

## I. ANOMALY CANCELLATION

From [1]: física por la mañana We use  $f$  ( $f$ ) to denote the general  $U(1)_X$  generation-independent charge assignments of the field  $f_R$  ( $F_L$ ). The three linear anomalies in  $U(1)_X$  [2]

$$\begin{aligned} [SU(3)_C]^2 U(1)_X : & \quad [3u + 3d] - [3 \cdot 2q] = 0, \\ [SU(2)_L]^2 U(1)_X : & \quad -[2l + 3 \cdot 2q] = 0, \\ [U(1)_Y]^2 U(1)_X : & \quad \left[ (-2)^2 e + 3 \left( \frac{4}{3} \right)^2 u + 3 \left( -\frac{2}{3} \right)^2 d \right] - \left[ 2(-1)^2 l + 3 \cdot 2 \left( \frac{1}{3} \right)^2 q \right] = 0, \end{aligned} \quad (1)$$

allows to express three  $X$ -charges in terms of the other two

$$u = -e + \frac{2l}{3}, \quad d = e - \frac{4l}{3}, \quad q = -\frac{l}{3}. \quad (2)$$

The quadratic anomaly condition is automatically satisfied, while the mixed gauge-gravitational and cubic anomalies depend of any extra singlet quiral fermions of zero hypercharge, like the right-handed counterpart of the Dirac neutrinos. For  $N$  extra quiral fields with  $X$ -charge  $n_\alpha$ , these conditions read

$$[\text{Grav}]^2 U(1)_X : \sum_{\alpha=1}^N n_\alpha + 3(e - 2l) = 0, \quad [U(1)_X]^3 : \sum_{\alpha=1}^N n_\alpha^3 + 3(e - 2l)^3 = 0. \quad (3)$$

We choose the solutions with  $r \equiv e - 2l$ , such that

$$\sum_{\alpha=1}^N n_\alpha = -3r, \quad \sum_{\alpha=1}^N n_\alpha^3 = -3r^3. \quad (4)$$

The full set of anomaly free SM  $X$ -charges in terms of two parameters [2–4] that we choose as  $l$  and  $r$ , is just

$$u = -r - \frac{4l}{3}, \quad d = r + \frac{2l}{3}, \quad q = -\frac{l}{3}, \quad e = r + 2l, \quad h = -r - l. \quad (5)$$

where the condition in the charged lepton Yukawa couplings have been used to fix  $h$ , and is automatically consistent with the conditions in the quark Yukawa couplings. By setting  $l = 0$  in the previous equations, we can define the Abelian symmetry in which only the right-handed charged fermions have non-vanishing  $X$ -charges as  $U(1)_R$ . Then the general anomaly free two-parameter solution can be written as

$$X(r, l) = rR - lY. \quad (6)$$

If we now change  $f \rightarrow f' = f/r$  for all the charged fermion  $X$ -charges [4], the first set of anomaly cancellation conditions Eq. (1) remains invariant, and without loss of generality it is always possible

Fields	SU(2) <sub>L</sub>	U(1) <sub>Y</sub>	U(1) <sub>X</sub>	U(1) <sub>B-L</sub>	U(1) <sub>R</sub>	U(1) <sub>D</sub>	U(1) <sub>G</sub>	U(1) <sub>D</sub>
$L$	<b>2</b>	-1	$l$	-1	0	-3/2	-1/2	0
$d_R$	<b>1</b>	-2/3	$1 + 2l/3$	1/3	1	0	2/3	0
$u_R$	<b>1</b>	+4/3	$-1 - 4l/3$	1/3	-1	1	-1/3	0
$Q$	<b>2</b>	1/3	$-l/3$	1/3	0	1/2	1/6	0
$e_R$	<b>1</b>	-2	$1 + 2l$	-1	1	-2	0	0
$H$	<b>2</b>	1	$-1 - l$	0	-1	1/2	-1/2	0
$N_\alpha$	<b>1</b>	0	$n_\alpha$	$n_\alpha$	$n_\alpha$	$n_\alpha$	$n_\alpha$	$n_\alpha$

