

# Global existence for quasilinear wave equations

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# Motivation and Examples

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# Motivation

Let  $(M, g)$  be an asymptotically flat  $d + 1$  dimensional Lorentzian manifold, and consider the equation

$$\begin{cases} g^{\alpha\beta} \partial_{\alpha\beta}^2 \phi = \mathcal{N}(\partial\phi, \partial\phi) \\ \phi(t=0) = f \\ \partial_t \phi(t=0) = g \end{cases}$$

where  $\alpha, \beta = 0, 1, 2, \dots, d$  range over Cartesian coordinates and  $\mathcal{N}$  is a quadratic nonlinearity.

What can we say about solutions to the equation?

# Motivation

Let  $(M, g)$  be an asymptotically flat  $d + 1$  dimensional Lorentzian manifold, and consider the equation

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What can we say about solutions to the equation?

Our primary motivations will be from the fields of **general relativity** and **compressible fluids**, though similar equations show up in e.g. electromagnetism or gauge theory.

Primary regime of interest is small data and  $d = 3$ , for reasons that will be expanded on later.

## Einstein vacuum

$$\text{Ric}_{\mu\nu}(g) = 0.$$

In appropriate coordinate system, takes the form

$$\begin{aligned} -\frac{1}{2}(g^{-1})^{\alpha\beta}\partial_{\alpha\beta}g_{\mu\nu} &+ \frac{1}{2}(g^{-1})^{\alpha\sigma}(g^{-1})^{\beta\rho}\partial_{\mu}g_{\sigma\rho}\partial_{\beta}g_{\alpha\nu} \\ &+ \frac{1}{2}(g^{-1})^{\alpha\sigma}(g^{-1})^{\beta\rho}\partial_{\nu}g_{\sigma\rho}\partial_{\alpha}g_{\beta\mu} \\ &- \frac{1}{2}(g^{-1})^{\alpha\sigma}(g^{-1})^{\beta\rho}\partial_{\mu}g_{\sigma\rho}\partial_{\nu}g_{\alpha\beta} \\ &+ F_{\mu\nu}(g, \partial g) = 0 \end{aligned}$$

with flat solution given by  $g = \text{diag}(-1, 1, 1, 1)$  (Minkowski).

## Compressible Euler

$$\begin{cases} D_t \rho = -\rho \nabla \cdot u \\ D_t u = -\frac{\nabla p}{\rho} \\ D_t s = 0 \end{cases}$$

where  $\rho$  is the density,  $p = p(\rho)$  is the pressure,  $u$  is the velocity,  $s$  is the entropy, and  $D_t := \partial_t + u \cdot \nabla$  is the material derivative. (Can be rewritten as a wave equation for the velocity and logarithmic density).

## The linear wave equation

$$\square \phi := (-\partial_t^2 + \Delta) \phi = 0$$

## 1. Stability of Minkowski:

- Christodulu, Klainerman '93 [1]
- Lindblad, Rodnianski '10 [2]
- Keir '18 [3]
- Shen '23 [4]
- Dafermos, Rodnianski '10 [5]
- Hintz, Vasy '20 [6]
- ...

## 2. Shock formation in Euler:

- Christodoulou '07 [7].
- Speck, Holzegel, Luk, Wong '16 [8]
- ...

## 3. Wave equations: Lindblad '08 [9], John '81 [10], Yu '24 [11],

...



# Review of the linear wave equation

Key properties of the linear wave equation:

1. Conservation of energy:  $\|\partial\phi(t, \cdot)\|_{L^2} = \|\partial\phi(0, \cdot)\|$ .
2. **Finite speed of propagation**: if initial data is  $\equiv 0$  in “backward light cone”, then solution is 0.
3. **Dispersive decay**: Near the “wave zone”  
 $\{r \approx t\}, |\partial\phi| \sim t^{-(d-1)/2}$

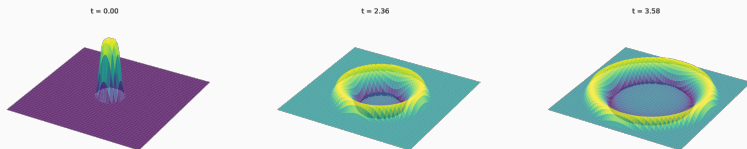
Can be read off from solutions using the fundamental solution, but for quasilinear problems, need robust methods of e.g. proving decay.

$d = 3$  corresponds to critical rate of  $t^{-1}$  decay.

# What does a free wave look like?

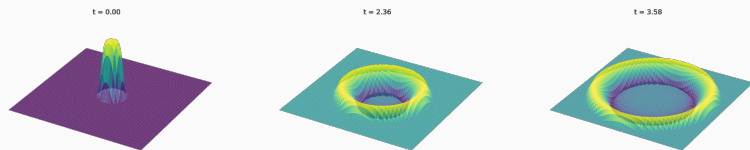
Why  $t^{-(d-1)/2}$  decay?

Heuristic picture: solution begins life with support  $\approx B(0, 1)$ , and is propagated along “forward light cone.”



After time  $t$ , solution is supported on annulus of radius  $\approx t$ , so  $|\text{supp } \phi(t, \cdot)| \approx t^{d-1}$ .

# Why $(d-1)/2$ decay?



After time  $t$ , solution is supported on annulus of radius  $\approx t$ , so  $|\text{supp } \phi(t, \cdot)| \approx t^{d-1}$ . On the other hand,  $\|\partial \phi(t, \cdot)\|_{L^2}$  is conserved, which is consistent with decay rate above.

Can get improved decay rate for derivatives in direction of propagation, which will be critical in analyzing the nonlinearity.

