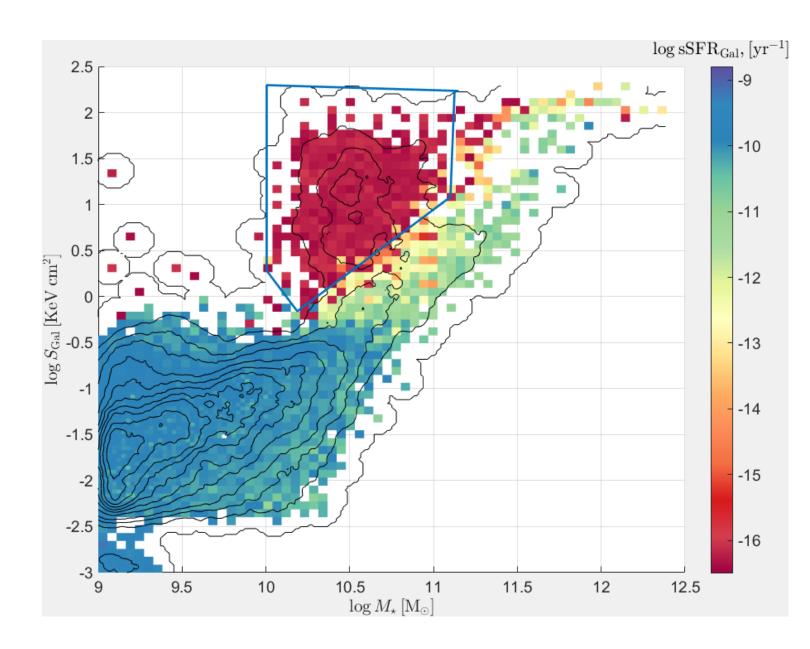
# Examining the importance of mass-matching

## set up the polygon which defines the 'quenched' group

```
10.0194];
       10.1379
                   9.8269
                            9.9108
                                     10.3255
                                               11.6090
                                                         11.0018
                                                                   10.0194
                                 -0.2410
%y=[ 2.3425 1.3433
                         0.0921
                                             2.1714
                                                       2.3425
                                                                 2.3785
                                                                           2.3785];
                             10.5476
                                                           10.0737
                                                                     10.1626 11.1202 ];
% X =[
         11.1202
                   11.0018
                                       10.3057
                                                 10.0293
                                       0.0291
        2.1624
                    0.9922
                             0.3531
                                                  0.3351
                                                         1.7034
                                                                     2.3695 2.1624 ];
% y =[
                       10.0046
                                                     11.1252];
x=[11.1499]
             10.0046
                                 10.1872
                                           11.1005
y = [2.2345]
             2.2975
                       0.2811
                                -0.1600
                                           1.0822
                                                     2.2345];
% define important parameters
galMass=tCoolStruct.galMass; % galaxy stellar mass
gasMass=tCoolStruct.inGal.gasMass(1,:);
gasEnt=tCoolStruct.inGal.meanEntMW(1,:);
%centralMask= subsInfo.isCentral(tCoolStruct.galMask);
sfr=subs.SubhaloSFRinRad; % sfr in galaxy
ssfr=sfr./galMass; % + 10^0.5*1e-17.*10.^(0.5.*rand(size(sfr)));
% plot the polygon
```



