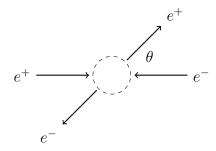
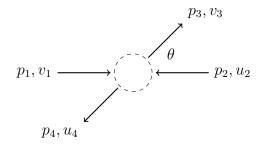
## BHABHA SCATTERING

Bhabha scattering is the interaction between positrons and electrons.



Here is the same diagram with momentum and spinor labels.



In a typical collider experiment the momentum vectors are

$$p_{1} = \begin{pmatrix} E \\ 0 \\ 0 \\ p \end{pmatrix} \qquad p_{2} = \begin{pmatrix} E \\ 0 \\ 0 \\ -p \end{pmatrix} \qquad p_{3} = \begin{pmatrix} E \\ p\sin\theta\cos\phi \\ p\sin\theta\sin\phi \\ p\cos\theta \end{pmatrix} \qquad p_{4} = \begin{pmatrix} E \\ -p\sin\theta\cos\phi \\ -p\sin\theta\sin\phi \\ -p\cos\theta \end{pmatrix}$$

where  $p = \sqrt{E^2 - m^2}$ . The spinors are

$$v_{11} = \begin{pmatrix} p \\ 0 \\ E+m \\ 0 \end{pmatrix} \quad u_{21} = \begin{pmatrix} E+m \\ 0 \\ -p \\ 0 \end{pmatrix} \quad v_{31} = \begin{pmatrix} p_3^z \\ p_3^x + ip_3^y \\ E+m \\ 0 \end{pmatrix} \quad u_{41} = \begin{pmatrix} E+m \\ 0 \\ p_4^z \\ p_4^x + ip_4^y \end{pmatrix}$$

$$v_{12} = \begin{pmatrix} 0 \\ -p \\ 0 \\ E+m \end{pmatrix} \quad u_{22} = \begin{pmatrix} 0 \\ E+m \\ 0 \\ p \end{pmatrix} \quad v_{32} = \begin{pmatrix} p_3^x - ip_3^y \\ -p_3^z \\ 0 \\ E+m \end{pmatrix} \quad u_{42} = \begin{pmatrix} 0 \\ E+m \\ p_4^x - ip_4^y \\ -p_4^z \end{pmatrix}$$

The last digit in a spinor subscript is 1 for spin up and 2 for spin down. Note that the spinors are not individually normalized. Instead, a combined spinor normalization constant  $N = (E + m)^4$  will be used where needed.

This is the probability density for Bhabha scattering. The formula is from Feynman diagrams.

$$|\mathcal{M}(s_1, s_2, s_3, s_4)|^2 = \frac{e^4}{N} \left| -\frac{1}{t} (\bar{v}_1 \gamma^{\mu} v_3) (\bar{u}_4 \gamma_{\mu} u_2) + \frac{1}{s} (\bar{v}_1 \gamma^{\nu} u_2) (\bar{u}_4 \gamma_{\nu} v_3) \right|^2$$

Symbol  $s_j$  selects the spin (up or down) of spinor j. Symbol e is electron charge. Symbols s and t are Mandelstam variables  $s = (p_1 + p_2)^2$  and  $t = (p_1 - p_3)^2$ .

Let

$$a_1 = (\bar{v}_1 \gamma^{\mu} v_3)(\bar{u}_4 \gamma_{\mu} u_2)$$
  $a_2 = (\bar{v}_1 \gamma^{\nu} u_2)(\bar{u}_4 \gamma_{\nu} v_3)$ 

Then

$$|\mathcal{M}(s_1, s_2, s_3, s_4)|^2 = \frac{e^4}{N} \left| -\frac{a_1}{t} + \frac{a_2}{s} \right|^2$$

$$= \frac{e^4}{N} \left( -\frac{a_1}{t} + \frac{a_2}{s} \right) \left( -\frac{a_1}{t} + \frac{a_2}{s} \right)^*$$

$$= \frac{e^4}{N} \left( \frac{a_1 a_1^*}{t^2} - \frac{a_1 a_2^*}{st} - \frac{a_1^* a_2}{st} + \frac{a_2 a_2^*}{s^2} \right)$$