# The semilinear problem

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# General strategy to prove global existence

General method for obtaining global existence for initial data of size  $\varepsilon^{3/2}$ :

1. Identify suitable set of **weighted** vector fields  $\Gamma \subseteq TM$  and commute to obtain equations for  $\Gamma^k \phi$ .
2. **Bootstrap:** assume energy estimates bounds of the form  $\|\Gamma^k \partial \phi\|_{L^2} \leq \mathcal{O}(\varepsilon)$  hold up to some time  $T_* < \infty$ .
3. Use weighted vector fields to obtain pointwise estimates **with improved decay**: e.g.  $|\partial \phi| \lesssim \varepsilon/t^{1+\delta}$ .
4. Use this to show  $\|\mathcal{N}\|_{L^1 L^2} = \mathcal{O}(\varepsilon^2)$ , hence **can improve energy estimates**:

$$\|\Gamma^k \partial \phi\|_{L^2} \leq \mathcal{O}(\varepsilon^{3/2}) + \mathcal{O}(\varepsilon^2) \ll \mathcal{O}(\varepsilon).$$

5. Conclude  $T_* = \infty$ .

Consider the equation

$$\square\phi = (\partial_t\phi)^2.$$

In [10], John showed that there are *no nontrivial global solutions* when  $d = 3$ .

What goes wrong?

## Closing (or not) the bootstrap

Functional estimate: if  $\square\phi = F$ , then

$$\|\partial\phi(T, \cdot)\|_{L^2} \leq \|\partial\phi(0, \cdot)\|_{L^2} + \|F\|_{L^1([0, T]; L^2)}.$$

Suppose  $\|\partial\phi\|_{L^2} \lesssim \varepsilon \implies |\partial\phi| \leq \varepsilon t^{-1}$ . Then

$$\left\|(\partial_t\phi)^2\right\|_{L^1L^2} \leq \|\partial\phi\|_{L^1L^\infty} \|\partial\phi\|_{L^\infty L^2} \lesssim \varepsilon^2 \log T \neq \mathcal{O}(\varepsilon).$$

Still can show “almost global” existence, e.g.  $T_* \approx e^{\mathcal{O}(1/\varepsilon)}$ .

Actually sharp; John’s proof shows  $T_* \leq e^{\mathcal{O}(1/\varepsilon)}$ .

Consider now Nirenberg's example: the equation

$$\square \phi = (\partial_t \phi)^2 - |\nabla \phi|^2.$$

All small data solutions are now global (via change of variable  $\psi = e^\phi - 1$ ), but why?



## A return to the linear wave equation

Consider choosing coordinates  $u := t - |x|$ ,  $v := t + |x|$ ,  $\theta \in \mathbb{S}^2$  on  $\mathbb{R}^{3+1}$ .

The vector fields  $\partial_v, \partial_\theta$  are tangential to the *forward light cone*  $\{u = u_0\}$ .

In this frame, the nonlinearity above *factors* as

$$4\partial_u\phi\partial_v\phi - |\nabla\phi|^2$$

which **decays at rate**  $t^{-3/2}$ .

Can extend this method to prove small data global existence for any equation satisfying **null condition** and some equations satisfying “**weak null condition**” say for data in  $C_c^\infty(B(0,1))$ . Weakest possible assumption is *energy class*; that is, solutions for which  $\|\Gamma^k \partial \phi\|_{L^2} < \infty$ .

Next question: what's the weakest type of decay for which global existence still holds?

Actually necessary for some applications: **positive mass theorem** says that any compactly supported perturbation of Minkowski is Minkowski.

(Usual working assumption is Schwarzschild or  $\mathcal{O}(1/t)$  tails).

We consider data satisfying the decay condition

$$|\partial^\alpha \phi(0, x)| \leq |x|^{-\delta-|\alpha|} \quad (1)$$

for any  $\delta > 0$ . Roughly speaking: decay like  $r^{-\delta}$ , and every derivative gains you a power. Now

$$\|\partial \phi\|_{L^2(B(0,R))} \approx \varepsilon R^{1/2-\delta},$$

which is far from bounded.

Now have to worry about following issues:

1. How do you do energy estimates? Local well posedness?
2. Weaker pointwise estimates:  $|\partial\phi| \lesssim t^{-1/2-\delta}$ ,  $|\bar{\partial}\phi| \lesssim t^{-1-\delta}$  decay, which makes the nonlinearity much worse than  $L^1L^2$ .

## Theorem

*Let  $d = 3$ ,  $N \geq 15$ ,  $\delta > 0$ , and fix an equation satisfying the null or weak null conditions. There exists  $\varepsilon_0 > 0$  such that if initial data satisfies the target decay condition with  $\varepsilon < \varepsilon_0$  for all  $|\alpha| \leq N$ , then the solution exists globally in time. Furthermore, we have the decay rates*

1.  $|\partial\phi| \lesssim \varepsilon t^{-1},$
2.  $|\bar{\partial}\phi| \lesssim \varepsilon t^{-1-\delta/2}.$

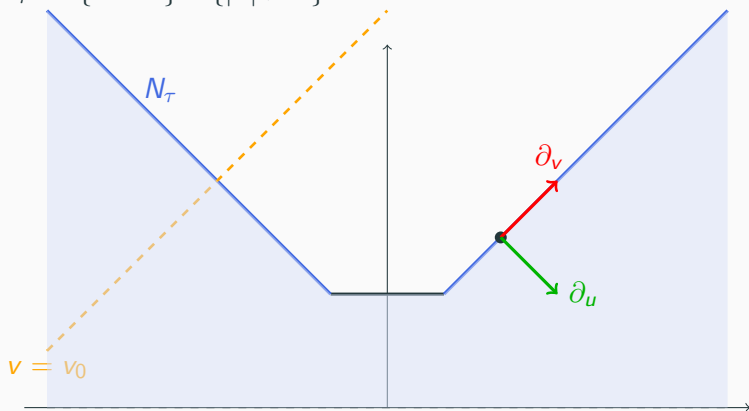
Remark: there is a much simpler Picard iteration-type proof for this result, but it requires  $N = \mathcal{O}(1/\delta)$  derivatives, and uses techniques very different than what we will be using for the quasilinear problem.

