

# Mass Sampling Documentation

## Functions

### Single Redshifts

#### Main function:

```
A. def mass_sampling(mass_range, redshift = 0.0, mdef = '200c', model = 'bocquet16', sample_num = 100000):
```

the function to give back a sample of single-redshift cluster mass distribution based on the halo mass function

#### Parameters:

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##### input:

**mass\_range:** a tuple of cluster masses, lower limit, and upper limit for sampling

**redshift:** a float, 0.0 by default

**sample\_num:** an integer of the number of samples, 100000 by default

**mdef:** The mass definition in which the halo mass  $M$  is given

**model:** the halo mass function model used by colossus

##### output:

**mass\_chain:** a NumPy array of length = sample\_num.

**test\_func:** the likelihood function

#### Library:

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- Colossus

#### step:

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- Initiate a NumPy array as cluster masses
- Use the halo mass function given by colossus, give back an array of number density
- Extract power from the mass array using helper 3, from  $10^{14} M_{\odot}$  to  $10^{15} M_{\odot}$
- Use helper 2 to calculate the final mass sampling

#### Helpers:

1. `lnpo(mass, min, max, test_fun):`

likelihood function used by MCMC

parameters:

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**mass:** a float or a 1d NumPy array of cluster mass in  $10^n M_\odot$  unit

2. `interpolate_MCMC(mass_array_p, mfunc_n, mass_range, sample_num, redshift):`

interpolate and normalize mfunc\_n, use the result as a likelihood function and perform MCMC method to get the sample.

parameters:

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input:

**mass\_arr\_p:** 1d NumPy array of cluster mass power (for example,  $10^{14} M_\odot$  represented as 14 in arr)

**mfunc\_n:** 1d NumPy array of halo number density \*  $10^5$

**mass\_range:** a tuple of cluster masses, lower limit, and upper limit for sampling

**sample\_num:** an integer of the number of samples

output:

**sample\_chain.flatten():** an 1d NumPy array of mass sampling, same unit as mass\_arr\_p

Library:

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- `scipy(interpolate & integrate)`
- `Emcee`

step:

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- Interpolate mass\_arr\_p & mfunc\_n, return a function object
- Calculate the integration of function between the mass range, normalize the function using integration, result in test\_func
- Use emcee package to do MCMC simulation, Helpers 1 (lnpo) used as the likelihood function, return test\_func and mass\_chain.flatten()

3. `extract_power(mass_arr):`

Function to extract the power of galaxy cluster mass array, switch from  $10^n$  to  $n$

**parameters:**

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**Input:**

**mass\_arr:** 1d NumPy array of cluster mass in  $10^n M_\odot$  unit

**output:**

**mass\_arr\_p:** 1d NumPy array of mass of power =  $n$

**step:**

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**mass\_arr\_p** = np.log10(mass\_arr)

## Multiple Redshifts

**Main function:**

```
B. def mul_redshift_mass_sampling(rs_dist = "skewnorm", rs_range =
(0.0, 1.5), mass_range = (14.0, 16.0), mdef = '200c', model = 'bocquet16',
sample_num = 100000, store = True):
```

the function to give back a sample of multi-redshift cluster mass distribution based on halo mass function

**Parameters:**

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**input:**

**rs\_dist:** a string, representing the distribution of cluster redshift, "skewnorm" by default

**rs\_range:** a tuple of redshift range, (0.0, 1.5) by default

**mass\_range:** a tuple of cluster masses, lower limit and upper limit for sampling, [min, max] in  $10^{\min} M_\odot$  unit

**mdef:** The mass definition in which the halo mass  $M$  is given; see colossus doc for more info ([https://bdiemer.bitbucket.io/colossus/lss\\_mass\\_function.html#lss.mass\\_function.massFunction](https://bdiemer.bitbucket.io/colossus/lss_mass_function.html#lss.mass_function.massFunction))

**model:** the halo mass function model used by colossus; see colossus doc for more info

**sample\_num:** an integer of the number of samples, 100000 by default  
**store:** a boolean, if True store mass array and redshift into a csv file and return a string of path to file if False returns None

**output:**

**fin\_cluster:** a Pandas dataframe with 2 columns of ["mass\_arr", "redshift"], NumPy array of cluster mass corresponding to redshift stored in each row

**filepath:** str of file path if store=True, else None

**Library:**

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- pandas

**Step:**

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- Call helper 2 to obtain the redshift interval (chop) and corresponding cluster number.
- Loop through every redshift in chop, call function A(mass\_sampling) to do single redshift mass sampling, store mass and redshifts in two pandas Series

## Helpers:

1. **skew\_sample(size = 10000):**

"""

the function to give back a sample of redshift based on skew gaussian distribution imitating SPT cluster data:

