 \end{aligned}$$

$$\Omega(1) = \frac{5!}{1!(5-1)!} = \frac{5!}{4!}$$
$$= \frac{5 \cdot \cancel{4} \cdot \cancel{3} \cdot \cancel{2} \cdot \cancel{1}}{1 \cdot \cancel{4} \cdot \cancel{3} \cdot \cancel{2} \cdot \cancel{1}}$$

$$= 5$$


$$\Omega(2) = \frac{5!}{2!3!} = 10$$

$$\Omega(N, n) = \frac{N!}{n!(N-n)!} \leftarrow \text{Notation: } \binom{N}{n}$$

$\uparrow$   
 # of coins

10 atoms each w/ 0 or 1 packets of energy (energy unit)


How many possible ways are there to distribute 4 energy units


 $\leftarrow$  microstate

4 energy packets  $\leftarrow$  macrostate

$$\Omega(10, 4) = \frac{10!}{4!6!} = 210$$

What if an atom can have more than one energy packet at a time?

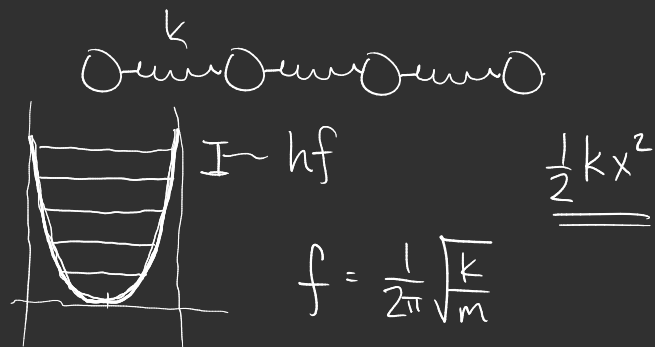

 $\leftarrow$  microstates

4 energy packets  $\leftarrow$  macrostate

$$\Omega(N, q) = \frac{(q + N - 1)!}{q!(N-1)!} \quad \binom{q+N-1}{q}$$

$\uparrow$  # of atoms       $\uparrow$  # of energy packets  
 $q = \text{macrostate}$

This model of a collection of atoms w/ equal size energy quanta distributed among them is the Einstein Solid.



→ Debye Model

• Large Number → addition of small numbers is not important

$$10^{23} + 23 = 10^{23}$$

• Very Large Number

$$10^{10^{23}} \times 10^{23} = 10^{10^{23} + 23} \approx 10^{10^{23}}$$

Stirling's Approximation

$$N! \approx N^N e^{-N} \sqrt{2\pi N} = \frac{N^N}{e^N} \sqrt{2\pi N}$$

$$\downarrow$$

$$\ln N! \approx N \ln N - N$$

very large numbers  
we leave this off

# Two Systems



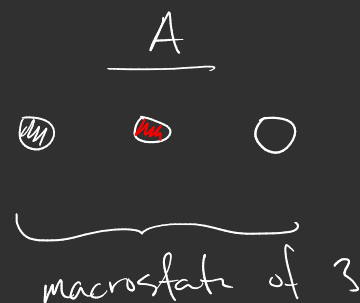
$q_A$	$\text{mult}_A$	$q_B$	$\text{mult}_B$	$\text{mult}_{\text{total}}$
0	1	6	28	28
1	3	5	21	63
2	6	4	15	90
3	10	3	10	100
4	15	2	6	90
5	21	1	3	63
6	28	0	1	28

macrostate w/  
the most microstates

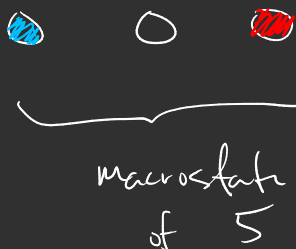
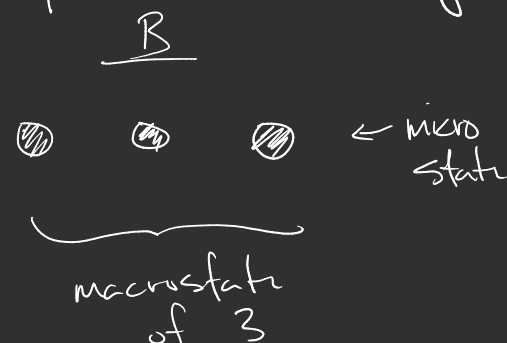
total number  
of microstates

(462)

Ex. Each system has 3 particles  
and 6 total quanta of energy



$$\Omega(3,1) = 3$$



$$\Omega(3,5)$$

$$= \frac{7!}{5!2!} = 21$$

Fundamental Assumption of Stat Mech:

all microstates are possible and equally probable

But that does not mean that every microstate will occur.