The expected probability density  $\langle |\mathcal{M}|^2 \rangle$  is computed by summing  $|\mathcal{M}|^2$  over all spin states and dividing by the number of inbound states. There are four inbound states.

$$\langle |\mathcal{M}|^2 \rangle = \frac{1}{4} \sum_{s_1=1}^2 \sum_{s_2=1}^2 \sum_{s_3=1}^2 \sum_{s_4=1}^2 |\mathcal{M}(s_1, s_2, s_3, s_4)|^2$$

$$= \frac{e^4}{4} \sum_{s_1=1}^2 \sum_{s_2=1}^2 \sum_{s_3=1}^2 \sum_{s_4=1}^2 \frac{1}{N} \left( \frac{a_1 a_1^*}{t^2} - \frac{a_1 a_2^*}{st} - \frac{a_1^* a_2}{st} + \frac{a_2 a_2^*}{s^2} \right)$$

Use the Casimir trick to replace sums over spins with matrix products.

$$f_{11} = \frac{1}{N} \sum_{\text{spins}} a_1 a_1^* = \text{Tr}\left((\not p_1 - m)\gamma^{\mu}(\not p_3 - m)\gamma^{\nu}\right) \text{Tr}\left((\not p_4 + m)\gamma_{\mu}(\not p_2 + m)\gamma_{\nu}\right)$$

$$f_{12} = \frac{1}{N} \sum_{\text{spins}} a_1 a_2^* = \text{Tr}\left((\not p_1 - m)\gamma^{\mu}(\not p_2 + m)\gamma^{\nu}(\not p_4 + m)\gamma_{\mu}(\not p_3 - m)\gamma_{\nu}\right)$$

$$f_{22} = \frac{1}{N} \sum_{\text{spins}} a_2 a_2^* = \text{Tr}\left((\not p_1 - m)\gamma^{\mu}(\not p_2 + m)\gamma^{\nu}\right) \text{Tr}\left((\not p_4 + m)\gamma_{\mu}(\not p_3 - m)\gamma_{\nu}\right)$$

Hence

$$\langle |\mathcal{M}|^2 \rangle = \frac{e^4}{4} \left( \frac{f_{11}}{t^2} - \frac{f_{12}}{st} - \frac{f_{12}^*}{st} + \frac{f_{22}}{s^2} \right)$$

Run "bhabha-scattering-1.txt" to verify the Casimir trick.

These formulas compute probability densities from dot products.

$$f_{11} = 32(p_1 \cdot p_2)(p_3 \cdot p_4) + 32(p_1 \cdot p_4)(p_2 \cdot p_3) - 32m^2(p_1 \cdot p_3) - 32m^2(p_2 \cdot p_4) + 64m^4$$

$$f_{12} = -32(p_1 \cdot p_4)(p_2 \cdot p_3) - 16m^2(p_1 \cdot p_2) + 16m^2(p_1 \cdot p_3) - 16m^2(p_1 \cdot p_4)$$

$$- 16m^2(p_2 \cdot p_3) + 16m^2(p_2 \cdot p_4) - 16m^2(p_3 \cdot p_4) - 32m^4$$

$$f_{22} = 32(p_1 \cdot p_3)(p_2 \cdot p_4) + 32(p_1 \cdot p_4)(p_2 \cdot p_3) + 32m^2(p_1 \cdot p_2) + 32m^2(p_3 \cdot p_4) + 64m^4$$

