

# PyAutoFit: A Classy Probabilistic Programming Language for Model Composition and Fitting

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## Software

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## Summary

A major trend in academia and data science is the rapid adoption of Bayesian statistics for data analysis and modelling, leading to the development of probabilistic programming languages (PPL). A PPL provides a framework that allows users to easily specify a probabilistic model and perform inference automatically. PyAutoFit is a Python-based PPL aimed at model fitting problems tackled by long-term software development projects. By interfacing with all aspects of the modelling (e.g. the model, data, fitting procedure, visualization, results, etc.), PyAutoFit provides a more complete management of modeling than other PPLs. This includes composing high-dimensionality models from individual model components, customizing the fitting procedure and performing data augmentation before a model-fit. Advanced features include database tools for analysing large suites of modeling results and exploiting domain-specific knowledge of a problem via transdimensional model-fitting pipelines. Accompanying PyAutoFit is the [autofit workspace](#), which includes example scripts and the HowToFit lecture series which introduces non-experts to model-fitting and provides a guide on how to write a software project using PyAutoFit. To get started readers should go to our [readthedocs](#) and contact us to join the [PyAutoFit Slack channel](#) where we are building our online community.

## Background of Probabilistic Programming

Probabilistic programming languages (PPLs) have enabled contemporary statistical inference techniques to be applied to a diverse range of problems across academia and industry. Packages such as PyMC3 (Salvatier, Wiecki, & Fonnesbeck, 2016), Pyro (Bingham et al., 2019) and STAN (Carpenter et al., 2017) offer general-purpose frameworks where users can specify a generative model and fit it to data using a variety of non-linear fitting techniques. Each package is specialized to problems of a certain nature, with many focused on problems like generalized linear modeling or determining the distribution(s) from which the data was drawn. For these problems the model is typically composed of linear equations which are easily expressed syntactically, such that the PPL API offers an expressive way to define the model and extensions can be implemented in an intuitive and straightforward way.

## Software Description

PyAutoFit is a PPL whose core design is providing a direct interface with the model, data, fitting procedure and results, allowing it to provide more complete management of the model-fitting task than other PPLs and making it suited to longer term software projects. Model components are written as Python classes, allowing PyAutoFit to define the model and associated parameters in an expressive way that is tied to the modeling software's API. A

model fit then only requires that a `PyAutoFit Analysis` class is written, which combines the data, model and likelihood function and defines how the model-fit is performed using a non-linear search (e.g. `dynesty` (Speagle, 2020), `emcee` (Foreman-Mackey, Hogg, Lang, & Goodman, 2013) or `PySwarms` (Miranda, 2018)).

The `Analysis` class provides a model specific interface between `PyAutoFit` and the modeling software, allowing it to handle the ‘heavy lifting’ that comes with writing model-fitting software. This includes interfacing with the non-linear search, outputting results in a structured path format and model-specific visualization during and after the non-linear search. Results are output in a database structure with metadata that allows the `Aggregator` tool to load results post-analysis via a Python script or Jupyter notebook. This includes methods for summarizing the results of every fit, filtering results to inspect subsets of model fits and visualizing results. Results are loaded as Python generators, ensuring the `Aggregator` can be used to interpret large result datasets in a memory efficient way. `PyAutoFit` is therefore suited to ‘big data’ problems where independent fits to large homogeneous data-sets using an identical model-fitting procedure are performed.

## Model Abstraction and Composition

For many modeling problems the model comprises abstract model components representing objects or processes in a physical system. For example, galaxy morphology studies in astrophysics where model components represent the light profile of stars (Häußler et al., 2013) (Nightingale et al., 2019). For these problems the likelihood function is typically a sequence of numerical processes (e.g. convolutions, Fourier transforms, linear algebra) and extensions to the model often requires the addition of new model components in a way that is non-trivially included in the fitting process and likelihood function. Existing PPLs have tools for these problems, for example ‘black-box’ likelihood functions in `PyMC3`. However, these solutions decouple model composition from the data and fitting procedure, making the model less expressive, restricting model customization and reducing flexibility in how the model-fit is performed.

By writing model components as Python classes, the model and its associated parameters are defined in an expressive way that is tied to the modeling software’s API. Model composition with `PyAutoFit` allows complex models to be built from these individual components, abstracting the details of how they change model-fitting procedure from the user. Models can be fully customized, allowing adjustment of individual parameter priors, the fixing or coupling of parameters between model components and removing regions of parameter space via parameter assertions. Adding new model components to a `PyAutoFit` project is straightforward, whereby adding a new Python class means it works within the entire modeling framework. `PyAutoFit` is therefore ideal for problems where there is a desire to compose, fit and compare many similar (but slightly different) models to a single dataset, with the `Aggregator` including tools to facilitate this.