Not all macrostates are equally probable.

$$P_A(q) = \frac{\Omega_{\text{total}}(q_A)}{\sum_{q_A} \Omega_{\text{total}}(q_A)}$$

We could find the total # of microstates:

$$\Omega(6, 6) = \frac{(6+6-1)!}{6! (6-1)!} = 462$$

$\uparrow \quad \uparrow$   
 $N \quad q$

So let's apply Stirling's Approx to Multiplicity  $[\ln N! = N \ln N - N]$

$$\Omega(N, q) = \frac{(q+N-1)!}{q! (N-1)!} \approx \frac{(q+N)!}{q! N!}$$

very large number

$$\ln \Omega = \ln(q+N)! - \ln q! - \ln N! \rightarrow N \ln N - N$$

$\underbrace{\ln(q+N)!}_{\ln(q+N)! = (q+N) \ln(q+N) - (q+N)}$     $\underbrace{\ln q!}_{q \ln q - q}$     $\underbrace{\ln N!}_{N \ln N - N}$

$$\ln \Omega = (q+N) \ln(q+N) - \cancel{q} - \cancel{N} - q \ln q + \cancel{q} - N \ln N + \cancel{N}$$

$$\boxed{\ln \Omega = (q+N) \ln(q+N) - q \ln q - N \ln N}$$

high temperature limit  $\rightarrow q \gg N$

$$\begin{aligned}
 \ln \Omega &= (q+N) \ln(q+N) - q \ln q - N \ln N \\
 &= \underbrace{q \ln(q+N)} + \underbrace{N \ln(q+N)} - q \ln q - N \ln N \\
 &= \ln \left[ q \cdot \left( 1 + \frac{N}{q} \right) \right] \\
 &= \ln q + \ln \left( 1 + \frac{N}{q} \right) \\
 &\quad \downarrow \qquad \qquad \downarrow \begin{array}{l} \ln(1+x) \approx x \text{ for small } x \\ \rightarrow \frac{N}{q} \end{array} \\
 &= \ln q + \frac{N}{q} \approx \ln(q+N)
 \end{aligned}$$

$$\begin{aligned}
 \ln \Omega &= \cancel{q \ln q} + N + \underbrace{N \ln \frac{q}{N}} + \frac{N^2}{q} - \cancel{q \ln q} - \underbrace{N \ln N}_{\text{small}} \\
 &= N \ln \left( \frac{q}{N} \right) + N + \frac{N^2}{q}
 \end{aligned}$$

$$\ln \Omega(q \gg N) \approx N \ln \left( \frac{q}{N} \right) + N$$

$$\Omega(q \gg N) = e^{N \ln \left( \frac{q}{N} \right) + N} = e^{N \ln \left( \frac{q}{N} \right)} \cdot e^N = \underbrace{\left( e^{\ln \left( \frac{q}{N} \right)} \right)^N} \cdot e^N$$

$$\Omega(q \gg N) = \left( \frac{q}{N} \right)^N e^N = \left| \left( \frac{eq}{N} \right)^N \right|$$