#### get mask from polygon

define the mass range according to the selcted group (via polygon) find the rest of the popultion in the mass range , thus defining two distinct groups in the same mass range

```
polyMask=inpolygon(log10(galMass),log10(gasEnt),x,y);
groupMask=polyMask & subsInfo.isCentral & tCoolStruct.galMask;
qGroupInd=find(groupMask);

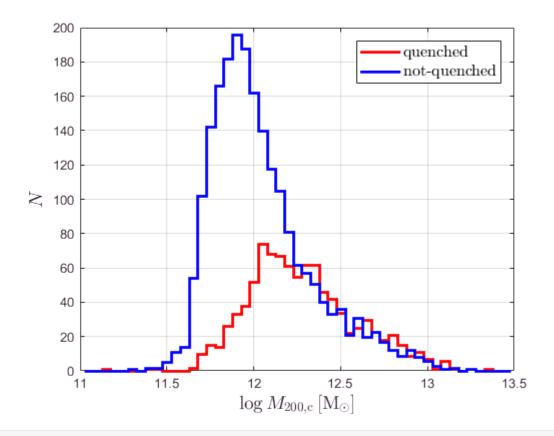
massRange=[min(galMass(groupMask)) max(galMass(groupMask))];
massMask=galMass>= massRange(1) & galMass<=massRange(2);

oGroupMask=massMask & subsInfo.isCentral & tCoolStruct.galMask & ~groupMask;
oGroupInd=find(oGroupMask);</pre>
```

#### show the distribution of host masses for the two groups

Here we show the distribution of the host masses for the galaxies in each gorup

```
qFofMass=fofs.Group_M_Crit200(subsInfo.hostFof(qGroupInd)+1).*illUnits.massUnit;
oFofMass=fofs.Group_M_Crit200(subsInfo.hostFof(oGroupInd)+1).*illUnits.massUnit;
nb=50;
binEdge=linspace(11,13.5,nb+1);
bins=binEdge(1:end-1)+0.5.*diff(binEdge);
qh=histcounts(log10(qFofMass),binEdge);
oh=histcounts(log10(oFofMass),binEdge);
figure
h(1)=stairs(bins,qh,'r','DisplayName','quenched','linewidth',2);
hold on
h(2)=stairs(bins,oh,'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log M_{\mathrm{200,c}}\\,[\mathrm{M_\odot}]$')
ylabelmine('$\log M_{\mathrm{200,c}}\\,[\mathrm{M_\odot}]$')
grid
hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthEast')
```

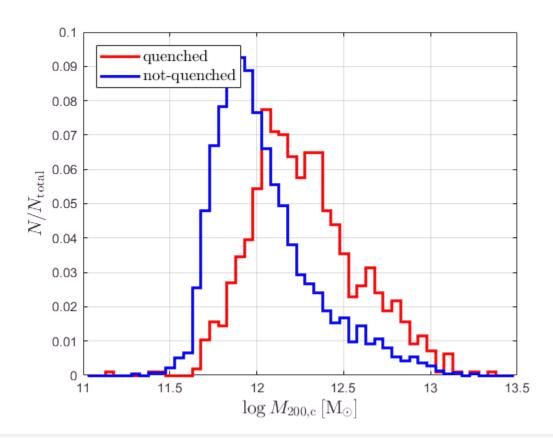


host-mass histogram for the two group which make up the mass range.

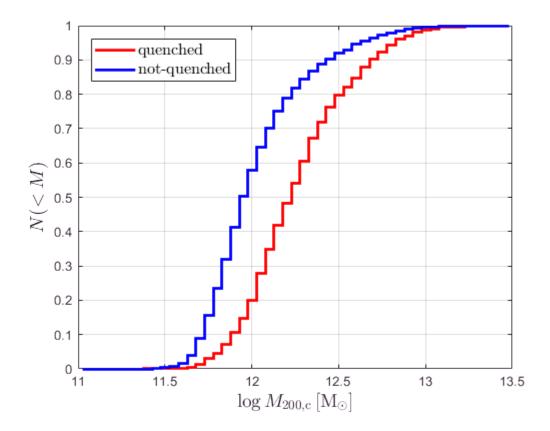
Since there is a different number of galaxies the histogram needs to be noramlized, as in the next figure.

```
figure
h(1)=stairs(bins,qh./sum(qh),'r','DisplayName','quenched','linewidth',2);
```

```
hold on
h(2)=stairs(bins,oh./sum(oh),'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log M_{\mathrm{200,c}}\,[\mathrm{M_\odot}]$')
ylabelmine('$N / N_{\mathrm{total}}$')
grid
hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthWest')
```



```
figure
h(1)=stairs(bins,cumsum(qh)./sum(qh),'r','DisplayName','quenched','linewidth',2);
hold on
h(2)=stairs(bins,cumsum(oh)./sum(oh),'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log M_{\mathrm{200,c}}\,[\mathrm{M_\odot}]$')
ylabelmine('$N(<M)$')
grid
hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthWest')</pre>
```