# Energy estimates

Geometric setup: recall we've chosen coordinates  $u := t - |x|$ ,  $v := t + |x|$ ,  $\theta \in \mathbb{S}^2$  on  $\mathbb{R}^{3+1}$ .

Consider foliating  $\mathbb{R}^{3+1}$  by surfaces of the form

$$N_\tau := \{u = \tau\} \cap \{|x| > 1\}.$$



# The energy identity

## Theorem

Suppose  $\phi$  is a solution to  $\square\phi = F$ . Then

$$\int_{Bulk} \partial_t \phi F = \int_{\partial} T_{\phi}(\partial_t, \nu).$$

Here  $\nu$  is the unit normal, and  $T_{\phi}$  is a 2-tensor given by

$$T_{\phi}(X, Y) = X\phi Y\phi - \langle X, Y \rangle \langle \nabla\phi, \nabla\phi \rangle / 2.$$

Key fact: when boundary component is  $\{t = t_0\}$ , RHS is

$$(\partial_t \phi)^2 + \sum_i (\partial_i \phi)^2,$$

and when boundary is  $N_{\tau}$ , RHS is

$$(\partial_{\nu} \phi)^2 + |\nabla \phi|^2.$$

*“The energy identity is the only method known to man that does not lose derivatives.” - J. Zhao.*



How do you actually obtain pointwise estimates?

Consider the following set of vector fields, which generate (conformal) isometries of Minkowski:

1. Translations:  $\partial_t, \partial_i$ .
2. **Scaling:**  $S := t\partial_t + x^i\partial_i = u\partial_u + v\partial_v$ .
3. Rotations:  $\Omega_{ij} := x_i\partial_j - x_j\partial_i$ .
4. Lorentz boosts:  $\Omega_{0j} := t\partial_j + x_j\partial_t$ .

## Theorem (Klainerman-Sobolev)

*Let  $\Gamma$  be the full set of commuting vector fields. For any sufficiently smooth  $\phi$ , we have the pointwise bounds*

$$|\partial\phi(u, v, \theta)| \lesssim u^{-1/2} v^{-1} \sum_{|\alpha| \leq 4} \|\Gamma^\alpha \partial\phi\|_{L^2}.$$

*Furthermore,*

$$\left| \bar{\partial}\phi(u, v, \theta) \right| \lesssim v^{-3/2} \sum_{|\alpha| \leq 4} \|\Gamma^\alpha \partial\phi\|_{L^2}.$$

We actually prefer to use an elliptic estimate that requires commuting with fewer vector fields.

## Theorem (Luk, Oh '23)

Set  $\Gamma := \{S, \partial_t, \Omega_{ij}\}$ . Fixing  $U \lesssim R$ , set  $A := \{u \sim U, r \sim R\}$  and set  $B$  to be a enlarged copy of  $A$ . Also let  $s := |\alpha| + |\beta|$ . Then we have the estimate

$$\begin{aligned} & \left\| (u\partial_u)^\alpha (r\bar{\partial})^\beta \phi \right\|_{L^2(A)} \\ & \lesssim \left\| \Gamma^{\leq s} \phi \right\|_{L^2(B)} + UR \left\| (u\partial_u)^{\leq s} (r\bar{\partial})^{\leq s} \square \phi \right\|_{L^2(B)}. \end{aligned}$$

# Ideas for the elliptic estimates

We combine the elliptic estimates above with the following rescaled Sobolev inequality:

## Theorem

$$\|\phi\|_{L^\infty(B_R)} \lesssim R^{-d/2} \sum_{|\alpha| \leq (d+1)/2} \|(R\partial)^\alpha \phi\|_{L^2(B_{2R})}$$

*in all odd space dimensions  $d$ .*

This implies the estimates

$$\begin{aligned} & \|\phi\|_{L^\infty(\{u \sim U, r \sim R\})} \\ & \lesssim R^{-3/2} U^{-1/2} \sum_{|\alpha|+|\beta|+|\gamma| \leq 5} \left\| (u\partial_u)^\alpha (r\partial_r)^\beta \Omega^\gamma \phi \right\|_{L^2(\{u \sim U, r \sim R\})}. \end{aligned}$$

## Finishing the pointwise estimates

Suppose we are in the region  $r \gg t$ . We estimate derivatives in the following way.

- **Good derivatives.** By conservation of energy, the integral of  $|\Gamma^k \bar{\partial} \phi|$  over any  $N_\tau$  is of size comparable to data. By the elliptic estimate and the rescaled Klainerman-Sobolev estimate on  $N_\tau$ , this implies  $|\bar{\partial} \phi| \lesssim r^{-3/2} \|\partial \phi\| \lesssim r^{-1-\delta}$ .
- **Bad derivatives.** Using the elliptic estimate over a  $u \sim U, r \sim R$  estimate gives  $|\partial \phi| \lesssim r^{-1} u^{-1/2} \lesssim r^{-1/2-\delta}$  since we lose a factor of  $r$  integrating over spacetime instead of just a surface.

Similar but more complicated elliptic estimates for region  $t \gg r$ .