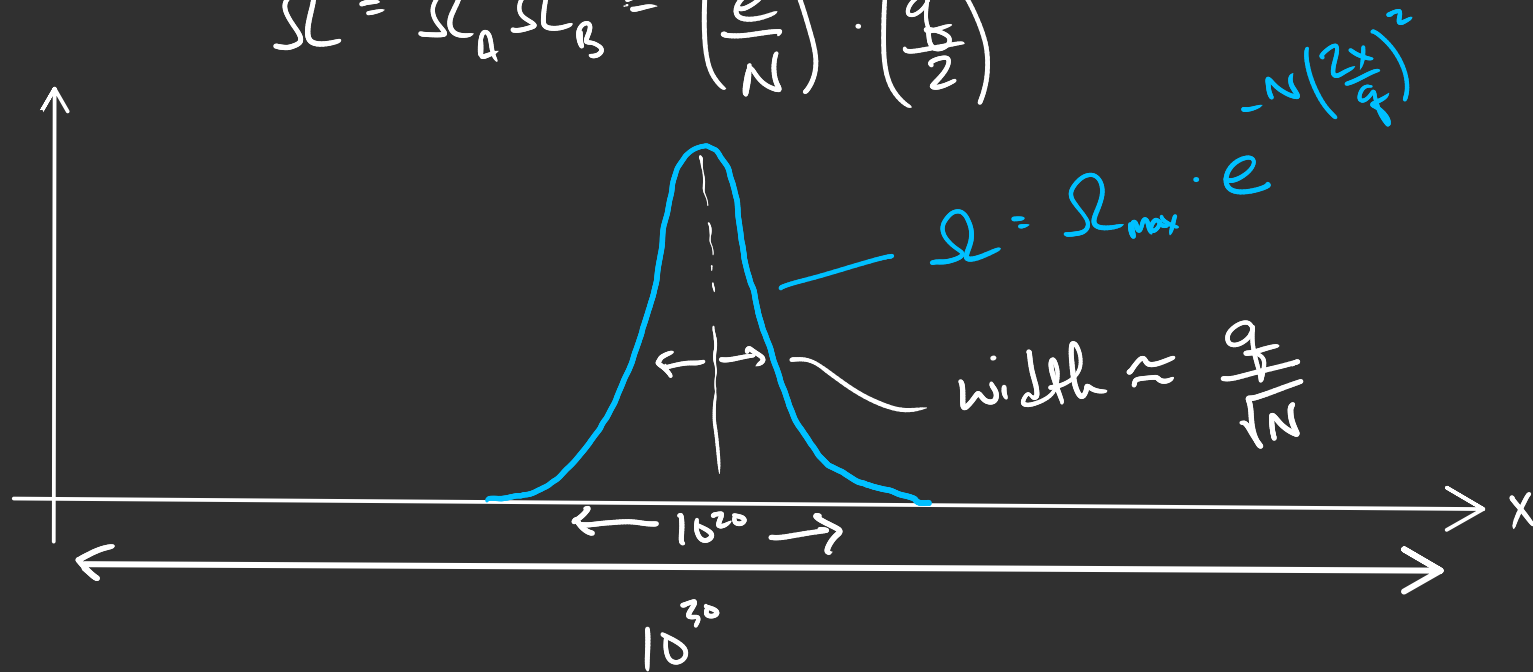
- 2 Einstein solids (high temperature limit) ( $q \gg N$ )

$$\Omega_A = \left( \frac{e q_A}{N_A} \right)^{N_A} \quad \Omega_B = \left( \frac{e q_B}{N_B} \right)^{N_B}$$

$$q = q_A + q_B$$

if  $N_A = N_B$ , multiplicity max will occur  $\frac{q}{2} = q_A = q_B$

$$\Omega = \Omega_A \Omega_B = \left( \frac{e}{N} \right)^{2N} \cdot \left( \frac{q}{2} \right)^{2N}$$



$$q = 10^{30}$$

$$N = 10^{20}$$

$$\frac{q}{\sqrt{N}} = 10^{20}$$

## Entropy + the 2nd Law:

Any large system in equilibrium will be found in the macrostate with the greatest ~~multiplicity~~  
entropy

### 2nd Law of Thermodynamics

~~Multiplicity~~ tends to increase.  
Entropy

Multiplicities are very large! Take the natural log of them.

$$\text{entropy} \rightarrow S = k_B \ln \Omega$$

$$\frac{S}{k_B} = \ln \Omega$$



Ex: Entropy of an Einstein solid

$$\Omega = \left(\frac{eq}{N}\right)^N \quad q \gg N \text{ Einstein solid}$$

homework  
( $q \approx N$ )

$$N = 10^{23} \quad q = 10^{25}$$

$$S = k_B \ln \left( \left( \frac{eq}{N} \right)^N \right) = N \cdot k_B \ln \left( \frac{eq}{N} \right) = N k_B \left( 1 + \ln \left( \frac{q}{N} \right) \right)$$