In Mandelstam variables  $s = (p_1 + p_2)^2$ ,  $t = (p_1 - p_3)^2$ ,  $u = (p_1 - p_4)^2$  the formulas are

$$f_{11} = 8s^2 + 8u^2 - 64sm^2 - 64um^2 + 192m^4$$
  

$$f_{12} = -8u^2 + 64um^2 - 96m^4$$
  

$$f_{22} = 8t^2 + 8u^2 - 64tm^2 - 64um^2 + 192m^4$$

When  $E \gg m$  a useful approximation is to set m=0 and obtain

$$f_{11} = 8s^2 + 8u^2$$
$$f_{12} = -8u^2$$
$$f_{22} = 8t^2 + 8u^2$$

For m = 0 the Mandelstam variables are

$$s = 4E^{2}$$

$$t = -2E^{2}(1 - \cos \theta) = -4E^{2}\sin^{2}(\theta/2)$$

$$u = -2E^{2}(1 + \cos \theta) = -4E^{2}\cos^{2}(\theta/2)$$

The corresponding expected probability density is

$$\langle |\mathcal{M}|^2 \rangle = \frac{e^4}{4} \left( \frac{8s^2 + 8u^2}{t^2} + \frac{16u^2}{st} + \frac{8t^2 + 8u^2}{s^2} \right)$$

$$= 2e^4 \left( \frac{s^2 + u^2}{t^2} + \frac{2u^2}{st} + \frac{t^2 + u^2}{s^2} \right)$$

$$= 2e^4 \left( \frac{1 + \cos^4(\theta/2)}{\sin^4(\theta/2)} - \frac{2\cos^4(\theta/2)}{\sin^2(\theta/2)} + \frac{1 + \cos^2\theta}{2} \right)$$

Run "bhabha-scattering-2.txt" to verify.

This is the differential cross section for Bhabha scattering.

$$\frac{d\sigma}{d\Omega} = \frac{\langle |\mathcal{M}|^2 \rangle}{64\pi^2 s} = \frac{\alpha^2}{8E^2} \left( \frac{1 + \cos^4(\theta/2)}{\sin^4(\theta/2)} - \frac{2\cos^4(\theta/2)}{\sin^2(\theta/2)} + \frac{1 + \cos^2\theta}{2} \right)$$

We can integrate  $d\sigma$  to obtain a cumulative distribution function.

Let

$$I(\xi) = \int_{\alpha}^{\xi} \frac{d\sigma}{d\Omega} \sin\theta \, d\theta, \quad \alpha \le \xi \le \pi$$

for some  $\alpha > 0$ . The support interval is restricted because  $d\sigma$  is undefined for  $\theta = 0$ .

The cumulative distribution function is

$$F(\theta) = \frac{I(\theta)}{I(\pi)}, \quad \alpha \le \theta \le \pi$$

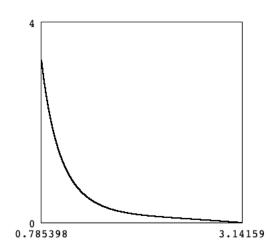
Hence

$$P(\theta_1 < \theta < \theta_2) = F(\theta_2) - F(\theta_1)$$

The probability density is

$$f(\theta) = \frac{dF(\theta)}{d\theta} = \frac{1}{I(\pi)} \frac{d\sigma}{d\Omega} \sin \theta, \quad \alpha \le \theta \le \pi$$

Run "bhabha-scattering-3.txt" to draw  $f(\theta)$  for  $\alpha = \frac{1}{4}\pi = 45^{\circ}$ .



The following table shows the corresponding probability distribution for three bins.

| $\theta_1$    | $\theta_2$    | $P(\theta_1 \le \theta \le \theta_2)$ |
|---------------|---------------|---------------------------------------|
| 0°            | $45^{\circ}$  | _                                     |
| $45^{\circ}$  | $90^{\circ}$  | 0.83                                  |
| $90^{\circ}$  | $135^{\circ}$ | 0.13                                  |
| $135^{\circ}$ | 180°          | 0.04                                  |

The following Bhabha scattering data is adapted from SLAC-PUB-1501.