We still need to recover  $t^{-1}$  decay for  $\partial_u \phi$ . Although the estimates above only yield a decay rate of  $t^{-1/2+\delta}$ , the key thing to note is that  $\partial_u$  satisfies a transport equation which allows us to improve this decay a posteriori. In particular, we have that

$$\partial_v(r\partial_u\phi) = r\Delta\phi + \partial_v\phi + r\mathcal{N}(\partial\phi, \partial\phi) = \mathcal{O}(r^{-1-\delta})$$

which is integrable.

## Tools for the quasilinear problem

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# Tools for the quasilinear problem

We move now to the realm of quasilinear equations.

## Model equation

$$-(1 + \phi)\partial_t^2\phi + \Delta\phi = 0.$$

(Corresponds to  $g = \text{diag}(-1/(1 + \phi), 1, 1, 1)$ ).

New issues: derivative loss, additional term  $\phi\partial_t^2\phi$ .



# Quasilinear equations

First attempt: just put the quasilinear term  $\phi \partial_t^2 \phi$  into error.  
Ignoring loss of derivatives, we have

$$\left\| \phi \partial_t^2 \phi \right\|_{L^1 L^2} \lesssim \left\| \phi \right\|_{L^1 L^\infty} \left\| \partial_t^2 \phi \right\|_{L^\infty L^2} \quad \text{or} \quad \left\| \phi \right\|_{L^\infty L^2} \left\| \partial_t^2 \phi \right\|_{L^1 L^\infty}$$

but we only have  $|\phi| \lesssim v^{-\delta}$ , and  $\left\| \partial_t^2 \phi \right\|_{L^2}$  is at best bounded.  
Similarly, the best pointwise decay for  $\partial_t^2 \phi$  we can get is  $t^{-1}$ , and  $\phi$  isn't bounded in energy.

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Similarly, the best pointwise decay for  $\partial_t^2 \phi$  we can get is  $t^{-1}$ , and  $\phi$  isn't bounded in energy.

Need derivatives on both terms in order to gain decay.

**Solution:** construct specialized vector fields  $\Gamma$  depending on the metric  $g$  that commute better with the equation.

## Theorem

*Let  $d = 3$ ,  $N \geq 20$ , and  $\delta > 0$ . There exists  $\varepsilon_0 > 0$  such that if the initial data satisfies the target decay condition with  $\varepsilon < \varepsilon_0$  for all  $|\alpha| \leq N$ , then the solution to the model equation exists globally in time (with similar decay rates to the semilinear case).*

## Theorem

*Let  $d = 3$ ,  $N \geq 20$ , and  $\delta > 0$ . There exists  $\varepsilon_0 > 0$  such that if “initial data” to the Einstein vacuum equations satisfies the target decay condition with  $\varepsilon < \varepsilon_0$  for all  $|\alpha| \leq N$ , then the solution to the EVE exists globally in time, is future geodesically complete, and decays asymptotically back to the Minkowski solution.*

How do we construct vector fields that “see” the geometry of  $g$ ?

- Construct an analogue of Minkowski  $t - r$  by solving eikonal equation  $\langle \nabla u, \nabla u \rangle = 0$  with initial data on appropriately chosen hypersurface.
- Let  $\mu^{-1} := \langle \nabla u, \nabla r \rangle$  (inverse foliation density) and define  $L := \mu \nabla u$  so that  $L(r) = 1$  (analogue of “good derivative”  $\partial_v$ ; note  $\mu \equiv 1$  in Minkowski).
- Use  $(u, r, \theta)$  coordinates, where  $\theta$  are angular coordinates propagated by  $L$ .

- The frame  $\{L, \partial_\theta\}$  spans the tangent space to constant  $u$  hypersurfaces.
- Complete the frame with vector field  $T$ , chosen to satisfy normalization/orthogonality conditions. (Analogue of Minkowski  $\partial_t$ ).
- Define “scaling vector field”  $S := rL + u\mu T$ , and commute with  $\Gamma := \{S, T, r^2 \nabla\}$ .
- In these coordinates,

$$\tilde{\square}_g \phi = -L(-L + 2T)\phi + \Delta \phi - \text{tr}_g \chi T\phi - \text{tr}_g \kappa L\phi - \zeta^\sharp \phi$$

which allows us to easily compute commutators.

Additional geometric quantities that appear in the estimates:

1. Second fundamental forms  $\chi, \kappa$  associated to constant  $u, r$  “spheres.”
2. Torsion  $\zeta$  associated to null hypersurfaces of constant  $u$  (0 on Minkowski).
3. Gauss curvature  $K$  used for elliptic estimates.
4. **Ricci curvature**  $\text{Ric}_{\mu\nu}$  appears when differentiating  $\chi, \kappa$  terms.

Variety of technical difficulties associated to losing derivatives, but eventually can get estimates to close.

**Thanks!**



**Thanks!**  
**Questions?**

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It is also useful to have estimates for bulk energy terms. We state this estimate below.

## Theorem (Morawetz)

$$\begin{aligned} \sup_{k \in \mathbb{N}} \int_{r \sim 2^k} \left( r^{-1} \phi^2 + r |\partial \phi|^2 \right) r^{-2} dx \\ \lesssim \|\partial \psi\|_{L^2(Future)}^2 + \|\partial \psi\|_{L^2(Past)}^2 + \int_{Bulk} |\partial_r \psi r F|. \end{aligned}$$

Takeaway: integral over all of time but a compact set in space scales like  $r^{1/2}$ .

We begin again with writing the wave equation as

$$-\partial_u \partial_v \psi + \Delta \psi = -\partial_t^2 \psi + \partial_r^2 \psi + \Delta \psi = rF$$

where  $\psi := r\phi$ . Let  $h(r)$  be a (bounded, smooth) function to be determined later. Multiplying by  $h(r)\partial_r \psi$  and commuting, we deduce that

$$h(r)\partial_r \psi (-\partial_t^2 \psi + \partial_r^2 \psi + r^{-1} \mathring{\Delta} \phi) = h(r)\partial_r \psi rF \quad (2)$$

(here we use the fact that  $\mathring{\Delta} = r^2 \Delta$  and commuted the multiplication of  $\phi$  by  $r$  with the angular derivatives).