$\uparrow$   
 $10^{23} \cdot 1.38 \cdot 10^{-23}$

$$= 1.38 \left( 1 + \ln(10^2) \right)$$

$$S = 7.7 \text{ J/K} \quad \underbrace{\quad}_{4.6}$$

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Entropy of a composite system

$$\Omega_{\text{total}} = \Omega_A \Omega_B$$

$$S = k_B \ln \Omega_A \Omega_B = k_B \ln \Omega_A + k_B \ln \Omega_B$$

$$S_{\text{t}} = S_A + S_B$$

# Chapter 3

Thermal equilibrium

For the Einstein solid

$$\frac{\partial S_{\text{total}}}{\partial q_A} = 0$$

↓ generalize

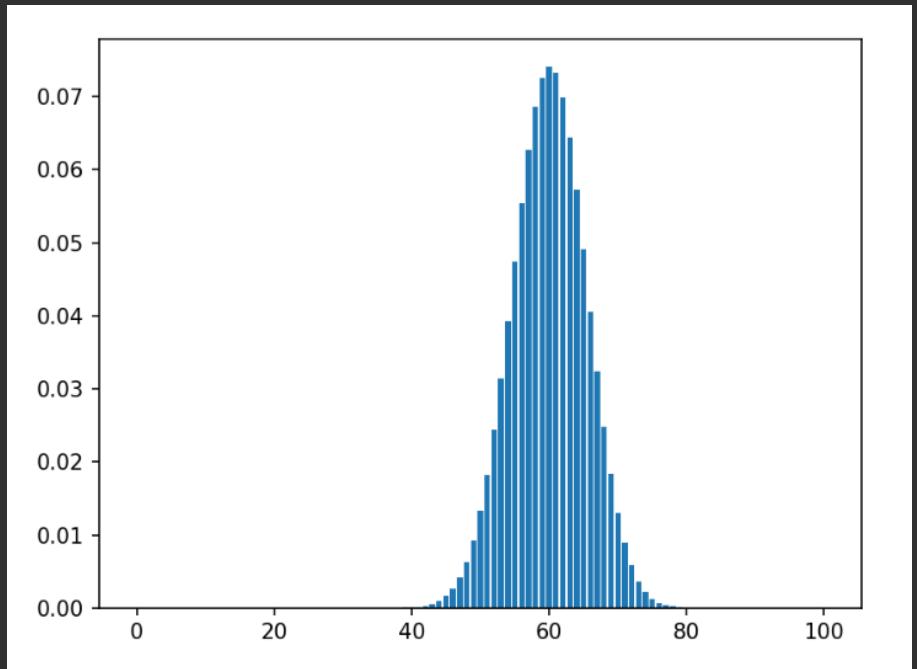
$$\frac{\partial S_{\text{total}}}{\partial U_A} = 0$$

$$\frac{\partial (S_A + S_B)}{\partial U_A} = 0$$

$$\frac{\partial S_A}{\partial U_A} + \frac{\partial S_B}{\partial U_A} = 0$$

$$U_B = U - U_A$$

$$\left\{ \begin{array}{l} dU_B = -dU_A \end{array} \right.$$



$$\frac{\partial S_A}{\partial U_A} - \frac{\partial S_B}{\partial U_B} = 0$$

$$\frac{\partial S_A}{\partial U_A} = \frac{\partial S_B}{\partial U_B}$$

$$\frac{1}{T} = \frac{\partial S}{\partial U}$$

$$T = \left( \frac{\partial S}{\partial U} \right)^{-1}_{N,N}$$

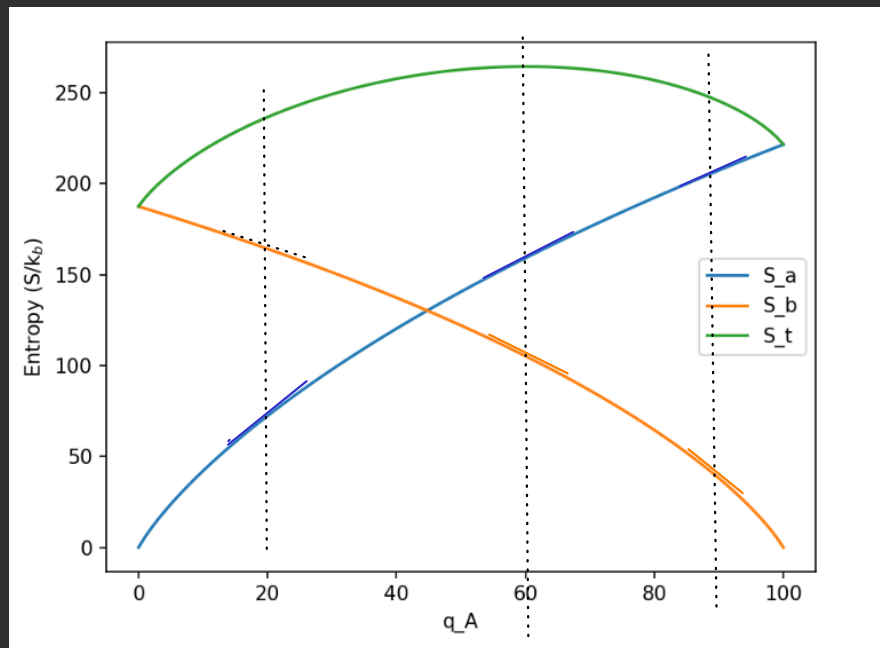
Now apply to our Einstein Solid ( $q \gg N$ )