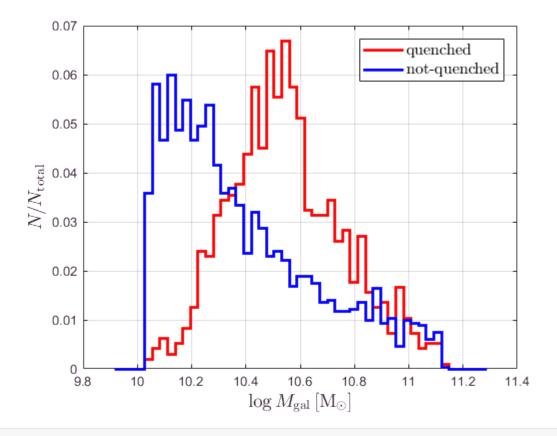


The cumulative function is also a good way to compare the distributions.

Sum-up: It appears that there is a distinct difference between the two distributions though, one must be carefule - the two groups do not sample the same mass range as can be seen in the next figure:

```
nb=50;
binEdge=linspace(9.9,11.3,nb+1);
bins=binEdge(1:end-1)+0.5.*diff(binEdge);
qh=histcounts(log10(galMass(qGroupInd)),binEdge);
oh=histcounts(log10(galMass(oGroupInd)),binEdge);

figure
h(1)=stairs(bins,qh./sum(qh),'r','DisplayName','quenched','linewidth',2);
hold on
h(2)=stairs(bins,oh./sum(oh),'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log M_{\mathrm{gal}}\,[\mathrm{M_\odot}]$')
ylabelmine('$N / N_{\mathrm{total}}\$')
grid
hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthEast')
```



It is this difference which may be responsible for the different host mass distributions (since they are known to be correlated). To examine this further we examine the mass-matched sample.

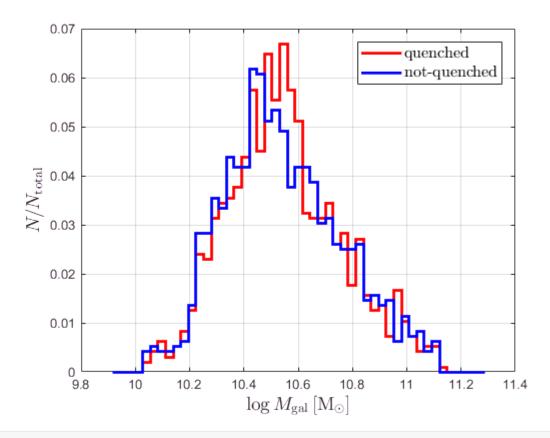
## examine mass-matching

```
matched=create_matched_sample(qGroupInd,oGroupInd,galMass,'range');

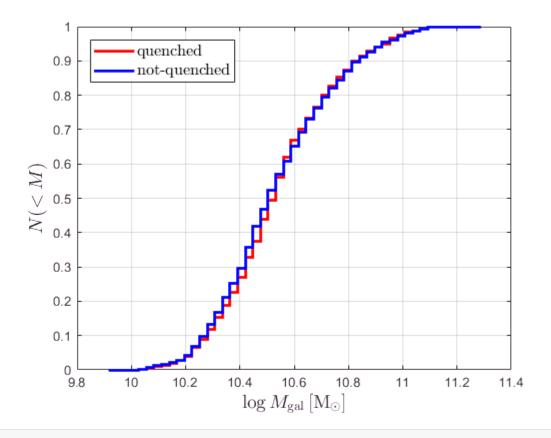
nb=50;
binEdge=linspace(9.9,11.3,nb+1);
bins=binEdge(1:end-1)+0.5.*diff(binEdge);

qh=histcounts(log10(galMass(qGroupInd)),binEdge);
oh=histcounts(log10(galMass(matched)),binEdge);

figure
h(1)=stairs(bins,qh./sum(qh),'r','DisplayName','quenched','linewidth',2);
hold on
h(2)=stairs(bins,oh./sum(oh),'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log M_{\mathrm{gal}}\,[\mathrm{M_\odot}]$')
ylabelmine('$\N / N_{\mathrm{total}}$')
grid
hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthEast')
```



```
figure
h(1)=stairs(bins,cumsum(qh)./sum(qh),'r','DisplayName','quenched','linewidth',2);
hold on
h(2)=stairs(bins,cumsum(oh)./sum(oh),'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log M_{\mathrm{gal}}\,[\mathrm{M_\odot}]$')
ylabelmine('$N(<M)$')
grid
hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthWest')</pre>
```