|                      | Bin | $\cos\theta$ (interval) | Count |
|----------------------|-----|-------------------------|-------|
| (Smallest $\theta$ ) | 1   | 0.6, 0.5                | 4432  |
|                      | 2   | 0.5, 0.4                | 2841  |
|                      | 3   | 0.4, 0.3                | 2045  |
|                      | 4   | 0.3, 0.2                | 1420  |
|                      | 5   | 0.2, 0.1                | 1136  |
|                      | 6   | 0.1, 0.0                | 852   |
|                      | 7   | 0.0, -0.1               | 656   |
|                      | 8   | -0.1, -0.2              | 625   |
|                      | 9   | -0.2, -0.3              | 511   |
|                      | 10  | -0.3, -0.4              | 455   |
|                      | 11  | -0.4, -0.5              | 402   |
| (Largest $\theta$ )  | 12  | -0.5, -0.6              | 398   |

"Count" is the number of Bhabha scattering events observed per bin. Let us see if the density function  $\langle |\mathcal{M}|^2 \rangle$  explains the distribution of counts in the table. Start by integrating  $\langle |\mathcal{M}|^2 \rangle$  over all the bins to obtain a normalization constant.

$$\int_{\text{bins}} \langle |\mathcal{M}|^2 \rangle \, d\Omega = \int_0^{2\pi} \int_{\arccos 0.6}^{\arccos -0.6} \langle |\mathcal{M}|^2 \rangle \sin \theta \, d\theta \, d\phi = 2\pi \times 9.3817 \times 2e^4$$

Let

$$f(\theta) = \frac{\langle |\mathcal{M}|^2 \rangle}{2\pi \times 9.3817 \times 2e^4} = \frac{1}{2\pi \times 9.3817} \left( \frac{1 + \cos^4(\theta/2)}{\sin^4(\theta/2)} - \frac{2\cos^4(\theta/2)}{\sin^2(\theta/2)} + \frac{1 + \cos^2\theta}{2} \right)$$

The probability of a scattering event occurring in an interval  $\theta_1$  to  $\theta_2$  is obtained by integrating  $f(\theta)$  over that interval.

$$P(\theta_1 < \theta < \theta_2) = \int_0^{2\pi} \int_{\theta_1}^{\theta_2} f(\theta) \sin \theta \, d\theta \, d\phi = 2\pi \int_{\theta_1}^{\theta_2} f(\theta) \sin \theta \, d\theta$$

The total number of counts in the table is 15773. To obtain a predicted distribution, multiply 15773 times the probability for each bin. For example, for the first bin we have

$$P(\arccos 0.6 < \theta < \arccos 0.5) \times 15773 = 4598$$

Repeat for all bins to obtain the following predicted distribution.

| Bin | $\cos\theta$ (interval) | Count | Predicted |
|-----|-------------------------|-------|-----------|
| 1   | 0.6, 0.5                | 4432  | 4598      |
| 2   | 0.5, 0.4                | 2841  | 2880      |
| 3   | 0.4, 0.3                | 2045  | 1955      |
| 4   | 0.3, 0.2                | 1420  | 1410      |
| 5   | 0.2, 0.1                | 1136  | 1068      |
| 6   | 0.1, 0.0                | 852   | 843       |
| 7   | 0.0, -0.1               | 656   | 689       |
| 8   | -0.1, -0.2              | 625   | 582       |
| 9   | -0.2, -0.3              | 511   | 505       |
| 10  | -0.3, -0.4              | 455   | 450       |
| 11  | -0.4, -0.5              | 402   | 411       |
| 12  | -0.5, -0.6              | 398   | 382       |

The coefficient of determination  $R^2$  measures how well predicted values fit the real data. Let y be observed counts per bin and let  $\hat{y}$  be predicted counts per bin. Then

$$R^2 = 1 - \frac{\sum (y - \hat{y})^2}{\sum (y - \bar{y})^2} = 0.997$$

The result indicates that the model  $\langle |\mathcal{M}|^2 \rangle$  explains 99.7% of the variance in the data.

Run "bhabha-scattering-4.txt" to verify.