$$U = q \cdot \epsilon \Rightarrow q = \frac{U}{\epsilon}$$

↑  
# of  
energy  
packets

↑ energy in  
one packet

$$\Omega = \left( \frac{e q}{N \theta} \right)^N$$



$$S = k_B \ln \Omega$$

$$S = k_B \ln \left( \left( \frac{eg}{N} \right)^N \right) = N \cdot k_B \ln \left( \frac{eg}{N} \right) = N k_B \left( 1 + \ln \left( \frac{g}{N} \right) \right)$$

$$S = N k_B \left( 1 + \ln \left( \frac{U}{N \epsilon} \right) \right)$$

$$S = N k_B + N k_B \ln U - N k_B \ln (N \epsilon)$$

$$\underbrace{\frac{1}{T} = \left( \frac{\partial S}{\partial U} \right)_{N,V} = \frac{\partial (N k_B \ln U)}{\partial U} = \frac{N k_B}{U}}$$

$$U = N k_B T \leftrightarrow U = \frac{f}{2} N k_B T$$

Heat Capacity of Einstein solid:  $f = 2$  } Einstein solid

$$C_v = \left( \frac{\partial U}{\partial T} \right)_{V,N} = N k_B \leftarrow \begin{array}{l} \text{experimentally} \\ \text{testable} \\ \text{quantity} \end{array}$$

↳ 1D kinetic energy

↳ 1D spring pot. energy

Review our process so far:

1. Used combinatorics and QM to find  $\Omega$  in terms of  $N, V, U$  etc. } could be impossible!

2.  $S = k_B \ln \Omega$  to get entropy

3.  $T = \left( \frac{\partial S}{\partial U} \right)^{-1}_{N, V, \text{etc}}$

4. Solve step #3 for  $U$  as a function of  $T$

5.  $C_v = \frac{\partial U}{\partial T}$

stat mech  
give an alternative  
to arriving  
at #4

But, going backward to measure the entropy is easy

$$dS = \underbrace{\frac{\partial S}{\partial U}}_{\text{mp}} \cdot dU$$

$$dS = \frac{dU}{T}$$

→ use constant volume process (isochoric)

$$\left. \begin{array}{l} dV = 0 \\ \pm W = 0 \end{array} \right\} dU = \pm Q + \pm W$$

$$dU = \pm Q$$

$$dS = \frac{\pm Q}{T}$$

→ important! used to be the definition of entropy!

→ T does not change much w/ a small addition of heat!

$$dS = \frac{C_v \cdot dT}{T}$$

$$\Delta S = \int_{T_i}^{T_f} \frac{C_v}{T} dT \rightarrow C_v \text{ can be constant, or a function of temperature}$$

$$S(T_f) - S(T_i) = \int_{T_i}^{T_f} \frac{C_v}{T} dT$$

$$S(T) - S(0) = \int_0^T \frac{C_v}{T} dT$$

but you need to know this function all the way to zero kelvin!  
 $C_v \rightarrow 0 \text{ as } T \rightarrow 0$