## Show host mass distributions from matched mass sample

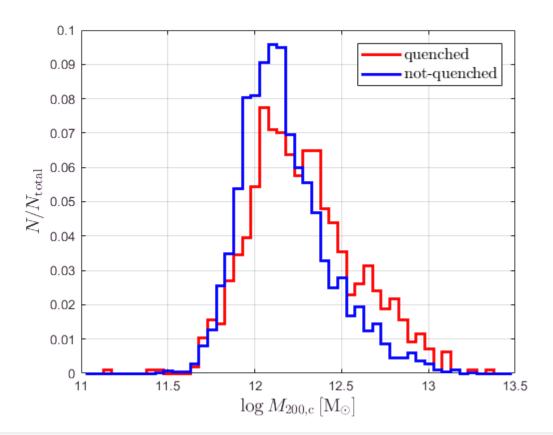
We perform mass matching 100 times and then average the resulting histogram.

```
nb=50;
binEdge=linspace(11,13.5,nb+1);
bins=binEdge(1:end-1)+0.5.*diff(binEdge);
qh=histcounts(log10(qFofMass),binEdge);

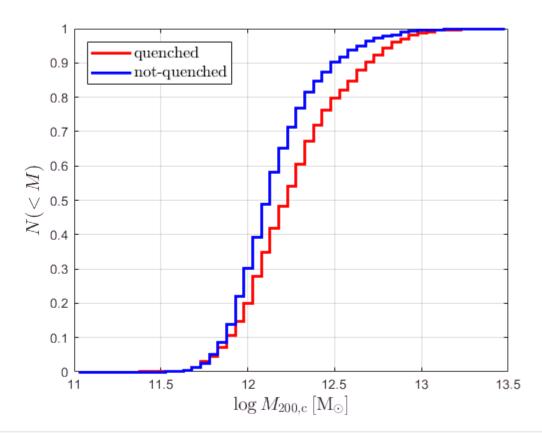
clear ohTemp
for i=1:100
    matched=create_matched_sample(qGroupInd,oGroupInd,galMass,'range');
    oFofMass=fofs.Group_M_Crit200(subsInfo.hostFof(matched)+1).*illUnits.massUnit;
    ohTemp(i,:)=histcounts(log10(oFofMass),binEdge);
end

oh=mean(ohTemp,1);
figure
h(1)=stairs(bins,qh./sum(qh),'r','DisplayName','quenched','linewidth',2);
hold on
```

```
h(2)=stairs(bins,oh./sum(oh),'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log M_{\mathrm{200,c}}\,[\mathrm{M_\odot}]$')
ylabelmine('$N / N_{\mathrm{total}}$')
grid
hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthEast')
```



```
figure
h(1)=stairs(bins,cumsum(qh)./sum(qh),'r','DisplayName','quenched','linewidth',2);
hold on
h(2)=stairs(bins,cumsum(oh)./sum(oh),'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log M_{\mathrm{200,c}}\,[\mathrm{M_\odot}]$')
ylabelmine('$N(<M)$')
grid
hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthWest')</pre>
```



```
%oh=
```

In the mass matched sample the difference between the host-mass distribution has become much smaller, and is probably not very significant.