We would like to write the left hand side as

$$\begin{aligned}
 & -\partial_t (h(r)\partial_r\psi\partial_t\psi) + \frac{1}{2}\partial_r \left( h(r) \left[ (\partial_t\psi)^2 + (\partial_r\psi)^2 \right] \right) \\
 & + \frac{1}{2}\partial_r (h(r)\partial_r\psi r^{-1}\partial_r\phi) + \text{Error}.
 \end{aligned}$$

Denote the terms in (2) by  $A, B, C$  and the terms in the above equation by  $I, II, III$ . We now calculate the error terms. First, we have

$$A = I + h(r)\partial_t\partial_r\psi\partial_t\psi.$$

(Note all  $t$  derivatives of  $h$  vanish identically by construction).

Now we compute that

$$II = B + h(r)\partial_r\partial_t\psi\partial_t\psi + \frac{1}{2} \left( h'(r) \left[ (\partial_t\psi)^2 + (\partial_r\psi)^2 \right] \right)$$

so

$$B = II - h(r)\partial_r\partial_t\psi\partial_t\psi - \frac{1}{2} \left( h'(r) \left[ (\partial_t\psi)^2 + (\partial_r\psi)^2 \right] \right)$$



Finally,

$$III = C + \frac{h(r)}{r} \langle \dot{\nabla} \partial_r \psi, \dot{\nabla} \phi \rangle \quad (3)$$

$$= C + h(r) \langle \dot{\nabla} \partial_r \phi, \dot{\nabla} \phi \rangle + \frac{h(r)}{r} \langle \dot{\nabla} \phi, \dot{\nabla} \phi \rangle \quad (4)$$

$$= C + \frac{1}{2} \left( \partial_r \left( h(r) |\dot{\nabla} \phi|^2 \right) - h'(r) |\dot{\nabla} \phi|^2 \right) + \frac{h(r)}{r} \langle \dot{\nabla} \phi, \dot{\nabla} \phi \rangle \quad (5)$$

Thus, the overall error term is equal to

$$\begin{aligned} & - \left[ \frac{1}{2} \left( h'(r) \left[ (\partial_t \psi)^2 + (\partial_r \psi)^2 \right] \right) + \frac{1}{2} \left( \partial_r \left( h(r) |\dot{\nabla} \phi|^2 \right) \right. \right. \\ & \quad \left. \left. - h'(r) |\dot{\nabla} \phi|^2 \right) + \frac{h(r)}{r} |\dot{\nabla} \phi|^2 \right] \end{aligned}$$

(note that we used the equality  $\partial_r \partial_t = \partial_t \partial_r$  to cancel the first error term).

Integrating by parts, we obtain the equality

$$\int_{Bulk} \frac{1}{2} \left( h'(r) \left[ (\partial_t \psi)^2 + (\partial_r \psi)^2 \right] \right) + \left( \frac{h(r)}{r} - \frac{h'(r)}{2} \right) |\dot{\nabla} \phi|^2 \quad (6)$$

$$= - \left( \int_{Future} h(r) \partial_r \psi \partial_t \psi - \int_{Past} h(r) \partial_r \psi \partial_t \psi + \int_{Bulk} h(r) \partial_r \psi r F \right). \quad (7)$$

Now set  $h(r) = \frac{r}{2^k + r}$ . Note that  $h(r) \leq 1$ , and we have the equality

$$h'(r) = \frac{2^k}{(2^k + r)^2} \implies \frac{h(r)}{r} - \frac{h'(r)}{2} = \frac{2^k + 2r}{2(2^k + r)^2}.$$

We obtain the estimate

$$\int_{Bulk} \frac{1}{2} \left( \frac{2^k}{(2^k + r)^2} [(\partial_t \psi)^2 + (\partial_r \psi)^2] \right) + \left( \frac{2^k + 2r}{2(2^k + r)^2} \right) |\dot{\nabla} \phi|^2 \quad (8)$$

$$\leq - \left( \int_{Future} h(r) \partial_r \psi \partial_t \psi - \int_{Past} h(r) \partial_r \psi \partial_t \psi + \int_{Bulk} h(r) \partial_r \psi r F \right). \quad (9)$$

We now quickly prove a functional estimate that will be used above. Observe that for any  $\alpha > 1$  we have

$$\int_{r_1}^{r_2} r^{-\alpha} f^2 dr = \frac{1}{\alpha - 1} \left( \int_{r_1}^{r_2} 2(r^{-\alpha+1}) f f' dr - r^{-\alpha+1} f^2 \Big|_{r_1}^{r_2} \right)$$

Applying Young's inequality gives

$$\begin{aligned} \int_{r_1}^{r_2} r^{-\alpha} f^2 dr &\leq \frac{1}{2} \int_{r_1}^{r_2} r^{-\alpha} f^2 dr \\ &\quad + \frac{1}{\alpha - 1} \left( \left( \int_{r_1}^{r_2} (r^{-\alpha+2}) f'^2 dr \right) - r^{-\alpha+1} f^2 \Big|_{r_1}^{r_2} \right) \end{aligned}$$

and hence

$$\begin{aligned} \int_{r_1}^{r_2} r^{-\alpha} f^2 dr + \left[ r^{-\alpha+1} f^2 \right] (r_2) \\ \lesssim_{\alpha} \left( \int_{r_1}^{r_2} (r^{-\alpha+2}) f'^2 dr \right) + \left[ r^{-\alpha+1} f^2 \right] (r_1). \end{aligned}$$