could be 0 or residual entropy but some constant

much experimental data has gone into this for many substances

3rd Law of Thermodynamics

$$\Omega = \frac{N!}{q!(N-q)!}$$

$$\begin{aligned} \Omega(1000, 500) &= \frac{1000!}{(500!)^2} \approx \frac{1000^{1000} e^{-1000} \sqrt{2\pi 1000}}{(500^{500} e^{-500} \sqrt{2\pi 500})^2} = \frac{1000^{1000} \sqrt{2\pi 1000}}{500^{1000} (\sqrt{2\pi 500})^2} \\ &= \left(\frac{1000}{500}\right)^{1000} \frac{\sqrt{2\pi 1000}}{\sqrt{2^2 \pi^2 500^2}} \\ &= 2^{1000} \cdot \sqrt{\frac{2 \cdot 1000}{(2 \cdot 500)^2 \pi}} \\ &= 2^{1000} \cdot \sqrt{\frac{2}{1000 \pi}} \\ &= \frac{2^{1000}}{\sqrt{500 \pi}} \end{aligned}$$

$$2^{1000}$$

$$= \sum_{q=1}^{1000} \frac{1000!}{q!(1000-q)!} \quad ??$$



## 2.5 Ideal Gas Multiplicity

Multiplicity is the number of microstate in a macrostate.

{ total energy, volume  
what defines macrostate  
for ideal gas

$\Omega$  = distribution of particles in space . distribution of energy among particles

For a single particle,

$\Omega_1 \propto V$  . combination of momentums that has some energy  
volume of momentum space

$$\Omega_1 \propto V \cdot V_p$$

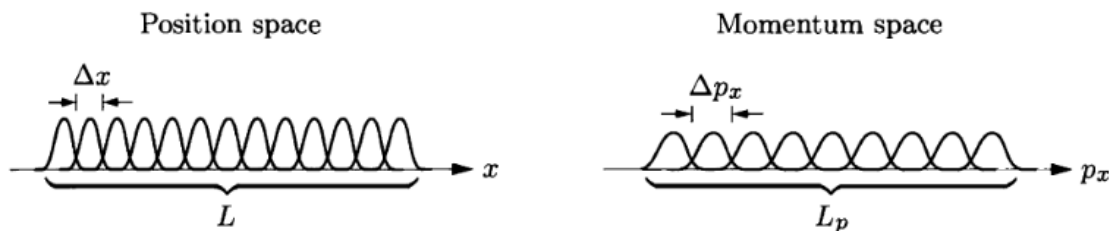
Heisenberg Uncertainty Principle

$$\Delta x \cdot \Delta p_x \approx h \leftarrow \text{Planck constant}$$

↑  
uncertainty  
in position

↑  
uncertainty  
in momentum

inversely proportional



**Figure 2.9.** A number of “independent” position states and momentum states for a quantum-mechanical particle moving in one dimension. If we make the wavefunctions narrower in position space, they become wider in momentum space, and vice versa.

$$\Omega_1 = \frac{L_x L_y L_z}{\Delta x \Delta y \Delta z} \cdot \frac{L_{p_x} L_{p_y} L_{p_z}}{\Delta p_x \Delta p_y \Delta p_z}$$

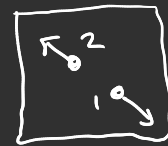
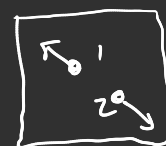
$$\Omega_1 = \frac{V \cdot V_p}{h \cdot h \cdot h} = \frac{V \cdot V_p}{h^3}$$

So what about two particles?

$$\Omega_2 = \Omega_A \cdot \Omega_B$$

$$= \frac{V^2 \cdot V_p^2}{(h^3)^2} \cdot \frac{1}{2}$$

cannot  
distinguish  
the  
particles



$$\Omega_N = \frac{1}{N!} \frac{V^N V_p^N}{(h^3)^N}$$

So lets go back to one particle that has all of the energy of the system,  $U$ ,

$$U = \frac{1}{2}mv^2 = \frac{1}{2}m(v_x^2 + v_y^2 + v_z^2) \cdot \frac{m}{m}$$

$$U = \frac{1}{2m}(p_x^2 + p_y^2 + p_z^2)$$

$$p_x^2 + p_y^2 + p_z^2 = 2mU = (\sqrt{2mU})^2$$

equation of a sphere  
in momentum space that  
has a radius,  $r = \sqrt{2mU}$

So the "volume" of momentum space that represents all of the possible momenta that the particle can have and still have  $U$  amount of energy is really the surface area of the sphere in momentum space  $\leadsto \underline{4\pi r^2} \leadsto \underline{4\pi(\sqrt{2mU})^2}$