## Show the distribution of BH masses for the two groups

Here we examine the difference in the BH mass distribution between the two groups. again, we begin with 2 groups which span the entire mass range, but are not mass-matched and then repear the analysis for a mass-matched sample.

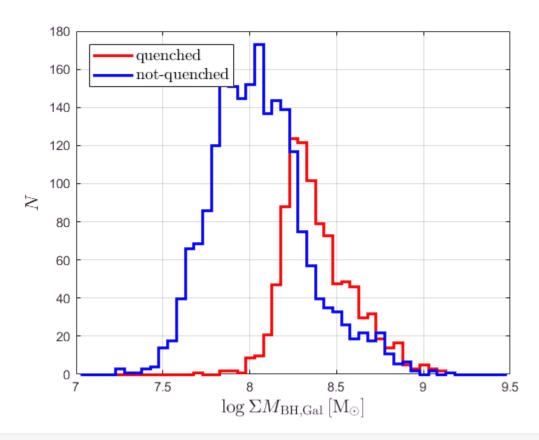
```
qBHMass=subs.SubhaloMassInRadType(illustris.partTypeNum('bh')+1,qGroupInd).*illUnits.massUnit;
oBHMass=subs.SubhaloMassInRadType(illustris.partTypeNum('bh')+1,oGroupInd).*illUnits.massUnit;

nb=50;
binEdge=linspace(7,9.5,nb+1);
bins=binEdge(1:end-1)+0.5.*diff(binEdge);

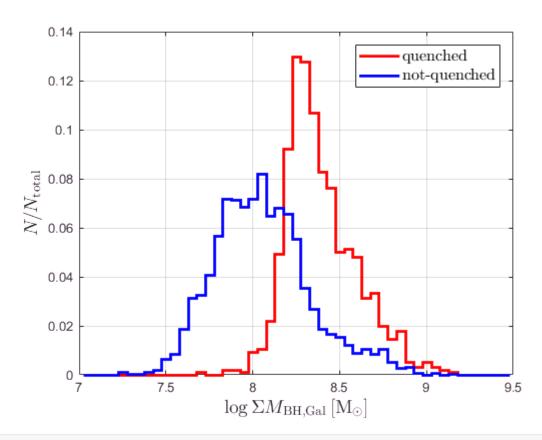
qh=histcounts(log10(qBHMass),binEdge);
oh=histcounts(log10(oBHMass),binEdge);

figure
h(1)=stairs(bins,qh,'r','DisplayName','quenched','linewidth',2);
hold on
h(2)=stairs(bins,oh,'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log \Sigma M_{\text{Mathrm}{BH,Gal}}\,[\mathrm{M_\odot}]$')
```

```
ylabelmine('$N$')
grid
hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthWest')
```

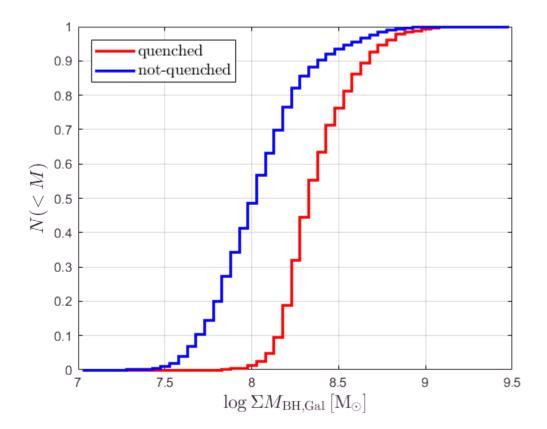


```
figure
h(1)=stairs(bins,qh./sum(qh),'r','DisplayName','quenched','linewidth',2);
hold on
h(2)=stairs(bins,oh./sum(oh),'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log \Sigma M_{\mathrm{BH,Gal}}\,[\mathrm{M_\odot}]$')
ylabelmine('$\ / \N_{\mathrm{total}}$')
grid
hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthEast')
```



```
figure
h(1)=stairs(bins,cumsum(qh)./sum(qh),'r','DisplayName','quenched','linewidth',2);
hold on
h(2)=stairs(bins,cumsum(oh)./sum(oh),'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log \Sigma M_{\mathrm{BH,Gal}}\,[\mathrm{M_\odot}]$')

ylabelmine('$N(<M)$')
grid
hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthWest')</pre>
```