For many model fitting problems, domain specific knowledge of the model can be exploited to speed up the non-linear search and ensure it locates the global maximum likelihood solution. For example, initial fits can be performed using simplified model parameterizations, augmented datasets and faster non-linear fitting techniques. Through experience users may know that certain model components share minimal covariance, meaning that separate fits to each model component (in parameter spaces of reduced dimensionality) can be performed before fitting them simultaneously. The results of these simplified fits can then be used to initialize fits using a higher dimensionality model. Breaking down a model-fit in this way uses `PyAutoFit`’s [transdimensional model-fitting pipelines](#), which granularize the non-linear fitting procedure into a series of linked non-linear searches. Initial model-fits are followed by fits that gradually increase the model complexity, using the information gained throughout the pipeline to guide each non-linear search and thus enable accurate fitting of models of arbitrary complexity.

## History

PyAutoFit is a generalization of [PyAutoLens](#), an Astronomy package developed to analyse images of gravitationally lensed galaxies. Modeling gravitational lenses historically requires large amounts of human time and supervision, an approach which does not scale to the incoming samples of 100000 objects. Domain exploitation enabled full automation of the lens modeling procedure (Nightingale & Dye, 2015) (Nightingale, Dye, & Massey, 2018), with model customization and the aggregator enabling one to fit large datasets with many different models. More recently, PyAutoFit has been applied to calibrating radiation damage to charge coupled imaging devices and a model of cancer tumour growth.

## Workspace and HowToFit Tutorials

PyAutoFit is distributed with the [autofit workspace](#), which contains example scripts for composing a model, performing a fit and using the *Aggregator*. Also included are the HowToFit tutorials, a series of Jupyter notebooks aimed at non-experts, introducing them to model-fitting and Bayesian inference. They teach users how to write a *phase* package in PyAutoFit, which interfaces with their modeling software and ensures the design follows the concepts of object-oriented design which cleanly separates different parts of the modeling problem. The lectures can be viewed on our [readthedocs](#).

## Software Citations

PyAutoFit is written in Python 3.6+ (Van Rossum & Drake, 2009) and uses the following software packages:

- `corner.py` <https://github.com/dfm/corner.py> (Foreman-Mackey, 2016)
- `dynesty` <https://github.com/joshspeagle/dynesty> (Speagle, 2020)
- `emcee` <https://github.com/dfm/emcee> (Foreman-Mackey et al., 2013)
- `matplotlib` <https://github.com/matplotlib/matplotlib> (Hunter, 2007)
- `NumPy` <https://github.com/numpy/numpy> (van der Walt, Colbert, & Varoquaux, 2011)
- `PyMultiNest` <https://github.com/JohannesBuchner/PyMultiNest> (Feroz, Hobson, & Bridges, 2009) (Buchner et al., 2014)
- `PySwarms` <https://github.com/ljvmiranda921/pyswarms> (Miranda, 2018)
- `Scipy` <https://github.com/scipy/scipy> (Virtanen et al., 2020)

## Related Probabilistic Programming Languages

- `PyMC3` <https://github.com/pymc-devs/pymc3> (Salvatier et al., 2016)
- `Pyro` <https://github.com/pyro-ppl/pyro> (Bingham et al., 2019)
- `STAN` <https://github.com/stan-dev/stan> (Carpenter et al., 2017)
- `TensorFlow Probability` <https://github.com/tensorflow/probability> (Dillon et al., 2017)