For two particles constrained to  $U$  total energy,

then

$$p_{ix}^2 + p_{iy}^2 + p_{iz}^2 + p_{2x}^2 + p_{2y}^2 + p_{2z}^2 = (\sqrt{2mU})^2$$

6 dimensional sphere? hypersphere!

$$\text{"surface area"} = \frac{2\pi^{d/2} r^{d-1}}{\left(\frac{d}{2} - 1\right)!}$$

3N dimensional sphere

$$\text{"surface area"} = \frac{2\pi^{3N/2} r^{3N-1}}{\left(\frac{3N}{2} - 1\right)!}$$

$$\sqrt{2mU} = (2mU)^{1/2}$$

TEST:  $d=2$

$$2\pi r \stackrel{?}{=} \frac{2\pi^{2/2} r^{2-1}}{\left(\frac{2}{2} - 1\right)!} = 2\pi r \checkmark$$

TEST:  $d=3$

$$4\pi r^2 \stackrel{?}{=} \frac{2\pi^{3/2} r^{3-1}}{\left(\frac{3}{2} - 1\right)!} = 2 \underbrace{\left[ \frac{\pi^{3/2}}{\left(\frac{1}{2}\right)!} \right]}_{2\pi} r^2$$

$$\frac{1}{\left(\frac{1}{2}\right)!} = \frac{2}{\sqrt{\pi}} \\ = 4\pi r^2 \checkmark \checkmark$$

$$\Omega_N = \frac{1}{N!} \frac{V^N}{h^{3N}} \cdot \frac{2\pi^{3N/2} \cdot (2mU)^{\frac{3N-1}{2}}}{\left(\frac{3N}{2} - 1\right)!}$$

So let's throw away what does not matter for large numbers

$$\Omega_N = \frac{1}{N!} \frac{V^N}{h^{3N}} \cdot \frac{\pi^{3N/2}}{(\frac{3N}{2})!} (2mU)^{3N/2}$$

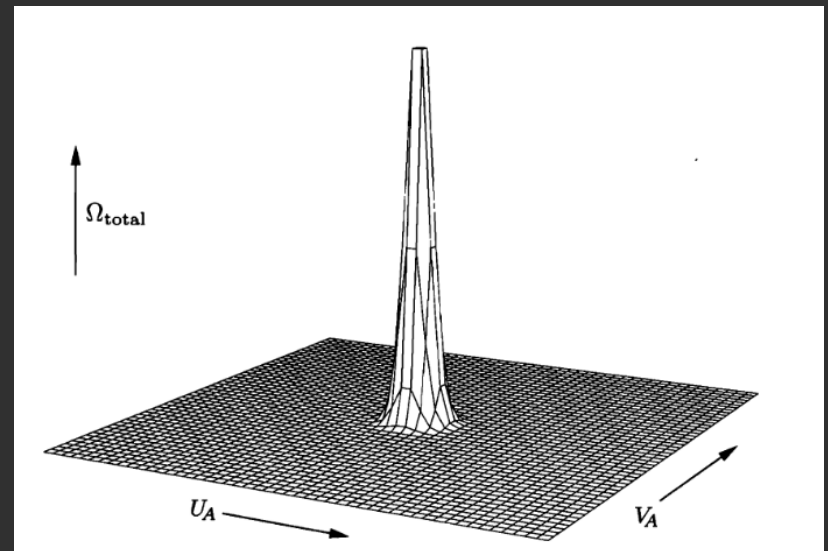
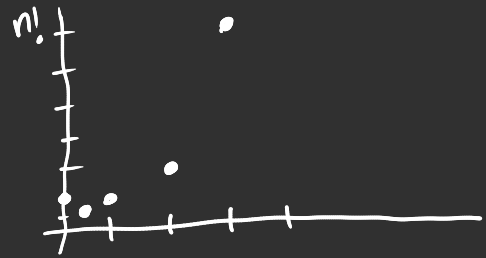
$$\Omega_N = f(N) \cdot V^N \cdot U^{3N/2}$$