TABLE I: General one-parameter solution with some examples of rational solutions ( $X = B - L, R, D, G$  and  $\mathcal{D}$ ) for the radiative type-I seesaw realization of the effective operator  $\mathcal{O}_{6D}$  for Dirac neutrino masses. The last column requires the condition  $\sum_{\alpha=1}^N n_\alpha = 0$ .

to normalize the solutions such that the last set Eq. (4) is just

$$\sum_{\alpha=1}^N n'_\alpha = -3, \quad \sum_{\alpha=1}^N n'^3_\alpha = -3. \quad (7)$$

For example, the solution with  $r = 3$ :  $n_\alpha = (-2, -2, -4, -1)$  [3] can be easily normalized to the form in Eq. (7) with  $f \rightarrow f/3$  to  $n'_\alpha = (-2/3, -2/3, -4/3, -1/3)$  as used in Ref. [5]. In this way, without loss of generality, we will work with the normalized solution in terms of a single parameter [6–8] that we choose to be  $l$ , by setting  $r = 1$  as summarized in column U(1)<sub>X</sub> of Table I, which is just

$$X(l) = R - lY. \quad (8)$$

In particular, this includes the solution  $n_\alpha = (-4, -4, +5)$  [3].

## II. DISCRETE SYMMETRIES

The possible charge assignments to a field  $\phi$  under a  $Z_N$  symmetry are

$$1, w, w^2, \dots, w^{N-1}, \quad \text{with } w = \exp(i2\pi/N). \quad (9)$$

We assume the existence of  $k$  scalar complex fields  $\phi_\alpha$ ,  $\alpha = 1, 2, \dots, k$ , each transforming as

$$\phi_\alpha \sim w^\alpha, \quad \text{with } k \leq N/2. \quad (10)$$

Note that  $(w^\alpha)^* = w^{-\alpha} = w^{N-\alpha}$ .

Consider one  $U(1)_X$  which is broken by a singlet scalar field  $S_N$  and one dark scalar field  $\phi_1$  where the subscript indicates the  $X$ -charge of the field. In addition to the self-conjugate terms the gauge symmetry allows the following term [9]

$$\mathcal{L} \propto S_N^\dagger \phi_1^N + \text{h.c.} \quad (11)$$

When  $S_N$  acquires a vev, the Lagrangian preserves a discrete  $Z_N$  subgroup of  $U(1)_X$  under which the dark scalar field transform as

$$\phi \rightarrow \phi' = e^{2\pi i k/N} \phi_1, \quad k=0, 1, 2, \dots, N-1. \quad (12)$$

Then the remnant symmetry will be exactly conserved at all orders including non-renormalizable terms. Note that the  $Z_N$  symmetry can also appear accidentally. In this case, in contrast to the discrete gauge symmetries, these accidental symmetries may be violated at the non-renormalizable level, and consequently, the DM stability can be only granted at the renormalizable level [9].

For the  $Z_2$  imposed usually by hand, we can identify this as a gauge discrete symmetry associated to the Lagrangian of the Singlet Scalar Dark Matter model (SSDM),  $\phi_1$ , supplemented by the interaction with the singlet scalar  $S_2$  which spontaneously breaks some  $U(1)_X$  gauge symmetry. The relevant term in the Lagrangian is

$$\mathcal{L} \subset \lambda S_2^* \phi_1 \phi_1 + \text{h.c.} \quad (13)$$

Note that in this case  $\phi_1$  needs to be complex. However, this Lagrangian term induces a mass splitting between the real and imaginary components of  $\phi_1$ . In this way we can have the usual SSDM with a real scalar.

For  $Z_4$  as remnant of the spontaneous breaking of some  $U(1)_X$ , we can have two-component dark matter depending on the mass hierarchy of the dark sector [9, 10]. For larger  $Z_N$  see [9]. In [11] it is claimed that the heterogeneous self-interaction is a natural consequence of any 2DM or nDM models. In nDM models there can be various dark matter (DM) annihilation processes that are different from the standard DM annihilation process [12]. In particular it is possible to have an assisted freeze-out which further reduces the direct detection signals [13]

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