The following table shows DESY-PETRA Bhabha scattering data obtained from HEP Data. 1

| x       | y       |
|---------|---------|
| -0.73   | 0.10115 |
| -0.6495 | 0.12235 |
| -0.5495 | 0.11258 |
| -0.4494 | 0.09968 |
| -0.3493 | 0.14749 |
| -0.2491 | 0.14017 |
| -0.149  | 0.1819  |
| -0.0488 | 0.22964 |
| 0.0514  | 0.25312 |
| 0.1516  | 0.30998 |
| 0.252   | 0.40898 |
| 0.3524  | 0.62695 |
| 0.4529  | 0.91803 |
| 0.5537  | 1.51743 |
| 0.6548  | 2.56714 |
| 0.7323  | 4.30279 |

Data x and y have the following relationship with the cross section model.

$$x = \cos \theta$$
  $y = \frac{d\sigma}{d\Omega}$ 

The differential cross section for Bhabha scattering is

$$\frac{d\sigma}{d\Omega} = \frac{\langle |\mathcal{M}|^2 \rangle}{64\pi^2 s} = \frac{\alpha^2}{2s} \left( \frac{s^2 + u^2}{t^2} + \frac{2u^2}{st} + \frac{t^2 + u^2}{s^2} \right)$$

The predicted cross section  $\hat{y}$  is computed from data x and beam energy E as

$$\hat{y} = \frac{\alpha^2}{2s} \left( \frac{s^2 + u^2}{t^2} + \frac{2u^2}{st} + \frac{t^2 + u^2}{s^2} \right) \times (\hbar c)^2 \times 10^{37}$$

where

$$s = 4E2$$
  

$$t = -2E2(1 - x)$$
  

$$u = -2E2(1 + x)$$

Factor  $(\hbar c)^2$  converts the result to SI and factor  $10^{37}$  converts square meters to nanobarns.

The following table shows  $\hat{y}$  for  $E = 7.0 \,\text{GeV}$ .

<sup>1</sup>www.hepdata.net/record/ins191231 (Table 3, 14.0 GeV)

| x       | y       | $\hat{y}$ |
|---------|---------|-----------|
| -0.73   | 0.10115 | 0.110296  |
| -0.6495 | 0.12235 | 0.113816  |
| -0.5495 | 0.11258 | 0.120101  |
| -0.4494 | 0.09968 | 0.129075  |
| -0.3493 | 0.14749 | 0.141592  |
| -0.2491 | 0.14017 | 0.158934  |
| -0.149  | 0.1819  | 0.182976  |
| -0.0488 | 0.22964 | 0.216737  |
| 0.0514  | 0.25312 | 0.264989  |
| 0.1516  | 0.30998 | 0.335782  |
| 0.252   | 0.40898 | 0.44363   |
| 0.3524  | 0.62695 | 0.615528  |
| 0.4529  | 0.91803 | 0.9077    |
| 0.5537  | 1.51743 | 1.45175   |
| 0.6548  | 2.56714 | 2.60928   |
| 0.7323  | 4.30279 | 4.61509   |

The coefficient of determination  $\mathbb{R}^2$  measures how well predicted values fit the real data.

$$R^{2} = 1 - \frac{\sum (y - \hat{y})^{2}}{\sum (y - \bar{y})^{2}} = 0.995$$

The result indicates that the model  $d\sigma$  explains 99.5% of the variance in the data.

Run "bhabha-scattering-5.txt" to verify.

Here are a few notes about how the scripts work. In component notation the trace operators of the Casimir trick become sums over the repeated index  $\alpha$ .

$$\begin{split} f_{11} &= \left( (\not\!p_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!p_3 - m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\alpha \right) \left( (\not\!p_4 + m)^\alpha{}_\beta \gamma_\mu{}^\beta{}_\rho (\not\!p_2 + m)^\rho{}_\sigma \gamma_\nu{}^\sigma{}_\alpha \right) \\ f_{12} &= (\not\!p_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!p_2 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\tau (\not\!p_4 + m)^\tau{}_\delta \gamma_\mu{}^\delta{}_\eta (\not\!p_3 - m)^\eta{}_\xi \gamma_\nu{}^\xi{}_\alpha \\ f_{22} &= \left( (\not\!p_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!p_2 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\alpha \right) \left( (\not\!p_4 + m)^\alpha{}_\beta \gamma_\mu{}^\beta{}_\rho (\not\!p_3 - m)^\rho{}_\sigma \gamma_\nu{}^\sigma{}_\alpha \right) \end{split}$$