Put two ideal gasses in thermal contact  
(barrier exchanges energy + volume)

$$\begin{aligned} \Omega_{\text{total}} &= f(N) V_A^N U_A^{3N/2} \cdot f(N) V_B^N U_B^{3N/2} \\ &= (f(N))^2 (V_A V_B)^N (U_A \cdot U_B)^{3N/2} \end{aligned}$$

$$2! = 2 \cdot 1$$

$$\frac{1}{2}! = \frac{\sqrt{\pi}}{2} = 0.8862$$



Entropy of an ideal gas

$$\Omega_N = \frac{1}{N!} \frac{V^N}{h^{3N}} \cdot \frac{(2\pi m U)^{\frac{3N}{2}}}{(\frac{3N}{2})!}$$

$$S = k_B \ln \Omega = k_B \ln \left[ \frac{1}{N!} \frac{V^N}{h^{3N}} \cdot \frac{(2\pi m U)^{\frac{3N}{2}}}{(\frac{3N}{2})!} \right]$$

$$S = k_B \left[ -\ln(N!) - \ln\left(\frac{3N}{2}!\right) + N \ln V + \frac{3N}{2} \ln(2\pi m U) - \frac{3N}{2} \ln(h^2) \right]$$

$$\rightarrow -N \ln N + N$$

$$N \ln \frac{1}{N} + N$$

$$\rightarrow -\frac{3N}{2} \ln\left(\frac{3N}{2}\right) + \frac{3N}{2}$$

$$\frac{3N}{2} \ln\left(\frac{2}{3N}\right) + \frac{3N}{2}$$

$$S = k_B N \left( \ln \left[ \frac{V}{N} \left( \frac{2\pi m U}{h^2} \cdot \frac{2}{3N} \right)^{\frac{3}{2}} \right] + 1 + \frac{3}{2} \right)$$

$$\boxed{S = k_B N \left( \ln \left[ \frac{V}{N} \left( \frac{4\pi m U}{3N h^2} \right)^{3/2} \right] + \frac{5}{2} \right)} \quad \text{Sackur-Tetrode Equation}$$

$\frac{V}{N} \rightarrow$  volume per particle

$\frac{U}{N} \rightarrow$  energy per particle

apply to one mole of  $N_2$  gas  
at room temp and atmospheric pressure

Find entropy.  $n_m = 1$ , or  $N = N_A = n_m$

$$U = \frac{5}{2} n_m R T = \frac{5}{2} N k_B T$$

$$V = \frac{n_m R T}{P} = \frac{N k_B T}{P}$$

$$\hookrightarrow V = \frac{R T}{P} = \frac{N_A k_B T}{P}$$

$$S = k_B N \left( \ln \left[ \frac{V}{N} \left( \frac{4\pi m U}{3N h^2} \right)^{3/2} \right] + \frac{5}{2} \right)$$

$$= k_B N_A \left( \ln \left[ \frac{\cancel{N_A} k_B T}{\cancel{N_A} P} \left( \frac{4\pi (4.65 \cdot 10^{-26} \text{ kg}) \cdot \frac{5}{2} \cancel{N_A} \cdot k_B \cdot 293 \text{ K}}{3 \cancel{N_A} h^2} \right)^{3/2} \right] + \frac{5}{2} \right)$$

$$= 8.31 \left( \ln \left[ \frac{1.38 \cdot 10^{-23} \cdot 293 \text{ K}}{10^5 \text{ Pa}} \left( \frac{4\pi (4.65 \cdot 10^{-26} \text{ kg}) \cdot \frac{5}{2} \cdot 1.38 \cdot 10^{-23} \cdot 293}{3 (6.626 \cdot 10^{-34})^2} \right)^{3/2} \right] + \frac{5}{2} \right) = 135.5 \text{ J/K}$$

$$N_2 = 28 \text{ g/mol} = 0.028 \frac{\text{kg}}{\text{mol}} = \frac{0.028 \text{ kg}}{6.022 \cdot 10^{23} \text{ atoms}} = 4.65 \cdot 10^{-26} \text{ kg}$$

HW: Ch 2.26, 32

HW: Ch 3.1, 10, 14 ←

$$\Omega(\text{all}) = \sum_{n=0}^N \frac{N!}{n!(N-n)!} = \cancel{N!} 2^N$$

2.26 |  $\Omega = \frac{A \cdot A_p}{h^2}$

so how  
do you →  $\Omega_N = \frac{1}{N!} \frac{V^N}{h^{3N}}$  x (area of the momentum hypersphere)  
adapt for  
2D gas

3.4 | Mechanical Equilibrium + Pressure