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## References

- Bingham, E., Chen, J. P., Jankowiak, M., Obermeyer, F., Pradhan, N., Karaletsos, T., Singh, R., et al. (2019). Pyro: Deep universal probabilistic programming. *Journal of Machine Learning Research*, 20(Xxxx), 0–5. Retrieved from <http://arxiv.org/abs/1810.09538>
- Buchner, J., Georgakakis, A., Nandra, K., Hsu, L., Rangel, C., Brightman, M., Merloni, A., et al. (2014). X-ray spectral modelling of the AGN obscuring region in the CDFS: Bayesian model selection and catalogue. *Astronomy and Astrophysics*, 564, A125. doi:[10.1051/0004-6361/201322971](https://doi.org/10.1051/0004-6361/201322971)
- Carpenter, B., Gelman, A., Hoffman, M. D., Lee, D., Goodrich, B., Betancourt, M., Brubaker, M. A., et al. (2017). Stan: A probabilistic programming language. *Journal of Statistical Software*, 76(1). doi:[10.18637/jss.v076.i01](https://doi.org/10.18637/jss.v076.i01)
- Dillon, J. V., Langmore, I., Tran, D., Brevdo, E., Vasudevan, S., Moore, D., Patton, B., et al. (2017). TensorFlow Distributions. *arXiv e-prints*, arXiv:1711.10604. Retrieved from <http://arxiv.org/abs/1711.10604>
- Feroz, F., Hobson, M. P., & Bridges, M. (2009). MultiNest: An efficient and robust Bayesian inference tool for cosmology and particle physics. *Monthly Notices of the Royal Astronomical Society*, 398(4), 1601–1614. doi:[10.1111/j.1365-2966.2009.14548.x](https://doi.org/10.1111/j.1365-2966.2009.14548.x)
- Foreman-Mackey, D. (2016). Corner.py: Scatterplot matrices in python. *The Journal of Open Source Software*, 1(2), 24. doi:[10.21105/joss.00024](https://doi.org/10.21105/joss.00024)
- Foreman-Mackey, D., Hogg, D. W., Lang, D., & Goodman, J. (2013). emcee : The MCMC Hammer. *Publications of the Astronomical Society of the Pacific*, 125(925), 306–312. doi:[10.1086/670067](https://doi.org/10.1086/670067)
- Häußler, B., Bamford, S. P., Vika, M., Rojas, A. L., Barden, M., Kelvin, L. S., Alpaslan, M., et al. (2013). Megamorph - multiwavelength measurement of galaxy structure: Complete Sérsic profile information from modern surveys. *Monthly Notices of the Royal Astronomical Society*, 430(1), 330–369. doi:[10.1093/mnras/sts633](https://doi.org/10.1093/mnras/sts633)
- Hunter, J. D. (2007). Matplotlib: A 2D graphics environment. *Computing in Science & Engineering*, 9(3), 90–95. doi:[10.1109/MCSE.2007.55](https://doi.org/10.1109/MCSE.2007.55)
- Miranda, L. J. V. (2018). PySwarms, a research-toolkit for Particle Swarm Optimization in Python. *Journal of Open Source Software*, 3(21). doi:[10.21105/joss.00433](https://doi.org/10.21105/joss.00433)
- Nightingale, J. W., & Dye, S. (2015). Adaptive semi-linear inversion of strong gravitational lens imaging. *Monthly Notices of the Royal Astronomical Society*, 452(3), 2940–2959. doi:[10.1093/mnras/stv1455](https://doi.org/10.1093/mnras/stv1455)
- Nightingale, J. W., Dye, S., & Massey, R. J. (2018). AutoLens: Automated modeling of a strong lens's light, mass, and source. *Monthly Notices of the Royal Astronomical Society*, 478(4), 4738–4784. doi:[10.1093/mnras/sty1264](https://doi.org/10.1093/mnras/sty1264)
- Nightingale, J. W., Massey, R. J., Harvey, D. R., Cooper, A. P., Etherington, A., Tam, S. I., & Hayes, R. G. (2019). Galaxy structure with strong gravitational lensing: Decomposing the internal mass distribution of massive elliptical galaxies. *Monthly Notices of the Royal Astronomical Society*, 489(2), 2049–2068. doi:[10.1093/mnras/stz2220](https://doi.org/10.1093/mnras/stz2220)

- Salvatier, J., Wiecki, T. V., & Fonnesbeck, C. (2016). Probabilistic programming in Python using PyMC3. *PeerJ Computer Science*, 2016(4), 1–24. doi:[10.7717/peerj-cs.55](https://doi.org/10.7717/peerj-cs.55)
- Speagle, J. S. (2020). dynesty: a dynamic nested sampling package for estimating Bayesian posteriors and evidences. *Monthly Notices of the Royal Astronomical Society*, 493(3), 3132–3158. doi:[10.1093/mnras/staa278](https://doi.org/10.1093/mnras/staa278)
- van der Walt, S., Colbert, S. C., & Varoquaux, G. (2011). The numpy array: A structure for efficient numerical computation. *Computing in Science Engineering*, 13(2), 22–30.
- Van Rossum, G., & Drake, F. L. (2009). *Python 3 reference manual*. Scotts Valley, CA: CreateSpace. ISBN: [1441412697](https://www.isbn-international.org/product/9781441412697)
- Virtanen, P., Gommers, R., Oliphant, T. E., Haberland, M., Reddy, T., Cournapeau, D., Burovski, E., et al. (2020). SciPy 1.0: Fundamental Algorithms for Scientific Computing in Python. *Nature Methods*, 17, 261–272. doi:<https://doi.org/10.1038/s41592-019-0686-2>