<https://pole.uchicago.edu/public/data/sptsz-clusters/>

**Parameters:**

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**input:**

**size:** integer, sample number

**output:**

**mass\_chain:** a NumPy array of length = sample\_num

**rs\_sample:** a 1d NumPy array of clusters' redshift sample with length = size

**Library:**

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- scipy.stats

**Step:**

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- Use `scipy.stats` to imitate redshift distribution
- Use `skewnorm.rvs` to draw a sample from it

2. `single_redshift_num(rs_range, sample_num, rs_dist_model):`

"""

the function to give back redshifts and `sample_num` per redshift for multi-redshift sampling

Parameters:

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input:

`rs_dist_model`: a string, representing the distribution of cluster redshift

`rs_range`: a tuple of redshift range, (0.0, 1.5) by default

`sample_num`: an integer of the number of samples, 100000 by default

output:

`chop`: a NumPy array of redshifts

`num_per_redshift`: a NumPy array of cluster num within the corresponding redshift interval of the same index number in chop

"""

Step:

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- Call helper 1 to obtain redshift distribution array
- Based on redshift range, divide it to several chops of redshift, each chop used as limit
- Loop through the redshifts array given by helper 1, approximate redshifts between two chops to the lower limit of it.
- Use another NumPy array of the same size to keep track of cluster number in each interval

# Procedure

Function A

[mass\_sampling]

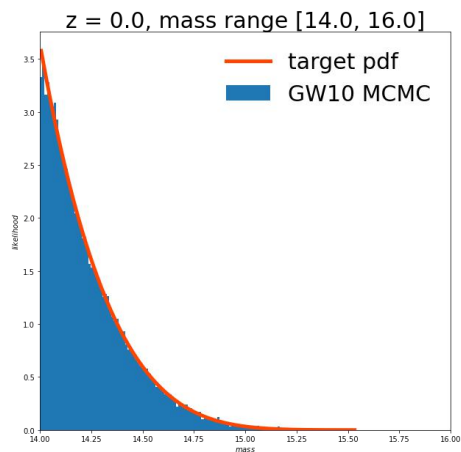
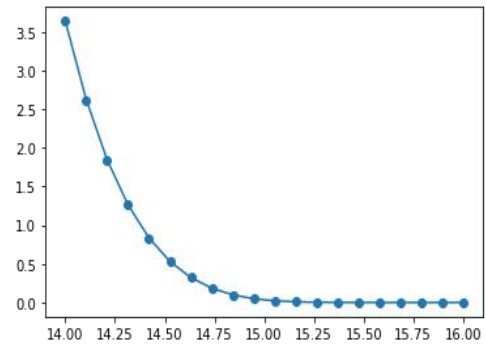
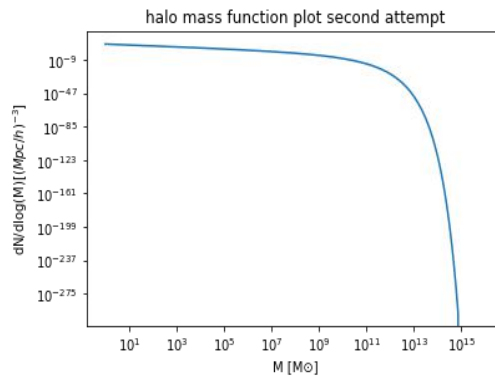
Single redshift mass sampling

Input:  
mass\_range, redshift

Colossus output:  
Mass array,  
number density

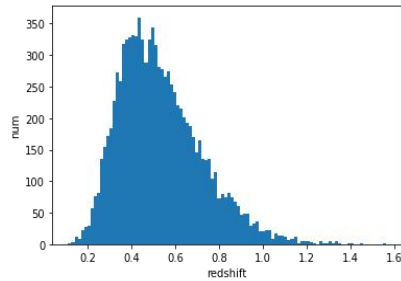
Interpolate &  
integrate output:  
Mass array,  
normalized likelihood

MCMC Output:  
function object and  
mass sample



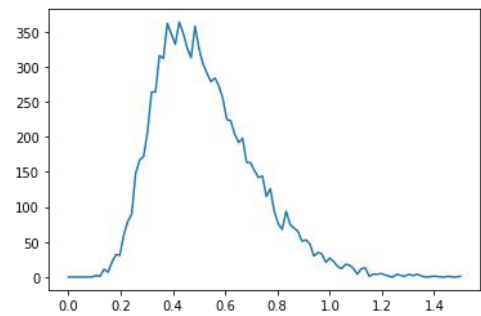
Input: mass\_range,  
redshift range,  
sample size

Function B  
[mul\_redshift\_mass\_sampling]  
Multiple redshift



[Skew\_sample] output:  
redshift array based on  
skewed gaussian  
distribution imitating

[single\_redshift\_num] output: a  
chop array of redshift interval, a  
num\_per\_redshift array of cluster  
number/redshift interval



Called [mass\_sampling] to  
sample for each redshift in  
chop

Output: a pandas dataframe  
with mass and redshift as  
two columns