Now note that when  $r \sim 2^k$  the coefficients in (8) control  $r^{-1}$ . Applying the inequality above with  $f = \int r\phi = \int \psi$ ,  $\alpha = 3$ , we deduce that

$$\sup_{k \in \mathbb{N}} \int_{r \sim 2^k} r^{-1} \phi^2 \lesssim \sup_{k \in \mathbb{N}} \int_{r \sim 2^k} r^{-1} |\partial \psi| \leq \sup_k (8).$$

Using the fact that  $\partial \psi = \phi + r\partial\phi$  and the triangle inequality, we have

$$\begin{aligned} \sup_{k \in \mathbb{N}} \int_{r \sim 2^k} r^{-1} \phi^2 + r |\partial \phi|^2 &\lesssim \sup_k (8) \\ &\leq \|\partial \psi\|_{L^2(Future)}^2 + \|\partial \psi\|_{L^2(Past)}^2 + \int_{Bulk} |\partial_r \psi r F|. \end{aligned}$$

## Pointwise estimates towards timelike infinity

So far, we've only discussed how to do pointwise estimates in the region  $u \lesssim r$ . It remains to prove decay in the regime  $u \gg r$ , or, equivalently,  $t \gg r$ . Since the weights attached to good derivatives only grow as  $r \rightarrow \infty$ , we will need to use the equation to exchange good derivatives for derivatives with weights in  $t$ . To do so, in this regime, we combine a variant of the spacetime elliptic estimate from above with an elliptic estimate involving  $\Delta$  to improve  $u$  decay in this region. The eventual goal is to show that

$$|\partial\phi(U, R)|^2 \lesssim U^{-3} R^\varepsilon$$

We will need the following functional estimate together with the obvious observation  $\Delta = \square + \partial_t^2$ .

### Theorem (Luk, Oh '23)

*For any  $\gamma \in (-3/2, -1/2)$ , we have the functional estimate*

$$\sum_{|\alpha| \leq 2} \|(\langle r \rangle \partial)^\alpha \phi\|_{L^{2,\gamma}} \leq C_\gamma \|\Delta \phi\|_{L^{2,\gamma+2}}. \quad (10)$$

*where*

$$\|f\|_{L^{2,\gamma}} = \|f\|_{L^2(\mathbb{R}^3, r^{2\gamma} dx)}$$

For any fixed  $U, R$ , we have

$$\begin{aligned}
|\partial\phi(U, R)|^2 &\lesssim U^{-1}R^{-\varepsilon} \int_{u \sim U} \left\| (u\partial_u)^\alpha (r\partial_r)^\beta \Omega^\gamma \partial\phi \right\|_{L^{2, -3/2+\varepsilon/2}}^2 \\
&\lesssim U^{-1}R^{-\varepsilon} \int_{u \sim U} \left\| \Gamma^{\leq 3} \partial\phi \right\|_{L^{2, -3/2+\varepsilon/2}}^2 \\
&\lesssim U^{-1}R^{-\varepsilon} \int_{u \sim U} \left\| \Delta \Gamma^{\leq 3} \partial\phi \right\|_{L^{2, 1/2+\varepsilon/2}}^2 \\
&\lesssim U^{-1}R^{-\varepsilon} \int_{u \sim U} \left\| \partial_{tt}^2 \Gamma^{\leq 3} \partial\phi \right\|_{L^{2, 1/2+\varepsilon/2}}^2 + \textit{better}
\end{aligned}$$

Use Morawetz to conclude this is enough decay to close.



## Lemma

*We have the equality*

$$\mathrm{Ric}_{LL} = L(\cdots) + L^\mu L^\nu \tilde{\square}_g g_{\mu\nu} + \bar{\partial} g \partial g$$

**Proof sketch** We recall that the Ricci tensor is given by

$$\text{Ric}_{\mu\nu} = \partial_\alpha \Gamma_{\mu\nu}^\alpha - \partial_\mu \Gamma_{\nu\alpha}^\alpha + \Gamma_{\alpha\beta}^\alpha \Gamma_{\mu\nu}^\beta - \Gamma_{\mu\beta}^\alpha \Gamma_{\alpha\nu}^\beta$$

and hence

$$\text{Ric}_{LL} = L^\mu L^\nu \partial_\alpha \Gamma_{\mu\nu}^\alpha - L^\nu L(\Gamma_{\nu\alpha}^\alpha) + \Gamma_{\alpha\beta}^\alpha \Gamma_{LL}^\beta - \Gamma_{L\beta}^\alpha \Gamma_{\alpha L}^\beta.$$

We first note that

$$\Gamma_{\alpha\beta}^\alpha = g^{\alpha\alpha'} (\partial_\alpha g_{\beta\alpha'} + \partial_\beta g_{\alpha'\alpha} - \partial_{\alpha'} g_{\alpha\beta}) = g^{\alpha\alpha'} (\partial_\beta g_{\alpha'\alpha})$$

since the first and third terms are antisymmetric with respect to switching  $\alpha$  and  $\alpha'$ . We can thus expand the third term as

$$L^\mu L^\nu g^{\alpha\alpha'} (\partial_\beta g_{\alpha'\alpha}) g^{\beta\beta'} (\partial_{\beta'} g_{\mu\nu})$$