To convert the above formulas to Eigenmath code, the  $\gamma$  tensors need to be transposed so that repeated indices are adjacent to each other. Also, multiply  $\gamma^{\mu}$  by the metric tensor to lower the index.

$$\gamma^{\beta\mu}_{\phantom{\beta}\rho}$$
  $\rightarrow$  gammaT = transpose(gamma)  $\gamma^{\beta}_{\phantom{\beta}\mu\rho}$   $\rightarrow$  gammaL = transpose(dot(gmunu,gamma))

Define the following  $4 \times 4$  matrices.

Then for  $f_{11}$  we have the following Eigenmath code. The contract function sums over  $\alpha$ .

$$(\not\!p_1-m)^\alpha{}_\beta\gamma^{\mu\beta}{}_\rho(\not\!p_3-m)^\rho{}_\sigma\gamma^{\nu\sigma}{}_\alpha \quad \rightarrow \quad \text{T1 = contract(dot(X1,gammaT,X3,gammaT),1,4)} \\ (\not\!p_4+m)^\alpha{}_\beta\gamma_\mu{}^\beta{}_\rho(\not\!p_2+m)^\rho{}_\sigma\gamma_\nu{}^\sigma{}_\alpha \quad \rightarrow \quad \text{T2 = contract(dot(X4,gammaL,X2,gammaL),1,4)}$$

Next, multiply then sum over repeated indices. The dot function sums over  $\nu$  then the contract function sums over  $\mu$ . The transpose makes the  $\nu$  indices adjacent as required by the dot function.

$$f_{11} = \operatorname{Tr}(\cdots \gamma^{\mu} \cdots \gamma^{\nu}) \operatorname{Tr}(\cdots \gamma_{\mu} \cdots \gamma_{\nu}) \quad o \quad exttt{f11} = exttt{contract(dot(T1,transpose(T2)))}$$

Follow suit for  $f_{22}$ .

$$(\not\!\!p_1-m)^\alpha{}_\beta\gamma^{\mu\beta}{}_\rho(\not\!\!p_2+m)^\rho{}_\sigma\gamma^{\nu\sigma}{}_\alpha \quad \rightarrow \quad \text{T1 = contract(dot(X1,gammaT,X2,gammaT),1,4)} \\ (\not\!\!p_4+m)^\alpha{}_\beta\gamma_\mu{}^\beta{}_\rho(\not\!\!p_3-m)^\rho{}_\sigma\gamma_\nu{}^\sigma{}_\alpha \quad \rightarrow \quad \text{T2 = contract(dot(X4,gammaL,X3,gammaL),1,4)}$$

Hence

$$f_{22} = \operatorname{Tr}(\cdots \gamma^{\mu} \cdots \gamma^{\nu}) \operatorname{Tr}(\cdots \gamma_{\mu} \cdots \gamma_{\nu}) \quad \rightarrow \quad \texttt{f22 = contract(dot(T1,transpose(T2)))}$$

The calculation of  $f_{12}$  begins with

$$(\not\!p_1 - m)^\alpha{}_\beta \gamma^{\mu\beta}{}_\rho (\not\!p_2 + m)^\rho{}_\sigma \gamma^{\nu\sigma}{}_\tau (\not\!p_4 + m)^\tau{}_\delta \gamma_\mu{}^\delta{}_\eta (\not\!p_3 - m)^\eta{}_\xi \gamma_\nu{}^\xi{}_\alpha$$

$$\rightarrow \quad T = \text{contract(dot(X1,gammaT,X2,gammaT,X4,gammaL,X3,gammaL),1,6)}$$

Then sum over repeated indices  $\mu$  and  $\nu$ .

$$f_{12}=\mathrm{Tr}(\cdots\gamma^{\mu}\cdots\gamma^{\nu}\cdots\gamma_{\mu}\cdots\gamma_{\nu}) \quad o \quad exttt{f12} = exttt{contract(T,1,3))}$$