Comparing the distributions shows a marked difference of the BH masses in the two groups. However, one must be cautious as we've seen before.

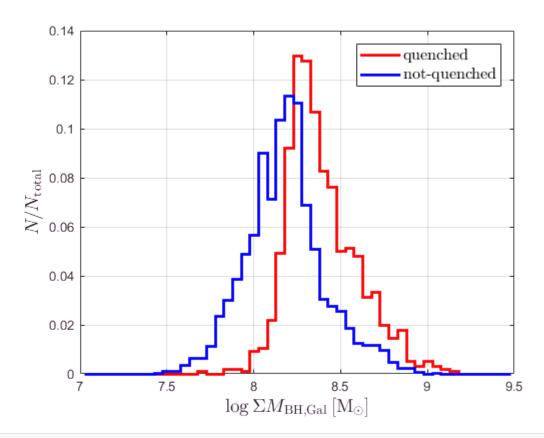
#### compare with mass-matched sample

```
for i=1:100

    matched=create_matched_sample(qGroupInd,oGroupInd,galMass,'range');
    oBHMass=subs.SubhaloMassInRadType(illustris.partTypeNum('bh')+1,matched).*illUnits.massUniohTemp(i,:)=histcounts(log10(oBHMass),binEdge);
end

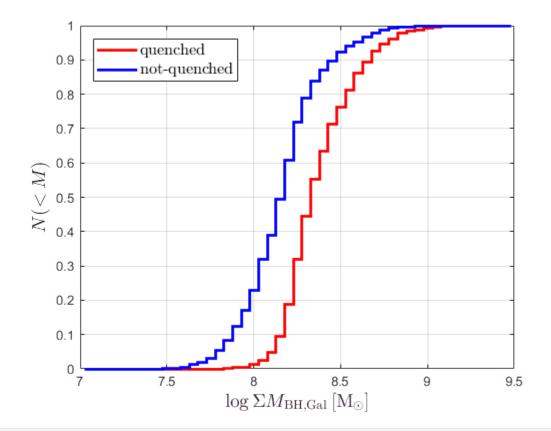
oh=mean(ohTemp,1);

figure
    h(1)=stairs(bins,qh./sum(qh),'r','DisplayName','quenched','linewidth',2);
hold on
    h(2)=stairs(bins,oh./sum(oh),'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log \Sigma M_{\{\mathrm{BH,Gal}}\,[\mathrm{M_\odot}]$')
ylabelmine('$\N / N_{\mathrm{total}}$')
grid
    hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthEast')
```



```
figure
h(1)=stairs(bins,cumsum(qh)./sum(qh),'r','DisplayName','quenched','linewidth',2);
hold on
h(2)=stairs(bins,cumsum(oh)./sum(oh),'b','DisplayName','not-quenched','linewidth',2);
xlabelmine('$\log \Sigma M_{\mathrm{BH,Gal}}\,[\mathrm{M_\odot}]$')

ylabelmine('$N(<M)$')
grid
hl=legend(h);
set(hl,'interpreter','latex','fontsize',12,'location','NorthWest')</pre>
```



Once again, comparing a mass-matched sample shows a smaller difference between the two distributions, but that difference is still there - though it may require some more rigorous analysis, it seems there is a significantly higher BH mass in the quenched galaxies versus a mass-matched sample.