Now we expand the first term as

$$\begin{aligned}
& L^\mu L^\nu \partial_\alpha (g^{\alpha\beta} (\partial_\mu g_{\nu\beta} + \partial_\nu g_{\beta\mu} - \partial_\beta g_{\mu\nu})) / 2 \\
&= L^\mu L^\nu (\partial_\alpha g^{\alpha\beta}) (\partial_\mu g_{\nu\beta} + \partial_\nu g_{\beta\mu} - \partial_\beta g_{\mu\nu}) / 2 \\
&\quad + L^\mu L^\nu g^{\alpha\beta} \partial_\alpha (\partial_\mu g_{\nu\beta} + \partial_\nu g_{\beta\mu} - \partial_\beta g_{\mu\nu}) / 2 \\
&= L^\nu (\partial_\alpha g^{\alpha\beta}) (L g_{\nu\beta}) + L^\mu L^\nu g^{\alpha\alpha'} \partial_\alpha g_{\alpha'\beta'} g^{\beta\beta'} (\partial_\beta g_{\mu\nu}) / 2 \\
&\quad + L^\nu g^{\alpha\beta} L (\partial_\alpha g_{\nu\beta}) - L^\mu L^\nu g^{\alpha\beta} (\partial_{\alpha\beta}^2 g_{\mu\nu}) / 2 \\
&= L^\nu (\partial_\alpha g^{\alpha\beta}) (L g_{\nu\beta}) + L^\mu L^\nu g^{\alpha\alpha'} \partial_\alpha g_{\alpha'\beta'} g^{\beta\beta'} (\partial_\beta g_{\mu\nu}) / 2 \\
&\quad + L^\nu g^{\alpha\beta} L (\partial_\alpha g_{\nu\beta}) - L^\mu L^\nu \tilde{\square}_g g_{\mu\nu} / 2
\end{aligned}$$

and the fourth term as

$$\begin{aligned}
& -(L^\mu L^\nu g^{\alpha\alpha'} (\partial_\mu g_{\beta\alpha'} + \partial_\beta g_{\mu\alpha'} - \partial_{\alpha'} g_{\beta\mu}) \\
&\quad g^{\beta\beta'} (\partial_\nu g_{\alpha\beta'} + \partial_\alpha g_{\beta'\nu} - \partial_{\beta'} g_{\nu\alpha})) / 4.
\end{aligned}$$

All of the terms above involving a contraction of the form  $L^\mu \partial_\mu$  are already admissible quadratic terms. Similarly, all the products of the form  $g^{\alpha\alpha'} \partial_\alpha(\cdots) \partial_{\alpha'}(\cdots)$  are admissible quadratic terms, as

$$g^{\alpha\alpha'} = -\frac{1}{2} L^\alpha \underline{L}^{\alpha'} - \frac{1}{2} \underline{L}^{\alpha'} L^\alpha + \not{g}^{AB} (X_A)^\alpha (X_B)^{\alpha'}$$

so there are no terms of the form  $\underline{L}(\cdots) \underline{L}(\cdots)$  above. The primary term lacking null structure is

$$-(L^\mu L^\nu g^{\alpha\alpha'} g^{\beta\beta'} (\partial_\beta g_{\mu\alpha'}) (\partial_\alpha g_{\beta'\nu}))/4$$

but this exactly cancels one of the terms above.

## Appendix: strong non-global existence

The goal of this appendix is to show that all global solutions to the equation

$$\begin{cases} \square u = (\partial_t u)^2 \\ u(t=0) = u_0 \\ \partial_t u(t=0) = u_1 \end{cases} \quad (11)$$

with  $u_i$  smooth and compactly supported are trivial, implying that all nontrivial solutions blowup in finite time. Following the Keir/Luk notes, we will deduce this via a reduction to spherical means and an ODE blowup type result. We begin with the Darboux equation. For  $h \in C^\infty(\mathbb{R}^n)$ , define

$$M_h(x, r) := \frac{1}{|B(x, r)|} \int_{B(x, r)} h(y) dy = \int_{\mathbb{S}^1} h(x + rz) dz.$$

We claim the following:

### **Theorem**

*With  $M_h$  defined as above, we have*

$$\Delta_x M_h(x, r) = \left( \partial_r^2 + \frac{n-1}{r} \partial_r \right) M_h(x, r).$$

**Proof.**

By definition, we have

$$|B(0, 1)| \int_0^R r^{n-1} M_h(x, r) dr = \int_{|y| \leq R} h(x + y) dy.$$

Taking  $\Delta_x$  on both sides and integrating by parts, we deduce that

$$\begin{aligned} |B(0, 1)| \int_0^R r^{n-1} \Delta_x M_h(x, r) dr &= \int_{|y| \leq R} \Delta_x h(x + y) dy \\ &= \int_{|y| \leq R} \partial^i \partial_i h(x + y) dy \\ &= \int_{|y|=R} \frac{y^i}{R} \partial_i h(x + y) dy. \end{aligned}$$



**Proof.**

Changing variables to  $z = y/R$ , this is further equal to

$$R^{n-1} \int_{\mathbb{S}^1} z^i \partial_i h(x + rz) dy = |B(0, 1)| R^{n-1} \partial_r M_h(x, r).$$

Now taking derivatives with respect to  $r$ , we deduce that

$$R^{n-1} \Delta_x M_h(x, r) = (n-1) R^{n-2} \partial_r M_h(x, r) + R^{n-1} \partial_r^2 M_h(x, r)$$

as desired. □



We will also need the following calculation, where all functions are now living in  $\mathbb{R}^{n+1}$ :

### Lemma

*If  $\square u = F$ , then*

$$M_F(0, r) = -\partial_t^2 M_u(0, r) + \left( \partial_r^2 + \frac{n-1}{r} \partial_r \right) M_u(0, r).$$

*where now  $M_F$  implicitly also may depend on time.*

### Proof.

For any fixed  $r$ , we have

$$(-\partial_t^2 + \Delta_x) M_u(x, r) = \square_x M_u(x, r) = M_{\square u}(x, r)$$

so using the previous equation and plugging in  $x = 0$  yields the result. □

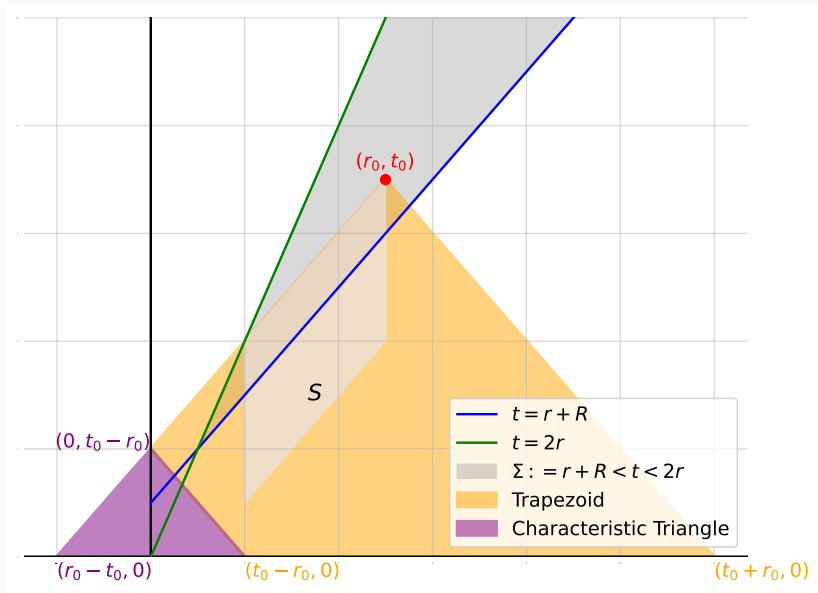
Finally, we will need the following explicit formula for solutions to the wave equation in  $1 + 1$  dimensions.

### Theorem

*The solution to the equation  $\square v = F$  with initial data  $v(t = 0) = v_0$  and  $\partial_t v(t = 0) = v_1$  is given by*

$$v(t, r) = \frac{1}{2} \left[ v_0(t - r) + v_0(t + r) + \int_{|r - r'| \leq t} v_1(r') dr' + \int_{T(t, r)} F(t', r') dt' r' \right]$$

*where  $T(t, r) := \{(t', r') \mid t' \leq t, |r - r'| \leq t - t'\}$  is the backward light cone from  $(t, r)$ .*



Now suppose we have a global  $C^2$  solution of (11), and take  $R$  to be such that the initial data is supported inside of  $B(x, R)$ . Define  $v(t, r) := M_u(0, r)$  and  $u := t - r$ . Note that  $\partial_r^2(rv) = r(\partial_r^2 + \frac{2}{r}\partial_r)v$ , so using 12, we know that  $rv$  satisfies the  $1 + 1$  dimensional wave equation

$$\partial_t^2(rv) - \partial_r^2(rv) = rF =: rM_{(\partial_t u)^2}$$

In particular, using 13 and dividing by  $r$ , we have that

$$v(t_0, r_0) = \frac{1}{2r_0} \left( \tilde{V} + \int_{T(r_0, t_0)} rFdrdt \right)$$

where  $\tilde{V}$  is a solution to  $\square \tilde{V} = 0$  with the correct data.

For  $(t_0, r_0) \in \Sigma := \{r + R < t < 2r\}$ , the contribution from the homogeneous solution vanishes, and hence

$$v(t_0, r_0) = \frac{1}{2r_0} \left( \int_{T(r_0, t_0)}^{t_0 + r_0} r F dr dt \right) \quad (12)$$

$$= \frac{1}{2r_0} \left( \int_{T(r_0, t_0) - T(0, u_0)} r F dr dt \right) \quad (13)$$

$$\geq \frac{1}{2r_0} \left( \int_{T^*(r_0, t_0)} r (\partial_t v)^2 dr dt \right) \quad (14)$$

where the last inequality follows from Jensen's inequality.

By positivity, we can further restrict the area of integration on the right hand side to the set

$$\{u_0 < r < r_0, -R < u < u_0\}$$

to replace the right hand side by

$$\frac{1}{2r_0} \int_{u_0}^{r_0} r dr \int_{r-R}^{r+u_0} (\partial_t v)^2 dt.$$

Now note that

$$|v(r, r + u_0)| = \left| \int_{r-R}^{r+u_0} \partial_t v(r, t) dt \right| \leq (u_0 + R)^{1/2} \left| \int (\partial_t v)^2 \right|^{1/2}$$

so plugging this into the previous equation yields

$$\begin{aligned} v(t_0, r_0) &\geq \frac{1}{2r_0} \int_{u_0}^{r_0} r dr \int_{r-R}^{r+u_0} (\partial_t v)^2 dt \\ &\geq \frac{1}{2r_0(u_0 + R)} \int_{u_0}^{r_0} r v(r, r + u_0) dr \end{aligned}$$

Now define

$$\beta(r_0) := \int_{u_0}^{r_0} r v(r, r + u_0)^2 dr$$

and note

$$\beta'(r_0) = r_0 v(r_0, r + u_0)^2 \geq \frac{1}{4(R + u_0)r_0} \beta^2$$

by the equation above.

Integrating this functional inequality implies that, if  $\beta(r_0) \neq 0$ , then

$$\frac{1}{\beta(r_0)} \geq \frac{1}{\beta(r_0)} - \frac{1}{\beta(r)} \geq \frac{1}{4} \frac{1}{(R + u_0)^2} \log \frac{r}{r_0}$$

for all  $r$ , which is impossible. We conclude that  $\beta = 0$  in  $\Sigma$ , hence  $v = 0$  in  $\Sigma$ . Now using 12, we deduce that  $v \equiv 0$  on a full slice, which concludes.